Near-surface characterization using Distributed Acoustic Sensing in an urban area: Granada, Spain

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10 Abstract

The Granada Basin in southeast Spain is an area of moderate seismicity. Yet, it hosts some of the 11 highest seismic hazards in the Iberian Peninsula due to the presence of shallow soft sediments 12 amplifying local ground motion. In urban areas, seismic measurements often suffer from sparse 13 instrumentation. An enticing alternative to conventional seismometers is the Distributed Acous-14 tic Sensing (DAS) technology that can convert fiber-optic telecommunication cables into dense 15 arrays of seismic sensors. In this study, we perform a shallow structure analysis using the am-16 bient seismic field interferometry method. We use a DAS array field test in the city of Granada 17 obtained on the August 26th and 27th, 2020, using a telecommunication fiber. In addition to the 18 existing limitations of using DAS with unknown fiber-ground coupling conditions, the complex 19 geometry of the fiber and limited data recording duration further challenge the extraction of surface-20 wave information from the ambient seismic field in such an urban environment. Therefore, we 21 develop an ad-hoc processing scheme in which we incorporate a frequency-wavenumber (f - f)22 k) filter to enhance the quality of the virtual shot gathers and related multi-mode dispersion im-23 ages. We are able to employ this dataset to generate several shear-wave velocity (V_S) profiles for 24 different sections of the cable. The shallow V_S structure shows a good agreement with different 25 geological conditions of soil deposits. This study demonstrates that DAS could provide insights 26 into soil characterization and seismic microzonation in urban areas. In addition, the results con-27 tribute to a better understanding of local site response to ground motion. 28

1 Introduction

The Granada basin, located in Andalusia in the southeast of Spain (Fig. 1), undergoes some of 30 the highest seismic hazards in the Iberian Peninsula. The Spanish Seismic Code (NCSE-02, 2002) 31 suggests that peak ground accelerations of 2.3 g are expected over a 500-years return period (J. P. Mon-32 tilla et al., 2001; Sanz de Galdeano et al., 2003; J. A. P. Montilla et al., 2003). This remarkably 33 strong ground motion for the region is due to a series of accumulating factors. In a regional con-34 text, Andalusia has a low-to-moderate seismicity resulting from the collision between the Eurasian 35 and African plates (Grimison & Chen, 1986; Hamdache, 1998; Serrano et al., 2002). Small seis-36 mic events are frequent, and moderate magnitude earthquakes (e.g., $M_w \leq 5.5$) are rather un-37 usual. Nonetheless, there are large earthquakes have been recorded: the 1910 M_w 6.2 Adra earth-38 quake (Stich et al., 2003), the deep 1954 M_w 7.8 Durcal earthquake (Chung & Kanamori, 1976; 39 Buforn et al., 1991), and the deep 2010 M_w 6.2 Nigüelas earthquake (Buforn et al., 2011). At the 40 local scale, the Granada Basin is surrounded by numerous faults, causing active microseismic-41

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ity ($M_w \ge 3.5$) in the Iberian Peninsula (Muñoz et al., 2002; Morales et al., 1997; Lozano et 42 al., 2022), but also catastrophic historical earthquakes: e.g., the 1884 December 25^{th} Andalu-43 sian earthquake (M \approx 6.7-7.1) that occurred in the south of the Granada basin, near Alhama de 44 Granada (Sánchez, 1987; Morales et al., 1996; Muñoz & Udias, 1992). 45 Besides, the seismic hazard in the basin is further exacerbated by the local site conditions, which 46 can lead to significant amplification of seismic ground motion. In the past, site effects in this re-47 gion have contributed to great damage during moderate and large earthquakes (Morales et al., 48 1991, 1993; García-García et al., 1996; Vidal & Castillo, 1994). Near-surface lithology can strongly 49 increase the amplitude and duration of earthquake ground motion and can also respond non-linearly 50 to incident seismic waves (Aki, 1998; Cruz-Atienza et al., 2016; Sanchez-Sesma, 1987; Viens 51 et al., 2022). In particular, sedimentary basins with soft sediments over hard basement rocks are 52 particularly prone to strong site effects as observed with the 1985 Michoacán earthquake in Mex-53 ico City (Campillo et al., 1989), the 1995 Kobe earthquake in Japan (Pitarka et al., 1998), and 54 the 2015 Gorkha earthquake in Nepal (Galetzka et al., 2015), among others. For these reasons, 55 site characterization is of great importance to seismic hazard analysis (Bommer et al., 2017; Aki, 56 1993), but generally challenging to obtain in densely populated areas. In urban centers, micro-57 zonation studies involving cabled or autonomous nodal acquisition are particularly onerous to 58 conduct because of the physical, legal, and logistical constraints inherent to seismic experiments. 59 In addition, seismic surveys are often expensive, preventing scientists and engineers from cov-60 ering larges areas. In some singular cases, the energy industry may provide assistance, and col-61 laborative efforts emerge (e.g., Castellanos & Clayton, 2021), as was the case for the Granada 62 basin (Morales et al., 1990). 63 Today, an alternative to heavy seismic surveys called Distributed Acoustic Sensing (DAS) is emerg-64 ing as a promising tool for urban microzonation (Dou et al., 2017; Z. J. Spica et al., 2020; Shragge 65 et al., 2021). DAS is a rapidly evolving technology that converts standard telecommunication fiber-66 optic cables into large and ultra-dense seismic vibration sensing arrays. In its simplest form, a 67 DAS interrogator is an optoelectrical unit probing a fiber with repeated laser pulses. A fraction 68 of the light is reflected back to the interrogator due to optical heterogeneities. External forcing, 69 such as seismic waves, generate phase shifts of the back-scattered Rayleigh light, which are mea-70 sured by the interrogator and are proportional to the total strain (or strain rate) along the direc-71 tion of the fiber and over a sliding spatial distance (i.e., the gauge length) (Grattan & Sun, 2000). 72 73 For a review of the DAS technology, we refer the reader to Hartog (2017).

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DAS offers new opportunities for urban microzonation by providing ultra-high measurements 74 in areas that are sometimes difficult to probe. Millions of kilometers of fiber-optic cables have 75 already been laid out around the world over the past decades to support our modern telecommu-76 nication infrastructures. Many of these cables are concentrated in urban centers and could there-77 fore compensate for the scarcity of available seismic stations. In addition, while providing ultra-78 dense measurements, DAS interrogators can acquire data within a wider frequency range (from 79 mHz to kHz) than standard exploration geophones (Lindsey, Rademacher, & Ajo-Franklin, 2020). 80 Yet, there are also known trade-offs and drawbacks to consider when designing a DAS survey 81 and setting up acquisition parameters (Z. J. Spica et al., 2023). For example, a larger gauge length 82 will lower the spatial resolution and decrease statistical uncertainty in measurements over the gauges 83 (E. R. Martin, 2018). Furthermore, the gauge length parameter affects the amplitude response 84 by generating zero strain notches at harmonic frequencies (Jousset et al., 2018; Lindsey, Rademacher, 85 & Ajo-Franklin, 2020). Except in special cases, DAS commonly has a lower signal-to-noise ra-86 tio (SNR) and a more limited angular sensitivity than standard geophones due to its broadside 87 insensitivity to incoming waves with particle motions oblique to the fiber axis (E. R. Martin et 88 al., 2018). Nonetheless, these limitations are often compensated by the fact that DAS can pro-89 vide large aperture and dense sampling of the seismic wavefield using only one signal power source. 90 DAS for urban seismology has proven to be successful in monitoring earthquakes (Lindsey et 91 al., 2017; Fang et al., 2020), tracking traffic signals (Lindsey, Yuan, et al., 2020; Yuan et al., 2020; 92 E. R. Martin et al., 2018), identifying different types of noise sources (Huot et al., 2018; Zhu & 93 Stensrud, 2019; X. Wang et al., 2020), monitoring the structural health of a wind turbine tower 94 (Hubbard et al., 2021) and providing geotechnical information of the shallow subsurface (Z. J. Spica 95 et al., 2020). In recent years, several studies focused on the near-surface characterization with 96 DAS and ambient seismic field (ASF) interferometry (Dou et al., 2017; Parker et al., 2018; Z. J. Spica 97 et al., 2020; Yang et al., 2022; Shragge et al., 2021; Viens et al., 2023; Jousset et al., 2018; Ajo-98 Franklin et al., 2019; Zeng et al., 2017; E. R. Martin et al., 2017; E. Martin et al., 2016; Lellouch 99 et al., 2019). 100 ASF interferometry is particularly suitable for urban near-surface imaging because it is cost-effective, 101 noninvasive, and does not require active sources. In addition, when using existing telecommu-102 nication infrastructure, researchers can get access to measurement sites that are normally diffi-103 cult to reach or even impossible to access. However, applying ASF interferometry with DAS also 104 comes with a series of practical and technical challenges that may prevent researchers from tak-105 ing full advantage of the methods. First, because researchers are not included in the fiber network 106

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design process, they often have very limited inputs on the deployment condition (including fiber 107 coupling with the ground) and the precise geometry of the cable. This frequently leads to prac-108 tical constraints forcing researchers to trade-off between accurate but time-consuming fiber ge-109 olocation (Biondi et al., 2023) and matching channel locations utilizing the approximate map-110 ping provided by the cable owner and the nominal channel spacing set during acquisition. Sec-111 ond, the broadside insensitivity to incoming waves recorded with DAS means that the wavefield 112 energy not aligned with the cable will have a weaker contribution during the interferometry pro-113 cess (E. Martin et al., 2016; Shragge et al., 2021). Therefore, in the case of a complex fiber ge-114 ometry, the interferometry can be limited to specific sections of the cable (Z. J. Spica et al., 2020). 115 Third, the fiber/ground coupling condition may vary substantially along the cable. In some ex-116 treme cases, the cable can be fully uncoupled, preventing the use of some sections of the cable. 117 In all cases, these issues must be taken into consideration prior to interpreting the data, and adapted 118 data processing should address these issues to obtain reliable results. 119 In this study, we expose the benefits and challenges of using DAS for microzonation in a densely 120 populated urban area. We present a case study in Granada using a 20-km section of an existing 121 telecommunication fiber, which recorded ASF with a DAS interrogator for less than one day. We 122 show that such a short data set can be used to infer the shallow velocity structure in several re-123 gions of the city. We first discuss the effect of data pre-processing on the retrieval of accurate dis-124 persion images at high frequency. We then perform a multi-mode inversion to constrain the lo-125 cal 1-D shallow shear-wave velocity (V_S) structure. Finally, we further discuss the challenges 126 of operating high-frequency ASF interferometry in an urban area with a limited amount of data 127 and compare our results with previous studies to validate them. 128

¹²⁹ 2 Geological setting

The Granada basin is one of the largest intramountainous Neogene–Quaternary basins of the Betic 130 Cordilleras (Morales et al., 1990; Banda & Ansorge, 1980), which resulted from the collision be-131 tween the Eurasian and African plates between the Cretaceous and Neogene periods. Jurassic 132 and Cretaceous carbonate sedimentary series bound the northern and western parts of the basin, 133 while the southern and eastern regions are bordered with metamorphic units (García-Dueñas & 134 Balanyá, 1986). 135 The Granada basin deep structure was first estimated through a combination of gravity data and 136 seismic-reflection profiles (Morales et al., 1990; Rodríguez-Fernández & De Galdeano, 2006). 137

¹³⁸ These results highlight a complicated basin geometry with four deep micro-basins (i.e., depocen-

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Figure 1. The GranaDAS array. (a) Map of the DAS fiber array located in the city of Granada. The background colors highlight the main superficial geological deposits (modified from Geological Survey of Spain, 2014). (b) The extent of the fiber probed during the experiment. The red box highlights the panel shown in (a). (c) Tectonic model derived from gravity data along profile A-A' shown in (b) (modified from Madarieta-Txurruka et al., 2021). (d) Map of Spain highlighting the location of Granada.

139	ters) with depths varying between \sim 1800-3000 m. The complexity of the basin geometry was
140	further evidenced through 3-D imaging of the V_S derived from Rayleigh wave dispersion anal-
141	ysis (Chourak et al., 2003). While the sedimentary layers in the basin have relatively low V_S ve-
142	locities (Banda & Ansorge, 1980; Gurria et al., 1997; Serrano et al., 2002), the basement under
143	the basin typically exhibits velocities around \sim 3.1 km/s (Banda & Ansorge, 1980). According
144	to the basement map of the Granada basin, the city of Granada lies on the edge of an 1800-m de-
145	pocenter (Morales et al., 1990; Gil-Zepeda et al., 2002). One interpretation of gravity and seis-
146	mic reflection data suggests that the depth of the basement beneath Granada city is approximately
147	1000 m (Morales et al., 1990), while another quantitative study of subsidence indicates a depth
148	of 500-900 m under the urban areas (Rodríguez-Fernández & De Galdeano, 2006). As a result,
149	the analysis of the shallow subsurface described in this contribution is unlikely to reach the in-
150	terface with the Basin's basement. Overall, these studies contributed to show that the overall struc-
151	ture of the basin contributes to the seismic wave amplification (Gil-Zepeda et al., 2002; Lee et
152	al., 2008). However, the lack of resolution in the shallow subsurface may lead to an underesti-
153	mation of the wave amplifications (Semblat et al., 2005; Narayan & Singh, 2006), as shallow stratig
154	raphy may impact local site effect's peak amplitude and frequency, and the dispersion behavior
155	of surface waves (e.g., Cruz-Atienza et al., 2016; Brissaud et al., 2020; Lee et al., 2008).
156	As such, the shallow basin structure has been assessed sporadically in Granada (Navarro et al.,
157	2010; Vidal et al., 2014). Navarro et al. (2010) got six shallow (\sim 40 m depth) V_S models through-

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out the city by means of inversion of Rayleigh wave dispersion curves obtained through the SPAC

- method. Vidal et al. (2014) obtained ten V_S models following a joint inversion of Rayleigh wave
- dispersion curves and horizontal-to-vertical spectral ratio from microtremor. They showed that
- the average V_S for the upper 30 m of soil (V_{S30}) varies from ~300 m/s to ~520 m/s, coincid-
- ¹⁶² ing with the Genil river deposit and Alhambra formation, respectively.
- In Granada, we can distinguish two main shallow geological formations (Fig. 1c). In the west-
- ern part of the Alhambra fault, we observe Quaternary sediments characterized by alluvial fa-
- cies. They are overlaid by a thin layer of clayey soils and flood plains (Fig. 1a), formed by re-
- cent sedimentary deposits of Holocene ages as a product of the Genil river draining. In the east-
- ¹⁶⁷ ern part, near the hill zones, we observe the Alhambra formation, which is characterized by sands,
- gravel, and carbonate conglomerates of Pliocene ages. Both the Alhambra formation and Granada
- depocenter shallow sedimentary deposits have a similar thickness (~200 m) near the Granada
- fault zone (Madarieta-Txurruka et al., 2021).

3 Data and methods

3.1 The GranaDAS array

The fiber-optic cable used in this study is operated by the IRAM (Instituto de Radioastronomía 173 *Milimétrica*) and provides continuous telecommunication between the radio-telescope at the top 174 of the Sierra Nevada and its headquarters in Granada (Fig. 1). The cable has no extra fiber (i.e., 175 dark) available during normal operation. Therefore, we connected a Febus Optics A1-R inter-176 rogator to the fiber during the maintenance of the radio-telescope that happened between the 26^{th} 177 and 27^{th} of August 2020. In total, we collected about 19 hours of strain rate data at a sampling 178 rate of 2000 Hz. Apart from the night interruption, during which the radio-telescope was oper-179 ating, the setup recorded data along the first 20 km of the fiber from the observatory, with 4167 180 channels spaced by 4.8 m, and with a 9.6-m gauge length. The raw dataset has a volume of up 181 to 4.56 TB. 182

¹⁸³ The GranaDAS array crosses different neighborhoods in Granada before climbing up the Sierra

184 Nevada in the east (Fig. 1). The fiber installation report provides a detailed location of the fiber,

- which is most often located underneath the side of the roads. The report does not mention any
- ¹⁸⁶ fiber slack loops placed in manholes, suggesting the fiber is outstretched between the IRAM and
- the radio-telescope. Therefore, the location of each channel has been interpolated based on the
- recording parameters, leaving a moderate degree of uncertainty in the assigned channel geolo-
- cations. Besides, the coupling conditions of the fiber with the ground are unknown. In addition,

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there are two three-component accelerometers in the city that are managed by the National Ge-

¹⁹¹ ographic Institute of Spain.

3.2 Ambient seismic field interferometry

In an urban environment, ambient seismic sources are generally strongly localized, moving, nar-193 row band, or even monochromatic (Jakkampudi et al., 2020; Bonnefoy-Claudet et al., 2006). For 194 example, these sources can originate from different traffic levels along the roads, the presence 195 of cooling and heating systems throughout the neighborhoods, construction activities, or even 196 the presence of pedestrians and bicycles near the fiber. All these sources make the assumption 197 of an equipartitioned ASF unrealistic (e.g., Snieder et al., 2010). Spectral and time normaliza-198 tions (Bensen et al., 2007) are generally unable to counterbalance the resulting heterogeneous 199 illumination. Additionally, spurious arrivals can appear in the cross-correlation functions (CCFs) 200 when the sources are between the receivers (Retailleau & Beroza, 2021). The specific case of DAS 201 with a shallow fiber placed along the road lines in a city is thus challenging for obtaining reli-202 able CCFs. As previously mentioned, the measurements have been made only during day time, 203 so quiet night hours cannot be selected. In this section, we describe the data processing work-204 flow utilized to obtain inverted V_S profiles. Specifically, we will demonstrate how the applica-205 tion of a frequency-wavenumber (f - k) filter to CCFs helps to mitigate spurious arrivals and 206 enhance the quality of the CCFs. 207

208 3.2.1 Cross-correlation functions

The entire dataset is first downsampled to 40 Hz, then windowed into short time series, demeaned, and bandpass filtered between 0.01 - 20 Hz. We also apply a 1-bit normalization to reduce the influence of non-stationary sources (Bensen et al., 2007). We then use the cross-coherence approach to emphasize the phase information, which is more suitable for noisy data that vary in amplitude among traces or have long and complex source wavelets (Nakata et al., 2011). The CCFs from cross-coherence interferometry are computed in the frequency domain as follows:

$$CCF(x_{\rm A}, x_{\rm B}, \tau) = \left\langle \mathcal{F}^{-1} \left[\frac{s(x_{\rm A}, \omega) s^*(x_{\rm B}, \omega)}{\{ |s(x_{\rm A}, \omega)| \} \{ |s(x_{\rm B}, \omega)| \} } \right] \right\rangle; \tag{1}$$

where $s(x_A, \omega)$ and $s(x_B, \omega)$ are the Fourier transform of windowed strain rate recordings at the receiver channels x_A and x_B , respectively. The asterisk indicates a complex conjugate and $\{\cdot\}$ represents a smoothing of the absolute amplitude spectrum $(|\cdot|)$ using a running absolute

mean algorithm. The inverse Fourier transform (\mathcal{F}^{-1}) is applied to retrieve the CCFs in the time 218 domain. $\langle \cdot \rangle$ expresses the stacking CCFs for all time windows. In general, the SNR increases by 219 stacking over longer periods. However, since we only have 19 hours of data, we optimized the 220 interval and overlap of CCF time windows. After testing parameters, CCFs are computed using 221 30-s windows with 50% overlap, and final CCFs are folded at zero time lag and stacked. In ad-222 dition, we use the phase-weighted stacking method to further enhance the signal quality (Schim-223 mel & Paulssen, 1997). CCFs are computed for virtual sources every five channels (i.e., every 224 24 m) with the corresponding 100 receiver channels (e.g., virtual source 610 with channels 610 225 to 710). These 100 CCFs form a virtual shot gather, which depicts the seismic wavefield prop-226 agating between the virtual source and the corresponding receivers (Fig. 2a). Finally, we only 227 compute virtual shot gathers for straight sections of the fiber because we focus on Rayleigh wave 228 (E. Martin et al., 2016). 229

3.2.2 Remove spurious arrival using f - k filtering

We utilize the high spatial density of the DAS array to transform the time-domain virtual shot gathers S(t, x) to its frequency-wavenumber (f - k) representation $\tilde{S}(\omega, k)$ (Fig. 2a & e, respectively). We then apply two filters to $\tilde{S}(\omega, k)$. The first one selects waves propagating in a certain velocity range (Embree et al., 1963) (Fig. 2b) and involves a phase velocity taper denoted by g:

1

$$\tilde{s}(\omega,k) = g(\omega,k,c_a,c_b,c_c,c_d) * \tilde{S}(\omega,k)$$
(2)

with
$$g(\omega, k, c_a, c_b, c_c, c_d) = \begin{cases} \sin \frac{\pi}{2} \frac{c-c_a}{c_b-c_a}, & \text{if } c_a \leq c \leq c_b \\ 1, & \text{if } c_b < c < c_c \\ 1-\sin \frac{\pi}{2} \frac{c-c_c}{c_d-c_c}, & \text{if } c_c \leq c \leq c_d \\ 0, & \text{otherwise.} \end{cases}$$
 (3)

Here $c = \omega/k$ is the observed phase velocity, and $c_b = 100$ m/s, $c_c = 2500$ m/s, and $c_a =$ 236 $0.95c_b$ and $c_d = 1.05c_c$ are the velocity parameters of the taper. The virtual shot gather result-237 ing from this first filter is shown in Figure 2b. The low-velocity waves already appear more clearly. 238 On the basis of the first filter, the second filter further helps remove reflected waves, waves due 239 to secondary sources, or aliasing, i.e. waves propagating in unexpected directions (Cheng et al., 240 2018). By noting the different quadrants of Fig.2e as Q1, Q2, Q3, and Q4, we observe the sym-241 metry: $\|\tilde{S}(\omega,k)\| = \|\tilde{S}(-\omega,-k)\|$ and $\|\tilde{S}(-\omega,k)\| = \|\tilde{S}(\omega,-k)\|$, i.e Q1=Q3 and Q2=Q4. 242 However, the absolute maximum amplitude (red and green lines) are all parallel, suggesting the 243 effect of aliasing (D.-Y. Wang & Ling, 2016). By selecting only the quadrant Q2 (i.e. $\tilde{S}(\omega > 0, k < 0)$) 244



Figure 2. Virtual shot gather for channel #610 and the f - k spectrum. (a) Initial virtual shot gather before f - k filtering. (b) Virtual shot gather after a f - k filtering, using the filter shown in f and following Eq. 2. (c) virtual shot gather using filtered f - k spectrum but only using Q1 and Q3 after a phase velocity taper in (f). (d) Similar to (c) but using energy from Q2 and Q4.

after a phase velocity taper, to construct a fully symmetric \tilde{S} , we obtain the virtual shot gather 245 shown in Figure 2c, where we mainly observe waves propagating toward the source. On the op-246 posite, by constructing a fully symmetric \tilde{S} from the unique selection of Q1, we retrieve surface 247 waves propagating from the source (Fig. 2d). This process clearly improves the quality of the 248 retrieved Rayleigh wave as well as the overall SNR of the filtered CCFs (Fig. 2d). 249 We also applied the f-k filtering to raw DAS data to investigate if we could improve the SNR 250 by removing moving sources. In particular, we focused on the locations exposed to vehicular traf-251 fic from two opposite directions (channels #3400-3600 located in western mountain areas). Af-252 ter filtering the waves outside the velocity range [5, 40] m/s, we enhance the passing of three cars 253 by making apparent three straight lines (fig. S1b) and further selected wave propagating either 254 in the positive or negative directions (fig. S1c, d). Even if waves propagating in one of the direc-255 tions are entirely removed, the straight lines associated with the traffic do not completely disap-256 pear after f - k filtering (fig. S1d). This is because cars are not only imposing stress changes 257 where they pass but also continuously generate waves in both directions due to their accelera-258 tion or their contact with road asperities. Consequently, the effect of coherent sources present be-259

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tween the channels cannot be fully removed. Additionally, the efficacy of the filtering requires the optical fiber section to be straight and applied through time-limited windows. Then, using this process to raw data prior to signal correlation is computationally expensive and may not guarantee the removal of in-between source effects. As a result, we only applied the f-k filtering directly to CCFs.

265 **3.2.3** *Phase velocity estimation*

We then calculate the SNR as the energy ratio of the signal in the time window of expected wave 266 arrivals, i.e., of s(t) with $t \in [d/v_{max}, d/v_{min}]$ with $v_{max} = 2500$ m/s and $v_{min} = 100$ m/s, 267 and the noise defined by s(t) with $t > [d/v_{min}]$. We then apply a slant stack algorithm (Chap-268 man, 1981) to the filtered CCFs with SNR larger than 10 to obtain dispersion images in the frequency-269 velocity domain (fig. S2). As observed in a previous study Viens et al. (2023), the number of re-270 ceivers considered in the slant stack process impacts the retrieval of the dispersion curves (DCs, 271 Fig. 3a, b & c). A higher number of receivers allows a better dispersion curve mode separation 272 by sampling the wavefield with a smaller wave number interval. Yet, it afflicts spatial resolution. 273 According to the tests shown in Fig. 3, we consider the number of 70 channels during the slant 274 stack process at source 610 a good compromise. However, this number also depends on local con-275 ditions (e.g., the shape of the array). Therefore, we optimize it for all the sources and obtain num-276 bers that range between 40 and 70 (Table 1). Next, we extract the dispersion points from the lo-277 cal energy maxima (energy maxima within subsets) observed in the dispersion images for each 278 frequency (Fig. 3a). Nonetheless, to avoid artifacts at low velocity, we further impose a selec-279 tion in the frequency-velocity domain, as shown in Fig. 3b, similarly to (Viens et al., 2023). Fi-280 nally, we select nine sources at which the dispersion pictures allowed obtaining enough infor-281 mation to realize the inversion of the dispersion points, including one source in mountainous ar-282 eas (see fig. S2). Table 1 displays the details of these virtual sources, including the number of 283 receivers and the retrieved modes of dispersion. 284

$3.2.4 V_S$ inversion

As explained in Viens et al. (2023), it is challenging to associate the selected dispersion points to a continuous curve for a given mode, particularly at high frequencies. Therefore, these points are treated individually without relating them to a specific mode and we minimize the distance between each of these dispersion points to the theoretical DCs. The forward theoretical DCs calculation follows the method presented in Perton & Sánchez-Sesma (2016). The distance is given

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Figure 3. Slant slack and dispersion images using virtual shot gather obtained for channel #610 and shown in Fig. 2d. (a-c) show the normalized spectral energy with 60, 70, and 80 channels. (d) shows the observed and theoretical DCs computed from the inverted V_S model #610 (Fig. 4).

by the misfit function between the observed and theoretical Rayleigh wave phase velocities (Eq.
4):

$$\epsilon_{\rm DC}^2 = \frac{1}{J_{\rm max}} \sum_{f=f_{\rm min}}^{f_{\rm max}} \sum_{j=0}^{j_{\rm max}} \sum_{n=0}^{n_{\rm max}} G\left(\left|c_j^{\rm obs}(f) - c_n^{\rm th}(f)\right|\right)^2 \tag{4}$$

where c_j^{obs} and c_n^{th} are the phase velocities of the observed and theoretical dispersion curves, re-

spectively. *n* represents the number of theoretical modes and *j* is the number of observed DC points

at a certain frequency f. J_{max} is the number of selected DC points in the frequency band from

 f_{\min} to f_{\max} . The function G evaluates if a selected dispersion point aligns with a theoretical dis-

²⁹⁷ persion point using the following equation:

$$G(x) = \begin{cases} x & \text{when} \quad |x| \le \delta \\ \delta & \text{when} \quad |x| > \delta \end{cases}$$
(5)

where δ represents the misfit threshold which is set to the average velocity difference between the observed DC points. The misfit function is minimized by a constrained nonlinear optimization method (Byrd et al., 1999; Perton et al., 2020).

We start the inversion process with an initial V_S profile built from two previous velocity mod-

els. The shallow layers (≤ 150 m) are similar to the ones presented inVidal et al. (2014), and the

half-space (\leq 400 m), which corresponds to the basement assessed in Chourak et al. (2003), has

a velocity of \sim 3 km/s. As surface wave DCs hold a higher sensitivity to V_S compared with den-

sity and the compression-wave velocity V_P (Yoshizawa & Kennett, 2005), these two latter pa-

rameters are determined from V_S by using the empirical relationships of (Brocher, 2005). Then,

the inversion process has only two free parameters by layer: the thickness and V_S .

308 4 Results

We present the 1-D V_S models at different virtual sources in Figure 4. The velocity models al-

low identifying two main seismic horizons in the upper 160 m. The shallowest seismic horizon

has a V_S ranging between ~350 and ~600 m/s, while the second horizon, identified as detritic

rocks, has a mean V_S of 2500-3000 m/s. The shallow seismic horizon has a minimum thickness

- of about 90 m, which increases to 125-135 m for models #595, 605, and 610. The interface be-
- tween these two seismic horizons has a significant V_S contrast, and the highest velocities for both
- ³¹⁵ horizons are observed at site #2850, on top of the Alhambra formation.



Figure 4. Shear-wave velocity inversion results. (a) V_{S30} for all inverted models. (b) 1-D V_S inversion results from all analyzed sections of the fiber cable. (c) All the inverted models in lines. The black line is the average of all velocity profiles.

To further investigate the shallow V_S properties, we extract the V_{S30} (average shear-wave velocity of the top 30-m depth). The variations in V_{S30} values across different virtual sources reflect the spatial heterogeneity of the subsurface. Our results reveal that the slowest V_{S30} of ~270 m/s is found at site #315 in the heart of the city, while the fastest (~530 m/s) is located near site #2850, in the mountain. The V_{S30} values are relatively constant in other sections of the array and are around 400 m/s.

322 4.1 Reliability of the results

To establish the reliability of our results, we conduct a sensitivity analysis of the inversion method 323 by examining the sensitivity kernel of each individual DC as well as the effect of all jointly in-324 verted modes on the V_S models. The sensitivity kernels enable us to assess the sensitivity of each 325 individual Rayleigh DC as a function of depth and frequency (Aki & Richards, 2002; Tanimoto 326 & Tsuboi, 2009; Campman & Dwi Riyanti, 2007). For instance, a certain mode may show more 327 sensitivity to shallow parts at high frequencies, while being more sensitive to deeper depths at 328 lower frequencies. At a given frequency, higher modes of Rayleigh waves exhibit more sensi-329 tivity to the deeper V_S structure (fig. S3). Therefore, multi-mode DCs can offer more constraints 330 to the deeper structure. The DCs exhibit the most sensitivities to the upper 100 m and moderate 331 sensitivities to depths between ~ 100 and ~ 150 m (fig. S3). However, these sensitivity kernels 332

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Sources	Receiver Numbers	Retrieved Modes	V _{S30} (m/s)
315	40	7	266
485	50	5	400
595	60	7	416
605	60	7	378
610	70	8	373
1140	60	7	384
1230	50	6	387
1270	50	4	410
2850	40	5	532

Table 1. Summary of virtual sources, number of receivers, retrieved modes, and corresponding V_{S30} values of each velocity model.

offer only partial information as they consider each mode independently, and the sensitivity might 333 be higher by considering these kernels jointly. As discussed in Perton et al. (2022), the joint sen-334 sitivity is better assessed by presenting all the velocity models within twice the lowest misfit. We 335 present these variations at virtual source #610 in fig. S4. In general, the shallowest regions (\leq 75 336 m) show less uncertainty thanks to higher Rayleigh wave sensitivity. Yet, the first two layers in 337 the upper 10 meters exhibit higher variability, suggesting a more complex superficial structure 338 with likely more lateral variations. There is also a large velocity variability above the two seis-339 mic horizons interface, i.e. between 75-100 m depth. 340

Another approach to estimating the site characteristics is the horizontal-to-vertical spectral ra-

tio computed from ASF, also known as the H/V method (Nakamura, 1989). Subsurface elastic

³⁴³ property contrasts along depth can have a significant impact on the resonant frequency (repre-

sented by the H/V peak frequency), making H/V analysis highly sensitive to identifying the depth

of the impedance contrasts. In addition, the H/V method has recently gained theoretical devel-

opments from the diffuse field theory as it corresponds to the horizontal-to-vertical energy den-

- sities ratio of the ASF (Perton et al., 2009; Sánchez-Sesma et al., 2011). To further validate our
- results, we compare the observed H/V from ASF data recorded at a nearby three-component ac-
- celerometer (EXCAB, Fig 1a) with the theoretical H/V computed from velocity profile #595. The
- observed H/V is calculated through the autocorrelation of the signal components as in Eq. 6 (Per-
- ton et al., 2018). The theoretical H/V calculation is made according to Eq. 7 and by using the Dis-

- crete Wave Number method to calculate the Green function components $G_{ii}(\omega)$ (Bouchon, 2003;
- Sánchez-Sesma et al., 2011):

$$\frac{H}{V}(\mathbf{x},\omega) = \sqrt{\frac{\left\langle |v_1(\mathbf{x},\omega)|^2 \right\rangle + \left\langle |v_2(\mathbf{x},\omega)|^2 \right\rangle}{\left\langle |v_3(\mathbf{x},\omega)|^2 \right\rangle}};$$
(6)

$$=\sqrt{\frac{\operatorname{Im}\left(\mathcal{G}_{11}(\mathbf{x},\omega) + \mathcal{G}_{22}(\mathbf{x},\omega)\right)}{\operatorname{Im}\left(\mathcal{G}_{33},(\mathbf{x},\omega)\right)}}\tag{7}$$

where $v_i(\omega)$ is the particular velocity spectrum components, with $i \in \{1, 2\}$ corresponds to horizontal directions, and i = 3 refers to the vertical direction. $\langle \cdot \rangle$ represents the average over several windows. $G_{ii}(\omega)$ is the displacement in the direction i caused by a unit point force applied

in the same direction. Im() denotes the imaginary part operator.

To further enhance our understanding of the velocity contrast, a simplified two-layer model of 358 V_S is designed. Figure 5 shows the results of the observed vs. the theoretical H/Vs, for the model 359 at channel #595 and the simplified velocity structure. All the H/V curves exhibit a dominant peak 360 with a frequency of 1-2 Hz and comparable amplitudes. By employing the simplified two-layer 361 model, this peak amplitude can more easily be associated with the velocity contrast at a depth 362 of approximately 80 m. Therefore, it confirms the reliability and reasonable constraints of the 363 obtained interface depth and velocity contrast for the two main seismic horizons across multi-364 ple V_S models. 365

366 5 Discussion

- **5.1 Shallow structure under Granada city**
- Taken together, our results provide a seismic characterization of two main seismic horizons with a strong velocity contrast at depths greater than 90 m. At the west of the Alhambra fault in the urban area, the deeper horizon shows consistent and high V_S values ranging between ~2300-3000 m/s, which is associated with Plio-Quaternary detritic rocks. In the mountainous region (model #2850), the velocities of this deeper seismic horizon jump to ~3800 m/s, associated with the Late-
- Miocene detritic rocks and gypsum (Fig. 1c). According to Madarieta-Txurruka et al. (2021),
- it is expected that these geological units have similar thicknesses in the vicinity of profile A-A'
- 375 (Fig. 1c).
- The shallower seismic horizon (<80-100 m) shows comparable average velocities of around 500
- m/s along the fiber. However, this seismic horizon is identified as Quaternary alluvial deposits



Figure 5. H/V Comparison and V_S models. (a) The green solid line depicts the H/V measurement from the accelerometer EXCAB. The dash and dash-dot lines are the theoretical H/V curves calculated from the inverted velocity model #595 and from a simplified two-layer model, respectively. (b) V_S models used to compute theoretical H/V shown in (a).

in the city area, and as the Alhambra formation conglomerates in the mountain area (MadarietaTxurruka et al., 2021) (Fig. 1c). Although their different geological origins, it is difficult to distinguish them in our inversion results. Finally, on top of these seismic horizons, stand different
kinds of soils (Fig. 1a).

As a typical index for soil conditions, V_{S30} is used for site classification in Uniform Building Code 382 in the US (Dobry et al., 2000) and in the Eurocode 8 in Europe (Code, 2005; Kanlı et al., 2006). 383 According to Spanish Seismic Code NCSE-02 (NCSE-02, 2002), soil types are classified under 384 Type II and III with a V_{S30} range of 400-750 m/s and 200-400 m/s, respectively. Most inverted 385 V_S models stand at the common limit of these two categories as they have V_{S30} between ~370-386 415 m/s. Models #315 ($V_{S30} = 266$ m/s), and #2850 ($V_{S30} = 532$ m/s) are classified with 387 less ambiguity as type III soil (i.e., medium-soft soil) and type II soil (i.e., medium-hard soil), 388 respectively. Additionally, #315 and #485 are located on top of clayey soils that are composed 389 of red clays, gravel, and sand (Fig. 1a, (Geological Survey of Spain, 2014)). Model #315 could 390 reflect the soft clayey soils with a low V_{S30} value of 266 m/s. These values align with a study from 391 Navarro et al. showing that V_{S30} for a site located in the urban area affected by the upper part of 392 the Genil river alluvial fan is ~294 m/s. Besides, model #485 results from several receivers over-393 lapping with the floodplain region. This may explain why this model has slightly higher veloc-394 ities even though it is located on clayey soils in Fig. 1a. Overall, other areas on top of the flood-395 plain region, which comprise sands, gravels, and carbonates, have typical V_{S30} values around 400 396

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³⁹⁷ m/s. Finally, model #2850 is close to alluvium with compact conglomerates, which show higher ³⁹⁸ V_{S30} (i.e., ~530 m/s) than clayey soils, comparable to V_{S30} observed on Alhambra formation ³⁹⁹ (Navarro et al., 2010; Vidal et al., 2014).

However, the sensitivity of inversion shows a larger uncertainty for the shallowest 10 m (fig. S4) 400 due to the intricate heterogeneous and complex nature of the urban shallow subsoil that may be 401 compacted, stiffened by constructions, or crossed by human infrastructure (tunnels, pipes etc.). 402 Due to this possible lack of resolution in the V_{S30} , it is preferable to discuss the general features 403 of the velocity models. In addition, the V_{S30} values only provide a first-order approximation for 404 liquefaction hazard and for estimating site amplification factors by considering a standard veloc-405 ity at the basement (Borcherdt, 2012). However, as observed with the H/V measurements, elas-406 tic resonances of surface waves may change the amplification behavior at a specific frequency 407 due to strong contrast at depth. For seismic hazard assessment, it is important that these frequen-408 cies do not correspond to the building's vibration frequency which depends roughly on their height. 409 The velocity profiles demonstrate a high impedance contrast at a depth of 90-140 m (Fig. 4) and 410 H/V peak frequencies ranging from 1-2 Hz (Fig. 5), as observed by Vidal et al. (2014) and Vi-411 dal & Feriche (2012). Therefore, this area is likely inadequate for tall buildings of about 10 sto-412 ries that could vibrate with the soil's dominant frequency. 413

5.2 Challenges of extracting surface wave information

The utilization of DAS provides a promising non-invasive method for conducting dense seismic 415 measurements in urban areas. However, despite the GranaDAS array consisting of 4167 chan-416 nels, we were able to obtain high-quality dispersion images at only nine locations along the ca-417 ble. Several reasons can explain the limited number of observations using an urban telecom fiber. 418 For instance, DAS data are affected by the often unknown and uneven cable-ground coupling, 419 which can reduce sensitivity to ground motion and cause several cable segments unsuitable for 420 seismic measurements (e.g., Ajo-Franklin et al., 2019; Yuan et al., 2020; Z. J. Spica et al., 2020; 421 Tribaldos et al., 2021). As each cable has its own set of characteristics, the coupling should there-422 fore be one of the first parameters assessed for seismological studies. 423 Low coupling regions can result in low-SNR CCFs and make it challenging to extract surface wave 424

- information. Nonetheless, the SNR can increase with longer time series stacking (Bensen et al.,
- ⁴²⁶ 2007). In Granada, we only recorded strain rate during 19 hours and during day time. Such a short
- ⁴²⁷ recording duration may hinder the Green's function retrieval, especially at lower frequencies where
- 428 ASF tends to be less coherent.

In addition, the crooked layout of the GranaDAS array in urban areas can pose further challenges 429 for extracting surface wave information. Indeed, computing CCFs along winding sections of the 430 fiber can result in the emergence of a combination of Love and Rayleigh waves (E. R. Martin et 431 al., 2021), making it difficult to separate both wavefields (Song et al., 2021). Typically, pure Rayleigh 432 waves are obtained through radial-radial cross-correlations, limiting their retrieval to straight sub-433 sections of the fiber. This physical requirement further restricts the areas where we can assess 434 the subsurface information (e.g., Shragge et al., 2021; Rodríguez Tribaldos & Ajo-Franklin, 2021; 435 Yang et al., 2022; Song et al., 2021). 436 Despite the combined challenges of coupling, short-duration recordings, and twisting array ge-437 ometry, DAS also offers unique opportunities to improve the resolution of the dispersion images. 438 For example, it is a common problem in seismology to acquire high-quality multi-mode DCs (Z. Spica 439 et al., 2017; Perton et al., 2020; Viens et al., 2023; Takagi & Nishida, 2022). In this case, one ad-440 vantage of using DAS for ASF interferometry is the ability to adjust the number of receivers and 441 channel spacing during the processing steps, improving modal content and avoiding wavefield 442 aliasing for a given virtual source (Viens et al., 2023). During our processing, we adjust these 443 parameters to improve the quality of the modal content (Table 1). Ultimately, we obtain Rayleigh 444 wave DCs with up to eight modes, providing more resolution to our models. 445

446 6 Conclusions

We present a case study to image the detailed shallow V_S structure (< 160 m) in the city of Granada 447 using a DAS acquisition on a telecommunication fiber cable. We process 19-h ASF data using 448 the cross-coherence correlation interferometric method to generate virtual shot gathers along the 449 fiber. We apply an ad-hoc f - k filter to remove waves propagating in unexpected and spuri-450 ous directions, which improves the quality of the CCFs. Then, we convert the CCFs into disper-451 sion diagrams by applying the slant stack technique. Despite the challenges of using telecom-452 munication fiber in an urban area, we are able to generate several dispersion images that exhibit 453 rich multi-mode Rayleigh wave energy within the frequency range of 2-18 Hz. Finally, we in-454 vert Rayleigh wave dispersion points into 1-D V_S models using a global optimization algorithm 455 and image the top 160-m V_S structure. V_S profiles suggest softer structures for the site on clayey 456 soils and stiffer structures for the site with conglomerates near the mountain areas. Our results 457 show that the shallow alluvial deposits and detritic rocks have V_S values of \sim 300-600 m/s and 458 \sim 2500-3000 m/s, respectively. The interface of these two units is generally located at depths of 459 \sim 100 m, while it is slightly deeper for sites with older alluvium. This study highlights the po-460

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- tential of utilizing ASF data obtained through urban DAS for near-surface characterization, notwith-
- standing the associated difficulties. The extensive sensor network provided by DAS makes it a
- ⁴⁶³ promising tool for seismic microzonation, which is a crucial aspect of seismic risk analysis in
- densely populated urban areas.

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471 DATA AVAILABILITY

- 472 All the GranaDAS array data will be available on PubDAS (Z. J. Spica et al., 2023) after revi-
- sion and prior potential acceptance of this article. Accelerometer data can be accessed upon re-
- quest to the National Geographic Institute of Spain. The velocity models at all sites are availableat GitHub.

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