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5 The Role of Normal Fault Growth History in Influencing Fault Seal Potential,

6 Samson Dome, Offshore Norway

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11 Key Points:

- Fault growth histories created differences in sealing potential within the studied fault network.
- Samson Dome experienced multiple phases of faulting and doming.
- The driving mechanism behind the development of the Samson Dome remains ambiguous.
- 17

18 Abstract

The growth of normal faults can influence subsurface fluid flow and entrapment within rift 19 20 basins. However, fault seal studies typically view faults as static structures, with their growth and the potential related temporal changes in hydraulic properties being ignored. In this study, we use 21 borehole data and a high-quality 3D full-stack depth migrated seismic reflection volume to 22 23 analyse the growth history of the normal fault network in the Samson Dome area, SW Barents Sea. We specifically focus on how the kinematic history of normal faults impacts their sealing 24 properties, whilst also considering their origin and implications for regional salt tectonics. We 25 show that the faults formed during two distinct phases in the Late Triassic and Middle Jurassic-26 to-Early Cretaceous, and two phases of dome growth occurred in the Late Triassic and Late 27 Cretaceous, challenging existing proposals for the timing of the development of structure. The 28 29 salt-tectonic origin of the Samson Dome itself remains enigmatic, although mechanical considerations suggest existing models require refinement. Our fault seal analysis reveals a 30 correlation between displacement patterns and sealing potential, as reflected in the Shale Gouge 31 Ratio (SGR) values of different fault groups. More specifically, faults that grew via vertical 32 linkage of initially isolated segments and experienced potential reactivation exhibit lower SGR 33 values, implying a higher likelihood of across and along-fault leakage. Conversely, faults lacking 34 evidence for reactivation or vertical linkage show higher SGR values, suggesting better sealing 35 36 potential. Notably, the accuracy of our fault seal analysis is greatly influenced by the calculation methods used for V_{shale}, particularly for faults with low displacement. Our study provides 37 valuable insights into the faulting, development, and sealing potential of the Samson Dome area. 38 We also highlight the importance of careful consideration of V_{shale} calculation methods when 39 conducting fault seal analysis. 40

41 Plain Language Summary

42 We examined how normal faults affect fluid movement underground and fluid trapping in the Samson Dome area in the SW Barents Sea. Instead of viewing faults as unchanging, we used 43 44 borehole data and advanced 3D seismic imaging to study how they grew over time. We focused on how the history of fault movement influences their ability to trap fluids and what this reveals 45 46 about salt movement. We found that faults formed during two different ancient periods, challenging previous beliefs about their timing. While the origin of Samson Dome is still 47 unclear, our research suggests current explanations might need adjustments. We discovered a 48 link between fault movement and fluid trapping. Faults that connected initially separate segments 49 and might have moved again had weaker trapping abilities, allowing fluids to potentially leak. In 50 contrast, faults that did not show signs of further movement had better trapping abilities, keeping 51 52 fluids underground more effectively. Overall, our study sheds light on fault growth, fluid trapping, and the importance of accurate value calculations in understanding these processes. 53

54 **1 Introduction**

- 55 It is widely accepted that faults can act as barriers to fluid flow by: i) juxtaposing permeable rocks
- against impermeable rocks (Allan, 1989), or ii) creating an impermeable fault rock that obstructs
- 57 lateral and vertical migration (e.g., Knipe, 1992; Yielding et al., 1997; Fossen & Bale, 2007; also
- see review by Manzocchi et al., 2010 and references therein). Alternatively, faults can facilitate

fluid flow by acting as high-permeability conduits (Hooper, 1991; Dockrill & Shipton, 2010; 59 Skurtveit et al., 2021). Reservoir simulation models thus need to incorporate faults in a 60 geologically realistic manner (Fisher & Jolley, 2007). Studying the sealing potential of a fault often 61 involves the use of quantitative algorithms that include Clay Smear Potential (CSP) (Bouvier et al. 62 1989; Fulljames et al. 1997), Shale Smear Factor (SSF) (Lindsay et al., 1993), and/or the most 63 64 commonly applied method, Shale Gouge Ratio (SGR) (Fristad et al., 1997; Yielding et al., 1997; Freeman et al., 1998). All these methods incorporate some level of analysis of clay content in the 65 faulted sequence (e.g., Freeman et al., 1998; Yielding, 2002; Manzocchi et al., 2010). SGR 66 calculations are influenced by uncertainty in the input parameters, and rely heavily on Vshale 67 calculations, which are susceptible to poorly constrained errors (Bailey et al., 2006). In fact, Bretan 68 et al. (2003) found that a 10% error range can exist in SGR calculations based on how Vshale is 69 estimated. SGR is also not the only factor that influences fault seal capacity, as other factors such 70 as sub-seismic strain (i.e., fractures and deformation bands), fault reactivation, and local variations 71 72 in fault plane continuity and throw can also impact a fault's sealing potential (Bailey et al., 2006). However, when calibrated using regional data (i.e., dynamic data such as reservoir pressure 73 measurements) and applied consistently, SGR calculations can be used in a relative sense to assess 74 75 seal potential and capacity across different faults within a given basin (e.g., Yielding, 2002; Bailey et al., 2006; Manzocchi et al., 2010; Yielding, 2015). A key aspect of almost all fault seal studies 76 is that they tend to only consider the present fault geometry, associated juxtapositions, and clay-77 78 related hydraulic properties, rarely considering the fault growth history and how seal properties may have changed through time (see Reilly et al., 2017 for an exception). For example, with 79 80 increasing displacement, SGR and sealing properties might increase. However, the fault may not 81 have been sealing if fluid or gas migration occurred during an earlier phases of fault development,

when displacement and related SGR values were lower. Such subtleties would be missed if
consider only present-day displacement and juxtaposition relationships.

The Barents Sea is being considered for the long-term geological storage of CO2 (e.g., Riis 84 & Halland, 2014), given it contains many dome-like structures and anticlines (e.g., Mitiku & 85 Bauer, 2013). The Samson Dome is a large, faulted, Mesozoic dome in the SW Barents Sea. 86 Although this structure contains good-quality reservoirs at multiple depths, an exploration well 87 88 proved that it did not contain hydrocarbons (Norwegian Petroleum Directorate, 2023), making it a possible candidate for future CO2 storage projects. However, the SW Barents Sea has a well-89 documented history of fault-related fluid leakage related to post-trap filling regional uplift (e.g., 90 91 Makurat et al., 1992; Doré & Jensen, 1996; Gabrielsen et al., 1997; Ostanin et al. 2013; Vadakkepuliyambatta et al., 2013; Hermanrud et al., 2014; Mohammedyasin et al., 2016; 92 Edmundson et al., 2020; Argentino et al., 2021). It is therefore vital to undertake a detailed 93 assessment of the hydraulic properties of the fault network above the Samson Dome, with explicit 94 recognition of its complex geometry and kinematic history. 95

The Samson Dome is a structural high on the Bjarmeland Platform, located between the 96 97 Hammerfest and Nordkapp basins (e.g., Gabrielsen et al., 1990; Breivik et al., 1995). The few 98 studies that have focused on the Samson Dome examined it at a regional scale using 2D seismic reflection (e.g., Gabrielsen et al., 1990; Vadakkepuliyambatta et al., 2013) and gravity data (e.g., 99 Breivik et al., 1995), interpreting it as a salt-related (i.e., inflated) anticline (e.g., Gabrielsen et al., 100 101 1990; Breivik et al., 1995; Mattos et al., 2016). Mattos et al. (2016) use regional 2D and 3D prestack time migrated seismic reflection and borehole data to investigate the salt-related structural 102 evolution of the Samson Dome by analysing the geometry and kinematics of the related normal 103 fault networks, suggesting doming and the main stage of faulting happened in the Late Cretaceous, 104

coincident, and therefore driven by, opening of the North Atlantic Ocean. In this study, we use 105 borehole data and a high-quality 3D full-stack depth seismic reflection volume that was generated 106 using a Full-Waveform inversion velocity model (Jones et al., 2013) to study the growth history 107 of the normal fault network in the Samson Dome area. Whereas Mattos et al. (2016) focus on using 108 fault geometries and kinematics to understand the halokinetic history of the Samson Dome 109 110 structure, we concentrate here on examining how the kinematic history and normal fault growth patterns might have influenced the faults' sealing potential and propose an alternative model for 111 the development of the Samson Dome and its related fault network. A subsequent study by 112 Alghuraybi et al. (2023b) provides a quantitative analysis of the velocity differences within the 113 fault zones of the fault network studied here and tests possible links between fault zone velocity 114 variations to the faults' kinematic history and fault seal potential. 115

116 **2 Geological Setting**

The Barents Sea is located in the northwest corner of the Eurasian tectonic plate south of the Arctic 117 Ocean (Figure 1) (e.g., Gabrielsen, 1984; Doré, 1995). The SW Barents Sea developed in response 118 to multiple phases of crustal extension, which formed predominately NNE-trending rift basins 119 (e.g., Nordkapp Basin, Hammerfest Basin) and basement highs (e.g., Loppa High, Norsel High; 120 Figures 1 and 2) (e.g., Faleide et al., 1984; 2008; 2015; Gabrielsen, 1984; Gabrielsen et al., 2016). 121 The oldest rocks in the SW Barents Sea are Late Cambrian to mid-Devonian (i.e., Caledonian) 122 igneous and metamorphic rocks, which contain fabrics and structures that influenced the evolution 123 and present structural framework of the area (e.g., Faleide et al., 1984; Ritzmann & Faleide, 2007). 124 The collapse of the Caledonian orogenic belt in the Devonian marked the onset of the first rift 125 phase, which lasted until the Carboniferous (e.g., Faleide et al., 2008). Late Devonian to 126 Carboniferous rifting created narrow basins that were initially filled during the latest 127

Carboniferous to Permian by evaporites and carbonates, followed by siliciclastic rocks (e.g., 128 Faleide et al., 1984). Significant accommodation was created during the second, Late Permian 129 rifting phase, which was subsequently infilled by further clastic sedimentation that continued into 130 the Triassic (e.g. Johansen et al., 1993; Larssen et al., 2002; Glørstad-Clark et al., 2010; 131 Harishidayat et al., 2015). The third rifting phase occurred during the Middle Jurassic – Early 132 133 Cretaceous and led to the formation of large, widespread basins between structural highs (e.g. Gabrielsen, 1984; Faleide et al., 1993; Doré, 1995; Faleide et al., 2008). Following the third rifting 134 phase, clastic sedimentation continued during the Late Cretaceous, with a major regional 135 unconformity forming at the base of the Paleogene in response to regional uplift (e.g., Faleide et 136 al., 2008). The opening of the Norwegian and Greenland seas during the Palaeocene – Eocene is 137 thought to be linked to the fourth rifting phase in the SW Barents Sea (e.g., Eldholm & Thiede, 138 1980; Faleide et al., 2008; Harishidayat et al., 2015). 139

Numerous salt diapirs have been described in the SW Barents Sea, in areas close to the 140 Samson Dome (e.g., the Nordkapp Basin; Nilsen et al., 1995; Rojo et al., 2016; Paoletti et al., 141 2020). However, the geometries and seismic expression of these structures and adjacent 142 depocentres (i.e., narrow, km-tall, seismic chaotic diapirs, flanked by seismically reflective 143 minibasins) in the Nordkapp Basin are clearly different from those characterising the Samson 144 Dome (see below). Thick (c. 2 - 4 km), evaporite-dominated layers were deposited across the 145 Barents Sea shelf following the Late Carboniferous - Early Permian rift phase (e.g., Nilsen et al., 146 1995; Gudlaugsson et al. 1998). More recent work estimates the average (depositional) salt 147 148 thickness within the Nordkapp Basin was c. 2.5 km (e.g., Grimstad, 2016). A combination of sediment loading (e.g., Grimstad, 2016; Rowan & Lindsø, 2017) and thick-skinned extension (i.e., 149

basement-involved normal faulting; e.g., Jensen & Sørensen, 1992; Nilsen et al., 1995) drove salt



151 mobilisation within the Nordkapp Basin.

Figure 1. (a) A map summarising the regional structural elements of the SW Barents Sea around the Samson Dome structure. The dashed lines denoted with X, X', Y, Y', Z, Z' are indicating the location of the regional seismic lines in Fig. 2. The map is modified after information found in the Norwegian Petroleum Directorate fact page <u>http://www.npd.no/en/</u>. (b) A subset location map highlighting the geographical location of the study area. (c) A summary of the present-day maximum horizontal stress orientations based on borehole breakout data retrieved from the World Stress Map (Heidback et al., 2016).

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161 The current model for the structural evolution of the Samson Dome suggests the main 162 phase of salt mobilisation was somehow triggered by extension caused by the opening of the North 163 Atlantic Ocean in the Early to Late Cretaceous. This is based on the observation that Late 164 Cretaceous strata are folded into a broad dome (e.g., Mattos et al., 2016). The proposed thickness

of salt required to inflate a structure the size of the Samson Dome is substantial (3.5 km; e.g., 165 Gabrielsen et al., 1990; Breivik et al., 1995). Consequently, we would expect to observe structures 166 related to complimentary salt withdrawal, e.g., minibasins, supra-salt faulting, welds (e.g., Jackson 167 & Talbot, 1986; Nilsen et al., 1995; Koyi, 1998; Jackson & Hudec, 2005; Hudec & Jackson, 2007; 168 Pichel et al., 2018; Rojo et al., 2019). Although such salt-related structures have been documented 169 regionally (i.e., in the Nordkapp Basin; e.g., Rojo et al., 2019), and despite supra-salt faulting being 170 locally intense above and around the Samson Dome (Mattos et al., 2016), they are notably absent 171 from the Samson Dome area (Figure 2). Seismic reflection data also show that the Samson Dome 172 is not underlain by the characteristic, chaotic, variable amplitude seismic facies that defines salt 173 structures in the Nordkapp Basin (e.g., Hassaan et al., 2021) and other salt-bearing sedimentary 174 basins (e.g., Jackson and Hudec, 2017) (Figure 2). This calls into question the interpretation of the 175 Samson Dome as a salt-tectonic structure purely related to halokinesis (i.e., salt flow induced 176 purely by gravity; Jackson and Hudec, 2017). 177



Figure 2. Un-interpreted (left) and interpreted (right) regional seismic sections showing the

182 overall structural style and stratigraphy around the Samson Dome (b). The section north of the

Samson Dome (a) shows features that might relate to carbonate build-ups or salt bodies. Similar
 features are not observed on the other two section.

187 **3 Data and Methods**

We used the BG1002 3D seismic reflection volume, which was acquired by CGG in 2010 188 and reprocessed by BG Geophysical Operations in 2013. We retrieved the data from the DISKOS 189 190 database (https://portal.diskos.cgg.com/whereoil-data/). The 3D survey covers an area of c. 1100 km² and is a full-stack depth migrated (PSDM) volume that was generated using a Full-Waveform 191 192 inversion (FWI) velocity model (Jones et al., 2013). The survey in-lines trend NNW, whereas the cross-lines trend ENE. The survey was acquired with 10 streamer arrays, each with 480 groups 193 194 and a 6 km cable length. The shot and group intervals were 25 m and 12.5 m, respectively. The total recording time was 5050 ms two-way time, and the data were processed with a 4 ms sampling 195 interval and a zero-phase wavelet. We displayed the data using the reverse SEG polarity 196 convention, where a downward increase in acoustic impedance is represented by a trough 197 198 (coloured blue) and a decrease by a peak (coloured red). The seismic resolution ranges from approximately 10 m to 70 m from depths of 500 to 3500 km, an estimate based on half the seismic 199 wavelength of the PSDM volume. A full processing report can be accessed from the DISKOS 200 201 database (https://portal.diskos.cgg.com/whereoil-data/) by searching for "BG1002 3D." The final PSDM volume generated using the FWI velocity model runs from 0 m to 6000 m and is sampled 202 every 10 m (Jones et al., 2013). We also used data from wellbore 7224/7-1 that drilled a total depth 203 of c. 3100 m. This wellbore provided age and lithological constraints down to Early Triassic 204 (Olenekian) strata (https://factpages.npd.no/en/wellbore/PageView/Exploration/All/1245). 205

Our seismic interpretation included horizon and fault mapping, and seismic attribute analysis (see below) that was performed using industry standard software. We interpreted a total of 16 horizons throughout the PSDM volume, the ages of which were constrained by well-log and biostratigraphic data from wellbore 7224/7-1 (Figure 3). We initially used a systematic grid interpretation with 32-line spacing (375 m), which we later used as an input for 3D auto-tracking.

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In areas where 3D auto-tracking was not possible or did not perform well due to poor data quality, we conducted manual interpretation every line (12.5 m). We used the interpreted horizons in the geometric and kinematic analysis of the fault network. We also used seismic attributes, namely variance, which helps to reveal the geometry of the fault network by highlighting discontinuities in the seismic signal (Figure 4; e.g., Randen et al., 2001).

The 3D nature of the PSDM volume allowed us to conduct a quantitative analysis of the 216 217 fault structure and displacement using arbitrary seismic lines oriented normal to local strike (e.g., displacement-length (T-x) plots; e.g., Cartwright et al., 1995; Jackson et al., 2017). We also 218 produced depth-structure and thickness maps to analyse (plan-view) fault geometries and 219 220 kinematics, respectively (e.g., Childs et al., 2003; Jackson & Rotevatn, 2013; Childs et al., 2017; Jackson et al., 2017). We also created (displacement) strike-projections (e.g., Walsh and 221 Watterson, 1991 and Alghuraybi et al., 2022) and calculated fault aspect ratios (i.e., fault trace 222 length divided by fault maximum height, e.g., Nicol et al., 1996; Alghuraybi et al., 2023a). These 223 various methods helped us describe the geometry and growth history of the fault network (see 224 review by Jackson et al., 2017). 225

226 Finally, we calculated SGR for each fault to investigate any potential relationships between fault geometry, growth history, and sealing properties (as expressed by SGR). To constrain the 227 clay content of the host sequence we used data from wellbore 7224/7-1, including Gamma Ray 228 (GR) well-log data and lithology descriptions from cuttings. However, we did not have access to 229 a volume of shale (V_{shale}) log or any dynamic data to QC or calibrate our calculated SGR results. 230 We therefore estimated the V_{shale} from GR log data using two approaches. We first calculated a 231 linear V_{shale} (see Appendix 1, Eq. 1) and secondly, we calculated a non-linear V_{shale} log using the 232 Clavier et al. (1971) equation (Appendix 1, Eq. 2). These equations only approximate the actual 233

formation V_{shale} as it is derived from GR data, which can be affected by the presence of radioactive 234 minerals, matrix density, and clay type (e.g., Clavier et al., 1971). Therefore, any derived 235 properties used here based on the V_{shale} calculations (i.e., SGR) should only be taken as an 236 approximation of the fault/rock property and not an exact measurement. There is significant 237 uncertainty in our fault seal analysis that arises from the limited wellbore control in the area. In 238 239 detail, the fact that we have only one wellbore that provided age and lithology control means that we are not accounting for any lateral facies or lithological changes across stratigraphic intervals. 240 We generated SGR strike-projections (similar to the displacement strike-projections mentioned 241 above) to analyse the 3D distribution of SGR values across fault planes. Additionally, we plotted 242 SGR values along the length of the fault across key stratigraphic levels to highlight any variability 243 in fault seal potential along the fault. Lastly, we used displacement backstripping to calculate SGR 244 through time and note any changes in fault seal potential through time (see review by Jackson et 245 al. 2017). 246

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Figure 3. Stratigraphic column for SW Barents Sea showing major tectonic events. The figure shows the major seismic horizons picked in the area and near the well location.





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Figure 4. A representative seismic section illustrating the structural and stratigraphic elements in the study area. (a) Un-interpreted section. (b) Interpreted section including the location and extent of wellbore 7224/7-1. (c) a variance slice along the H3 horizon highlighting the fault

network as seismic discontinuities in white. The seismic section is indicated by the yellow

260 dashed line and the blue and white star shows the location of the wellbore.

261 **4 Results**

262 4.1 Strucutral elements

The Samson Dome is elliptical in map view, with a N-trending major radius of c. 13 km 263 and an E-trending minor radius of c. 9 km (Figure 5) The folding appears prominent across all 264 265 strata between the Lower Cretaceous and the acoustic basement (Figure 4). Despite the poor seismic quality at depths >2 km, below the main, seismically well-imaged part of the dome, we do 266 not see an obvious salt body (Figure 2). Instead, we observed moderately reflective, mounded 267 bodies that may be the seismic expression of Late Carboniferous to Early Permian carbonate 268 269 mounds (Figures 2 and 3) (e.g., Ahlborn et al., 2014; Di Lucia et al., 2017; Elvebakk et al., 2002). The interpretation of carbonate mounds is consistent with regional observations in the SW Barents 270 Sea and is constraint by seismic reflection and wellbore data regionally (e.g., Larssen et al., 2002; 271 272 Rafaelsen et al., 2008; Di Lucia et al., 2017; Hassaan et al., 2020). The top of the Samson Dome exhibits radial faulting, whereas the study area generally appears to be dominated by NW-striking 273 and EW- to WNW-striking faults (Figures 4 and 5). In detail, we mapped 48 faults and identified 274 four main groups of normal faults. 275

The first fault group is restricted to the Cretaceous interval (Figures 6. a. i and 7. a, b) and is largely restricted to the dome flanks (Figure 6. a. i). The displacement on these faults is typically <50 m, with individual faults being 0.5-4 km long (Figure 7. a, b). South of the Samson Dome, the faults exhibit a more consistent strike direction of E-W to ESE-WNW and generally dip southward, whereas north of the structure they are more polygonal (i.e., they show no preferred strike or dip direction; Figures 6. a. i and 7. a, b).



Figure 5. Depth structure maps for horizons H3 (a), H7 (b), H10 (c) and H14 (d).

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The second fault group includes 11 faults that offset Upper Triassic to Lower Cretaceous stratigraphy (Figures 6. a. ii, b and 7. b, c). Six faults (F1, F37, F40, F42, F46, and F48). In contrast to the first fault group, these faults are restricted to the NE of the study area, on the crest of the dome. Like the first group, however, they also strike ESE-WNW, but are shorter (<500 m long) F4, F6, F7, and F32) are longer (maximum lengths of c. 8 km) and have higher displacements (c.

291 50 m). (Figures 6. a. ii, b and 7. b, c).

The third fault group consists of 17 faults that offset Middle Triassic to Early Cretaceous 292 stratigraphy, which we sub-divided into two subgroups. The first subgroup includes two WNW-293 striking faults (F2 and F5) located on the NE flank the dome. They have a maximum length and 294 295 displacement of c. 7 km and c. 80 m, respectively (Figures 6. a. ii, b and 7. b, d). The other two faults (F25 and F26) are found on the SW side of the study area towards the flank of the dome, 296 and show a maximum trace length and displacement of c. 6 km and c. 45 m, respectively (Figures 297 6. a. ii, b and 7. e). The second subgroup comprises 13 predominately NW-SE-striking faults that 298 occur on the dome crest (F13, F18, F19, F29, F33, F43, and F44) and flanks (F8, F14, F15, F41, 299 F45, and F47) (Figure 6. a. ii, b). These faults are 0.5-5.5 km long and have displacements of <10-300 85 m (Figure 7. a, e, f, g). 301

The fourth and final fault group is composed of 20 faults that offset Lower Triassic to Lower Cretaceous stratigraphy. These faults strike mainly NW-SE and occur on the NW and SE flanks of the dome (F9, F12, F16, F21, F22, F23, F24, F28, F39) or on the dome apex (F10, F11, F17, F20, F27, F30, F31, F34, F35, F36, F38) (Figure 6. a. ii, b). These faults have variable trace lengths (0.35-12 km), with an average trace length of c. 4 km, and an average maximum displacement of c. 55 m (ranging from c. 5-90 m) (Figure 7).

An alternative way to classify the fault network is to separate the faults according to their strike, dip, and dip direction. By plotting these parameters on a stereonet and rose diagram, we can subdivide the fault network into four subsets (Figure 6. c. i, ii). These four subsets generally correlate well with the fault grouping described above. For instance, subsets 1 and 3 broadly 312 correlate to the second and third fault groups, whereas subsets 2 and 4 capture faults that are part 313 of the fourth fault group. The fourth fault group includes faults that show evidence of growth by 314 vertical linkage (see below) and strikes that are parallel to the present-day maximum horizontal 315 stress direction based on borehole breakout data (Heidbach et al., 2016) (Figure 6. c. iii).

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319 Figure 6. (a) A variance map taken at a depth of 700 m (i) and along the top of Knurr FM – H3 (ii). These two variance slices illustrate the different faulting styles with depth where the shallow 320 slice (i) shows evidence of polygonal faulting in the north to northwest while (ii) highlights the 321 predominance of the NW-trending faulting. (b) Th spatial distribution of the studied fault 322 network colour-coded by evidence of vertical linkage (dark blue), no evidence of vertical linkage 323 with base fault tip along H10 (green) and H12 (red). (c) The spatial distribution of the studied 324 fault network colour-coded by strike orientations (i) and stereonet and rose diagram with the 325 colours corresponding to the ones shown on the map (ii). (iii) strike-orientations of the faults 326 trending NW (purple lines) and the light blue shading shows the direction of the present-day 327 maximum horizontal stress. 328

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Figure 7. Eight seismic sections taken perpendicular to strike of different fault within the

network. The locations of the seismic sections are shown on the accompanying variance map of

H3. These seismic sections try to capture the faulting styles and geometries of the studied fault

- 333 network.
- 334





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By closely examining the geometry of the fault network, we note that the studied faults 338 lack any clear along-strike bends (Figure 6. a) and are generally planar (Figure 7). Nearly all the 339 studied faults (n=44) have their maximum displacement across the Lower Cretaceous to Lower 340 Jurassic intervals (H3 and H7) and show broad, bell-shaped displacement-length profiles (Figure 341 8. a. i, ii). These faults are similar to a suite of faults studied by Alghuraybi et al. (2023a) over the 342 Troms-Finnmark Fault Complex. However, we note some internal variability within the Samson 343 Dome fault network, with evidence of displacement partitioning between faults as shown by 344 345 multiple displacement maxima and local minima across the same stratigraphic level (e.g., F21 in Figure 8. a. i). These multiple displacement maxima could indicate that the faults grew via the 346 linkage of smaller, previously isolated segments (e.g., Cartwright et al., 1995; Childs et al., 2003). 347

The variable nature of the displacement is most clearly seen across the H10 (i.e., Middle 348 Triassic) stratigraphic interval (Figure 8. a. iii). For example, some faults show bi-modal 349 displacement patterns along H10 (that is, two displacement maxima separated by a displacement 350

Displacement across the H14 stratigraphic level (i.e., Lower Triassic) is only present on 354 faults with multiple displacement maxima along different stratigraphic intervals separated by a 355 displacement minimum (i.e., the fourth fault group; Figure 9.b). We infer these faults to have 356 experienced growth via vertical linkage (e.g., Nicol et al., 1996). The displacement-length patterns 357 for these faults across H14 are also variable, being similar to the patterns described across H10 358 (Figure 8. a. iv). However, in terms of their maximum displacement (D_{max}) and length (L_{max}), 359 360 regardless of the structural level of observation (i.e., H10 or H14), the fault network plots within the scatter of a global dataset of normal faults (Lathrop et al., 2022) (Figure 8. c). The network 361 shows a D_{max}/L_{max} scaling relationship between 0.1 and 0.001 (Figure 8. c). Additionally, all faults 362 have an aspect ratio <5, with a mean ratio of 1.6 for the 48 faults we studied in detail (Figure 8. 363 364 d).

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Figure 8. A summary of the geometric properties of 12 fault examples representing the fault

network. (a) Normalised displacement-length for (i) the Early Cretaceous top Knurr FM (H3),

(ii) the Early-Middle Jurassic Stø FM (H7), (iii) the Middle Triassic Kobbe FM, and (iv) the

Early Triassic Havert FM (H14). (b) Spatial distribution of the selected faults across the study

area with each fault highlighted with a different colour. (c) D_{max} - L_{max} plot from a global database

of normal faults (modified after Lathrop et al., 2022). (d) A box plot of the aspect ratio values for

the studied fault network. The fault network has typical aspect ratio values of natural normal

faults with a maximum aspect ratio of 5 and a mean of 1.6.



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Figure 9. Strike-projected displacement distribution along the fault surface. The four fault

examples shown in (a) Fault 2 (a.i), Fault 3 (a.ii), Fault 25 (a.iii), and Fault 26 (a.iv) represent

faults in subset 3. The three fault examples shown in (b) Fault 23 (b.i), Fault 12 (b.ii), and Fault 16 (b.iii) are part of fault subset 4. Three key horizons (H3, H7, and H10) are indicated by light

381 grey lines on the strike-projections (a) & (b).

383 4.3 Fault network kinematics

384 4.3.1 Observations

Our detailed mapping shows that during the Paleozoic, deposition was regionally 385 isopachous, with local thinning across the top of what we interpret as carbonate mounds (Figure 386 387 10. a). The regionally isopachous deposition appears to have continued during the Early and Middle Triassic, with minor, highly localised, across-fault thickening (c. 200 - 250 m) around the 388 crest of the dome along NW-SE-striking faults (Figure 10. b). In contrast, the overlying Upper 389 390 Triassic interval thins towards the dome apex (Figure 10. c. i). The thickness difference between 391 the dome crest and distal flanks is up to 600 m. Whilst this regional, dome-related thickness pattern is most apparent within the Snadd Formation (Figure 10. c. ii-iii), it also occurs in the overlying 392 Fruholmen Formation, which also shows more localised, across-fault thickening (c. 100 - 150 m) 393 associated with NW-SE-striking structures (Figure 10. d). Note that we do not observe any clear 394 395 thickness variations across or along other faults within the network (i.e., those not striking NW-SE) (Figure 10. b - d). 396

In contrast to Triassic strata, whose thickness is primarily related to the Samson Dome and 397 during the latter stages, NW-SE faulting, deposition during the Jurassic appears to be controlled 398 only by faulting, with notable across-fault thickening (c. 100 - 200 m) along most faults within 399 the studied fault network and isopachous deposition elsewhere (Figure 11. a). Fault-controlled 400 deposition likely continued into the Early Cretaceous during deposition of the Knurr Formation, 401 402 which also shows regional isopachous deposition with minor across-fault thickening (up to c. 90 m) in association most of the faults within the network (Figure 11. b). The Late Cretaceous marked 403 a shift in deposition from across-fault thickening to regional isopachous deposition with significant 404

thinning (c. 500 m) towards the crest of the dome (Figure 11. c). The Cenozoic thickness map
 reveals largely isopachous deposition with local evidence of erosion (Figure 11. d).

407

4.3.2 Summary and interpretation

In summary, the overall structure of the Samson Dome developed over multiple phases 408 following deposition of Late Carboniferous to Early Permian carbonate mounds. First, some of the 409 NW-SE-striking faults initiated during the Early to Middle Triassic, as is evident by the 410 development of depocenters along segments of these faults (Figures 10. B and 12. b). However, 411 most of the faults within the fault network developed during the Middle Jurassic to Early 412 Cretaceous, with faults that initiated earlier, in the Triassic, remaining active (Figures 10. d; 11. 413 a-b and 12. c-f). The cessation of across-fault thickening and the position of the fault upper tips 414 indicate that faulting stopped during the Late Cretaceous (Figure 11. b; see also Appendix 2 for 415 backstripped displacement vs. length profiles). 416

In contrast to faulting, which commenced in the Early to Middle Triassic, growth of the 417 Samson Dome appears to have initiated in the Late Triassic, during the deposition of the Snadd 418 Formation (Figures 10. C and 12. c). We do not observe any evidence of doming during the 419 Jurassic and Early Cretaceous (Figures 11. a-b and 12. e, f), although thinning and onlap of age-420 related strata towards the dome suggest resurgence of the dome during the Late Cretaceous, after 421 a period of quiescence (Figures 11. C and 12. g). Therefore, by combining all of the observations 422 presented above, we propose that the structures at the Samson Dome area developed through two 423 phases of faulting in the Late Triassic and Middle Jurassic - Early Cretaceous and two phases of 424 doming that occurred during the Late Triassic and Late Cretaceous. No clear faulting or dome 425 growth occurred during the Cenozoic (Figure 12. h). 426



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Figure 10. Isopach (thickness) maps for the Paleozoic interval between the interpreted acoustic

- 432 basement (H15) and the top Havert FM (H14) (a), Early Triassic interval between the Havert
- 433 (H14) and Klappmyss FM (H12) (b.i), Middle Triassic interval between Klappmyss FM (H12)
- 434 and Kobbe FM (H10) (b.ii), Late Triassic interval between Kobbe FM (H10) and Fruholmen FM
- 435 (H8) (c.i). The Late Triassic interval is divided into two sections a lower interval between Kobbe
- 436 FM (H10) and Snadd FM (H9) and an upper interval between Snadd FM (H9) and Fruholmen
- 437 FM (H8). A schematic interpretation of each interval is shown in (a.ii), (a.iii), (c.iii) and (d.ii).
- 438 These isopachs show temporal variations in depositional styles across these time intervals.
- 439



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Figure 11. Isopach (thickness) maps for the Middle Jurassic interval between Top Fruholmen
FM (H8) and Top Stø FM (H7) (a.i), Late Jurassic interval between Top Stø FM (H7) and Top
Hekkingen FM (H4) (a.ii), Early Cretaceous interval between Top Hekkingen FM (H4) and Top
Knurr FM (H3) (b.i), Late Cretaceous interval between Top Knurr FM (H3) and Top Kolmule

447 FM (H2) (c.i), and the Cenozoic interval between Top Kolmule FM (H2) and present-day seabed

448 surface (seabed) (D.i). A schematic interpretation of each interval is shown in (a.iii), (b.ii), (c.ii)

and (d.ii). These isopachs show temporal variations in depositional styles across these time

450 intervals.

Figure 12. A proposed structural evolution model of the Samson Dome. The model includes (a)
carbonates build-up, (b) first phase of faulting, (c) first phase of doming, (d) second phase of
faulting, (e) third phase of faulting, (f) second phase of doming and (g) isopachous deposition.

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459 4.4 Fault seal analysis

Having documented the geometry and kinematic evolution of the Samson Dome and related fault network, we calculated SGR values for the latter using their present-day displacement and V_{shale} values estimated from GR logs in wellbore 7224/7-1 (see methods). We then used strikeprojections to visualise 3D variations in SGR across the fault surface. Therefore, the only factors that influence our SGR calculations are displacement variations within the fault network and any seismically defined changes in stratal thickness. However, in this study we focus on potential long-

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term (i.e., geological timescale) changes in fault seal caused by fault growth history and not
 relatively shorter-term (i.e., production timescale) changes in the probability of cross-fault flow.

We used seven examples to demonstrate the potential impact of fault growth history on 468 temporal changes in sealing characteristics and likelihood. These seven examples were divided 469 into two groups. The first group includes four faults that show a single displacement maximum 470 and do not display any evidence of growth via vertical linkage (F2, F3, F25 and F26; Figure 9. a). 471 472 In contrast, the second group of three faults exhibits multiple displacement maxima downdip of the fault, possibly indicating growth by vertical linkage (F23, F12, and F16; Figure 9. b). Our 473 results show that, regardless of total displacement and/or evidence of growth via vertical linkage, 474 the interval between H3 and H7 has the highest SGR values (SGR >> 20%) for all seven studied 475 faults, whereas the lowest SGR values (SGR $\leq 20\%$) are found in the Middle to Lower Jurassic 476 interval (between H7 and H8; Figure 13). The Upper to Middle Triassic section shows SGR values 477 of c. 20% (interval between H8 and H10; Figure 13). Below H10, the average SGR value is < 20% 478 479 across all seven faults (Figure 13). Closer to the lower tip of the fault near H14, we notice a slight increase (to c. 20%) in SGR values (Figure 13). 480

481 By comparing the SGR strike-projections of the two groups, we can see that faults in the first (Figure 13. a) have higher overall values than those in the second (Figure 13. b). The 482 differences between the two fault groups are evident by the colours of the strike projections, with 483 faults in the first (Figure 13. a) showing darker colours (SGR >> 20%) than those in the second 484 (SGR << 20%) for a corresponding stratigraphic interval. This visual comparison provides only 485 qualitative evidence for the variability of SGR values between the two fault groups, thus, to further 486 investigate and quantify the differences in fault seal potential between the two fault groups, we 487 analysed SGR values plotted along the length of each fault along the Middle Jurassic Stø 488

Formation (H7) and Middle Triassic Kobbe Formation (H10). The first four examples (F2, F3, F25 and F26) do not reach or offset Middle Triassic strata, so we only show the SGR analysis along the Middle Jurassic Stø Formation. In contrast, the other three examples (F23, F12 and F16) offset both Middle Triassic and Middle Jurassic rocks. The reason we focus on the Stø and Kobbe formations is because these sandstone-dominated intervals are considered potential CO₂ reservoirs (NPD, 2023).

495

496 **Figure 13.** Strike-projected SGR value distribution along the fault surface. The four fault

497 examples shown in (a) (a.i) Fault 2, (a.ii) Fault 3, (a.iii) Fault 25, and (a.iv) Fault 26 represent

faults with no evidence of vertical linkage. The three fault examples shown in (b) (b.i) Fault 23,

(b.ii) Fault 12, and (b.iii) Fault 16 represent faults with evidence of vertical linkage. Three key

horizons (H3, H7, and H10) are indicated by light grey lines on the strike-projections (a) & (b).

All four faults (F2, F3, F25 and F26) are likely presently leaking (i.e., SGR << 20%) along 502 their entire lengths (Figure 14. a. i - d. i). The only exception to this is near the faults' lateral tips 503 where displacement approaches zero and SGR values increase (Figure 14. a. i - d. i). However, by 504 examining SGR during the Middle Jurassic (i.e., by using the backstripped displacement), we 505 notice increased lateral variability in fault seal potential between faults and within individual faults 506 507 (Figure 14. a. ii - d. ii). For example, only half of F2 was likely leaking at this time, with the remaining portion of the fault being likely sealing or exhibiting SGR values transitional between 508 20 and 40% (Figure 14. a. i, a. ii). In contrast, F26 was likely sealing during the Middle Jurassic 509 along its entire length, compared to being entirely leaking at present-day (Figure 14. d. i, d. ii). 510

511 Similar to the first four fault examples, the second group of faults (F23, F12, and F16) are all likely presently leaking (i.e., SGR $\leq 20\%$) across the Middle Jurassic Stø Formation, along 512 their entire lengths (Figure 15. a. i - c. i). However, unlike the first group, the second group (F23, 513 F12, and F16) appear to be presently sealing across the Middle Triassic Kobbe Formation (Figure 514 15. a. iii – c. iii). The sealing potential of these fault (F23, F12, and F16) also varied through time, 515 with faults having leaking and sealing patches along length of the fault in the Middle Jurassic 516 (Figure 15. a. ii - c. ii) and being likely entirely sealing during the Middle Triassic (Figure 15. a. 517 iv - c. iv). 518

As mentioned earlier in the methods section (3), we calculated V_{shale} using linear and nonlinear equations (Appendix. 1). The differences in present-day SGR value arising from these two methods were comparable (i.e., SGR values fell within the same range of sealing or leaking potential; Figures 14. a. i - d. I and 15. a. i - c. i; 15. a. iii-c. iii). However, these two V_{shale} estimation methods resulted in large variabilities in the backstripped SGR values as captured by the uncertainty range in our results (Figures 14 and 15). We note that the absolute SGR values are 525 perhaps less relevant to our study than the overall variations in SGR behaviour between the 526 presented fault examples given the limited wellbore control in the area.

Likely sealing Likely leaking Sealing/leaking transition 🛛 Uncertainty range 👾 Upper V_{shale} Linear V_{shale} Non-linear V_{shale}

527

Figure 14. Sealing potential along strike of Fault 2 (a), Fault 3 (b), Fault 25 (c), and Fault 26 (d)

as expressed by shale gouge ratios (SGR) at present-day (i) and at time of deposition (ii) using

backstripped displacement. The SGR values are estimated using the linear and non-linear V_{shale}

equations to represent the uncertainty in calculating SGR values. We use common cut-offs for

532 SGR to represent sealing (SGR > 40), leaking (SGR < 20) potential.

Figure 15. Sealing potential along strike of Fault 23 (a), Fault 12 (b), and Fault 16 (c) as
expressed by shale gouge ratios (SGR) at present-day along the H7 (Top Stø FM) structural level

(i), at time of deposition of Stø FM during Middle Jurassic (ii) using backstripped displacement,

at present-day along the H10 (Top Kobbe FM) structural level (iii), at time of deposition of

Kobbe FM during Middle Triassic (iv) using backstripped displacement. The SGR values are
 estimated using the linear and non-linear V_{shale} equations to represent the uncertainty in

- estimated using the linear and non-linear V_{shale} equations to represent the uncertainty in calculating SGR values. We use common cut-offs for SGR to represent sealing (SGR > 40),
- $_{540}$ leaking (SGR < 20) potential.
- 541

The way normal faults grow shapes their geometry and internal structure, which control 542 fault rock distribution and juxtaposition relationships, and thus influence the fluid flow properties 543 of fault zones (Yielding et al., 1997; Knipe et al., 1998; Walsh et al., 1998;). Therefore, accurately 544 modelling how faults grow and how they develop through time can significantly contribute 545 towards increasing the predictability of basin models during the exploration stages. It can also 546 improve site characterisation of water aquifers, CO₂ storage and nuclear waste disposal locations, 547 geothermal fields (e.g., Main, 1996; Allmendinger et al., 2000; Sorkhabi & Tsuji, 2005; Rutqvist, 548 549 2012; Kaldi et al., 2013; Brune et al., 2017).

550 5.1 Fault network kinematics and sealing potential

In this study, we focused on the faulting style and development of the fault network in the 551 Samson Dome area, offshore Norway. Overall, we identify four groups of faults. These include, i) 552 faults restricted to the Cretaceous interval (Figures 6. a. i and 7. a, b), ii) faults offsetting Upper 553 Triassic to Lower Cretaceous stratigraphy (Figures 6. a. ii, b and 7. b, c), iii) faults offsetting 554 Middle Triassic to Lower Cretaceous stratigraphy (Figures 6. a. ii, b and 7. b, d, e) and iv) faults 555 offsetting Lower Triassic to Lower Cretaceous (Figures 6. a. ii, b and 7). Isopach analysis indicates 556 557 that faulting in the Samson Dome developed through multiple phases. The first phase of faulting initiated during the Early to Middle Triassic and formed several NW-SE-striking faults (Figure 10. 558 b). Additionally, we see clear evidence of across-fault thickening by the Late Triassic (Figure 10. 559 d), indicating that the NW-SE-striking faults were active at that time. The second phase of faulting 560 started during the Middle Jurassic and continued into the Early Cretaceous, forming most of the 561

faults in the Samson Dome area (Figure 11. a-b). The multiple phases of fault growth in the Samson
Dome are generally consistent with the model proposed by Mattos et al. (2016), where they
interpreted an initial phase of faulting in the Middle to Late Triassic and a second phase of faulting
to have occurred in the Early to Late Cretaceous.

Another important observation we make is the variable ways in which faults grew. For 566 example, some faults grew via the vertical linkage of segments, something that is defined in our 567 568 strike-projections by sub-horizontal bands of low displacement (Figure 9). The differences in vertical linkage could be related to the fact some faults experienced reactivation and others did 569 not. Specifically, our results show that faults that experienced vertical linkage represent faults in 570 subset 4 (Figure 6. c), which strike NW-SE and were active since the Early to Middle Triassic, and 571 that were subsequently reactivated during the Late Triassic and Middle Jurassic to Early 572 Cretaceous. The other fault population that shows no evidence of vertical linkage (subset 3; Figure 573 9. a) corresponds to faults that initiated during the Middle Jurassic to Early Cretaceous (Figure 11. 574 a-b) and that tend to strike almost E-W (Figure 6. c). North Atlantic Ocean opening-related fault 575 reactivation has been proposed by previous studies (e.g., Faleide et al., 1993; 2008; Mattos et al., 576 2016; Figure 12). However, the opening of the North Atlantic Ocean occurred in the Paleocene – 577 Eocene (e.g., Faleide et al., 2008), some time after the main, Middle Jurassic to Early Cretaceous 578 579 phase of faulting we document here (Figures 11 and 12). Our results are, however, consistent with 580 those of Mattos et al. (2016), who also concluded faulting commenced in the Late Triassic (Figures 10 and 12), noting that fault initiation might have occurred earlier, during the Early to Middle 581 582 Triassic (Figure 10. b). The main difference between our results and those of Mattos et al. (2016) is that we show the main period of faulting occurred during the Middle Jurassic to Early 583 Cretaceous, in response to the main rifting phase documented elsewhere in the SW Barents Sea 584

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The value of considering the fault growth history in assessing the potential of faults to seal 588 or leak fluids has recently received increased interest (e.g., Reilly et al., 2017; Song et al., 2020; 589 Michie & Braathen, 2023). While assessing present-day juxtaposition relationships and fault 590 591 sealing potential is a valuable approach when considering reservoir behaviour over relatively short (i.e., a few tens of years) production timescales, it does not capture the changes that may occur to 592 fault hydraulic properties over longer, geological timescales (e.g., Manzocchi et al., 2010; Reilly 593 594 et al., 2017). For example, Reilly et al. (2017) show that, despite being likely sealing at present, the Cape Egmont Fault, offshore New Zealand experienced sand-against-sand juxtaposition and 595 was likely leaking at an earlier phase of its development. Our fault seal analysis results reveal that 596 relative differences in SGR values correlate with fault growth patterns. Specifically, we show that 597 the faults associated with vertical linkage and possible reactivation show lower SGR values on 598 average, indicating that they are more likely to leak than other faults within the network (Figure 599 13). The reason for this is that SGR is inversely proportional to fault throw (Fristad et al., 1997; 600 601 Yielding et al., 1997; Freeman et al., 1998). Hence, if we keep the thickness of the faulted strata 602 and their V_{shale} content constant, faults with more throw (or displacement) will have lower SGR values. In fact, these reactivated faults (subset 4) are predominately found at the crest of the 603 Samson Dome (Figure 6. c.i), above which there is seismic reflection evidence for fluid leakage 604 605 and the accumulation of shallow gas (e.g., Vadakkepuliyambatta et al., 2013). In contrast, the faults that appear *not* to have been reactivated and that do *not* show evidence of vertical linkage are 606 characterised by higher SGR values relative to the other faults (Figure 13). In fact, fault 607

reactivation is one of the factors that is thought to influence the fault seal potential (e.g., Bailey et 608 al., 2006), which might explain the variability we observe here. 609

V_{shale} calculation methods have major implications for fault seal analysis algorithms since 610 differences in those methods will cause inconsistency in the fault seal analysis results (e.g., 611 Yielding, 2002). Our results show that the variations between V_{shale} calculation methods are more 612 significant for low displacement (i.e., < 20 m) faults. The Middle Jurassic and Middle Triassic 613 614 SGR values for our fault examples demonstrate that depending on method used, a given fault may be potentially sealing or leaking (Figures 14 and 15). In contrast, for faults with larger 615 displacement (i.e., > 20 m), the variations between V_{shale} calculation methods are less significant 616 on the resulting SGR value and inferred sealing potential. These results likely reflect the 617 uncertainty and poorly constrained nature of the V_{shale} calculations. Higher displacement faults are 618 expected to have wider fault damage zones (e.g., Torabi et al., 2020). However, the lithology of 619 the faulted rocks has been shown to have an impact on the width, mechanical and petrophysical 620 621 properties of their related fault zone (e.g., Torabi et al., 2020). Therefore, in our study, higher displacement could be associated with higher or lower V_{shale} value (i.e., higher, or lower SGR 622 values). In the absence of better constraints on V_{shale} calculations, our SGR results are thus largely 623 driven by changes in displacement and are less sensitive to changes in lithology or V_{shale} values. 624 625 Hence, the SGR results presented here are highly uncertain and need to be better constrained by additional wellbore and lithology control. 626

627

5.2 Timing and origin of the Samson Dome

Previous studies of the Samson Dome use gravity data, 2D and 3D time-migrated seismic 628 reflection data to infer that it is underlain by a several km-thick, at least 12 km-wide, lenticular 629 body of Permian salt (e.g., Breivik et al., 1995; Mattos et al., 2016). This body has a moderately 630

convex-up upper surface, which is parallel to overlying strata within the dome, whereas its basal 631 surface is convex-down, and seemingly discordant and possibly even displaying an erosional 632 relationship with underlying and adjacent strata. In terms of the timing and evolution of the 633 structure, Mattos et al. (2016) suggest that the (moderately well-layered, low-amplitude) Permian 634 salt body was isolated, being somehow encased in (well-layered, high-amplitude) carbonate 635 636 platform deposits, and that the main period of salt mobilisation in the Early-to-Late Cretaceous resulted from the reactivation of basement faults (not shown in their Figure 13) and to a lesser 637 degree, differential loading imposed by NW-prograding shelf deposits. Similar pockets of low-638 amplitude seismic facies, flanked by high-amplitude reflections, are observed in seismic reflection 639 data northeast of the Samson Dome near the Swaen Graben (Figure 2, X to X'), where they which 640 might document salt pillows or remnant salt deposits related to salt withdrawal (e.g., Jackson & 641 Talbot, 1986; Nilsen et al., 1995; Kovi, 1998; Jackson & Hudec, 2005; Hudec & Jackson, 2007; 642 Pichel et al., 2018; Rojo et al., 2019). Mattos et al. (2016) then suggest that subsequent deflation 643 644 and collapse of the dome, during the latest Cretaceous and following a period of quiescence, was driven by a combination of salt dissolution and sediment loading. The preferred model in the 645 literature for the formation of the domes in the SW Barents involves contractional buckling 646 647 triggered by tectonic inversion during the Late Triassic (Hassaan et al., 2020).

Our study suggest that such a salt body is absent below the Samson Dome, at least based on the apparent lack of a chaotic, poorly reflective seismic facies that might characterise salt that if not diapiric, had at least flowed to inflate beneath and elevate the overburden (e.g., Clark et al., 1998; Jackson & Lewis, 2012; Jones & Davison, 2014; Hassaan et al., 2021). The top of the inferred salt body is convex-up and parallel to overlying strata, consistent with the interpretation of Mattos et al. (2016). However, we cannot infer the geometry of the basal surface of the salt

body because of the poor data quality and imaging extent at the depth of the Late Permian - Early 654 Triassic strata (c. 5.5 - 6 km depth). Instead, we observe a low-amplitude, moderately well-layered 655 seismic facies (consistent with Figures 5 and 6 in Mattos et al. 2016), which may be the 656 geophysical expression of layered evaporite and carbonate, rather than pure halite (e.g., Fiduk & 657 Rowan, 2012; Rowan et al., 2019). Indeed, Hassaan et al. (2020) document potentially similar, 658 659 halite-poor, layered evaporite bodies at the core of several geometrically similar domes in the SW Barents Sea (e.g., Happet dome, Alpha Dome, and Veslekari Dome). However, the basal surfaces 660 of those bodies are flat and not convex-down (Hassaan et al., 2020). The same flat basal surface 661 is often associated with other known salt bodies like the Permian salt deposits in the North Sea 662 (e.g., Clark et al., 1998; Davison et al., 2000; Jackson & Lewis, 2012; Tvedt et al., 2016; Hansen 663 et al., 2021). Even if sufficient halite (the most mobile lithology within layered evaporite 664 sequences; e.g., Jackson et al., 2014, 2019; Rowan et al. 2020) was present in the Permian deposits 665 beneath the Samson Dome, there are some geometric and more critically, mechanical challenges 666 to this model. First, it is unclear why salt might spontaneously flow and inflate, having not moved 667 between the Late Carboniferous/Early Permian and Early Cretaceous (i.e., the time between salt 668 deposition and the main stage of halokinesis proposed by Mattos et al., 2016). This would imply 669 670 that for 150 Myr, in a tectonically active basin characterised by complex patterns of sediment dispersal (e.g., Faleide et al., 1984; 2008; 2015; Gabrielsen, 1984; Gabrielsen et al., 2016), the salt 671 672 was not subjected to any gravitational instabilities driven by tilting (i.e., gravity) or differential 673 loading. Related to this, it is highly unlikely that pure salt inflation, in the absence of coeval saltdetached shortening (see below), might lift a very thick (4-5 km), rigid overburden, such as that 674 characterised by the Triassic to Cretaceous succession. Active diapirism, of the type described by 675 676 Mattos et al. (2016) can only occur below a relatively thin (≤ 500 m) overburden, where salt can

rise buoyantly due to it being less dense than its overburden. Salt flow below a thick overburden 677 is however possible during horizontal compression and shortening, with salt anticlines and pillows 678 forming in the core of buckle folds (e.g., Hudec & Jackson, 2007; Dooley et al., 2009). However, 679 in the case of the Samson Dome it is not clear which regional tectonic events might have driven 680 shortening during the Late Triassic or Late Cretaceous, nor why shortening might lead to formation 681 682 of a sub-circular structure (i.e., a salt pillow), when more commonly elongate structures form (i.e., salt anticlines) (e.g., Hudec & Jackson, 2007; Dooley et al., 2009). Likewise, progradational 683 loading, of the type suggested by Mattos et al. (2016), would likely lead to the formation of 684 elongate rather than sub-circular structures (e.g., Precaspian; e.g., Fernandez et al., 2017; Jackson 685 et al., 2020; Paradox Basin; e.g., Trudgill, 2011; Santos Basin; e.g., Fiduk & Rowan, 2012). It is 686 also important to note that salt-detached shortening and progradational loading are most commonly 687 associated with the formation of salt structures (anticlines or pillows) that have flat rather than 688 convex-down, seemingly erosional base of the type described by Mattos et al. (2016). It is unclear 689 690 how slip or pre-existing relief along linear basement-involved, sub-salt faults (Mattos et al. 2016) would trigger the differential flow of overlying salt and the formation of a sub-circular structure 691 such as a salt pillow. Finally, we see no evidence in the seismic reflection data for dissolution of 692 693 salt underlying the Samson Dome. It is also unclear how sufficient volumes of NaClundersaturated (i.e., meteoric-water might percolate downwards through a very thick (c. 5 km) 694 695 overburden to dissolve the salt (see discussion by Jackson and Lewis, 2013). As such, rather than 696 dissolution, we infer that the supra-salt faults record regional extension and perhaps to a lesser degree, outer-arc bending above the growing dome. 697

In terms of the timing of formation of the Samson Dome, current models describe the mainphase of salt mobilisation to have occurred in the Early to Late Cretaceous, as evidenced by broad

anticlines deforming Late Cretaceous strata (e.g., Mattos et al., 2016). However, our thickness
maps clearly show thinning from as early as the Late Triassic (Figure 10. c), which suggests that
the doming or formation of the anticline began much earlier, during the Late Triassic.

We propose an alternative model for the timing of development of the Samson Dome. More 703 specifically, our results support an earlier timing (i.e., Late Triassic) of dome initiation than 704 previously proposed (i.e., Late Mesozoic or Early to Late Cretaceous; Mattos et al., 2016). Our 705 706 Late Triassic timing of dome initiation is consistent with regional observations of the formation of similar structures in the SW Barents Sea (e.g., Hassaan et al., 2020). However, the mechanism or 707 mechanisms driving dome formation remain ambiguous. Our observations suggest that the Samson 708 709 Dome formed through two phases involving a possible ductile layer that caused the inflation of the observed anticline (Figures 10; 11; 12). The nature of the ductile material remains unknown 710 but might consist of a layered evaporite sequence or stratified salt based on the reported negative 711 gravity anomaly (e.g., Breivik et al., 1995; Barrère et al., 2009). While our results support the Late 712 713 Triassic timing of the dome formation proposed by Hassaan et al. (2020), this model still does not explain the formation of an elliptical dome with radial faults at the crest of the dome. An alternative 714 mechanism that could lead to the development of a dome with a similar geometry to the Samson 715 716 Dome is what is known as a salt-wing intrusion (sensu Hudec & Jackson, 2006). Salt-wing 717 intrusion involves the placement of a salt wedge along a stratigraphic layer or a thin bed in the 718 adjacent rocks (e.g., Hudec & Jackson, 2006). As the sheet inflates, the roof is lifted, often resulting in the onlap or erosion of the roof strata (e.g., Hudec & Jackson, 2006). This process accounts for 719 720 the observed onlapping in the Late Triassic and Early Cretaceous along with the Upper Cretaceous - Base Eocene unconformity (Figure 12). However, it is unlikely that a salt-wing intrusion led to 721 the formation of the Samson Dome because of the absence or lack of evidence for a deep thick 722

autochthonous salt layer. Therefore, whilst our results provide an updated timing of the SamsonDome formation, further research into the genesis of the dome is needed.

725 6 Conclusions

We examine the faulting style, development, and sealing potential of the fault network in 726 the Samson Dome area, SW Barents Sea. The fault network has four groups, each offsetting 727 different stratigraphy from the Cretaceous, Late Triassic-Early Cretaceous, Middle Triassic-Early 728 Cretaceous, and Early Triassic-Early Cretaceous. We propose the fault network evolved through 729 multiple phases. NW-SE-striking faults likely initiated during the Early to Middle Triassic, but 730 establishing their timing and initiation is challenging because of data limitations. Most faults 731 developed during the Middle Jurassic to Early Cretaceous, with Triassic faults remaining active. 732 Faulting ceased in the Late Cretaceous, while doming likely began in the Late Triassic, with 733 evidence of thinning and onlapping patterns. Our observations suggest the structures in the Samson 734 Dome area resulted from two faulting (Late Triassic and Middle Jurassic to Early Cretaceous) and 735 two doming (Late Triassic and Late Cretaceous) phases. However, existing halokinetic models do 736 not account for the absence of salt tectonic structures, indicating the need for further research on 737 dome genesis. Fault seal analysis reveals that SGR values correlate with fault growth patterns. 738 Faults associated with vertical linkage and potential reactivation show lower SGR values, 739 suggesting a higher likelihood of leakage compared to other faults. Faults without evidence of 740 reactivation or vertical linkage display higher SGR values, indicating better sealing potential. 741 742 Variations in V_{shale} calculation methods have significant implications for fault seal analysis, particularly for low displacement faults. In summary, our study provides insights into faulting, 743 development, and sealing potential in the Samson Dome area. It proposes an updated timing for 744 dome formation and highlights the need for further research into the driving mechanism behind 745

the development of the Samson Dome. Consideration of V_{shale} calculation methods is crucial for reliable fault seal analysis.

748

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756 **Open Research**

- The seismic and wellbore data are openly available in the Norwegian national data repository for
- 758 petroleum data at https://portal.diskos.cgg.com/whereoil-data/.

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Appendix 9. Volume of shale (Vsh) estimations from Gamma Ray (GR) log responses.

1300 Depth 750 Z

Appendix 11. IL and XL seismic lines crossing the Samson Dome that show the full depth coverage of the seismic survey used in Chapters 4 and 5.