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# 11 **Evidence of subduction, collision, and** 12 **extension in northern Borneo: Constraints from** 13 **receiver functions** 14

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## 29 **Highlights**

- 30 • A new Vs model for northern Borneo is obtained from P-receiver  
31 functions jointly inverted with a global surface wave model.
- 32 • Crustal thickness is found to vary from 24km beneath central northern  
33 Borneo, to 60km beneath the Crocker Range.
- 34 • We image Dangerous Grounds crustal material underthrust beneath  
35 western Sabah following subduction and collision.

## 36 **Abstract**

37

38 Northern Borneo (Sabah) has a complex geological history, experiencing  
39 multiple episodes of subduction, magmatism, uplift, subsidence, and  
40 extension since the Mesozoic. This includes the subduction of the proto-South  
41 China Sea beneath Sabah, terminating ~21 Ma; a postulated later phase of  
42 subduction of the Celebes Sea plate, terminating ~9 Ma; extension in central  
43 Sabah ~9-10Ma; rapid emplacement and exhumation of a granite intrusion  
44 ~7Ma, and the development of a fold-thrust belt offshore during the last 5 Myr.  
45 While these events have left imprints in the surface rock record, it has not  
46 been possible, until recently, to investigate deeper lithospheric processes that  
47 have shaped Sabah. The installation of 46 broadband seismometers - the  
48 northern Borneo Orogeny Seismic Survey (nBOSS) - between 2018 and 2020  
49 means that it is now possible to constrain the architecture of the crust and  
50 uppermost mantle beneath Sabah. We use two years of passive seismic data  
51 recorded by the nBOSS network, and an additional 24 Malaysian  
52 Meteorological Service broadband seismometers in Sabah to calculate P-  
53 wave receiver functions. We then use these in a joint inversion with surface  
54 wave data to obtain shear velocity models of crustal structure. The thickest  
55 crust (60km) occurs beneath the Crocker Range, while the thinnest crust  
56 (24km) is found in central Sabah, potentially recording Miocene extension.  
57 The crust beneath the 4095m high Mt Kinabalu is also comparatively thin.  
58 Distinct, low-velocity, dipping anomalies identified in our shear wave velocity  
59 models provide clear evidence for underthrusting of Dangerous Grounds  
60 continental crust following subduction and collision.

61

62 **Plain language summary**

63

64 The processes that happen because of the movement of tectonic plates leave  
65 an imprint on the interior of the Earth, and to fully understand these processes  
66 we need to make images of the subsurface. Over the last 65 Million years, the  
67 northern part of the island of Borneo (Sabah) has been affected by multiple  
68 tectonic events including potentially two episodes of subduction – where one  
69 tectonic plate descends beneath the other - and mountain building due to the  
70 collision between the Dangerous Grounds and Sabah. Until recently there  
71 were insufficient instruments to make the measurements necessary to obtain  
72 images of the interior of the Earth beneath Sabah. In 2018 this situation was  
73 transformed with the installation of 46 seismometers – instruments that detect  
74 earthquakes. We use the records from distant earthquakes recorded by these  
75 and pre-existing seismometers to build up a picture of the Earth ~100km  
76 below the surface of Sabah. Our work reveals that the thickness of the crust  
77 beneath Sabah varies from 24km in central Sabah, to 60km beneath the  
78 Crocker Range. We also image part of the Dangerous Grounds crust now  
79 situated beneath Sabah, emplaced as a result of the collision between these  
80 two tectonic blocks.

81

82

83

## 84 **Introduction**

85

86 Subduction is fundamental to the growth of continents (e.g. Foley et al.,  
87 2002), driving plate motion (Forsyth and Uyeda, 1975), and long-term climate  
88 regulation (e.g. Johnston et al., 2011). Eventually subduction will come to an

89 end (e.g., via continent-continent collision), and this may result in magmatism,  
90 exhumation, rapid uplift, and subsidence (e.g. Zandt et al., 2004, Levander et  
91 al., 2011, Li et al., 2016). The processes occurring in these post-subduction  
92 settings remain, at present, poorly understood. Given that subduction has  
93 been happening on Earth for at least 1.8 Ga (Weller and St Onge, 2017),  
94 explaining post-subduction processes is vital, not just for our understanding of  
95 present-day tectonics, but also for interpreting the deep geological record.

96

97 Northern Borneo is an ideal location for studying post-subduction processes.  
98 It is thought to be the site of two subduction systems that have terminated  
99 since the start of the Neogene: the subduction of the proto-South China Sea  
100 (pSCS) until ~21 Ma (Lai et al., 2021, Hall 2013, Morley and Back, 2008,  
101 Tongkul 1994, Tongkul 1991) along the present-day NW coast of Sabah, and  
102 the subduction of the Celebes Sea along the present-day SE coast of Sabah,  
103 which terminated ~9 Ma (Lai et al., 2021). In this study we use passive  
104 seismic data recorded by a network of broadband seismometers deployed  
105 across the Malaysian state of Sabah, situated on the northern end of the  
106 island of Borneo, between 2018-2020 (Figure 1) to image the crust and  
107 mantle lithosphere to both improve our understanding of the tectonic setting of  
108 northern Borneo and provide new insight into subduction termination and  
109 post-subduction processes.

110

### 111 ***Geology and tectonic setting of Sabah***

112

113 The diverse surface geology of Sabah is testament to the rich range of  
114 tectonic processes that have affected the northern part of Borneo since the

115 Mesozoic. Some of the oldest dated rocks are those of the Segama Valley  
116 Felsic Intrusions (250 and 241 Ma) (Burton-Johnson et al., 2020) in eastern  
117 Sabah, which are intruded into ophiolitic rocks, with some subsequent,  
118 mineralogically distinct, felsic intrusions in the same area being dated at ~178  
119 Ma. It has been proposed (Burton-Johnson et al. 2020, Balaguru and Nichols,  
120 2004) that these formed in an extensional basin in a suprasubduction setting,  
121 which was then uplifted and eroded in the latest Cretaceous or earliest  
122 Paleocene, around 66 Ma (Balaguru and Nichols, 2004). A recent study by  
123 Tian et al., (2024) has dated, using Zircon U-Pb ages, ophiolites in Sabah to  
124 248-244 Ma, which they argue formed in a small extensional setting related to  
125 the subduction of Paleo-Tethys and the start of subduction of the Paleo-  
126 Pacific plate.

127

128 By the Paleocene (66-56Ma) subduction of the proto-South China Sea  
129 towards the south-east beneath what is now the western edge of Sabah had  
130 begun (e.g., Hutchison et al., 2000, Rangin et al., 1999). Cumulate gabbros,  
131 part of the Sabah Ophiolite, in the Tongod-Telupid area have been dated to  
132  $42.65 \pm 0.51$  Ma (Lai et al., 2021), and from their geochemical signature are  
133 thought to have formed in a back-arc basin. Similarly, the geochemistry of the  
134 Sandakan andesitic tuff ( $33.9 \pm 7.7$  Ma, Bergman et al., 2000) could also  
135 suggest a back-arc basin setting (Lai et al., 2021, Hutchison et al., 2000). In  
136 the fore-arc, thick (~9000 m) sedimentary successions of deep marine  
137 sandstones, shales and minor conglomerates, are present in the NE-SW  
138 trending Crocker Basin (Balaguru and Nichols, 2004).

139

140 Opening of the South China Sea (~33-32 Ma, (Franke, 2013, Barckhausen et  
141 al., 2014, Li et al., 2014)), driven by subduction of the proto-South China Sea,  
142 pushed continental slivers, including the Dangerous Grounds and Reed Bank  
143 blocks, towards Borneo (e.g., Tongkul 1991, Tongkul 1994, Hutchison et al.,  
144 2000, Hall 2013, Rangin et al., 1999). Subduction of the proto-South China  
145 Sea continued in the earliest part of the Miocene (24-21 Ma, Lai et al., 2021  
146 and references therein), however, at around 21 Ma the Dangerous Grounds  
147 block collided with and then underthrust northern Borneo (Lai et al., 2021, Hall  
148 2013, Morley and Back, 2008, Tongkul 1994, Tongkul 1991), which ultimately  
149 caused subduction to cease. This collision led to uplift above sea level  
150 (Burton-Johnson et al, 2020, Hall 2013, Morley and Back, 2008). However, by  
151 the end of the Early Miocene, most of Sabah was at or below sea level once  
152 more, with low hills where the Crocker Range is today (Hall 2013, Cottam et  
153 al, 2013).

154

155 Subduction of the Celebes Sea beneath eastern Sabah began at a similar  
156 time to the termination of subduction of the proto-South China Sea and may  
157 have been a result of changes in regional stresses due to the Sabah-  
158 Dangerous Grounds collision (Lai et al., 2021, Linang et al., 2022). In the Dent  
159 Peninsular, rocks that are the product of arc magmatism have been dated to  
160 18.8-17.8 Ma and in the Semporna Peninsula to 18.2-14.4 Ma (Macpherson  
161 et al., 2010). It is, however, important to note that the idea that there was  
162 northwards subduction of the Celebes Sea is contested (Burton-Johnson and  
163 Cullen, 2023). A slab from this subduction event has yet to be imaged in the  
164 mantle.

165

166 Roll-back of the Celebes Sea subduction from 19 Ma led to extension in the  
167 Sulu Sea and in Sabah (Hall, 2013). Thick (~6 km) successions of  
168 carbonates, shallow marine, and fluvio-deltaic sediments, including coals,  
169 were deposited in a basin in Central Sabah (Tongkul and Chang, 2003,  
170 Balaguru and Nichols 2004, Burton-Johnson et al., 2021). The coastal/shelf  
171 environments for all these sediments, including coal that was buried to 3 km  
172 depth (Baluguru and Nichols, 2004), means that subsidence must have  
173 continued over a prolonged period. Tsikouras et al., (2021) argue that the  
174 major increase in extension suggested by Huang (1991) between 9 and 11  
175 Ma led to rifting in Ranau area, and suggest that sea-floor spreading took  
176 place in the Telupid area. This is disputed by Cullen and Burton-Johnson  
177 (2021) who argue that while extension took place, the Sulu Sea rift did not  
178 extend into Sabah. In a recent review Lai et al., (2021) suggest that Celebes  
179 Sea subduction beneath Borneo terminated ~9 Ma.

180

181 The Kinabalu pluton, which forms the 4095 m high Mt Kinabalu, was intruded  
182 into peridotites and the Crocker formation between 7.85 and 7.22 Ma, at a  
183 depth of 3-8 km (Cottam et al., 2013). Between 6.6 and 5.8 Ma it was rapidly  
184 cooled and exhumed with rates of up to 7 mm/yr (Cottam et al., 2013). The  
185 emplacement and exhumation of the Kinabalu pluton likely occurred in an  
186 extensional setting (Hall et al., 2013, Burton-Johnson et al., 2019).

187

188 Sabah only became fully emergent above sea level by the end of the Miocene  
189 to early Pliocene (~5 Ma), and uplift has occurred since (Roberts et al., 2018,

190 Hall 2013, Morley and Black 2008). This includes uplift of the circular basins,  
191 such as the Maliau Basin, in central Sabah (Tongkul and Chang, 2003).  
192 During the Pliocene large-scale gravitational collapse occurred, seen in mass  
193 transport slumps, megalides and extensional faults (Cottam et al., 2013), and  
194 as result of this, a fold and thrust belt has developed offshore of western  
195 Sabah (e.g., Spain et al., 2013, Franke et al., 2008, King et al., 2010). Around  
196 5 Ma a change in the composition of volcanic rocks in eastern Sabah also  
197 occurs from calc-alkaline to a similar composition to ocean island basalts  
198 (OIB), (Macpherson et al., 2010). Volcanism in eastern Sabah has continued  
199 into the Holocene, potentially as recently as 24-27 ka based on radiocarbon  
200 dating of carbonised material (Kirk, 1968; Bellwood, 1988; cited in Tjia et al.,  
201 1992), although Takashima et al., (2004) date the youngest volcanics in their  
202 study using thermo-luminescence to 90 ka.

203

#### 204 ***Previous geophysical work***

205

206 Regional-scale tomographic studies of South-East Asia have observed  
207 anomalously high seismic velocities in the upper mantle beneath Sabah at  
208 depths of ~100-300 km (e.g., Amaru, 2007, Tang and Zheng, 2013, Hall and  
209 Spakman, 2015, Zenonos et al., 2019, Wehner et al., 2022). These high  
210 velocities are attributed to the presence of slab remnants in the upper mantle.  
211 While the earlier body-wave studies (e.g., Amaru, 2007, Hall and Spakman  
212 2015, Zenonos et al., 2019) had limited resolution beneath Sabah, thus  
213 bringing the existence of higher velocities into question, the full-waveform  
214 model of Wehner et al., (2022), SASSY21, uses data from the same dense



215 seismic network in Borneo used in this study, and so has improved resolution  
216 in this region. Other results from this dense seismic network – the nBOSS  
217 network – are described below.

218

219 In a Sabah-focused P- and S-wave tomographic study, also using nBOSS  
220 data, Pilia et al., (2023a), observe two distinct fast velocity anomalies in the  
221 upper mantle beneath Sabah. One, an elongate anomaly at depths >250 km  
222 underlying most of the Crocker Range, is attributed to the proto-South China  
223 Sea Slab, while the other, a relatively narrow (<100 km) anomaly between  
224 ~150 and 300 km depth in central Sabah, is interpreted to be a lithospheric  
225 drip from the volcanic arc root beneath the Semporna Peninsular. Pilia et al.,  
226 (2023b) perform thermo-mechanical modelling and suggest that the  
227 downwelling drip can cause extension and crustal thinning, resulting in  
228 melting and exhumation of sub continental material. As such, the ‘Semporna  
229 drip’ may play an important role in the emplacement of the Kinabalu pluton, as  
230 well as explaining subsidence and uplift, and the lavas with an OIB  
231 composition in eastern Sabah.

232

233 Bacon et al., (2022) investigate anisotropy beneath Sabah using XKS splitting  
234 measurements extracted from nBOSS teleseismic data. Their results  
235 demonstrate that fossil anisotropy in the lithosphere is the main control on  
236 anisotropic properties in this post-subduction setting. They observe fast  
237 directions parallel to the strike of the Crocker range in western Sabah, likely  
238 imparted when the Dangerous Grounds block collided with Sabah. In the east  
239 of Sabah, fast directions are sub-parallel to the direction of spreading in the

240 Sulu Sea, suggesting that the anisotropic fabric may have developed as a  
241 result of extension, while the null results they observe in the southeast may  
242 arise due to the lithospheric drip observed by Pilia et al., (2023a).  
243  
244 Roberts et al., (2018) suggest that removal of the lithosphere and  
245 replacement by hot asthenospheric material could explain the relatively rapid  
246 uplift and erosion rates observed in Sabah (~0.1-0.3 mm/yr, Morely and Back,  
247 2008). They base their estimates of thin lithosphere on regionally extensive  
248 slow shear wave velocities at 100-200 km depth in the global tomographic  
249 model of Schaeffer and Lebedev (2014). However, in recent 2-plane-wave  
250 tomography of Sabah from Greenfield et al., (2022), average lithosphere  
251 thickness beneath Sabah is found to be ~100 km, with the lithosphere only  
252 being thin (<50km) beneath the Semporna Peninsula, consistent with the work  
253 of Pilia et al., (2023b) that suggests that the lithosphere here has dripped off.  
254  
255 Until recently, estimates of crustal thickness in Sabah had been limited. Holt  
256 (1998) modelled gravity data from Sabah and suggested that the whole of  
257 Sabah was underlain by crust >30 km thick, that crustal thicknesses beneath  
258 the Crocker Range were ~50 km, and 39 km beneath central Sabah.  
259 Estimates of  $27\pm 3$  km and  $33\pm 2$  km beneath seismometers KKM and LDM  
260 near Kota Kinabalu and Lahad Datu respectively have been made by Lipke  
261 (2008) from H-k stacking of receiver functions, while Syuhada et al., (2022)  
262 invert receiver functions from KKM and LDM using the Neighbourhood  
263 algorithm and find crustal thicknesses of 30 km and 26 km, respectively. Latiff  
264 and Othman (2020) invert receiver functions from KKM and find a crustal

265 thickness of 40km. A regional crustal thickness map derived from surface  
266 wave data made by Tang and Zheng (2013) estimates crustal thickness  
267 beneath Sabah to be 27.5-32.5 km. The deployment of the nBOSS seismic  
268 network between 2018-2020 has allowed for more detailed studies of crustal  
269 thickness across the whole of Sabah to be conducted. Greenfield et al.  
270 (2022) use the 4.1 km/s velocity contour in their shear wave velocity model as  
271 a proxy for the Moho and suggest that crustal thicknesses vary from 25-55  
272 km, with the thickest crust beneath the Crocker Range and the Dent  
273 Peninsula, and the thinnest crust in north east Sabah. Linang et al. (2022) use  
274 Virtual Deep Seismic Sounding (VDSS) to estimate crustal thickness in the  
275 range 21 - 46 km, with a similar pattern of thicker and thinner crust.

276

277 Until now, due to a lack of seismic instrumentation in the region, it has not  
278 been possible to derive a detailed model of the seismic velocity structure of  
279 Sabah's crust. Consequently, debates have continued to emerge (e.g.,  
280 Milsom et al. 2001, Cullen and Burton-Johnson, 2021) about the nature of the  
281 crust beneath this part of Borneo and the processes that have shaped it. Fone  
282 et al., (2024) use data from the nBOSS network to develop an ambient noise  
283 phase velocity model of the crust to 36s period, and invert their surface wave  
284 data for a shear velocity model. This model is able to well resolve lateral  
285 features greater than 50km well, highlighting the heterogenous nature of the  
286 crust in Sabah, but is limited in its resolution of vertical contrasts.

287

288 In order to further understanding of crustal and upper mantle structure,  
289 particularly vertical discontinuities, we calculate radial P-wave receiver

290 functions at 70 seismic stations, including the recent nBOSS deployment  
291 across Sabah, and jointly invert these with surface wave data to develop a  
292 detailed shear velocity model of the crust in Sabah, allowing us to map the  
293 Moho geometry beneath. These results provide important constraints on  
294 processes that have shaped the region.

295

## 296 **Data and Methods**

297

298 Broadband teleseismic data in this study come from two seismic networks  
299 deployed in Sabah (Figure 1, Supplementary Table 1). The temporary nBOSS  
300 network of 46 seismometers, installed mostly on a ~40x40 km grid, with a  
301 mean interstation distance of 37.5km (Bacon, 2021), between March 2018  
302 and January 2020, consisted of 18 Güralp 3ESPD instruments and 28 Güralp  
303 6TD instruments (Rawlinson, 2018, Pilia et al., 2019). We also used data from  
304 the Malaysian Metrological Service permanent seismic network. In Sabah, this  
305 consists of 24 permanently installed Streckeisen STS2/2.5 and SS1-Ranger  
306 seismometers, predominantly located in regions of elevated seismicity around  
307 Mt Kinabalu and Darvel Bay.

308

309 Calculation of radial P receiver functions from 3-component seismograms of  
310 teleseismic (30-90° epicentral distance) earthquakes allows us to investigate  
311 the structure of the crust, including determining Moho depth and identifying  
312 layering within the crust.

313

314 We performed the initial quality control of the seismograms in two stages.

315 First, a total of 27,660 3-component seismograms from March 2018-

316 September 2018 for earthquakes  $M_w > 5$  that met the distance criteria were  
317 visually inspected. Where the P-wave signal-to-noise ratio was high (e.g., a P  
318 arrival could clearly be identified) on all 3 components, these seismograms  
319 were classified as 'good' and were taken forward for further analysis. All other  
320 seismograms were rejected and classified as 'bad'. Using this classified data  
321 set we developed a deep learning algorithm to determine the probability of a  
322 3-component seismogram being suitable for further analysis. The annotated  
323 (good or bad) seismograms were converted into spectrograms and 80% of the  
324 data were used to train an image classification convolutional neural network,  
325 ResNet50, pretrained on ImageNet (He et al., 2016). The data classification  
326 algorithm was then tested using the remaining 20 per cent of the data and had  
327 a 92.7 per cent accuracy. A total of 57,858 3-component seismograms from  
328 September 2018-January 2020 were then used with the classification  
329 algorithm, and only those classified with a greater than 50 per cent probability  
330 of being 'good' were visually inspected. This significantly reduced the time  
331 needed for this stage of data quality control, while resulting in a similar  
332 proportion of events being taken forward for further analysis.

333

334 After initial quality control, 14,447 seismograms were used to calculate  
335 receiver functions using the time-domain iterative deconvolution method of  
336 Ligorria and Ammon (1999), with a gaussian width of 1.6, corresponding to a  
337 frequency of 0.9Hz. A gaussian width of 1.6 provides a good balance of  
338 resolving detail in the crust, while avoiding the risk of overinterpreting noise.  
339 Further quality control steps included removal of receiver functions with a poor  
340 fit (<70 per cent), and those which appeared noisy, oscillatory or anomalous

341 to other receiver functions from a similar distance and backazimuth on visual  
342 inspection. This left a remaining dataset of 3338 receiver functions. Eight  
343 stations, one from the nBOSS network, and seven from the Malaysian  
344 Metrological Service permanent seismic network, had no usable receiver  
345 functions. For stations with usable receiver functions, the number of receiver  
346 functions at individual seismometers ranges from 9 at SPM to 183 at SBA8  
347 (Supplementary Table 1, Supplementary Figure 1). This is due to variations in  
348 the amount of data available from individual stations and the noise levels at  
349 the installation sites.

350

351 Receiver functions at an individual station are stacked together to reduce  
352 noise, and these stacks are used in an inversion for crustal velocity structure.  
353 While useful, the interpretation of receiver functions on their own is inherently  
354 non-unique (Ammon et al., 1990). Therefore, to ensure that shear velocities  
355 we obtain from inversions of receiver functions are realistic for this region, we  
356 jointly invert the receiver functions with fundamental mode Rayleigh wave  
357 group velocity dispersion curves extracted from the GDM52 global compilation  
358 (Ekström, 2011) for each station location for a period range 25–250 s. This  
359 model is relatively coarse ( $1^\circ \times 1^\circ$ ), does not see much variation in group  
360 velocities beneath Sabah (Supplementary Figure 2), and with a minimum  
361 period of 25s, is most sensitive to depths of 25km and below, corresponding  
362 to the mid to lower crust and upper mantle in this area. We are primarily  
363 concerned in this study that the inversions result in velocity models that fit the  
364 receiver function data.

365

366 The radial P receiver function stacks, together with the dispersion curves, for  
367 each station were inverted for shear velocity structure using joint96  
368 (Herrmann, 2013), an iterative linearised least squares inversion method.  
369 Several starting models were tested including constant values of 4.48 km/s,  
370 (mantle velocity in the ak135 model - see Kennett et al., 1995), 4.28 km/s, and  
371 3.70 km/s and a  $V_p/V_s$  value of 1.74 down to 100 km depth, parameterised  
372 into 2 km thick layers, overlying ak135 (Supplementary Figure 3). While there  
373 are some small variations in the absolute shear velocities due to differences in  
374 the starting models, they are sufficiently small to not alter the interpretation of  
375 the structure. We test different relative weights ( $p$  value in joint96) of surface  
376 waves to receiver function data in the inversion: 0.5, 0.1, 0.05 and 0.01.  
377 Models with  $p=0.5$  are smoother than those with a lower  $p$  value, reflecting the  
378 greater contribution of surface wave data. Overall, the models show little  
379 variation in structure with  $p$  value, indicating that the recovered features are  
380 robust (Supplementary Figure 4).

381 In order to test the stability of the results, the receiver functions at each station  
382 were divided into 3 time periods (March 2018-Sept 2018, Sept 2018-March  
383 2019, March 2019-Jan 2020) and were stacked and inverted separately.

384 There was very little difference between the resulting models and those  
385 obtained from the complete dataset (Supplementary Figure 5). This can be  
386 considered a form of bootstrapping, and gives us further confidence that the  
387 results are robust

## 388 **Results**

389

390 *Individual receiver functions*

391 Plotting individual receiver functions with respect to backazimuth (e.g  
392 Supplementary Figure 1, Supplementary Figure 6), shows that for some  
393 stations there is a degree of variability for events at different backazimuths.  
394 There are a number of potential causes of this, including short-length scale  
395 variation in crustal structure, anisotropy in the crust, and dipping layers in the  
396 crust. Unfortunately the backazimuthal range of events in this study is limited,  
397 with events being mainly to the north east or south east of the seismometers,  
398 meaning it is not possible to model the cause of the backazimuthal variation  
399 effectively. Because of this limitation, we consider all receiver functions at a  
400 station in a signal stack.

401

402 At some stations we observe receiver functions that do not fit with the typical  
403 idea of a receiver function, which in some automated QC approaches would  
404 have been rejected. One of the most extreme examples of this are the  
405 receiver functions computed for events recorded by the station SBG3  
406 (Supplementary Figure 6). Here, the amplitude of the first arrival is relatively  
407 low, merging with a large amplitude positive arrival at ~2. There is a large  
408 amplitude negative arrival at ~4s, and a large amplitude positive arrival at ~5s.  
409 At this station there are 52 individual receiver functions that have this pattern  
410 (i.e all those with north easterly or south easterly backazimuths). Further, the  
411 Malaysian Metrological Service seismometer PTM, located ~20km from  
412 SBG3, also shows similar pattern of positive and negative arrivals. PTM is a  
413 different type of seismometer, deployed in a different way to SBG3. The  
414 consistency between multiple events, and similarities between receiver



415 functions at different, but relatively close, locations suggests that, while these  
416 receiver functions are not necessarily typical, they do reflect real crustal  
417 structure.

418

#### 419 *Stacked receiver functions*

420

421 Stacked receiver functions at each station are plotted on cross-sections  
422 across Sabah (Figure 2). Cross-section A (Figure 2 (a)) cuts through the  
423 highest topography in Sabah in the region around Mt Kinabalu. Heading SE  
424 from the NW coast there is a positive arrival that decreases from ~4.5 s to ~3  
425 s delay time at the stations immediately beneath the highest topography.  
426 Moving further SE, the delay time of this prominent positive arrival then  
427 increases to ~7 s at station SBE4 (dark grey dashed line). At stations to the  
428 SE of Mt Kinabalu, this positive arrival is preceded by a large amplitude  
429 negative arrival that similarly shows an increase in arrival time from ~3 s at  
430 SBF2 to ~5 s at SBE4 (light grey dashed line). There is a clear change in the  
431 character of the receiver functions from ~200 km along the cross section, with  
432 the portion of cross-section A between SBD5 and SBD7 having a large  
433 amplitude positive arrival at ~6-6.5 s (dark grey dashed line).

434

435 In cross-section B (Figure 2 (b)), the stations in the Crocker Range have a  
436 relatively consistent large amplitude positive arrival (dark grey dashed line) at  
437 ~4 s, while at SBB3 and SBD4, immediately to the SE, the largest amplitude  
438 positive arrival appears to be at ~6.5 s. Between these stations and those in  
439 the vicinity of the Maliau Basin there is a strong positive arrival at ~3.5 s, while

440 at the stations near to the Maliau Basin there is an arrival at ~5 s. At the  
441 south-east end of the cross-section, in the Semporna Peninsula, the largest  
442 positive arrivals, after the direct P arrival, is again at a shorter delay time of  
443 around 4 s.

444

445 The peak at ~0 s should correspond to the direct P arrival; however if there  
446 are low-velocity sediments in the uppermost crust the P-to-S conversion from  
447 the base of these may interfere with the direct P resulting in the first positive  
448 arrival being shifted away from 0 s. This is observed at several sites,  
449 e.g., SBD6 (Figure 2 (a)) and MALB (Figure 2 (b)) and is anticipated given the  
450 thick sedimentary basins (>6 km sediments, Hall, 2013) in Sabah.

451

452 *Shear velocity structure*

453

454 Using the 1-D shear wave velocity model beneath each station, 2D,  
455 composite velocity cross sections have been constructed using bicubic  
456 interpolation for several lines across Sabah, taking the velocity models for  
457 stations within 50km of the line. The models shown in Figure 3 are derived  
458 from the inversions that used  $p=0.1$  and the 3.7 km/s starting model; however,  
459 the features remain consistent with the various weightings and starting  
460 models tested. Given the high weight of the receiver functions in the models  
461 shown, the descriptions and interpretations of the models are primarily  
462 concerned with changes in velocity, and relative velocities, rather than  
463 absolute velocities. The orientations of cross-sections A and B are chosen to  
464 be approximately perpendicular to the strike of the Crocker range, while

465 cross-sections C, D, and E are chosen to help further elucidate the 3D crustal  
466 structure.

467 In cross-section A (Figure 3(a)), low velocity sedimentary basins, labelled as 1  
468 in the cross-section, are illuminated offshore to the NW (SBG1) and in the  
469 south-eastern half of the cross-section (SE of SBE4), confirming the  
470 observation of a broadened, delayed P arrival in the receiver function stacks.  
471 Low velocities extend to depths of ~10 km. The most striking feature of this  
472 cross-section is a dipping high velocity layer, labelled 2, extending from SBE3  
473 to SBE5 from ~5 km to 50 km depth, with a dip in the cross section to the SE,  
474 which overlies a low velocity layer, labelled 3, with a similar dip. This fits the  
475 pattern of arrivals seen in the receiver function cross-sections: a seemingly  
476 dipping transition from a high velocity to a low velocity layer resulting in a  
477 negative arrival, followed by a low to high velocity discontinuity with increasing  
478 depth.

479 In cross-section B (Figure 3(b)) low velocities, labelled 4, are also observed to  
480 ~10 km depth in the vicinity of known sedimentary basins between SBC4 and  
481 SBA7. The transition to mantle velocities ( $\sim >4.2$  km/s in this model) occurs at  
482 ~30-35 km depth in the north-western part of the cross section but deepens to  
483 greater than 40 km beneath SBC4. It shallows to ~25 km beneath SBB4 and  
484 SBC5, before deepening to ~35 km again beneath the Maliau Basin. This  
485 agrees with the pattern of positive arrivals observed in the stacked receiver  
486 functions, with those for stations between the Crocker Range and the Maliau  
487 Basin experiencing the shortest delay times. At the southeast end of cross-  
488 section B, the crust beneath SBA8 and SBA9 has lower velocities (~3.9 km/s

489 in this model) at the gradient interpreted to be the Moho, labelled 5, than  
490 beneath stations elsewhere in the section.  
491  
492 The differences in crustal structure from south to north through western  
493 Sabah is highlighted in cross-section C (Figure 3(c)). The crustal structure at  
494 the north-east end of this section, to the north of Mt Kinabalu, has a different  
495 character to that in the central portion of the section (between SBD2 and  
496 SBE3). In the north east, very high velocities ( $>4.2$  km/s in this model),  
497 labelled 6, are observed at  $\sim 20$  km, while in the central portion they are  
498 generally low ( $<3.4$  km/s in this model) at this depth in a somewhat  
499 discontinuous layer, labelled 7, likely the dipping low velocity layer observed  
500 in cross-section A. Cross-section D (Figure 3(d)) cuts to the east of the  
501 Crocker Range, through the Maliau Basin. In the south west of this section  
502 there is a northeasterly dipping transition from crust to mantle velocities ( $\sim 4.2$   
503 km/s in this model) from 25 to 45 km depth between SBA3 and the Maliau  
504 Basin. At SBC5 it decreases sharply to a depth of 25 km, the depth it is also  
505 observed to be at between SBE4 and SBF4. Beneath SBD4 relatively high  
506 velocities are observed in the upper crust, labelled 8, and low velocities are  
507 observed between 25 and 45km depth. This location corresponds with where  
508 ophiolitic material is found on the surface around Telupid (e.g. Hall 2013) . In  
509 the south of cross-section E (Figure 3(e)), which cuts through the Semporna  
510 and Dent Peninsulas, the transition between crustal and mantle velocities at  
511  $\sim 35$  km depth is relatively gradual. This contrasts with further north, where this  
512 transition is sharper, and upper mantle velocities are faster.  
513

514 *Moho depth*

515

516 The depth of the Moho beneath each station is picked from its corresponding  
517 1-D shear velocity model at the depth that corresponds to the base of the  
518 steepest positive velocity gradient where shear velocity exceeds 4 km/s  
519 (Figure 4). The depth of the Moho is found to vary from 24 km at SBC6 and  
520 SBF4 to 60 km at MTM, although most other measurements are <48 km. The  
521 deepest Moho is found in a SW-NE trending band on the eastern edge of the  
522 Crocker Range, where the Moho depth exceeds 40 km. It is also relatively  
523 deep (40-44 km) beneath the Maliau Basin and beneath other circular basins  
524 to the north of the Maliau Basin, and to the west of the Segama ophiolite. The  
525 shallowest Moho depth (24-26 km) is found in a band between the Crocker  
526 Range and the circular basins, with changes in Moho depth of ~15-25 km  
527 occurring over short lateral distances (~20 km).

528

529 At the stations marked with white hexagons in Figure 4, it was not clear where  
530 the Moho should be picked. For instance, at a subset of stations (e.g., SBD6,  
531 SBD7, and SBC8) there is a very gradual increase in velocities over a wide  
532 (~40 km) depth range (Supplementary Figure 7), while at other stations (e.g.,  
533 SBD5, SBE4, and SBG3) the models have two steep velocity gradients, both  
534 of which could represent plausible Moho locations given those found  
535 elsewhere in Sabah, for example at 28km and 58km for SBD5.

536

537 **Discussion**

538

539 *Limitations of the models*

540

541 The 1D velocity models obtained in the study have a number of limitations,  
542 which should be acknowledged in order to avoid over interpretation of the  
543 results. The receiver functions used in the inversions are single stacks for  
544 each station. This is done to reduce noise and source effects, however it may  
545 mask real complexity in crustal structure such as short-length scale lateral  
546 variations, anisotropy and dipping layers, which we have not be able to model  
547 due to the limitations imposed by the backazimuthal range of the events.  
548 Therefore it is likely that crustal structure beneath Sabah is more complex  
549 than shown in the models here.

550

551 In the inversions we use the relatively coarse, global model, GDM52  
552 (Ekstrom, 2011), to help ensure the shear velocities are reasonable for the  
553 regional context. While the fits to the dispersion curves are reasonably good  
554 (Supplementary Figure 4), the focus of this study is the constraints provided  
555 by the radial P wave receiver functions, and so we prioritise the fit to the  
556 receiver functions for the models that we interpret. As the models are strongly  
557 weighted to the receiver functions, this means that the best constrained  
558 features in the crustal structure will be the velocity discontinuities, such as the  
559 Moho, rather than absolute velocities. In not using the dispersion  
560 measurements from Fone et al., (2024) and Greenfield et al., (2022), we are  
561 able to obtain an independent image of crustal structure, and thus make  
562 comparisons between the models, giving insight into the robustness of  
563 individual features and their potential interpretations.

564

565 *Processes affecting crustal thickness*

566

567 It is important to account for the effect of the interference between  
568 conversions and multiples (e.g., Gilligan et al., 2014): a consequence is that  
569 the largest signal on a receiver function should not necessarily be interpreted  
570 as being due to the velocity increase at the Moho. However, there is a  
571 consistent pattern between the positive arrivals observed in the stacked  
572 receiver functions and the velocity changes at the Moho observed in the  
573 models from joint inversion, suggesting that the receiver functions can, in this  
574 instance, provide an interpretable picture for trends in Moho depth.

575

576 The variations in crustal thickness across Sabah observed in this study are in  
577 general agreement with the estimates made by recent studies using 2-plane-  
578 wave tomography (Greenfield et al., 2022) and virtual deep seismic sounding  
579 (Linang et al., 2022): central Sabah appears to have significantly thinner crust  
580 than that beneath the Crocker Range and the Circular Basins (Supplementary  
581 Figure 8). One notable difference with the Greenfield et al., (2022) Moho  
582 estimate is beneath the Semporna Peninsular they observe thick crust  
583 (>55km), while in this study we observe crustal thickness of ~34km. This  
584 difference is likely to arise due to Greenfield et al., (2022) using the 4.1km/s  
585 shear velocity contour as a proxy for Moho depth, and this, as discussed  
586 below, may not be an appropriate velocity proxy for the lower crust/upper  
587 mantle beneath the Semporna Peninsular.

588

589 The pattern of thicker and thinner crust broadly agrees with the estimates of  
590 Holt (1998) using gravity data; however, between the Crocker Range and  
591 thickened crust beneath the circular basins we observe a significantly thinner  
592 crust (e.g., 25 and 24 km at SBC5 and SBC6 respectively) than the 32 km  
593 suggested by Holt (1998). Modelling gravity data is notoriously non-unique,  
594 and Holt (1998) uses a very simple model for crustal densities. Given the lack  
595 of other constraints on the properties of the crust at the time this may have  
596 been appropriate; however, the lateral and vertical heterogeneity of the crust  
597 demonstrated in this study indicates a more complex model is required, which  
598 may alter the estimates of crustal thickness from the gravity data.

599

600 It should be noted that the crust in this study is thicker throughout much of  
601 Sabah than was shown in interpretative cross-section of Hall (2013), which  
602 bases Moho depth off the results of Holt (1998), modified for denser material.  
603 Hall (2013)'s cross-section shows a maximum Moho depth of 40 km beneath  
604 the Crocker Range, and around 20 km beneath both the circular basins and  
605 the Dent and Semporna peninsulas. Further, our estimates of crustal  
606 thickness shows a significantly different pattern and depths to the estimates  
607 made by Tang and Zheng (2013). They report crustal thicknesses of ~27.5 to  
608 32.5 km, increasing southward across Sabah, based on the depth of the 4  
609 km/s velocity contour in their shear velocity model. This model encompasses  
610 the whole of the South China Sea and surrounding region and, as such, has  
611 more limited resolution in Sabah compared to this study and others  
612 (Greenfield et al., 2022, Linang et al., 2022) that have used the data from the  
613 nBOSS network. The thicker crust we observe, compared to earlier estimates,



614 may mean that there is a larger contribution from regional tectonic shortening  
615 to the regional uplift observed by Roberts et al., (2018).

616

617 The relatively thick crust (>40 km) beneath the Crocker Range, particularly on  
618 the eastern side, is likely to have been thickened during the Sabah Orogeny  
619 (~23 Ma), when the Dangerous Grounds block collided with the western edge  
620 of northern Borneo at the final stage of the subduction of the proto-South  
621 China Sea (e.g., Hutchison et al., 2000, Hall 2013, Rangin et al., 1999). The  
622 velocity discontinuity picked as the Moho for many stations in the Crocker  
623 Range is the base of the SE dipping slow velocities seen in cross-section A,  
624 which we interpret as the base of the underthrust Dangerous Grounds crust.

625

626 Extension, related to the roll-back of the Celebes Sea slab (e.g., Hall, 2013),  
627 could have thinned the crust in central Sabah to 20-25 km. Indeed, Tsikouras  
628 et al. (2021) argue that the basalts they date to 9-10 Ma in the Telupid area,  
629 are rift-related, thus implying significant extension and crustal thinning,  
630 although this is disputed by Cullen and Burton-Johnson (2021). A double  
631 discontinuity is observed in our study in the 1-D velocity models at some  
632 stations (e.g., SBE4), and could indicate that it may not necessarily be  
633 appropriate to simply interpret the velocity gradient at ~20-25 km as a single  
634 Moho. If the crust is 20-25 km thick, the question as to the extent to which this  
635 crust may have been thinned remains, i.e., what was the pre-extensional  
636 crustal thickness? Greenfield et al. (2022) assume that it was 40-50 km, as is  
637 observed beneath the Crocker Range and circular basins, and thus calculate  
638 a stretching factor of 1.3-2. However, if this area was not significantly  
639 thickened during the Sabah Orogeny, which is plausible given the lack of

640 underthrust Dangerous Grounds material observed in this study, then pre-  
641 extensional thickness may have been less to begin with.

642

643 Related to the question of pre-extensional thickness is whether the ~45 km  
644 thick crust observed beneath the Maliau Basin and other circular basins is a  
645 result of thickening during the Sabah orogeny (~23 Ma). After thickening it  
646 may have been separated from other thickened crust beneath the Crocker  
647 Range as a result of extension (e.g., in a crustal scale boudinage process as  
648 suggested by Linang et al. (2022)). Alternatively, the crust may have been  
649 thickened at a later point in time. Tongkul and Chang (2003) suggest that  
650 eastern Sabah experienced N-S compression in the mid Middle Miocene (~13  
651 Ma), which led to the segmentation of the large basin in eastern Sabah that  
652 had been active in the Early Miocene, and NW-SE compression in the late  
653 Upper Miocene (~7-5 Ma), which enhanced the circular shape of the basins,  
654 with a period of sediment deposition in between these two compressional  
655 events. It may be that during these compressional episodes, potentially  
656 associated with Celebes Sea subduction, some crustal thickening occurred  
657 beneath the circular basins. More recently, Meju et al., (2024), using results  
658 from a magnetotelluric study, argue for thick-skinned deformation and deep  
659 crustal flow being involved with the evolution of Neogene 'mini-basins'  
660 offshore, which may be directly related to the onshore circular basins such as  
661 the Maliau Basin.

662

663 While areas of the highest topography may be anticipated to have some of the  
664 thickest crust, intriguingly beneath the stations in the vicinity of the 4095m  
665 high Mt Kinabalu the crust is only 30-35km thick. The Kinabalu Granite was

666 emplaced between 7.2-7.8Ma (Cottam et al., 2013), well after the termination  
667 proto-South China Sea subduction, thus it would be expected that the crust in  
668 this region would have been thickened as a result of this collision. The  
669 thermomechanical modelling of Pilia et al., (2023b) shows that as a result of a  
670 downwelling drip, e.g the Semporna drip, a region of initially thick crust can be  
671 thinned. This thinning could facilitate melting of the lower crust, thus it may be  
672 that both the presence of the Kinabalu pluton and the thinner-than-anticipated  
673 crust we observe can both be explained by part of the lithosphere having  
674 dripped off beneath the Semporna peninsula.

675

#### 676 *Crustal structure*

677

678 Given the diversity of the surface geology in Sabah, it is unsurprising that the  
679 crust shows considerable variation. The key elements of our interpretation are  
680 shown in Figure 5.

681

682 With the caveat that we have projected 1D velocity models from 50km around  
683 the lines of section, and thus may be masking some of the lateral variation  
684 perpendicular to this line, we interpret the low velocity (<3.4 km/s in our  
685 model) layer seen dipping to the south east from the west coast of Sabah to  
686 the eastern edge of the Crocker Range in Cross-section A (Figure 3(a)) as  
687 Dangerous Grounds material that has been underthrust beneath Sabah.  
688 Underthrusting of attenuated Dangerous Grounds crust has been proposed as  
689 the mechanism by which subduction of the proto-SCS stopped (e.g., Hall,  
690 2013, Morley and Back, 2008, Hutchison, 2000). Fone et al., (2024) also

691 observe low shear velocities at ~30km depth along a similar line of section,  
692 but this is arguably the first time it has been imaged as a dipping structure.  
693 Cross-section B (Figure 3(b)), which cuts to the south of Cross-section A, also  
694 has a low velocity layer at depths of 20-25 km. In this instance this layer does  
695 not seem to dip. We consider this to also be underthrust Dangerous Grounds  
696 crust, although this suggests potential along strike variation in the nature of  
697 the collision between Sabah and the Dangerous Grounds. Rangin et al.  
698 (1999), considering the whole of the proto-SCS, argue that the proto-SCS  
699 basin was narrower off the coast of Borneo than the Sulu Sea, and it may be  
700 that the differences we observe in underthrust Dangerous Grounds crust are a  
701 manifestation of this. Furthermore, Greenfield et al., (2022) note that the  
702 lithosphere is thinner in the southwest of Sabah, again suggesting that  
703 different processes may have influenced this area compared to those further  
704 north.

705

706 Overlying the low-velocity layer in Cross-section A is a high velocity layer,  
707 also dipping to the south east, with velocities exceeding 4 km/s in our model  
708 in what is interpreted as the upper- to mid-crust. High velocities overlying the  
709 lower velocities are also observed by Fone et al., (2024). The velocities we  
710 observe are consistent with this being mafic to ultramafic material, although,  
711 as noted above, the velocities in our models should primarily be interpreted  
712 relatively due to the high weight of receiver function observations in the  
713 inversion. However, as this high velocity layer appears to lie beneath areas of  
714 peridotite rocks near Ranau and ophiolitic rocks near Telupid, we interpret the  
715 layer as obducted ophiolitic material. It is not possible to constrain the timing

716 of the emplacement from this study, e.g., it may be the result of late Mesozoic  
717 rifting (Tsikouras et al., 2021), or it could have been emplaced earlier (e.g.,  
718 Cullen and Burton-Johnson 2021).

719

720 The crustal structure of the northern tip of Sabah (to the north of SBF2) is  
721 distinct from areas to the south, as is particularly seen in Cross-section C  
722 (Figure 3(c)), suggesting that distinct geological processes have shaped this  
723 region. As similar structure is seen by Fone et al., (2024), this gives us  
724 confidence that this is real structure, given that two independent, but  
725 complementary, methods have obtained similar results. The change in the  
726 character of the crustal structure is in the same place where there is a change  
727 in strike of the surface geology, from ~SW-NE to the south to ~WNW-ESE in  
728 the north (Tongkul 1990), in the vicinity of Mt Kinabalu. Moreover, it is  
729 approximately coincident with where the fast velocity anomaly in the upper  
730 mantle that Pilia et al. (2023a) associate with the proto-South China Sea slab  
731 terminates. Tongkul (1994) suggests, based on the relationships between  
732 sedimentary rocks in this region, that the basement here - Mesozoic oceanic  
733 crust - is uplifted relative to the area to the south. Tongkul (1994) further  
734 suggests that this region was affected by the collision with the Reed Bank,  
735 resulting in N-S compression, while further south the collision was with the  
736 Dangerous Grounds block. Gozzard et al., (2018) observe that the crust  
737 beneath the Reed Bank has not been thinned in the same way as the  
738 Dangerous Grounds block. Franke et al., (2008) also note the presence of the  
739 Kudat block off the eastern shore of northernmost Sabah, which active source  
740 seismic data suggests has a different crustal structure. It may be that the

741 different properties of blocks colliding with Sabah, as well as the orientation of  
742 the collisions, resulted in the contrasting crustal structure we observe today:  
743 underthrust material to the south but not at the northern tip of Sabah.

744

745 In the east of Sabah, the lower crust and upper mantle beneath stations in the  
746 Semporna Peninsula (SBA8 and SBA9) is relatively slow (~3.8-4 km/s for the  
747 uppermost mantle compared to 4.3-4.5 km/s elsewhere). This is similar to the  
748 results from the two-plane wave tomography of Greenfield et al. (2022).

749 Volcanism in this area occurred until at least 0.2 Ma (Lai et al., 2021) and  
750 potentially as recently as 24-27 ka (Kirk, 1968; Bellwood, 1988; cited in Tjia et  
751 al., 1992) with hot springs found in the vicinity of Tawau today, with water  
752 temperatures of up to 75°C (Siong et al., 1991). Pilia et al. (2023b) and  
753 Greenfield et al. (2022) propose that part of the lithosphere has been removed  
754 beneath the Semporna Peninsula and has been replaced by hot  
755 asthenospheric material. This would mean that the remaining crust and  
756 mantle would be expected to be warm and thus seismically slow, as we  
757 observe here.

758

759

760

## 761 **Conclusion**

762

763 We present a high-resolution crustal shear velocity model of Sabah, northern  
764 Borneo, from the joint inversion of P receiver functions and surface wave  
765 data. We image, for the first, time dipping Dangerous Grounds crust

766 underthrust beneath most of the Crocker Range. This has had the effect of  
767 thickening the crust beneath the present-day mountain range, with crustal  
768 thicknesses exceeding 40 km. However, beneath Mt Kinabalu, crustal  
769 thicknesses are only in the range 30-35km, supporting earlier ideas (Cottam  
770 et al., 2013, Sapin et al., 2013, Tsikouras et al., 2021, Pilia et al., 2023b) that  
771 some degree of crustal thinning may have been involved in its emplacement.  
772 Thinner crust (~25 km) between the Crocker Range and the Circular Basins  
773 may be due to extension related to the rollback of the Celebes Sea slab (Hall,  
774 2013), although the amount of extension remains unclear given that pre-  
775 extensional crustal thickness remains unknown. Thicker crust (>40 km)  
776 beneath the Maliau and other circular basins suggests that these areas have  
777 experienced some degree of crustal thickening, which given the late-mid  
778 Miocene age of the sediments that have been deformed is likely to have  
779 occurred later than the ~21 Ma Sabah Orogeny. Relatively slow velocities in  
780 the lower crust and upper mantle beneath the Semporna Peninsula support  
781 work by Pilia et al. (2023b) and Greenfield et al. (2021) that lithospheric  
782 delamination has occurred here.

783

784 Overall, we observe a high degree of heterogeneity in the crustal structure  
785 beneath Sabah, on length scales of 10s of kilometres. This highlights the  
786 complexity of subduction, collisional, post-subduction, and extensional  
787 processes that have shaped Sabah over the Cenozoic, and reinforces the  
788 importance of dense instrumentation in order to better understand tectonic  
789 activity that has occurred in similar settings.

790

791 **Global Research Collaboration Statement**

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793 recovery of the nBOSS network between March 2018 and January 2020.  
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800 Continental Margin.”

801

802

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804

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818 2019). We thank Robert Herrmann for making the Computer Programs in  
819 Seismology freely available.

820

## 821 **Data availability**

822

823 The nBOSS dataset is accessible through the EarthScope Data Management  
824 Center ([https://www.fdsn.org/networks/detail/YC\\_2018/](https://www.fdsn.org/networks/detail/YC_2018/)). Data from the  
825 Malaysian national seismic network  
826 (<https://www.fdsn.org/networks/detail/MY/>) are restricted but may be obtained  
827 by contacting the Malaysian Meteorological Department. The exceptions to  
828 this are stations KKM and LDM which are also available through the  
829 Earthscope Data Management Center. The Open Science Framework site for  
830 this project, which includes figures for each station analysed, and SAC files  
831 for the receiver functions used can be found at: <https://osf.io/2zvcg/>

832

## 833 **Author contributions**

834

835 **A.G.:** Formal analysis, conceptualisation, funding acquisition, investigation,  
836 resources, visualisation, writing – original draft; **D.C.:** Investigation, resources,  
837 writing - review and editing; **N.R.:** Conceptualisation, funding acquisition,  
838 resources, investigation, writing – review and editing; **F.T.:** Conceptualisation,  
839 resources, investigation; **S.P.:** Investigation, writing – review and editing,

840 funding acquisition; **T.G.:** Investigation, writing – review and editing; **C.B.:**

841 Data curation, investigation.

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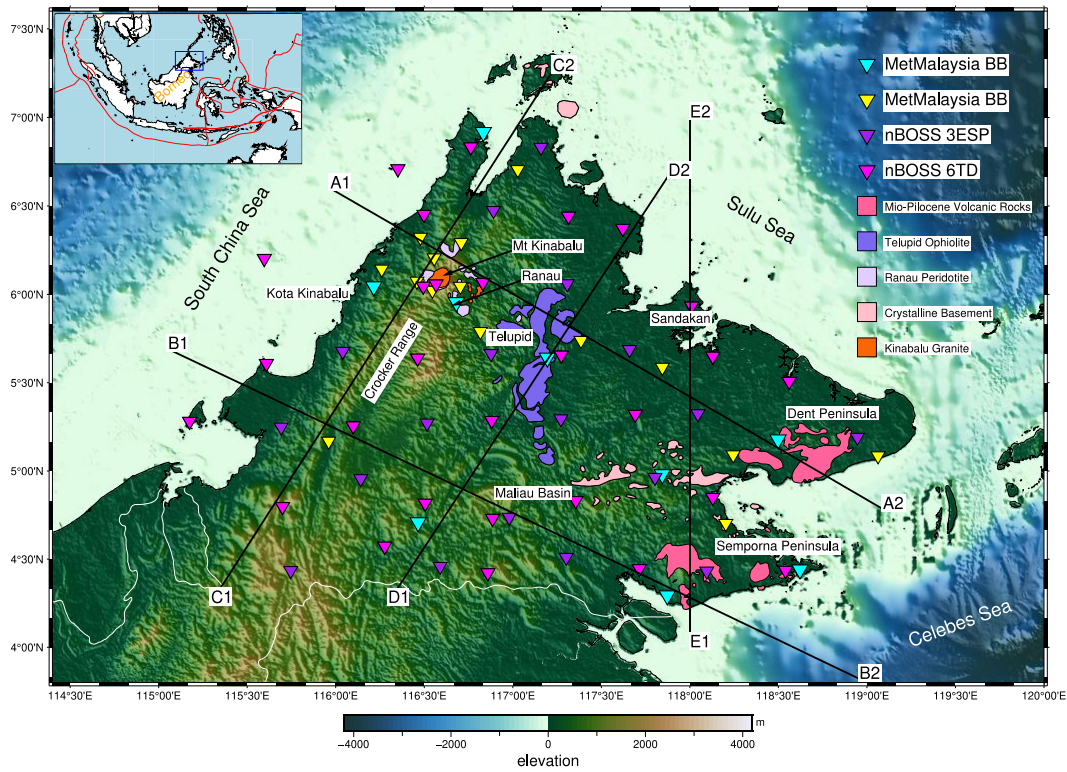
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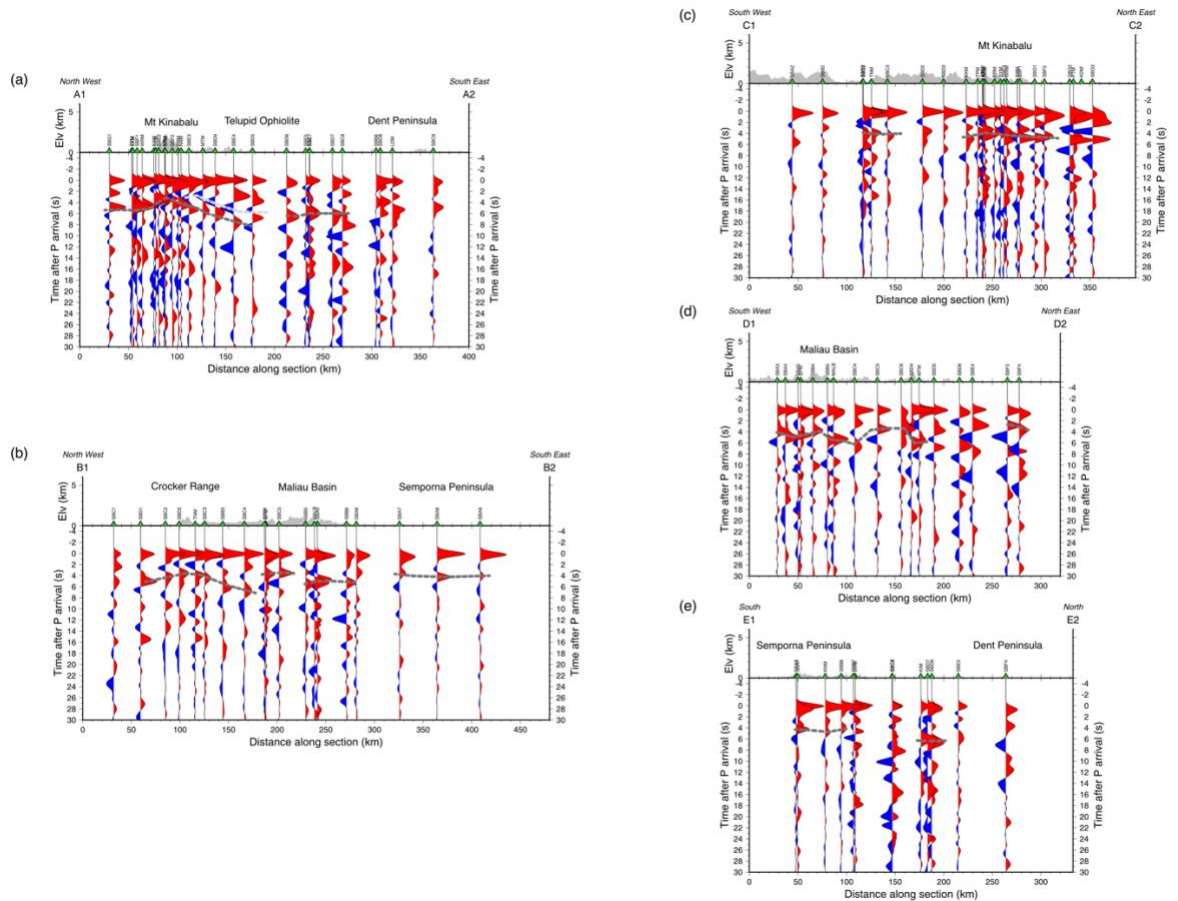
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1109 **Figures**

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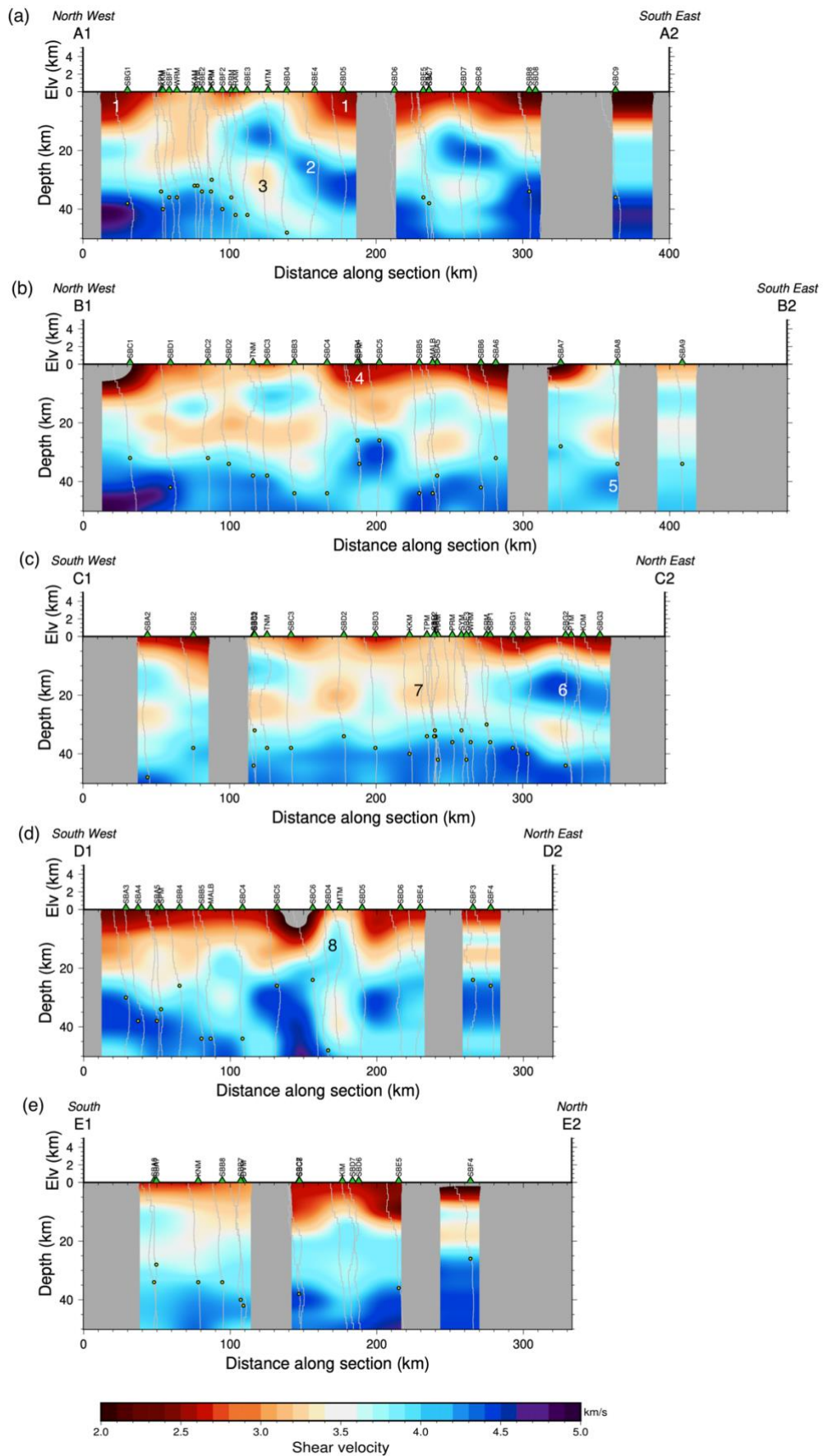
**Figure 1:** Map of seismometer stations in Sabah used in this study. Blue triangles are MetMalaysia seismometers deployed before 2017, yellow triangles are MetMalaysia seismometers deployed after 2017. Pink (6TD) and Purple (3ESP) triangles are seismometers deployed as part of the nBOSS project. Lines of section are shown: A1-A2 (6.56°N 115.97°E - 4.78°N 119.09°E), B1-B2 (5.691°N 115.05°E - 3.82°N 118.96°E), C1-C2 (4.25°N 115.30°E - 7.26°N 117.23°E), D1-D2 (4.25°N 116.30°E - 6.67°N 117.88°E) E1-E2 (4°N 118°E - 7°N 118°E). Geological units are plotted after Hall (2013). The inset map shows the wider geographical area, with the area of the main map highlighted by the blue box, and plate boundaries after Bird (2003), are shown by red lines.



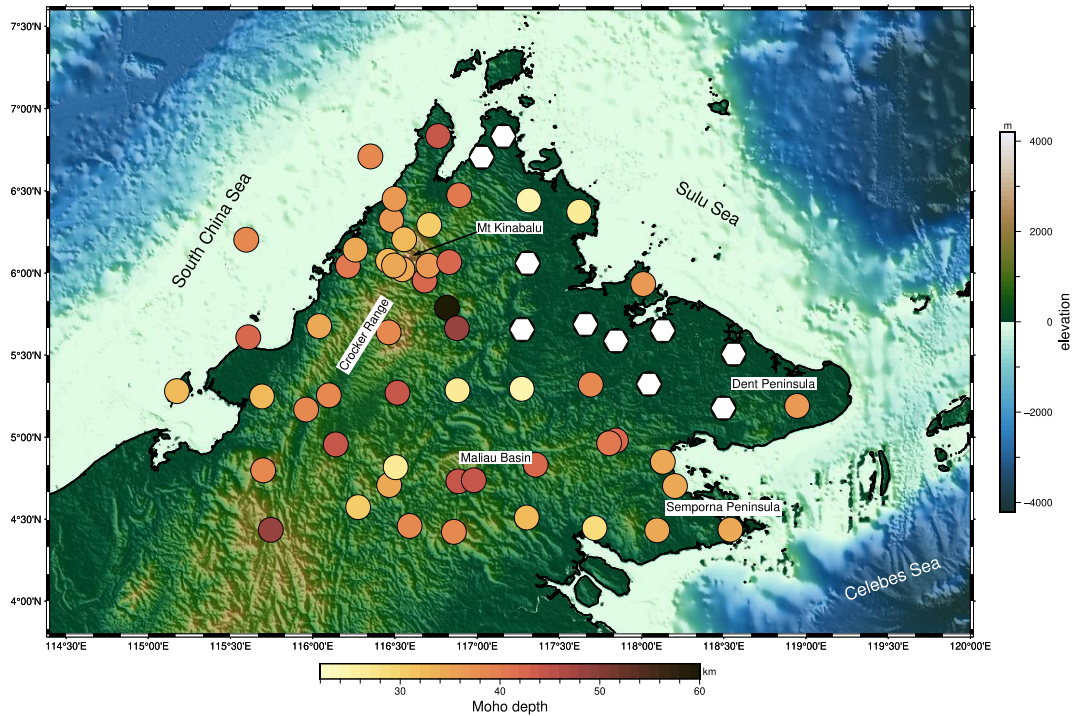
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1133 **Figure 2:** Stacked receiver functions along the lines (a) A1-A2, (b) B1-B2, (c)  
1134 C1-C2, (d) D1-D2, and (e) E1-E2. Positive arrivals are filled red, and negative  
1135 arrivals are filled blue. In both cases receiver functions from stations within 50  
1136 km of each line have been projected onto the section, along with their  
1137 respective station (green triangles), and topography is plotted above. The  
1138 dark grey dashed line highlights positive arrivals, likely from the P-to-S  
1139 conversion at the Moho. The light grey dashed line highlights negative arrivals  
1140 corresponding to a velocity increase with depth in the crust.

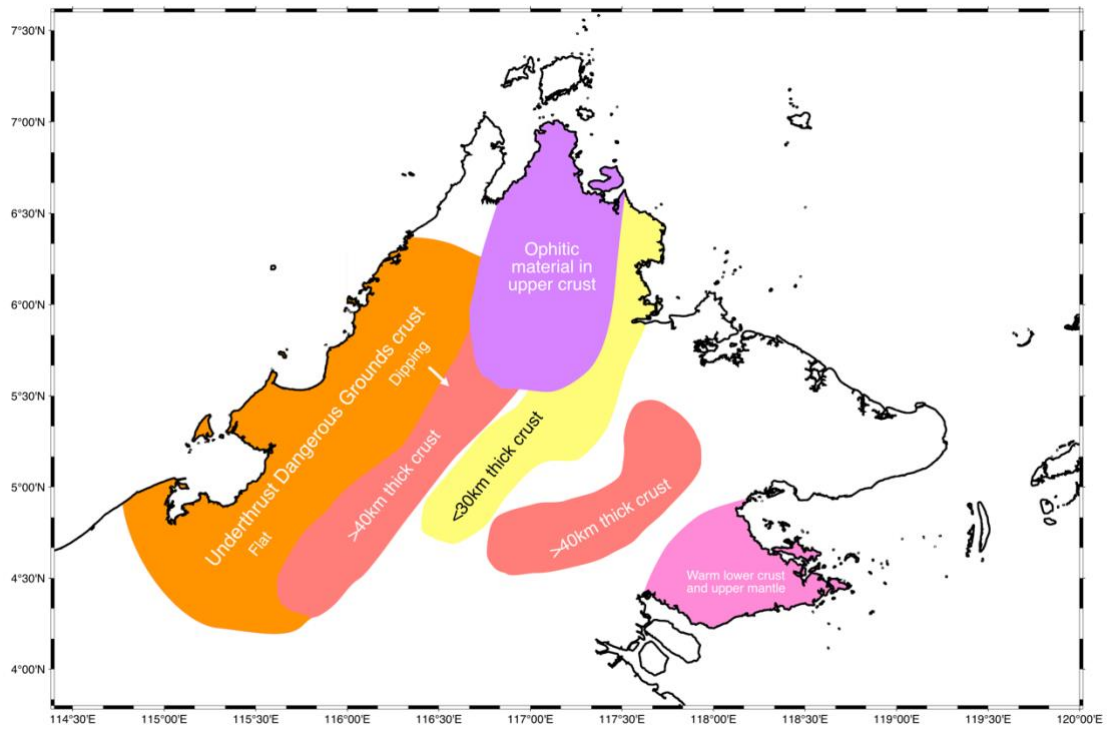




1142 **Figure 3:** Shear velocity vs depth along lines (a) A1-A2, (b) B1-B2, (c) C1-C2,  
 1143 (d) D1-D2, (e) E1-E2 from the joint inversion of receiver function and surface  
 1144 wave data. 1-D models from stations within 50 km of the line of section are  
 1145 interpolated to make the cross-sections. Grey areas indicate areas with no  
 1146 station coverage. Green triangles mark the location of stations. Topography  
 1147 along the line of section is plotted above. Labelled velocity anomalies 1-8  
 1148 are discussed in the text.  
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 1153 **Figure 4:** Moho depths at seismometer stations in Sabah picked from 1D  
 1154 shear velocity models from the joint inversion of receiver function and surface  
 1155 wave data. The colour of the circle indicated Moho depth for the station  
 1156 located at that point, as shown in the scale. White hexagons are locations  
 1157 where there was no clear Moho to be picked or where there were multiple  
 1158 plausible velocity discontinuities that could be the Moho.  
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**Figure 5:** Summary map highlighting the key interpretations from this study from the shear velocity models derived from the joint inversion of receiver function and surface wave data.

1190 **Imaging subduction, collision, and extension in**  
1191 **northern Borneo: Constraints from receiver**  
1192 **functions**

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1210 **Supplementary Material**

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**Supplementary table 1:** The name, instrument type, and location of the seismometers used in this study, together with the number of good receiver functions after quality control, and the crustal thickness estimated from the joint inversion of receiver function and surface wave data. Where the crustal thickness is N/A this is because there were no good receiver functions for that station. Where crustal thickness is 'X' these are stations where it was not possible to estimate the crustal thickness from the velocity model.

Network	Station	Instrument type	Latitude (N)	Longitude (E)	# of receiver functions	Crustal thickness (km)
YC	SBA2	3ESP	4.43506	115.74560	76	48
YC	SBA3	6TD	4.57347	116.27660	74	30
YC	SBA4	3ESP	4.45879	116.58977	73	38
YC	SBA5	6TD	4.42271	116.85881	84	38
YC	SBA6	3ESP	4.51025	117.30176	18	32
YC	SBA7	6TD	4.44587	117.71442	47	28
YC	SBA8	3ESP	4.43208	118.09522	183	34

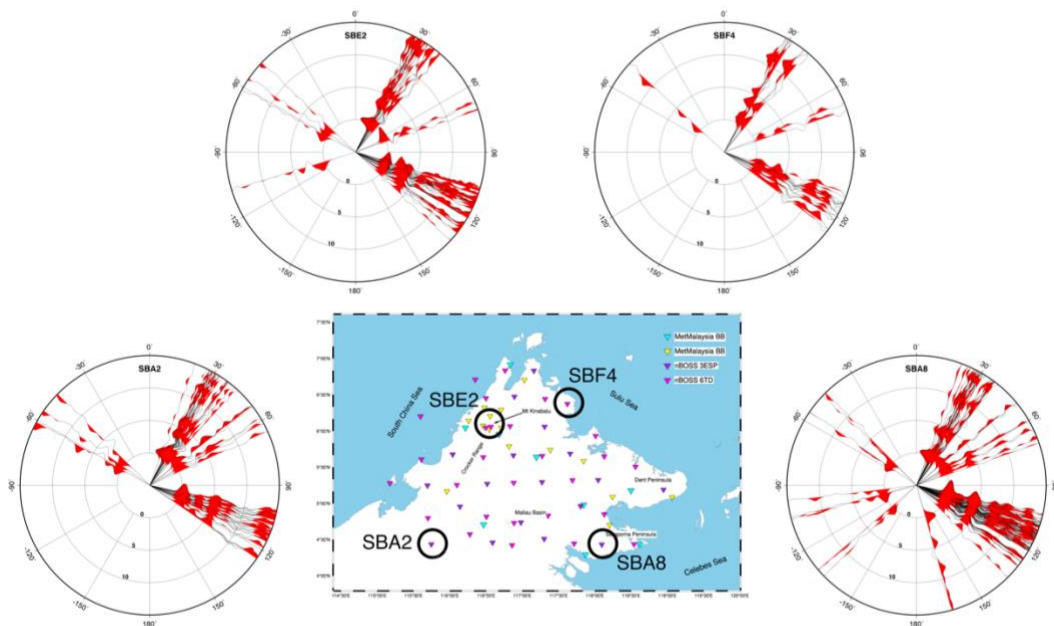
YC	SBA9	6TD	4.43637	118.53992	76	34
YC	SBB2	6TD	4.79788	115.69913	62	38
YC	SBB3	3ESP	4.95640	116.14096	126	44
YC	SBB4	6TD	4.81717	116.50497	36	26
YC	SBB5	6TD	4.73101	116.88574	82	44
YC	MALB	3ESP	4.73740	116.97997	51	44
YC	SBB6	6TD	4.83194	117.35575	54	42
YC	SBB7	3ESP	4.96355	117.80282	70	40
YC	SBB8	6TD	4.85014	118.12941	79	34
YC	SBC1	6TD	5.28108	115.17476	24	32
YC	SBC2	3ESP	5.24880	115.69165	114	32
YC	SBC3	6TD	5.25540	116.09856	48	38
YC	SBC4	3ESP	5.27107	116.51573	76	44
YC	SBC5	6TD	5.28637	116.88076	73	26
YC	SBC6	3ESP	5.29518	117.27165	19	24
YC	SBC7	6TD	5.32075	117.69134	23	42
YC	SBC8	3ESP	5.32373	118.04523	13	X
YC	SBC9	3ESP	5.19098	118.94610	54	36
YC	SBD1	6TD	5.60898	115.60830	42	42
YC	SBD2	3ESP	5.67735	116.03960	55	34
YC	SBD3	6TD	5.63900	116.46225	89	38
YC	SBD4	3ESP	5.66416	116.87691	86	48
YC	SBD5	6TD	5.65637	117.27355	28	X
YC	SBD6	3ESP	5.68750	117.65900	28	X
YC	SBD7	6TD	5.64967	118.12955	80	X
YC	SBD8	6TD	5.50699	118.56041	34	X

YC	SBE1	6TD	6.20282	115.5963	33	38
YC	SBE2	6TD	6.04611	116.49462	83	34
YC	KINA	6TD	6.05826	116.56593	0	N/A
YC	SBE3	6TD	6.06708	116.83097	63	42
YC	SBE4	3ESP	6.05975	117.30715	48	X
YC	SBE5	6TD	5.93328	118.01012	37	36
YC	SBF1	6TD	6.45216	116.49845	48	36
YC	SBF2	3ESP	6.47376	116.89069	42	40
YC	SBF3	6TD	6.44177	117.31431	95	24
YC	SBF4	6TD	6.37312	117.62083	30	26
YC	SBG1	6TD	6.70950	116.35092	22	38
YC	SBG2	6TD	6.83352	116.76262	49	44
YC	SBG3	3ESP	6.83170	117.15904	55	X
MY	DVM	STS-2.5	4.98038	117.84421	11	42
MY	FSM	STS-2.5	5.0855	119.0627	0	N/A
MY	KAM	STS-2.5	6.0745	116.4583	45	32
MY	KDM	SS-1 Ranger	6.9167	116.8333	0	N/A
MY	KIM	STS-2.5	5.587083	117.844717	48	X
MY	KKM	STS-2	6.0443	116.2147	61	40
MY	KNM	STS-2.5	4.7026	118.203	82	34
MY	KPM	STS-2.5	6.0227	116.545417	21	34
MY	LDM	STS-2.5	5.1777	118.498	35	X
MY	MTM	STS-2.5	5.789333	116.81665	37	60
MY	PRM	STS-2.5	6.0455	116.70375	51	36
MY	PTM	STS-2.5	6.70523	117.0283	39	X

MY	RAM	STS-2.5	5.9546	116.681	64	42
MY	SDM	SS-1 Ranger	5.6409	117.195	0	N/A
MY	SGM	STS-2.5	5.0912	118.2446	0	N/A
MY	SMM	SS1- Ranger	4.439838	118.622028	0	N/A
MY	SPM	STS-2	4.7083	116.465	9	34
MY	SRM	STS-2.5	6.29265	116.708383	36	30
MY	SYM	STS-2.5	6.20585	116.5559	42	32
MY	TLM	STS-2.5	5.7391	117.385	0	N/A
MY	TNM	STS-2.5	5.168633	115.960183	27	38
MY	TPM	STS-2.5	6.1427	116.2596	19	34
MY	TSM	SS-1 Ranger	4.2936	117.8725	0	N/A
MY	WRM	STS-2.5	6.3229	116.47825	29	36

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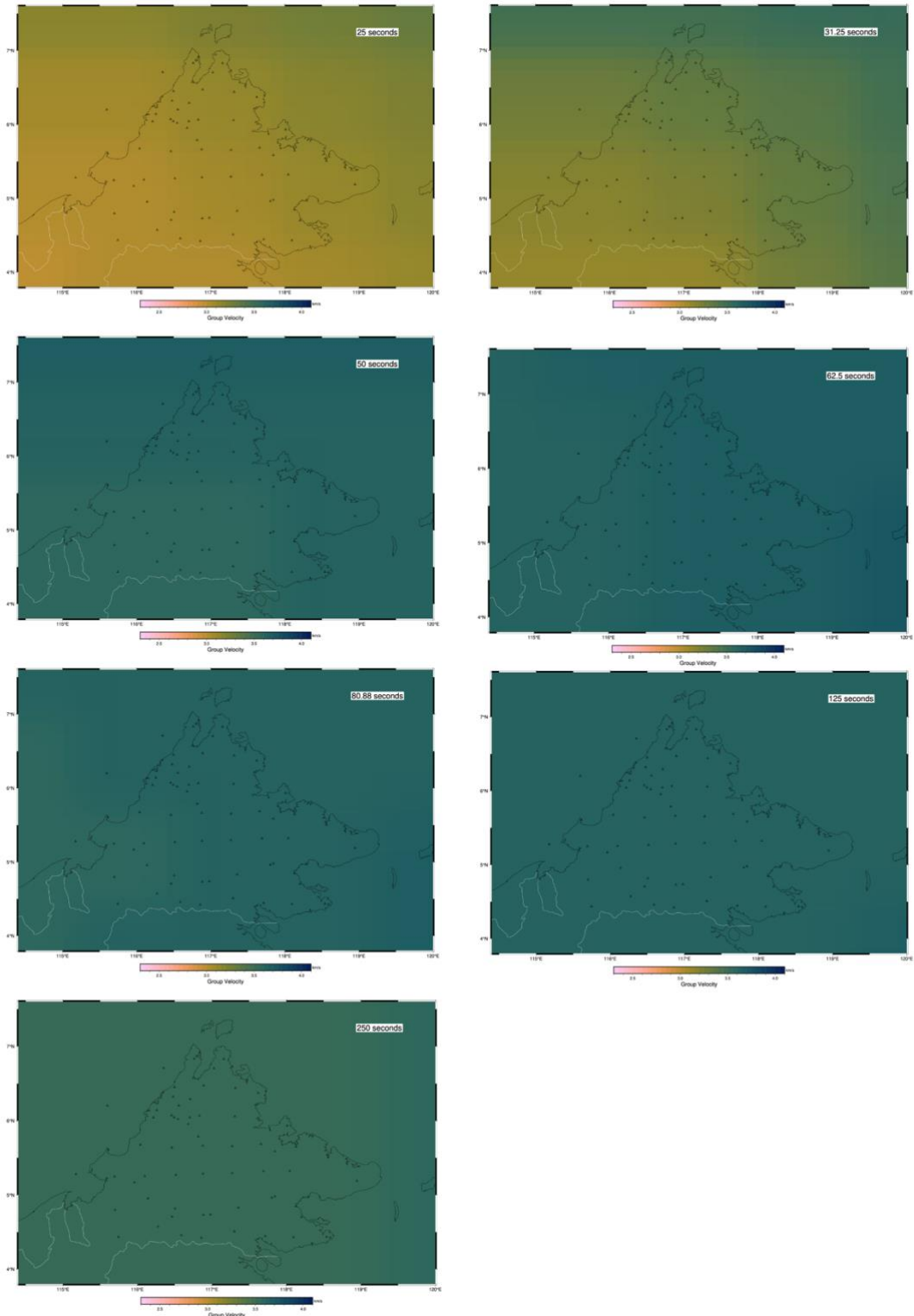
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1223 **Supplementary figure 1:** Examples of the receiver functions of individual  
1224 events used in the station stacks for four stations across Sabah: SBA2,  
1225 SBE2, SBF4, and SBA9, and a map indicating the station locations. Receiver  
1226 functions are plotted with respect to backazimuth with positive amplitudes  
1227 filled red.

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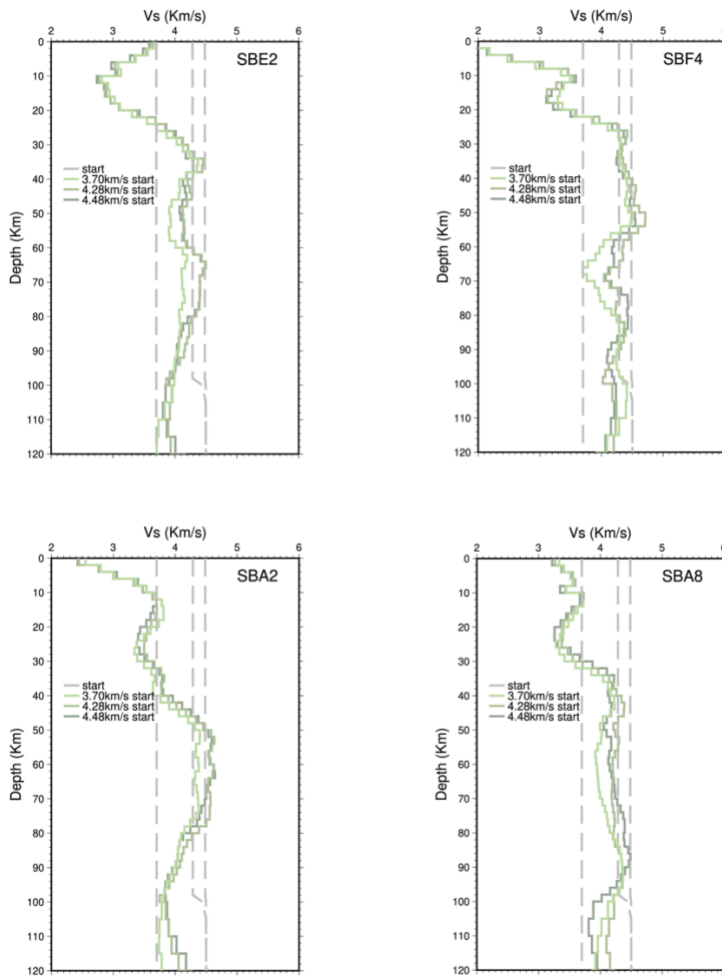


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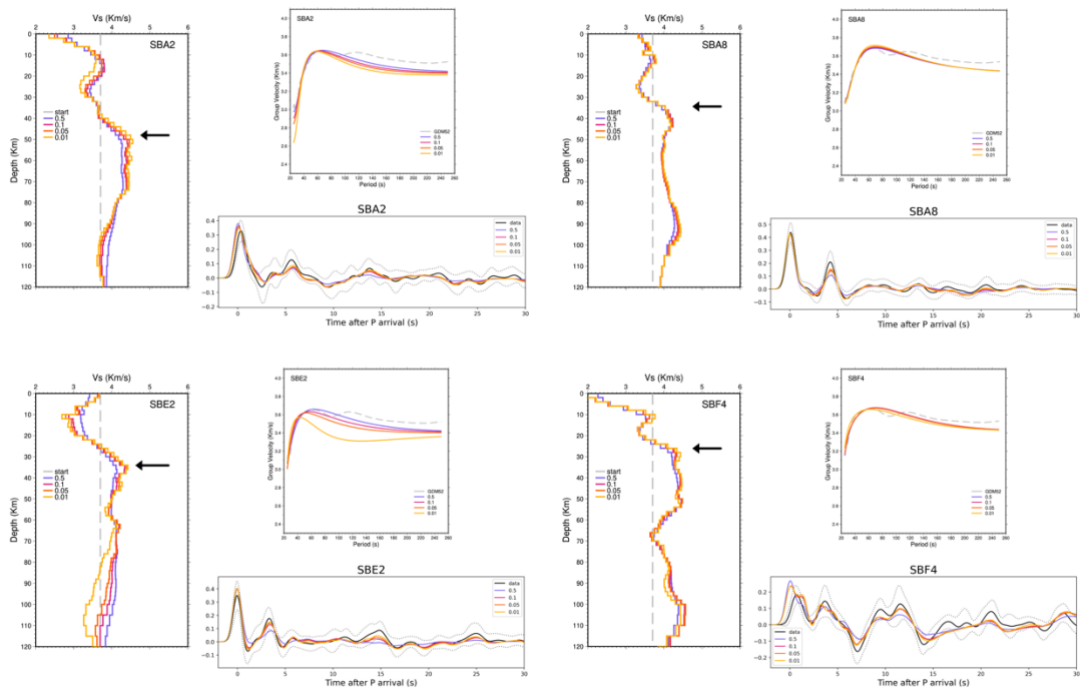
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1231 **Supplementary figure 2:** Group velocity maps for the GDM52 model  
1232 (Ekstrom, 2011), for the periods 25s, 31.25s, 50s, 62.5s, 80.88s, 125s, and  
1233 250s.



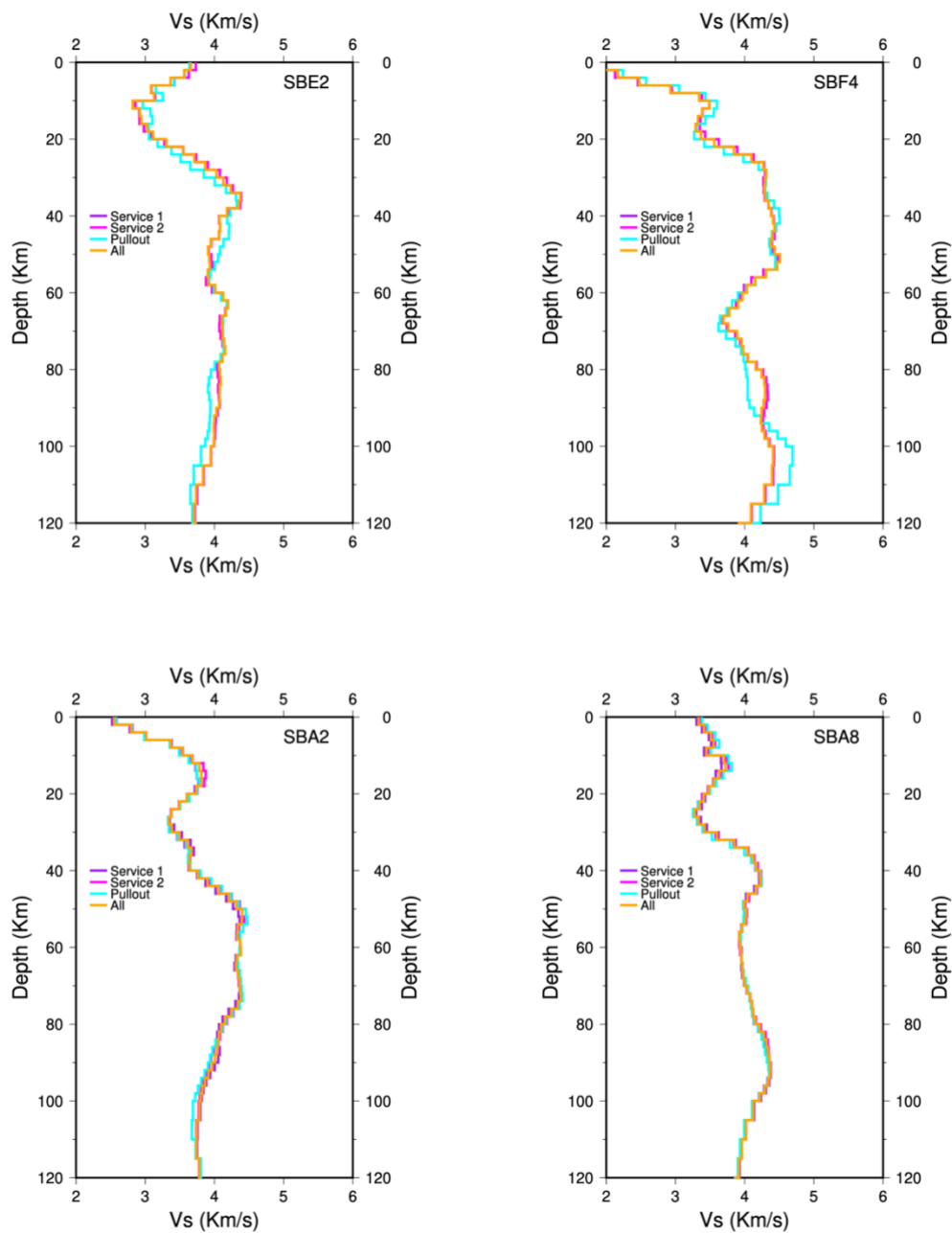
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1235 **Supplementary figure 3:** Examples of models of shear velocity vs depth from  
1236 the joint inversion of surface wave and receiver function data four stations  
1237 SBA2, SBE2, SBF4, and SBA8 with a p value of 0.1. On each shear velocity  
1238 model the coloured lines show the different models that result from testing  
1239 different starting models: 4.48 km/s, 4.28 km/s, and 3.70 km/s.  
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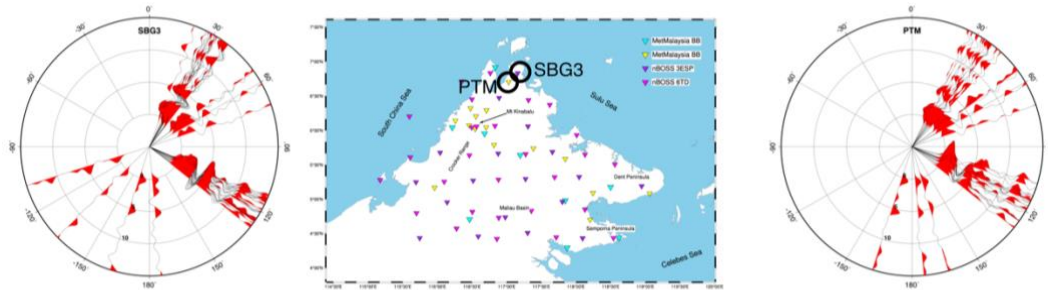
1242 **Supplementary figure 4:** Examples of models of shear velocity vs depth from  
 1243 the joint inversion of surface wave and receiver function data, and the receiver  
 1244 functions for these models for four stations across Sabah: SBA2, SBE2,  
 1245 SBF4, and SBA9. On each of the shear velocity and receiver function plots the  
 1246 coloured lines show the results from testing different weights of receiver  
 1247 function and surface wave data. On the shear velocity plots, the black arrow  
 1248 indicates the depth that is picked for the Moho in each example. On the  
 1249 receiver function plots, the receiver function data are shown in black.  
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**Supplementary figure 5:** Examples of models of shear velocity vs depth from the joint inversion of surface wave and receiver function data four stations SBA2, SBE2, SBF4, and SBA8 with a p value of 0.1. On each shear velocity model the coloured lines show the different models that result from testing different subsets of the receiver function data: Service 1 (purple) is from the inversion of stacked receiver functions for events between March 2018-Sept 2018, Service 2 (magenta) is from the inversion of stacked receiver functions from events between Sept 2018-March 2019, Pullout is from the inversion of

1260 stacked receiver functions from events between March 2019-Jan 2020, and  
1261 All is from the inversion of stacked receiver functions for the whole time  
1262 period, as shown in Supplementary figure 1. For each station, the same  
1263 surface wave dispersion data was used for each of the inversions.  
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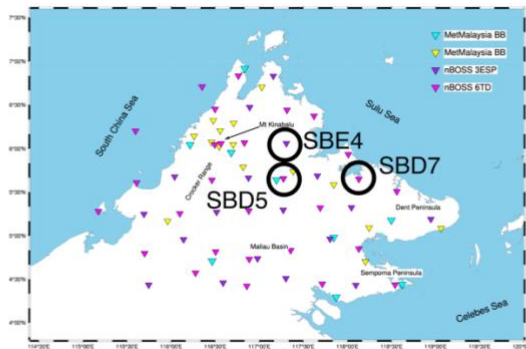
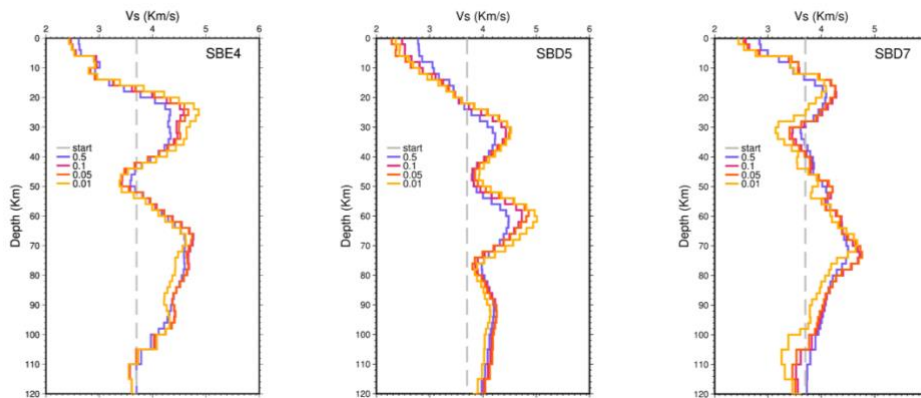
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1267 **Supplementary figure 6:** Examples of the receiver functions of individual  
1268 events used in the station stacks for SBG3 (left) and PTM (right) and a map  
1269 indicating the station locations. Receiver functions are plotted with respect to  
1270 backazimuth with positive amplitudes filled red.

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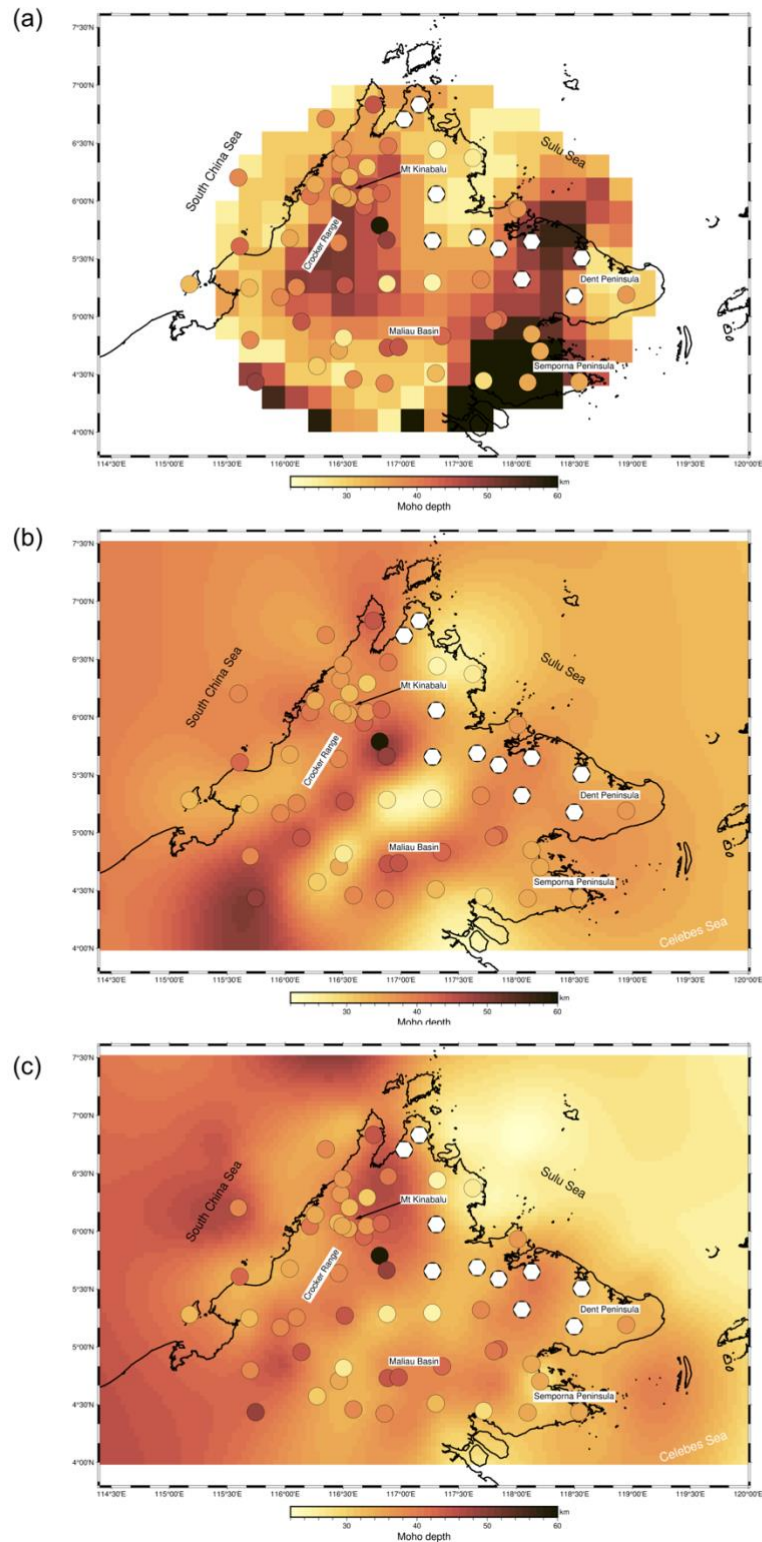
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**Supplementary figure 7:** Examples of models of shear velocity vs depth from the joint inversion of surface wave and receiver function data for stations where a Moho was challenging to identify in this study. SBE4 and SBD5 are examples of stations with two potential discontinuities, while SBD7 shows a gradual increase in velocities over a wide depth range. On each shear velocity model the coloured lines show the different models that result from testing different weights of receiver function and surface wave data.



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**Supplementary figure 8:** Comparison of Moho depths from this study (circles) with other Moho depth estimates from other studies of the region. (a) comparison with the Moho depth from Greenfield et al., (2022) based on the 4.1km/s velocity contour in their shear velocity model, (b) comparison with the Moho depth from Linang et al, (2022) from the interpolation of crustal depths obtained from stacked VDSS traces, assuming the depths reflect the Moho beneath stations, (c) comparison with the Moho depth from Linang et al,

1295 (2022) from the interpolation of crustal depths at reflection points in the VDSS  
1296 method.  
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