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5 **Mesozoic drift of the Wrangellia superterrane revisited:** 6 **the way forward from paleomagnetic data**

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8 **Goran Andjić^{1,2*}, Bram Vaes¹, Suzanna H.A. van de Lagemaat¹, Lydian M. Boschman¹,**
9 **Mark J. Dekkers¹, Stephen T. Johnston³, and Douwe J.J. van Hinsbergen¹**

10
11 ¹*Department of Earth Sciences, Utrecht University, Princetonlaan 8A, 3584 CB Utrecht, the*
12 *Netherlands*

13 ²*Institute of Earth Sciences, University of Lausanne, Géopolis, 1015 Lausanne, Switzerland*

14 ³*Department of Earth and Atmospheric Sciences, University of Alberta, Edmonton, AB T6G*
15 *2E3, Canada*

16
17 *Corresponding author: Goran Andjić (goran.andjic@unil.ch)

18 19 **Key Points:**

- 20
- 21 • New and compiled paleomagnetic data from Mesozoic rocks of the Wrangellia
22 superterrane show consistent inclinations and declinations
 - 23 • Two new tectonic scenarios describe Wrangellia's drift from a Triassic subequatorial
24 to a mid-Cretaceous, mid-latitude accreted position
 - 25 • Robust and reproducible paleomagnetic data of Wrangellia are useful for solving the
26 Baja-British Columbia controversy

27 28 **Abstract**

29
30 The allochthonous origin of the Wrangellia superterrane relative to North America has been
31 established in the early days of plate tectonics using paleomagnetic and geological data.
32 However, long-standing disagreement between paleomagnetic and structural studies on

33 magnitude of northward translation of the Wrangellia superterrane during the Latest
34 Cretaceous–earliest Cenozoic has cast doubt on the validity of the paleomagnetic data of the
35 superterrane, including data from Paleozoic and Mesozoic rocks. Here, we compile all
36 paleomagnetic data from the superterrane and present new results from uppermost Triassic
37 limestones and lowermost Jurassic lavas of the Bonanza arc, which confirm that the Wrangellia
38 superterrane was at those times at a much lower latitude than today, either $\sim 25\text{--}35^\circ$ North or
39 $\sim 25\text{--}35^\circ$ South. Moreover, declinations reveal a coherent, major clockwise or
40 counterclockwise rotation, depending on hemispheric origin. When correcting for previously
41 documented true polar wander—the wholesale rotation of the solid Earth relative to the Earth’s
42 spin axis—at the approximate longitude of the Wrangellia superterrane, new and existing
43 paleomagnetic data allows for two possible scenarios of Mesozoic kinematic evolution: from
44 190 Ma to 80 Ma, the Wrangellia superterrane was either transported ~ 5000 km northward
45 while rotating $\sim 110^\circ$ clockwise at a north-dipping subduction zone or remained at northern
46 middle latitudes while rotating $\sim 70^\circ$ counterclockwise at a south-dipping subduction zone. The
47 robust and reproducible Triassic–lowermost Jurassic and Cretaceous paleomagnetic data make
48 previously speculated systematic artifacts unlikely solutions for the kinematic debate on the
49 Late Cretaceous to Eocene tectonic history of the Wrangellia superterrane.

50

51 **1. Introduction**

52

53 The Cordilleran orogen of northwestern North America is among the largest
54 accretionary orogenic complexes that formed since Mesozoic time. The orogen developed at a
55 series of subduction zones that collectively accommodated plate convergence between the
56 Farallon and other oceanic Panthalassa plates, and the continental North America Plate
57 (Engelbreton et al., 1985; Nokleberg et al., 2000; Sigloch & Mihalynuk, 2013; van der Meer
58 et al., 2012). Kinematic reconstruction of the Cordilleran orogen of North America is important
59 for connecting surface geological processes to mantle dynamics (Sigloch & Mihalynuk, 2017;
60 van der Meer et al., 2010, 2012), exploring economic resources (Nokleberg et al., 2005), and
61 modelling the Earth’s paleogeography (Scotese, 2021) and paleoclimate (Caruthers et al., 2021;
62 Dal Corso et al., 2020). Restoring the kinematic history of the Cordilleran orogen requires
63 quantifying the amount, timing, and direction of past displacement of the fault-bounded crustal
64 fragments that compose the orogen. This appears at first glance straightforward: the Cordilleran
65 orogen exposes terranes that sutured against the North American margin in the Mesozoic, and

66 that were transported northward since Late Cretaceous times along margin-parallel strike-slip
67 faults, accommodated by complex deformation and oroclinal bending in Alaska (e.g., Johnston,
68 2001). Such northward motions may be quantified from paleomagnetic data demonstrating
69 paleolatitudinal motions, and structural estimates from fault displacements. However, this led
70 to a long-standing, and unsolved problem: terrane displacements deduced from paleomagnetic
71 data are much larger than those obtained from structural data on faults that quantify relative
72 displacements between exposed rock units (Gabrielse et al., 2006).

73 The paleomagnetic point of view has been summarized as the Baja-British Columbia
74 (Baja-BC) hypothesis. Paleomagnetic data from the western terrane of the Cordilleran orogen
75 that currently makes up much of British Columbia—the so-called Wrangellia superterrane—were
76 during the Late Cretaceous (~80 Ma) at the same latitude as Baja California (~30°N), implying
77 a ~2000 km northward motion since the Late Cretaceous (Irving, 1985; Umhoefer, 1987).
78 Field-based, structural geological correlations, however, so far cannot account for more than
79 800 km of displacement since that time (Gabrielse et al., 2006). If the conceptual solution is to
80 be found in the geological estimates, a plate boundary-scale fault has remained unrecognized
81 in the geological record (e.g., Cowan et al., 1997; Johnston, 2008). If the solution lies in the
82 estimate of paleolatitude based on paleomagnetic data, then the Wrangellia superterrane should
83 have suffered from a systematic, previously unrecognized regional paleomagnetic bias—such
84 as tectonically induced remagnetization of igneous and sedimentary rocks and inclination
85 flattening of sedimentary rocks unaccounted for (Butler et al., 2001b)—or there may have been
86 flaws in the statistical procedures or the reference apparent polar wander path that constrains
87 the paleolatitudinal position of North America (e.g., Kent & Irving, 2010; Torsvik et al., 2012).
88 Finally, the use of paleomagnetic data obtained from intrusive rocks (e.g., Beck & Noson,
89 1972; Rusmore et al., 2013), from which interpretation is complicated by lack of control on
90 paleohorizontal at the time of magnetization, has cast further doubt on the pertinence of using
91 paleomagnetic data to quantify the northward transport of segments of the North American
92 Cordillera (Butler et al., 2001b).

93 In this paper, we aim to re-assess the coherence and consistency of paleomagnetic data
94 from the Wrangellia superterrane, and whether a paleomagnetic artifact or misinterpretation
95 may lie at the heart of the Baja-BC problem. We therefore compile paleomagnetic data from
96 the Wrangellia superterrane, which has a long history that predates the Upper Cretaceous:
97 paleomagnetic data are available from rocks dating back to the early Paleozoic. The basement
98 of the Wrangellia superterrane consists of a continental arc basement and sedimentary rocks

99 that are overlain by a thick pile of Triassic lavas interpreted as a large igneous province, which
100 is in turn overlain by the Late Triassic to Middle Jurassic “Bonanza” arc (Nokleberg et al.,
101 2000). East, north, and south of the Wrangellia superterrane are the Intermontane terrane of the
102 Canadian Cordillera and the Franciscan accretionary prism of Oregon and California,
103 respectively, which both formed as part of the North American margin and consist of
104 metamorphosed continental margin rocks, ophiolites, and accreted complexes that were
105 amalgamated in a subduction zone between the Wrangellia superterrane and the North
106 American continent since the Early Jurassic (Nokleberg et al., 2000; Wakabayashi, 2015). The
107 Wrangellia superterrane is therefore thought to have been located in an upper plate position
108 relative to a subduction zone that formed the Bonanza arc, but in a downgoing plate position
109 relative to North America, until its collision with North America, sometime in the middle to
110 Late Cretaceous (Nokleberg et al., 1994, 2000; Plafker et al., 1989; Tikoff et al., 2023). Such
111 a Mesozoic double subduction system is consistent with slab remains imaged by seismic
112 tomography that revealed parallel belts of mid-mantle slabs below western North America and
113 the eastern Pacific (Clennett et al., 2020; Fuston & Wu, 2021; Sigloch & Mihalynuk, 2013,
114 2017; Sigloch et al., 2008; van der Meer et al., 2010, 2012, 2018). The extensive rock records
115 of the Wrangellia superterrane dating back to the Paleozoic thus provides the opportunity to
116 constrain the paleolatitudinal and rotational components of the pre-collisional plate motion
117 history relative to North America.

118 Interestingly, paleomagnetic data from middle Paleozoic to Lower Jurassic (~450–180
119 Ma) rocks of the Wrangellia superterrane show that it was in a subequatorial position at these
120 times (~0–25°; Bazard et al., 1995; Irving & Yole, 1987; Kent & Irving, 2010; and references
121 therein), whereby northern and southern hemisphere options are both possible (Panuska &
122 Stone, 1981). Both options suggest significant northward motion in the ~190–80 Ma interval,
123 preceding the controversial Baja-BC problem. In this study, we compiled all paleomagnetic
124 data from the Wrangellia superterrane, covering the Paleozoic to Cenozoic, and collected a
125 new, large paleomagnetic dataset from uppermost Triassic sedimentary rocks and Lower
126 Jurassic Bonanza arc lavas of the southern Wrangellia superterrane. We use these to test the
127 reproducibility of paleomagnetic data from the youngest Mesozoic age interval for which
128 previous data revealed a subequatorial position. We compare these data to an updated global
129 apparent polar wander path (APWP; Vaes et al., 2023) and use data comparison methods (Vaes
130 et al., 2021, 2022) that overcome statistical issues of classic paleomagnetic approaches
131 (Rowley, 2019).

132 Not only do we test whether the inclinations, constraining paleolatitudinal position
133 through time, provide a coherent pattern, but we also evaluate whether the declinations,
134 constraining vertical axis rotations, provide a coherent history. We test rates of reconstructed
135 northward motion of the Wrangellia superterrane against recent reconstructions of minimum
136 oblique subduction components obtained from kinematic evidence from the Californian forearc
137 ophiolites (Arkula et al., 2023), and from Ocean Plate Stratigraphy (OPS) accreted in the
138 Californian Franciscan subduction complex (Alvarez et al., 1980; Courtillot et al., 1985;
139 Tarduno et al., 1985, 1986). Finally, we evaluate whether reconstructed positions of the
140 Bonanza arc at the time of the end of subduction coincide with the presence of slab remnants
141 in mantle tomography. The updated statistical paleomagnetic procedures and reference frame,
142 as well as the multiple independent tests, will determine whether paleomagnetic artifacts lie at
143 the heart of the Baja-BC problem, allowing us to propose ways forward in reconciling
144 paleomagnetic and structural data.

145

146 **2. Geological context**

147

148 *2.1. The Wrangellia superterrane*

149

150 From western Alaska to southeastern British Columbia, the Wrangellia superterrane is
151 subdivided into the Peninsular, Wrangellia, and Alexander terranes (Figure 1a), which overall
152 share a multi-stage history of arc magmatism that spans the Late Paleozoic and most of the
153 Mesozoic (~360–100 Ma; Alberts et al., 2021; Nokleberg et al., 1994; Plafker & Berg, 1994).
154 The amalgamation of the three terranes is thought to have occurred during the Paleozoic, prior
155 to their Mesozoic accretion to North America. Below, we summarize the main geological
156 features of each terrane and key moments in their tectonic evolution.

157 The core of the Wrangellia superterrane is the Alexander terrane, which is interpreted
158 as a microcontinental fragment that was intruded by late Neoproterozoic to early Paleozoic arc-
159 related rocks (~600–400 Ma; Gehrels & Saleeby, 1987; White et al., 2016). Volcanic and
160 plutonic rocks dominate the southern part of the Alexander terrane, whereas the northern part
161 consists mainly of Paleozoic shelf strata as old as the upper Cambrian (Beranek et al., 2012).
162 Magmatic activity in the Alexander terrane was interrupted during two orogenic events, the
163 Cambrian Wales orogeny and the lowermost Devonian Klakas orogeny (Gehrels & Saleeby,
164 1987).

165 The Alexander terrane is bordered to the south and the north by the younger Wrangellia
166 terrane, which consists mainly of upper Devonian to lower Permian and upper Triassic to
167 middle Cretaceous arc-related igneous and sedimentary rocks (Alberts et al., 2021; Nokleberg
168 et al., 1994; Plafker & Berg, 1994). Although basement rocks underlying the Paleozoic arc
169 have not been observed, detrital zircon data from the southern Wrangellia terrane suggest that
170 it incorporated fragments of the Alexander terrane (Alberts et al., 2021). Geological ties
171 between the Alexander and Wrangellia terranes dating back as early as ~360 Ma (Late
172 Devonian) are corroborated by coeval gabbro complexes intruding both terranes (Israel et al.,
173 2014).

174 An episode of plume-related volcanism occurred between the Paleozoic and Mesozoic
175 arc phases of the Wrangellia terrane, forming the up to 6 km-thick Wrangellia large igneous
176 province (Upper Triassic, ~232–226 Ma; Greene et al., 2010), which spans the whole length of
177 the Wrangellia terrane. In our study area on Vancouver Island, the Bonanza arc was built on
178 these flood basalts and associated sedimentary rocks (Canil et al., 2010, 2013; DeBari et al.,
179 1999; D’Souza et al., 2016). The Bonanza Group consists here of a basal Upper Triassic–
180 lowermost Jurassic volcanic-sedimentary succession (Parson Bay Formation and
181 Volcaniclastic-Sedimentary unit; ~226–200 Ma, based on detrital zircon U-Pb geochronology
182 and biostratigraphy) overlain by the Lower Jurassic Le Mare Lake Volcanic Unit (~201–190
183 Ma, based on zircon U-Pb geochronology and biostratigraphy) and the Lower to Middle
184 Jurassic Holberg Volcanic Unit (~201–164 Ma, based on detrital zircon U-Pb and amphibole
185 Ar-Ar geochronology; Nixon & Orr, 2007; Nixon et al., 2011a, 2011b, 2011c, 2011d). The
186 latter volcanic units are interbedded with minor marine and non-marine epiclastic rocks and
187 limestone. The bulk of the Bonanza arc complex was intruded by the Lower to Middle Jurassic
188 Island Plutonic Suite (~201–164 Ma, based on zircon U-Pb and amphibole Ar-Ar
189 geochronology; D’Souza et al., 2016; Nixon et al., 2011a, 2011b, 2011c, 2011d). The Le Mare
190 Lake Volcanic Unit and the Island Plutonic Suite represent the main phase of growth of the
191 Bonanza arc. On Vancouver Island, Triassic and younger units have experienced burial
192 metamorphism to zeolite and prehnite-pumpellyite facies, implying temperatures below 350°C
193 and pressures below 2.5 kbar (<10 km depth; Kuniyoshi & Liou, 1976; Lei et al., 2020; Morris
194 & Canil, 2021; Stewart & Page, 1974).

195 Except for the absence of flood basalts, the stratigraphic architecture of the Peninsular
196 terrane is similar to that of the Wrangellia terrane (Figure 1). The Peninsular terrane mainly
197 consists of Upper Triassic to Middle Jurassic volcanic, plutonic, and volcanoclastic rocks of the

198 Talkeetna arc, overlain by Upper Jurassic to Lower Cretaceous volcanoclastic basinal strata
199 (McClelland et al., 1992; Plafker et al., 1989; Rioux et al., 2007, 2010). The Talkeetna arc,
200 considered as an archetypal intraoceanic arc, is lithologically, temporally, and geochemically
201 correlative to the Bonanza arc of the Wrangellia terrane (D'Souza et al., 2016; Plafker et al.,
202 1989; Rioux et al., 2007). Scarce outcrops of metamorphosed upper Paleozoic mafic to
203 intermediate volcanic rocks, limestones, and quartz-rich sedimentary rocks provide a glimpse
204 into the basement of the Talkeetna arc (Plafker et al., 1989). Zircon xenocrysts from the
205 Talkeetna arc and younger volcanic products suggest ties with the Alexander and Wrangellia
206 terranes from at least the early Carboniferous (~310 Ma; Amato et al., 2007; Bacon et al., 2012;
207 Beranek et al., 2014).

208 Based on stratigraphic, structural, geochemical, and geochronological data from arc and
209 accretionary complex rocks, the polarity of the subduction beneath the Wrangellia superterrane
210 is thought to have been northward in western Alaska to eastward in British Columbia (in
211 present-day coordinates) at least since ~200 Ma (Amato et al., 2013; Clift et al., 2005a, 2005b;
212 Plafker & Berg, 1994; Trop & Ridgway, 2007). Notably, blueschist-facies metamorphic rocks
213 were formed during the Early Jurassic in a subduction complex now juxtaposed against the
214 seaward margin of the superterrane (Roeske et al., 1989; Sisson & Onstott, 1986).

215 The timing and latitude, and possible diachroneity of accretion of the Wrangellia
216 superterrane to North America are not yet resolved. Basinal records in the suture zone between
217 the northern Wrangellia superterrane and the Intermontane superterrane, which is located east
218 of it, have been used to suggest timings of collision that include the Middle Jurassic (e.g.,
219 McClelland et al., 1992), the Late Jurassic (e.g., Trop & Ridgway, 2007), the Early Cretaceous
220 (e.g., Hampton et al., 2010), the middle Cretaceous (e.g., Amato et al., 2013; Plafker et al.,
221 1989), and the Late Cretaceous (e.g., Hulst et al., 2013). Proponents of diachronous accretion
222 suggest that the southern part of the Wrangellia superterrane in British Columbia collided
223 during the Jurassic, whereas the northern part of the superterrane in Alaska collided during the
224 Late Cretaceous (e.g., Manselle et al., 2020; Trop & Ridgway, 2007). These widely differing
225 accretion scenarios mostly differ from each other in the amount of post-subduction translation
226 that the Wrangellia superterrane experienced that has been interpreted based on paleomagnetic
227 studies (Kent & Irving, 2010).

228 Notably, estimates of post-accretion translation of the superterrane based on detrital
229 zircon data from northern Washington, southern British Columbia, and southern Alaska have
230 not reached any consensus, with offsets ranging from ~500 km (Mahoney et al., 1999, 2021),

231 through ~1000 km (Yokelson et al., 2015), to >1500 km (Boivin et al., 2022; Housen & Beck,
232 1999; Matthews et al., 2017; Sauer et al., 2019). Among these studies, the main disagreement
233 revolves around the source of detrital zircons from the Upper Cretaceous Nanaimo Group of
234 Vancouver Island, which may have been sourced from rocks outcropping in Idaho (~300–600
235 km south from Vancouver Island) and/or southern California (~1500–1800 km south from
236 Vancouver Island; Boivin et al., 2022, and references therein). The latter option is compatible
237 with paleomagnetic data from the Upper Cretaceous Nanaimo Group, which have consistently
238 yielded paleolatitudes ~1600–2500 km south of its present-day location (Kim & Kodama,
239 2004; Krijgsman & Tauxe, 2006; Ward et al., 1997).

240

241 *2.2. Jurassic ophiolites and OPS of California*

242

243 Westernmost California and southern Oregon display a Mesozoic–Cenozoic subduction
244 complex that formed by episodic accretion of Ocean Plate Stratigraphy (OPS), which records
245 the history of a subducting oceanic plate from its formation to arrival at the trench (Isozaki et
246 al., 1990). This subduction complex—the Franciscan Complex—accreted below Middle to Upper
247 Jurassic (~170–155 Ma) supra-subduction zone ophiolites that are preserved as isolated klippen
248 (Wakabayashi, 2015). The oldest accreted rocks are high-temperature, high-pressure
249 metabasites interpreted as metamorphic sole rocks with Lu/Hf garnet ages of 180 Ma that mark
250 the (minimum) age of subduction initiation (Mulcahy et al., 2018) and show that the ophiolites
251 formed above an active subduction zone (e.g., Guilmette et al., 2018), consistent with their
252 geochemical composition (Snortum & Day, 2020). Arkula et al. (2023) reconstructed western
253 North American deformation to restore the relative positions of the Californian ophiolites in
254 the Jurassic. They also showed paleomagnetic results that imply that the forearc paleo-ridges
255 that generated the Jurassic ophiolites had near-perpendicular orientations to that of the
256 Franciscan subduction zone. The kinematic restoration of the ophiolite belt shows that the
257 paleo-ridges may have accommodated spreading rates of ~6 cm/yr, suggesting that the plate
258 subducting obliquely beneath California in the Jurassic had a northward motion relative to
259 North America of up to ~6–7 cm/yr.

260 The Franciscan Complex consists of rocks accreted from oceanic crust that formed
261 since the Early Jurassic (Wakabayashi, 2015). The youngest accretion may have happened as
262 recently as 12 Ma (McLaughlin et al., 1982). Among the accreted rock assemblages, two
263 localities with middle Cretaceous limestone blocks provided paleomagnetic results that allow

264 computing a paleolatitudinal journey of the OPS prior to accretion. The Laytonville limestone
265 (103–90 Ma) yielded paleolatitudes of $17^{\circ} \pm 7^{\circ}$ (Alvarez et al., 1980) and $14^{\circ} \pm 5^{\circ}$ (Tarduno et
266 al., 1986). The Calera limestone (129–90 Ma), located ~300 km southeast of Laytonville,
267 yielded paleolatitudes of $24^{\circ} \pm 4^{\circ}$ (Courtilot et al., 1985) and 18–25° at 105–90 Ma (Tarduno
268 et al., 1985). Because the blocks may have rotated during accretion to the Franciscan Complex,
269 it is not obvious whether they formed on the southern or northern hemisphere. Also, the
270 accretion age of the limestone blocks to the Franciscan Complex has uncertainties. All
271 scenarios show that in Early to Late Cretaceous times, the OPS sequences that accreted to the
272 Franciscan Complex were derived from lower latitudes, moving north at rates varying from ~8
273 cm/yr for the northern hemisphere options (Courtilot et al., 1985; Tarduno et al., 1985) to \geq
274 15 cm/yr for the southern hemisphere options (Alvarez et al., 1980; Tarduno et al. 1986).
275 Additionally, OPS sequences with ages ranging from 180 to 110 Ma in the Santa Elena
276 complex—part of the Chortis block that was connected to the North American continent (Andjić
277 et al., 2019)—of western Costa Rica which accreted around 100 Ma at 11°N yielded
278 paleolatitudes of 8–20° north or south (Boschman et al., 2021a).

279

280 **3. Paleomagnetic sampling and measurements**

281

282 To test the robustness of the paleomagnetic data from the youngest pre-Cretaceous
283 rocks of the Wrangellia superterrane, we collected a total of 295 cores for paleomagnetic
284 analysis from four localities in the Bonanza Group of northern Vancouver Island, Canada
285 (Figure 2; Nixon et al., 2011a, 2011d). Locality YM (“Yreka Main”) was sampled in a
286 continuous ~50 m section of impure limestones of the Norian–Rhaetian Parson Bay Formation,
287 from which 119 samples were collected (average dip direction = 230°, average dip = 50°, n =
288 8). Basaltic to andesitic lava flows of the Hettangian–Sinemurian Le Mare Lake Volcanic Unit
289 were sampled in three sections (MD = “Main Drive”, TT = “Teeta Creek”, VL = “Victoria
290 Lake”) that are ~5 km apart, with each section consisting of ~15 m (TT: average dip direction
291 = 226°, average dip = 38°, n=3) to ~50 m (MD: average dip direction = 273°, average dip =
292 29°, n = 14 ; VL: average dip direction = 196°, average dip = 33°, n = 12) of lavas. We used a
293 gasoline-powered motor drill to sample 2.5 cm-diameter paleomagnetic cores, the orientation
294 of which was measured with a magnetic compass with an inclinometer attached. We followed
295 procedures recommended by Gerritsen et al. (2022) and drilled one core per limestone bed or
296 volcanic flow (Figure S1 in Supporting Information S1) to optimize the amount of individual

297 spot readings of the magnetic field, and as a field test selected a total of seven lava sites where
298 5 cores per lava flow were drilled to evaluate whether within-site scatter of paleomagnetic data
299 is low (i.e., k values typically exceeding 50; e.g., Johnson et al., 2008). Measurements were
300 corrected for the local declination ($16^{\circ}47'$ E to $16^{\circ}48'$ E).

301 The cores were processed at the Paleomagnetic Laboratory Fort Hoofddijk at Utrecht
302 University, The Netherlands. The cores were cut into 2.2 cm-long samples using a double-
303 blade circular saw. To determine the nature of magnetic carriers for both types of sampled
304 lithology (lavas and impure limestones), thermomagnetic analyses were performed using a
305 horizontal translation-type Curie balance with a sinusoidally cycling applied magnetic field,
306 usually 100–300 mT (Mullender et al., 1993). Several heating-cooling cycles were applied to
307 detect magneto-mineralogical alterations during heating. We used the following temperature
308 scheme (in $^{\circ}\text{C}$): 150, 75, 225, 150, 300, 225, 375, 300, 450, 375, 525, 450, 600, 20 (for lavas);
309 250, 150, 350, 250, 450, 350, 520, 420, 620, 500, 700, 20 (for limestones). Stepwise thermal
310 (TH) demagnetization was applied to 117 limestone samples and 30 lava samples, whereas
311 stepwise alternating field (AF) demagnetization was applied to 72 limestone samples and 174
312 lava samples, the latter processed with a robotized magnetometer (Mullender et al., 2016).
313 Natural remanent magnetizations (NRM) were measured on a 2G DC SQUID magnetometer.
314 Temperature steps of 100, 180, 210, 240, 270, 300, 320, 340, 360, 380, and 400°C were used
315 for TH treatment of 87 limestone samples. Temperature steps of 100, 180, 210, 240, 270, 300,
316 330, 360, 390, 420, 450, 480, and 510°C were used for TH treatment of 30 other limestone
317 samples. Temperature steps of 100, 180, 210, 240, 270, 300, 330, 360, 390, 420, 450, 480, 510,
318 540, 570, and 600°C were used for TH treatment of the 30 lava samples. Demagnetization steps
319 of 5, 10, 15, 20, 25, 30, 35, 40, 45, 50, 60, 70, 80, 90, 100, and 120 mT were used for AF
320 treatment of all samples. To improve the resolution of AF demagnetization results of
321 limestones, the latter were heated to 150°C in a thermal demagnetizer prior to AF treatments
322 (van Velzen & Zijdeveld, 1995).

323 Sample interpretation and statistical analysis were conducted using the online portal
324 Paleomagnetism.org (Koymans et al., 2016, 2020). All results can be imported into the portal
325 from data files (.col) available in the Supporting Information (Data Sets S1 and S2), as well as
326 in the Paleomagnetism.org 2.0 database (Koymans et al., 2020) and the MagIC database
327 (Jarboe et al., 2012). Demagnetization diagrams were plotted on orthogonal vector diagrams
328 (Zijdeveld, 1967), and the magnetic components were determined through principal
329 component analysis (Kirschvink, 1980). Great circle solutions were determined using the

330 method of McFadden and McElhinny (1988). The fold test (Tauxe & Watson, 1994) and the
331 bootstrapped coordinate reversal test (Tauxe, 2010) were used when applicable. The
332 elongation-inclination (E/I) correction for inclination shallowing (Tauxe & Kent, 2004; Tauxe
333 et al., 2008) was applied to the sedimentary locality, but its result was not used in our
334 interpretations, as discussed in section 4.2. A maximum angular deviation cut off (i.e., $MAD \leq$
335 15°) was not applied to our dataset because it only reduces the number of samples, which
336 decreases paleopole precision (Gerritsen et al., 2022). Paleomagnetic pole positions were
337 calculated using Fisher (1953) statistics on virtual geomagnetic poles (VGPs)—whereby each
338 VGP is derived from a single site—following statistical procedures described in Deenen et al.
339 (2011), providing a measure of the VGP dispersion (Fisher precision parameter, K) and a 95%
340 confidence ellipse on the pole position (A_{95}). The mean paleomagnetic direction and the 95%
341 confidence regions on the declination (ΔD_x) and inclination (ΔI_x) were computed from the
342 paleomagnetic pole and its A_{95} . A 45° cutoff was applied to the VGPs (Johnson et al., 2008) at
343 the group level for the volcanic localities (TT, MD, VL). We did not apply this cutoff to the
344 sedimentary dataset (YM), as that the application of this cutoff would not lead to the exclusion
345 of any of the data points.

346

347 **4. Results**

348

349 *4.1. Volcanic rocks (samples VL-TT-MD)*

350

351 Thermomagnetic curves show that lava samples have Curie temperatures close to
352 580°C (Figure 3), indicating that magnetite is the main magnetic carrier. In orthogonal vector
353 plots, the majority of the VL (58/80) and MD (32/54) samples reveal an overprint at low
354 temperature/coercivity steps (up to $\sim 100\text{--}240^\circ\text{C}$ or $5\text{--}15\text{ mT}$). VL samples (46/80) have an
355 additional overprint at mid-range temperature/coercivity steps ($\sim 420\text{--}570^\circ\text{C}$ or $15\text{--}90\text{ mT}$),
356 most of which (40/46) did not yield characteristic remanent magnetization (ChRM) directions.
357 In contrast, TT samples show a linear decay towards the origin of demagnetization diagrams
358 without overprints. Initial intensities ranged $7\text{--}550\text{ mA/m}$ for TT, $0.2\text{--}38\text{ mA/m}$ for MD, and
359 $0.6\text{--}1650\text{ mA/m}$ for VL. ChRM values were generally interpreted between $420\text{--}570^\circ\text{C}$ or 10--
360 120 mT for TT, $\sim 270\text{--}510^\circ\text{C}$ or $\sim 25\text{--}70\text{ mT}$ for MD, and $\sim 240\text{--}510^\circ\text{C}$ or $\sim 20\text{--}70\text{ mT}$ for VL.
361 Interpreted ChRM directions from most samples yielded eastward declinations with shallow
362 upward inclinations (in tilt-corrected coordinates; Figure 3); one TT sample and ten MD

363 samples yielded westward, down directions that suggest opposite polarity. Which of these two
 364 directions represent normal or reversed is not a priori known, and depends on the hemisphere
 365 of origin. Opposite-polarity ChRM directions of locality MD share a common true mean
 366 direction (CTMD); reversals are located throughout the section at samples MD1.9, MD1.23,
 367 MD1.24, and MD1.39–MD1.44. Mild differences in bedding orientation among the three
 368 localities were used for a regional fold test, but bedding differences are insufficient to yield a
 369 conclusive result (best clustering between 59 and 78% unfolding).

370 The lava sites from which we collected multiple samples per site returned low
 371 dispersion results (e.g., $k = 395.9$ for TT 1.31–1.35 and $k = 54.8$ for MD 1.16–1.2; Figure 3),
 372 confirming that lava sites may be treated as spot readings of the field. The paleomagnetic pole
 373 computed from the lava sites is located at latitude = 35.0°S , longitude = 340.8°E ($N = 117$, K
 374 = 18.0 , $A_{95} = 3.2^\circ$), corresponding to a paleomagnetic direction of *Declination* (D) $\pm \Delta D_x =$
 375 $102.6^\circ \pm 3.9^\circ$, *Inclination* (I) $\pm \Delta I_x = -54.2^\circ \pm 3.2^\circ$ (Figure 3, Table 1). The D and I values of
 376 the combined volcanic localities differ significantly from the recent GAD field ($D/I =$
 377 $000^\circ/70^\circ$). The A_{95} value ($A_{95\text{min}} = 1.8^\circ < A_{95} = 3.2^\circ < A_{95\text{max}} = 4.1^\circ$) satisfies the N -dependent
 378 reliability envelope of Deenen et al. (2011), suggesting that the observed VGP scatter can be
 379 straightforwardly explained by paleosecular variation (PSV).

380

381 4.2. Sedimentary rocks (samples YM)

382

383 Thermomagnetic curves of the limestones show two types of magnetic carriers.
 384 Samples YM 1.91 and YM 1.118 have very low initial magnetization, even diamagnetic, i.e.,
 385 the signal is dominated by the quartz glass sample holder. The curves are completely reversible
 386 after cycling to 250°C and show minor decay after the 350°C cycle, with a pyrite-to-magnetite
 387 oxidation signature developing after 420°C (e.g., Passier et al., 2001). On cooling to room
 388 temperature after the 700°C cycle, a discontinuity at $\sim 320^\circ\text{C}$ is observed, the Curie temperature
 389 of pyrrhotite (Figure 4 and Figure S2). In contrast, samples YM 1.52 and YM 1.61 present a
 390 much higher (~ 50 – 100 times) initial magnetization, with smooth curves that are essentially
 391 reversible up to 500°C , which supports the presence of titanomagnetite with a variable Ti
 392 content (Figure 4 and Figure S2). Interestingly, both groups of samples have similar values of
 393 their initial NRM (YM 1.52 = 3.3 mA/m, YM 1.61 = 4.1 mA/m vs. YM 1.91 = 6.0 mA/m, YM
 394 1.118 = 5.1 mA/m), which suggests that the efficiency of the NRM acquisition mechanism
 395 differs widely between the two groups. YM1.91 and YM1.118 feature a chemical remanent

396 magnetization (CRM) which is an efficient NRM acquisition mechanism while the NRM of
 397 YM1.52 and YM1.61 is much less efficient (a lot more magnetic material is required for a
 398 similar NRM intensity) which would imply a detrital remanent magnetization (DRM).
 399 Nonetheless, after thermal demagnetization at 360°C, usually <10–15% of the initial NRM
 400 intensity remains for both types of samples. This implies that ChRM directions were interpreted
 401 from the same temperature steps for both groups.

402 Initial NRM intensities ranged 0.3–115 mA/m, with values mostly restricted to 2–6
 403 mA/m for samples from which ChRMs were interpreted. In orthogonal vector plots, magnetic
 404 components could not be determined from a subset of samples (49/189) because of erratic
 405 demagnetization behavior. The maximum applicable alternating field (120 mT) was not high
 406 enough to fully demagnetize limestone samples with AF demagnetization. Nevertheless,
 407 ChRM directions interpreted from AF demagnetization diagrams agree well with those
 408 obtained from TH demagnetization diagrams. A few samples (5/140) show an overprint at low
 409 temperature/coercivity steps (~100–210 °C or 5–20 mT). ChRM values were generally
 410 interpreted at ~210–360 °C or ~30–70 mT, yielding down, westward directions. The
 411 paleomagnetic pole computed for locality YM is located at latitude = 23.0°N, longitude =
 412 151.9°E ($N = 138$, $K = 52.7$, $A_{95} = 1.7^\circ$), providing a direction with $D \pm \Delta D_x = 278.1^\circ \pm 1.8^\circ$, I
 413 $\pm \Delta I_x = 40.2^\circ \pm 2.3^\circ$ (Figure 4, Table 1). The D and I values of YM differ significantly from
 414 that of the recent GAD field ($D/I = 000^\circ/70^\circ$).

415 Overall, we find that the results of the YM limestones are consistent with them
 416 dominantly carrying a CRM, that was acquired soon after sediment deposition, and that resulted
 417 in ChRM acquisition that integrated secular variation over a longer duration, possibly over
 418 $\geq 10^3$ yr. This is compatible with the following aspects: (i) No polarity reversals have been
 419 measured in the YM limestones; (ii) The relatively low dispersion of the VGPs of the
 420 limestones yield an A_{95} that coincides with the $A_{95\text{min}}$ of the reliability envelope of Deenen et
 421 al. (2011); (iii) After removing great circle solutions ($n = 9$) from the dataset (following Vaes
 422 et al., 2021), we performed the E/I correction of Tauxe and Kent (2004) using
 423 Paleomagnetism.org (Koymans et al., 2016, 2020), which yielded a flattening factor of 0.8 and
 424 a slightly higher inclination of 46.6° (95% bootstrapped confidence bounds of 41° to 62°;
 425 Figure S2). We note that the elongation of the YM directions (~1.68) is close to that predicted
 426 by the TK03.GAD field model (~1.78), which is compatible with a DRM contribution in a
 427 portion of the samples and/or CRM acquisition during (early) diagenesis.

428

429 *4.3. Summary of the results*

430

431 There is no independent control on the hemispheric origin and direction of rotation
432 (clockwise or counterclockwise) for the YM and combined MD–TT–VL localities, since both
433 normal and reversed polarities are common in the Norian–Rhaetian and Hettangian–
434 Sinemurian, respectively. Therefore, for each paleomagnetic dataset, there are two possible
435 solutions corresponding to either the northern or southern hemisphere, with opposite directions
436 of rotation. The northern hemisphere option for the sediment-derived dataset (YM) yields a
437 paleolatitude of $22.9^\circ \pm 1.7^\circ$ N and a $\sim 82^\circ$ counterclockwise rotation, whereas the southern
438 hemisphere option yields a paleolatitude of $22.9^\circ \pm 1.7^\circ$ S and a $\sim 98^\circ$ clockwise rotation. The
439 combined result of the igneous localities of the Bonanza Group (MD-TT-VL) yields a northern
440 hemisphere solution with an estimated paleolatitude of $34.7^\circ \pm 3.2^\circ$ N and a $\sim 77^\circ$
441 counterclockwise rotation, with a southern hemisphere solution providing a paleolatitude of
442 $34.7^\circ \pm 3.2^\circ$ S and a $\sim 103^\circ$ clockwise rotation.

443

444 **5. Updated paleomagnetic database for the Wrangellia superterrane**

445

446 We combine our new data with a database of available paleomagnetic data from the
447 Wrangellia superterrane that we compiled from the literature. We only chose data from the
448 NW-SE striking part of the Wrangellia superterrane, from southern British Columbia to eastern
449 Alaska, and left the western and northern parts of Alaska out of the compilation, as these were
450 likely strongly rotated in Late Cretaceous to Paleocene times during oroclinal bending
451 (Johnston, 2001). Our database includes datasets that contain at least eight individual directions
452 of the magnetic field (either from eight individual cooling units in magmatic rocks, or eight
453 sedimentary beds, as suggested by Meert et al., 2020), and of which the distribution of magnetic
454 directions passes the Deenen et al. (2011) criterium of representing paleosecular variation
455 ($A_{95\min} < A_{95} < A_{95\max}$). We excluded datasets from the compilation when the authors of the
456 original study interpreted the magnetic signal to represent a remagnetization (Hillhouse &
457 Grommé, 1980; Irving & Massey, 1990; Symons, 1985), in cases in which rocks with an
458 unknown paleohorizontal were sampled (Butler et al., 2001a; Irving et al., 1985; Irving &
459 Massey, 1990; Rusmore et al., 2013; Symons, 1973), or in which shearing was interpreted to
460 have influenced the magnetic directions (Butler et al., 2002).

461 Our compilation contains datasets for Early Jurassic and older times, and for Late
462 Cretaceous and younger times. There are two small datasets from the Lower Cretaceous of
463 Alaska that were collected by Stone et al. (1982), and some preliminary results provided by
464 Panuska et al. (1984). However, the original data and descriptions are not available, and the
465 datasets were considered unreliable by subsequent studies (Butler et al., 1997; Harbert, 1990;
466 Hillhouse, 1987). In addition, Butler et al. (1997) revisited Ordovician, Silurian, Devonian, and
467 Carboniferous rocks of the Wrangellia superterrane from which Van der Voo et al. (1980)
468 reported paleolatitudes. Butler et al. (1997) argued that these Paleozoic rocks were
469 remagnetized in the Triassic, and that none of these provide useful paleogeographic
470 information. We show the data of Van der Voo et al. (1980) in our compilation but we consider
471 them as potentially unreliable, as discussed further in section 6.1.

472 The final compilation consists of a total of 39 collections from 18 studies, 23 for the
473 Late Cretaceous and younger, and 16 for pre-Late Cretaceous times (Figure 6; Tables S1 and
474 S2 in Data Set S3). For Late Cretaceous and younger times, paleolatitudes are all $\sim 30^\circ$ or higher
475 and interpreted as northern hemispheric. For the older datasets, paleolatitudes can be either
476 southern or northern hemispheric, with the exception of rocks from the upper Carboniferous-
477 Permian Kiaman superchron (~ 320 – 260 Ma; Opdyke & Channell, 1996), which must be
478 reversed if they carry a primary magnetization. In Figure 6, we present both hemispheric
479 options.

480

481 **6. Discussion**

482

483 *6.1. Paleomagnetic data from the Wrangellia superterrane: primary or secondary*
484 *magnetizations?*

485

486 To evaluate the paleomagnetically permissible plate motions of the Wrangellia
487 superterrane since the Triassic, we now use the updated paleomagnetic database, which
488 includes our two new poles that are the largest paleomagnetic datasets for the Wrangellia
489 superterrane to date and whereby our sediment-based dataset is corrected for inclination
490 shallowing (except for our new sedimentary pole). First, we briefly re-evaluate whether the
491 new global APWP of Vaes et al. (2023) for the last 320 Ma provides significant modifications
492 to the paleolatitude and declination of the North American continent. This new path is based
493 on directional paleomagnetic data rather than on compilations of paleomagnetic poles that were

494 previously used (e.g., Kent and Irving, 2010; Torsvik et al., 2012; see Vaes et al., 2022, 2023,
495 for details). As illustrated in Figure 5, the global APWP of Vaes et al. (2023) gives a smaller
496 uncertainty but is mostly within error of previous APWPs (in North American coordinates).
497 We compare the data of the Wrangellia superterrane with the APWP of Vaes et al. (2023) for
498 the last 320 Ma, and with the moving average of Laurentia APWP of Torsvik et al. (2012) for
499 earlier parts of the Paleozoic (Figure 5).

500 Plotting the database against a North American reference curve, in northern and
501 southern hemisphere scenarios, allows us to re-evaluate previous hypotheses on the primary or
502 secondary nature of paleomagnetic data. First, we note that both declinations and inclinations
503 from Paleozoic to Lower Jurassic rocks of the Wrangellia superterrane are systematically
504 different from those of the Upper Cretaceous to Cenozoic rocks. If a regional unrecognized
505 remagnetization in Upper Cretaceous or younger time had occurred, creating apparently low
506 paleolatitudes by erroneously correcting for bedding tilt (Butler et al. 2001b; Hollister et al.,
507 2004; Housen & Beck, 1999; Monger & Price, 1996; Nelson & Colpron, 2007), then that
508 remagnetization would also have affected older rocks. The coherent, systematically deviating
509 paleomagnetic directions of pre-Upper Cretaceous rocks makes a regional remagnetization in
510 Late Cretaceous or younger times unlikely.

511 On the other hand, previous arguments for remagnetization of Paleozoic rocks (Butler
512 et al., 1997) cannot be excluded. Rocks older than the Triassic have declinations as well as
513 inclinations that are similar to those from the Triassic Wrangellia large igneous province
514 (Figure 6). If not remagnetized during the eruption of the large igneous province, these rocks
515 would suggest that for a period of ~250 Ma, from the Devonian to the Triassic, Wrangellia was
516 part of a plate that was not moving in paleolatitude much, nor undergoing systematic vertical-
517 axis rotations. We therefore do not interpret the pre-Triassic history of Wrangellia here in
518 detail. We note, however, that the northern hemisphere-counterclockwise rotation scenario
519 permits that the Paleozoic data are primary, whereas the southern hemisphere, clockwise
520 rotation scenario requires that at least the rocks from the Carboniferous-Permian Kiaman
521 superchron do not carry a primary magnetization. Future detailed paleo- and rock magnetic
522 study of these rocks may thus be helpful in evaluating the possibility of this scenario.

523

524 *6.2 Paleomagnetic constraints on Wrangellia plate motion since the Triassic*

525

526 When inspecting the paleomagnetic dataset of the Wrangellia superterrane, compared
527 to the APWP of North America as reference, we first note that the two hemispheric options
528 yield consistent declinations (Figure 6). These cluster in either $\sim 100^\circ$ clockwise or $\sim 80^\circ$
529 counterclockwise rotation relative to the magnetic north pole since the Early Jurassic. Previous
530 workers have mostly focused on paleolatitudes and assumed that declinations were unreliable
531 because of local rotations related to orogenic deformation (e.g., Kent & Irving, 2010). Such
532 local rotations would of course be easily explained, given that the Cordilleran orogen has been
533 folded, faulted, and transported northward along the North American margin, possibly along
534 major strike-slip faults (Beck, 1976, 1980; Irving & Yole, 1987). However, inspection of the
535 database does not suggest that local rotations play a major role. The scatter in declinations for
536 collections of Upper Cretaceous rocks of the Wrangellia superterrane is several tens of degrees
537 (Figure 6). Much of these data collections are small, based on a dozen or so datapoints (Tables
538 S1 and S2 in Data Set S3). Even the datasets behind global APWPs, collected from stable plate
539 interiors, are scattered over $30\text{--}40^\circ$ mostly because of un-averaged paleosecular variation
540 (Rowley, 2019; Vaes et al., 2022). The declination scatter of the Wrangellia superterrane is
541 larger than that, suggesting that local tectonic rotations slightly enhanced it. Nonetheless, the
542 declinations of Jurassic rocks are not chaotic: Lower Mesozoic and older rocks yield
543 declinations that are systematically much larger than those of the Upper Cretaceous rocks, but
544 their scatter is similar to the latter (Figure 6). We therefore infer that the Wrangellia
545 superterrane underwent a coherent rotation between the Early Jurassic and the Late Cretaceous,
546 i.e., during the pre-collisional period when it was part of a plate converging with North
547 America. In a southern hemisphere scenario (Scenario S), this corresponds to a $\sim 110^\circ$
548 clockwise rotation relative to North America, whereas in a northern hemisphere scenario
549 (Scenario N) it represents a counterclockwise rotation of $\sim 70^\circ$ (Figure 6).

550 Previous data from Triassic and older rocks of the Wrangellia superterrane yielded
551 paleolatitudes close to the equator. Our new dataset shows a rapid paleolatitudinal motion on
552 the order of $20\text{--}30^\circ$ in the Late Triassic to Earliest Jurassic, which is northward in Scenario N,
553 and southward in Scenario S (Figure 6). In Scenario S, this is followed by a northward motion
554 of $\sim 60^\circ$ in ~ 120 Ma. In Scenario N, paleolatitudes would remain fairly constant until the Late
555 Cretaceous.

556 To interpret what such paleolatitudinal drifts would mean for plate motion rates we
557 compare these trajectories with the paleolatitudinal motion of North America. Moreover, it is
558 important to note that the paleolatitude of North America is affected by a major phase of True

559 Polar Wander (TPW) that occurred in Late Triassic to Jurassic times (Steinberger & Torsvik,
560 2008; Torsvik et al., 2012; Vaes, 2023). Comparison between the global APWP of Vaes et al.
561 (2023) with the recent mantle reference frame based on a series of tectonic ‘rules’ of Müller et
562 al. (2012) suggested that the pole of TPW was in the Atlantic Ocean (Vaes, 2023), close to the
563 earlier inferred pole based on the shared rotation of all plates in the paleomagnetic reference
564 frame of Steinberger & Torsvik (2008). At the longitude of western North America, this TPW
565 phase caused a southward and then northward shift in latitude on the order of 15° with a peak
566 magnitude of TPW around 200 Ma. Farther west, in the eastern Panthalassa Ocean where the
567 Wrangellia superterrane must have been, the magnitude of TPW increased up to a maximum
568 of $\sim 20^\circ$.

569 Scenario S gives a southward shift of the Wrangellia superterrane in the Triassic,
570 followed by a northward motion. Much of this southward shift could be the result of TPW.
571 Taking the TPW reconstruction of Torsvik et al. (2012) into account, approximately half of the
572 southward latitude change of the Wrangellia superterrane between ~ 220 and 190 Ma may have
573 been caused by TPW, suggesting $\sim 10\text{--}15^\circ$ absolute southward motion at subequatorial
574 latitudes. Because Jurassic motion of North America was northward, northward motion of
575 Wrangellia relative to North America was $\sim 50^\circ$ between ~ 190 and ~ 80 Ma ago, a period of 110
576 Ma (Figure 6). There are currently no high-quality paleomagnetic data to further specify how
577 this relative northward motion was distributed through time.

578 Scenario N requires a rapid northward shift of $\sim 20^\circ$ of the Wrangellia superterrane in
579 the Late Triassic to Earliest Jurassic. It is important to note that this motion is opposite to the
580 TPW-induced southward motion, which means that the plate tectonic motion in Scenario N is
581 larger, by about $\sim 15^\circ$, than the paleomagnetically determined motion, followed by a net
582 paleolatitudinal standstill until the Late Cretaceous. Because North America kept moving
583 northwards in the Jurassic, this would require that the plate carrying the Wrangellia
584 superterrane moved southwards relative to North America, i.e., with a left-lateral strike-slip
585 component. Below, we place both scenarios in further plate tectonic context.

586

587 *6.3. Wrangellia in context of eastern Panthalassa plate kinematic history*

588

589 We now explore the plate tectonic consequences and feasibility of the paleomagnetic
590 scenarios above. A detailed kinematic restoration of Cordilleran orogenic architecture is
591 required to fully justify a final choice between the two options, which is beyond the scope of

592 the current paper. We therefore restrict ourselves here to outlining the implications, solutions,
593 and problems that the two scenarios generate.

594 We present a simplified reconstruction in Figure 7. The paleolatitudinal position of a
595 simple, straight Wrangellia superterrane (which does not include the portions bent into Alaska;
596 Johnston, 2001) relative to North America was determined in the paleomagnetic reference
597 frame of Vaes et al. (2023), whereby we assumed a 190 Ma-position of $\sim 30^\circ$ south or north.
598 Because we now aim to develop a plate tectonic scenario and compare it to seismic
599 tomographic constraints, we placed the reconstruction in the TPW-corrected paleomagnetic
600 reference frame of Torsvik et al. (2012). We note that this frame has no paleolongitudinal
601 constraints, so we restrict ourselves in the comparison to tomography and latitudinal fits only.

602 The paleolongitudinal position of the Wrangellia superterrane relative to the Americas
603 is not constrained by paleomagnetism, but options are restricted given the estimated position
604 of the Farallon-Phoenix-Izanagi triple junction at which the Pacific Plate formed at 190 Ma
605 (Boschman & van Hinsbergen, 2016). The position of this triple junction must have been
606 located in the eastern Panthalassa Ocean to maintain convergence of the major Panthalassa
607 plates with the surrounding continents, and we use the approximated position of Boschman et
608 al. (2021a). This gives the Wrangellia superterrane a possible paleolongitudinal range of ~ 5000
609 km at 190 Ma.

610 From Late Triassic (~ 210 Ma) to Middle Jurassic times (165 Ma), when the Bonanza
611 arc was active, the Wrangellia superterrane was in an upper plate position of a subduction zone.
612 In present-day coordinates, the trench is thought to have been located to the west of the
613 Wrangellia superterrane (Clift et al., 2005b; Trop & Ridgway, 2007). This means that in the
614 southern hemisphere, clockwise rotation Scenario S, the Bonanza arc was underlain by a
615 northward dipping subduction zone. In the northern hemisphere, counterclockwise rotating
616 Scenario N, the arc was underlain by a southward dipping subduction zone.

617 We evaluate the scenarios against the evidence for oblique subduction below the
618 California forearc between 170 and 160 Ma with a N-S component of ~ 6 -7 cm/yr, concluded
619 from reconstructing the Californian ophiolites (Arkula et al., 2023), and the evidence for a
620 northward motion component throughout the Cretaceous suggested by the paleolatitudes of the
621 accreted seamounts of the Franciscan accretionary complex of California (Courtillot et al.,
622 1985; Tarduno et al., 1985, 1986). In addition, a northward motion component is also permitted
623 in line with the OPS that accreted to the western margin of the Chortis block in the Early

624 Cretaceous, although a stable eastward relative plate motion is also permitted there (Boschman
625 et al., 2021a).

626 We use two additional constraints in the 190–80 Ma time window to illustrate the
627 implications of the two scenarios. First, the Caribbean Plate contains Jurassic crust that formed
628 around ~160–150 Ma, and that was part of the Farallon Plate prior to the ~100 Ma onset of
629 subduction at the western Caribbean subduction zone (Boschman et al., 2019). This crust likely
630 formed at the Farallon-Phoenix spreading ridge, and paleomagnetic data reveals that this crust
631 was located around the equator during the Late Jurassic (Boschman et al., 2019). This requires
632 that by ~160 Ma, the Wrangellia superterrane must have been located to the north of the
633 Caribbean lithosphere, and that it has likely always been located to the north of the Phoenix-
634 Farallon spreading ridge.

635 Second, the Bonanza arc subduction must have been associated with a subducting slab,
636 which likely broke off around or shortly after the time of arc cessation, ~165 Ma. Slabs may
637 during their subduction be dragged horizontally through the mantle over 1000 km or more,
638 driven by absolute plate motion of the subducting plate (Parsons et al., 2021; Qayyum et al.,
639 2022; Spakman et al., 2018; van de Lagemaat et al., 2018), and the present-day position of the
640 slab remains of the Bonanza arc subduction zone may no longer reflect the location at which
641 subduction started. However, tomography-plate reconstruction correlations suggest that slabs
642 undergo no major horizontal motion after their break-off (Domeier et al., 2016; van der Meer
643 et al., 2010, 2018), which suggests that the Bonanza arc-related slab is likely located in the
644 lower mantle beneath the location where it broke off. This slab is likely one of the western belt
645 slabs that have been identified below and west of North America (Clennett et al., 2020; Sigloch
646 & Mihalynuk, 2013, 2017; van der Meer et al., 2010, 2012). The southernmost of these, the
647 Malpelo slab west of Colombia, was previously considered a candidate to be linked to the
648 Bonanza arc (van der Meer et al., 2018), but because the Caribbean and western South
649 American Jurassic to Lower Cretaceous arcs have been located at that paleolatitude (Boschman
650 et al., 2019), these may provide a better candidate to explain the Malpelo slab. The
651 southernmost slab of the western belt that cannot be explained by Caribbean arcs is the Socorro
652 slab (van der Meer et al., 2010; 2018), at a latitude of ~15°N (Figure 1). It is also not likely that
653 this slab correlates to subduction records of the Guerrero terrane (in Mexico), which was built
654 since Triassic time on accretionary prism rocks formed at the North American margin and was
655 only temporarily separated from North America by a short-lived and likely narrow back-arc
656 basin (Boschman et al., 2018a, 2018b; Busby et al., 2023; Martini et al., 2014). Because the

657 Bonanza arc is the southernmost intra-oceanic arc complex in the Cordilleran orogen, we
658 therefore discuss in our scenarios S and N whether this slab may have been linked to the
659 Bonanza arc.

660 Both scenarios share the post-80 Ma history in which paleomagnetism places the
661 Wrangellia superterrane approximately 20° south of its present latitude—the basis for the Baja-
662 BC hypothesis that remains difficult to reconcile with structural geological correlations. The
663 declinations are scattered and permit a rotated or a non-rotated position relative to today (Figure
664 6). In our schematic reconstructions, we place the Wrangellia superterrane in a non-rotated
665 position, parallel to and close to the North American margin. We are well aware of the
666 controversy around this reconstruction (e.g., Johnston, 2001; Trop & Ridgway, 2007), or even
667 around the timing of the Wrangellia superterrane accretion (e.g., Gehrels et al., 2009; Hampton
668 et al., 2010; McClelland et al., 1992; Monger, 2014; Nokleberg et al., 2000; Saleeby, 2000;
669 Trop et al., 2002; Stevens Goddard et al., 2018; Tikoff et al., 2023), or diachroneity of accretion
670 (e.g., Manselle et al., 2020; Nokleberg et al., 2000; Pavlis et al., 2019; Trop & Ridgway, 2007).
671 At this stage, we have no satisfactory solution for where the northward motion that remains
672 unaccounted for in the structural record was accommodated. The position of the Wrangellia
673 superterrane in our simple reconstruction is based on paleolatitude only, and the general
674 agreement that by 80 Ma, the Wrangellia superterrane was located along the North American
675 margin. We note, however, that the ~190–80 Ma history discussed below, during which the
676 northern or southern hemisphere options play a role, would not change if we assumed a more
677 northerly position for the Wrangellia superterrane at 80 Ma.

678 Scenario N requires a few thousand kilometers of northward motion of the Wrangellia
679 superterrane relative to the mantle in the Late Triassic–Early Jurassic, which with southward
680 subduction requires roll-back of the slab below the Bonanza arc. Between 190 and 80 Ma, the
681 Wrangellia superterrane remains at middle latitudes until its collision with North America, in
682 which case the Bonanza arc was not associated with subduction of the Socorro slab. Another
683 southward dipping subduction zone must therefore have existed between the Wrangellia
684 superterrane and the Farallon Plate, which in Late Jurassic to Early Cretaceous time must have
685 experienced a northward motion component (Arkula et al., 2023; Courtillot et al., 1985;
686 Tarduno et al., 1985, 1986). Reconstructing the Wrangellia superterrane farther south around
687 the mid-Jurassic, as depicted in Figure 7, would require slab advance of the Bonanza
688 subduction zone, combined with another southward dipping subduction zone to accommodate
689 the convergence between the Wrangellia superterrane and the Caribbean lithosphere which

690 remained around equatorial latitudes (Boschman et al., 2019). Finally, the arrival of the
691 Wrangellia superterrane at the North American margin must have involved a complex,
692 counterclockwise plate rotation so that the western margin of the Wrangellia superterrane faced
693 the Panthalassa Ocean. This counterclockwise rotation cannot be reconstructed without
694 generating Farallon-Wrangellia convergence, requiring a syn-rotation, westward dipping
695 subduction zone between these plates.

696 Scenario S also requires roll-back of the subduction zone below the Wrangellia
697 superterrane in the Late Triassic–Earliest Jurassic, but southward, over $\sim 15^\circ$ latitude. This must
698 have been followed by a northward shift over as much as 30° until the Middle Jurassic, by
699 which time Caribbean lithosphere was forming to the south of the Wrangellia superterrane
700 around the equator (Boschman et al., 2019). This northward Wrangellia motion occurred during
701 activity of the Bonanza arc, and thus occurred during subduction. Hence, it must have been
702 associated with slab advance and dragging over large distances. In Scenario S, it is possible
703 that the Socorro slab represents the lithosphere that detached below the Bonanza arc upon its
704 cessation in the Middle Jurassic. Following this cessation, northward motion of the Wrangellia
705 superterrane continued while the plate carrying it rotated clockwise. This rotation is faster than
706 that reconstructed for the Farallon Plate from Pacific anomalies (e.g., Seton et al., 2012), so it
707 requires that the Wrangellia superterrane was part of a separate plate. However, no subduction
708 zones are required between the Wrangellia superterrane and the Farallon Plate: its clockwise
709 rotation could have been accommodated by subduction between the Wrangellia superterrane
710 and North America, and the systematic northward motions are consistent with the
711 paleomagnetic evidence from accreted OPS units (Courtillot et al., 1985; Tarduno et al., 1985,
712 1986) and North American upper plate ophiolites (Arkula et al., 2023). Subduction reactivation
713 west of the Wrangellia superterrane is only required upon its accretion to North America (and
714 may thus serve as a constraint on collision age).

715

716 *6.4 The way forward*

717

718 The two scenarios discussed above satisfy paleomagnetic constraints and illustrate the
719 complexity in reconstructing the plate kinematic history of the northeastern Panthalassa Ocean.
720 We consider it premature to make a firm choice between the northern and southern scenarios,
721 but our analysis shows possible ways forward. First, our analysis shows that paleomagnetic
722 data provide coherent and consistent results, not only in paleolatitude but also in declination.

723 A key step forward is filling the paleomagnetic data gap for the Wrangellia superterrane in the
724 Late Jurassic and Early Cretaceous. Second, advances have been made in reconstruction of lost
725 lithosphere of the Panthalassa Ocean and the Caribbean regions back to Jurassic times and
726 before from geological records of accretionary prisms (Boschman et al., 2019, 2021a, 2021b;
727 van de Lagemaat et al., 2023; Wright et al., 2016). Third, the rich seismic tomographic imagery
728 of the upper and lower mantle below the former northeastern Panthalassa Ocean demonstrates
729 where remains of subduction are currently residing (Sigloch & Mihalynuk, 2013; Sigloch et
730 al., 2008; van der Meer et al., 2010, 2012). Reconstructions using those tomographic images
731 as evidence for plate kinematic evolution has so far not led to reconstructions that also satisfy
732 geological observations from the Cordillera of North America (Pavlis et al., 2019, 2020;
733 Sigloch & Mihalynuk, 2020), but slab remnants may be helpful in determining where
734 subduction terminated in the geological past. We foresee that a holistic analysis of geological
735 architecture of the Cordillera, in which paleomagnetic constraints and seismic tomographic
736 analyses are cast in context of Panthalassa and western ‘Pangean’ plate reconstructions,
737 providing kinematically feasible scenarios. Those scenarios may serve to identify the key
738 assumptions and interpretations that underpin the long-lasting controversy of Wrangellia’s
739 motion history relative to North America, both before and after collision. Solving this
740 controversy is important, because the error (or errors) which must exist in our thinking of
741 tectonics or paleomagnetism, or both, are difficult to find and identifying them may provide
742 fundamental lessons with repercussions in orogenic and plate reconstructions elsewhere.

743

744 **7. Conclusions**

745

746 Based essentially on paleomagnetic constraints, the Wrangellia superterrane is widely
747 thought to have been located in an intraoceanic setting during most of the Mesozoic. In this
748 study, we show that new and previous Triassic to Lower Jurassic paleomagnetic data of the
749 Wrangellia superterrane are coherent in terms of inclination and declination throughout a 2000
750 km-long stretch of the superterrane that lies south of the Alaskan orocline. To address the
751 hemispheric ambiguity of these paleomagnetic data, we propose two scenarios in which the
752 Wrangellia superterrane was located either in the southern hemisphere (Scenario S) or in the
753 northern hemisphere (Scenario N) at ~190 Ma. From ~190 to 80 Ma, the Wrangellia
754 superterrane moved northward while rotating 110° clockwise at a north-dipping subduction
755 zone (Scenario S) or remained at northern middle latitudes while rotating 70° counterclockwise

756 at a south-dipping subduction zone (Scenario N). The main conclusion that can be drawn from
757 comparing these scenarios is that scenario S represents the simplest solution to transport the
758 Wrangellia superterrane from the equator to a position alongside North America allowing its
759 accretion after a significant (clockwise) rotation. In contrast, scenario N requires at least two
760 additional steps in the overall motion of the Wrangellia superterrane towards North America:
761 (1) after a rapid northward shift of $\sim 20^\circ$ in the Late Triassic to Earliest Jurassic, the superterrane
762 moved southward during the Early Jurassic; (2) the superterrane rotated away from North
763 America during the Middle Jurassic before moving eastward to be accreted to North America.
764 Although new paleomagnetic data from Early Cretaceous rocks are required to test both
765 scenarios, the clockwise rotation in Scenario S fits existing models of Early Cretaceous
766 paleogeography in which the Wrangellia superterrane accreted to the Intermontane
767 superterrane through a northward zipper closure of the ocean between them.

768 In both scenarios N and S, the accretion of the Wrangellia superterrane to the
769 Intermontane superterrane must have taken place >1500 km south of their present location.
770 Interestingly, Upper Cretaceous to lower Cenozoic paleomagnetic data suggest a common
771 northward motion of the amalgamated Wrangellia and Intermontane superterrane from $\sim 35^\circ\text{N}$
772 (at ~ 70 Ma) to $\sim 50^\circ\text{N}$ (at ~ 50 Ma), which is at odds with the estimates of northward motion
773 (<800 km) obtained from major strike-slip faults located east of the Intermontane superterrane.
774 By combining Triassic to Cretaceous paleomagnetic data with geological constraints on
775 subduction polarity and mantle tomography, we found that Late Cretaceous paleomagnetic data
776 yielding middle latitudes ($\sim 35^\circ\text{N}$) should not be discarded when reconstructing the tectonic
777 history of the Wrangellia superterrane. Overcoming the fact that paleomagnetic and structural
778 datasets yield apparently robust, yet contradicting estimates of northward motion of the
779 Wrangellia superterrane will require enhanced collaboration across disciplinary specialties. In
780 future studies, to obtain overlapping values of northward motion from both paleomagnetic and
781 structural data will require: (i) identifying hidden faults or suture zones in the eastern Cordillera
782 that may have contributed to overall higher Mesozoic northward motions than presently
783 acknowledged; (ii) processing paleomagnetic data with new statistical standards, as this could
784 lead to paleolatitude results that would require Mesozoic northward motions of lower
785 amplitude than currently thought. This represents a clear opportunity for specialists of different
786 disciplines to guide future joint research into understanding the source of discrepancies
787 between their respective datasets. Such a community effort may be beneficial when

788 paleomagnetic and structural data underpin research questions in other tectonic settings and in
789 deeper time periods.

790

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792

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798

799 **Data availability statement**

800

801 New paleomagnetic data can be imported into the Paleomagnetism.org 2.0 portal from data
802 files (.col) available in the Supporting Information (Data Sets S1 and S2), as well as in the
803 Paleomagnetism.org (link to be provided) and the MagIC (link to be provided) databases.

804

805 **References**

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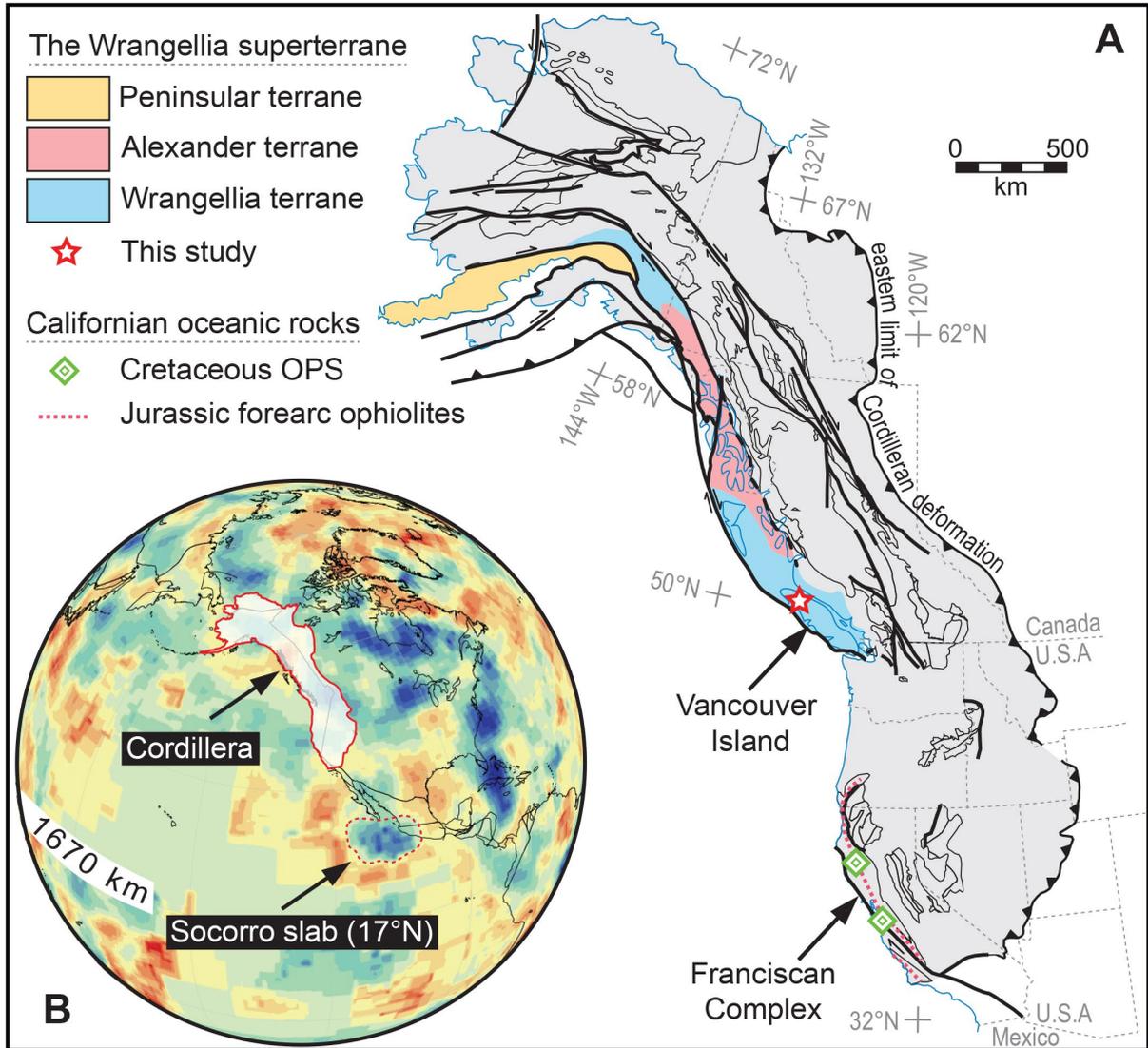
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1537 **Captions**

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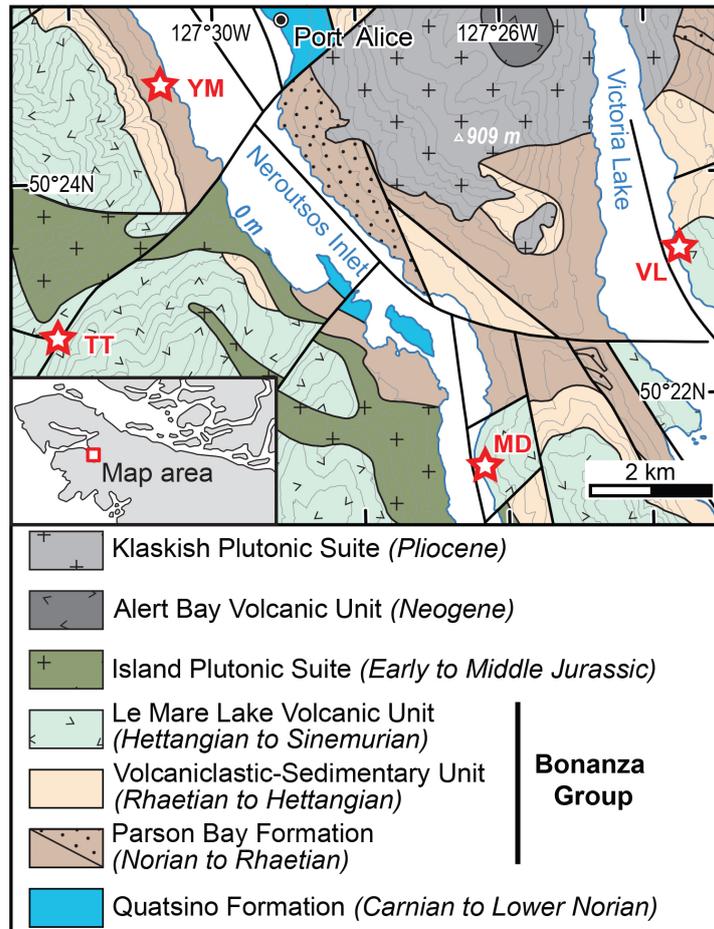
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1541 **Figure 1.** Tectonic setting of the Wrangellia superterrane and other geological records
 1542 discussed in the text. (a) Tectonic map of the Cordillera of western North America (modified
 1543 after Colpron & Nelson, 2009). (b) UUP07 P-wave tomographic model at 1670 km depth
 1544 (Amaru, 2007) showing the location of the Socorro slab (offshore western Mexico) and the
 1545 outline of the Cordillera of North America. Positive seismic wave-speed anomalies, such as
 1546 the Socorro slab, are in blueish colors (up to +0.5%), whereas negative seismic wave-speed
 1547 anomalies are in reddish colors (down to -0.5%).

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1552 **Figure 2.** Geological map of the studied area in northern Vancouver Island (100 m contour
 1553 interval; modified after Nixon et al., 2011a, 2011d), which is shown in the lower left inset. The
 1554 red stars on the geological map indicate the four localities where paleomagnetic samples were
 1555 collected. Localities VL-TT-MD are in the Le Mare Lake Volcanic Unit and locality YM is in
 1556 the Parson Bay Formation. Parts of the Parson Bay Formation with higher volcaniclastic
 1557 contents are indicated with a sandy lithological pattern.

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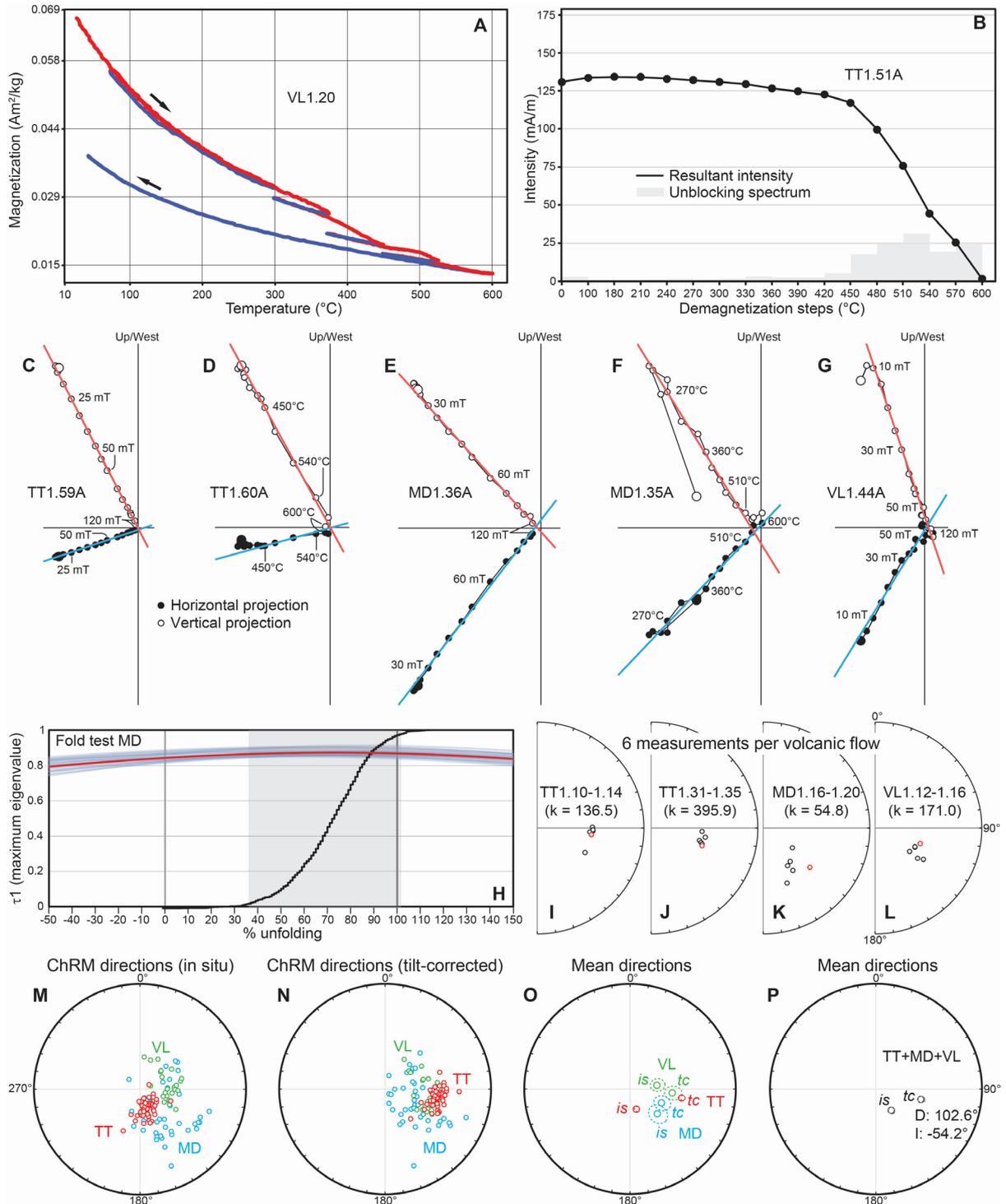
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Figure 3. Rock magnetic and paleomagnetic results from volcanic localities TT, MD, and VL. (a) Thermomagnetic curves measured on a Curie balance. Heating segments are in red and cooling segments are in blue. (b) Intensity decay curve measured during TH treatment. (c)–(g) Orthogonal vector diagrams in geographic (= in situ) coordinates, where closed (open) symbols indicate declination (inclination). (h) Bootstrapped fold test of locality MD, with cumulative distribution function (confidence interval in light grey) based on 1000 bootstrap samples (mean

1575 shown in red). (i)–(l) 6 characteristic remanent magnetization (ChRM) directions per volcanic
1576 flow in tectonic (= tilt-corrected) coordinates. For each flow, 5 directions from AF treatments
1577 (black circles) and 1 direction from TH treatment (red circle). k values correspond to the
1578 average (not shown) of the 6 directions. (m), (n) ChRM values from localities TT (red circles),
1579 MD (blue circles), and VL (green circles). (m) In situ coordinates. (n) Tilt-corrected
1580 coordinates. (o) Mean directions, including confidence intervals (dashed lines), of localities
1581 TT, MD, and VL. is = in situ coordinates, tc = tilt-corrected coordinates. (p) Same as in (o)
1582 with all localities combined. Declination and inclination values are for the tilt-corrected mean.

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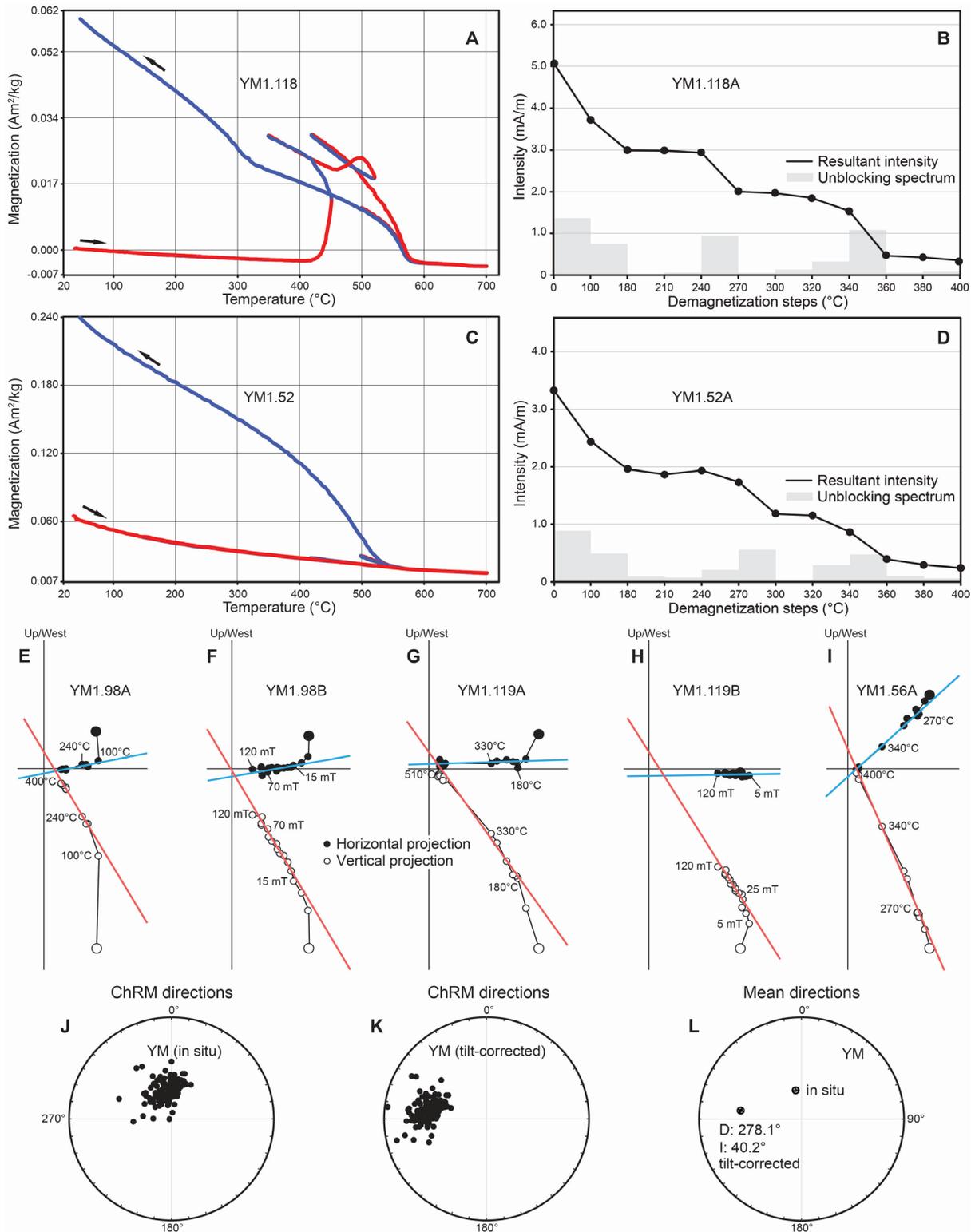
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Figure 4. Rock magnetic and paleomagnetic results from sedimentary locality YM. (a), (c) Thermomagnetic curves measured on a Curie balance. Heating segments are in red and cooling segments are in blue. (b), (d) Intensity decay curves measured during TH treatment. (e)–(i) Orthogonal vector diagrams in geographic (= in situ) coordinates, where closed (open) symbols indicate declination (inclination). (j), (k) ChRM values from locality YM. (j) In situ

1614 coordinates. (k) Tilt-corrected coordinates. (l) Mean directions, including confidence intervals
1615 (dashed lines), of locality YM. Declination and inclination values are for the tilt-corrected
1616 mean.

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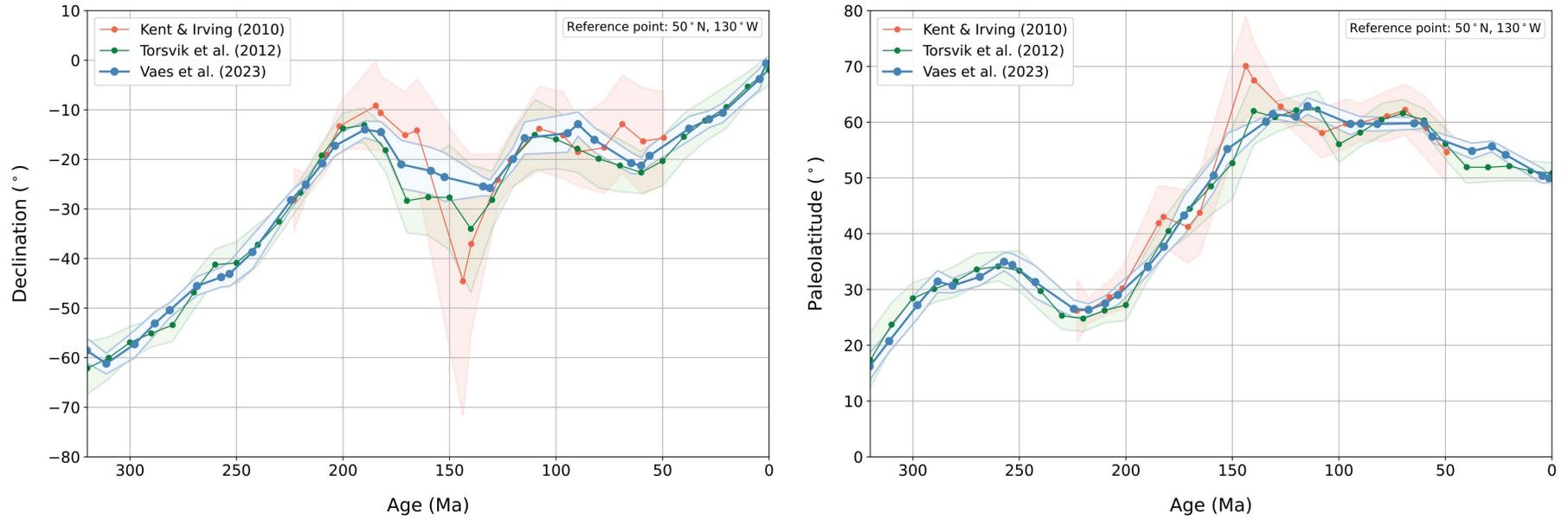
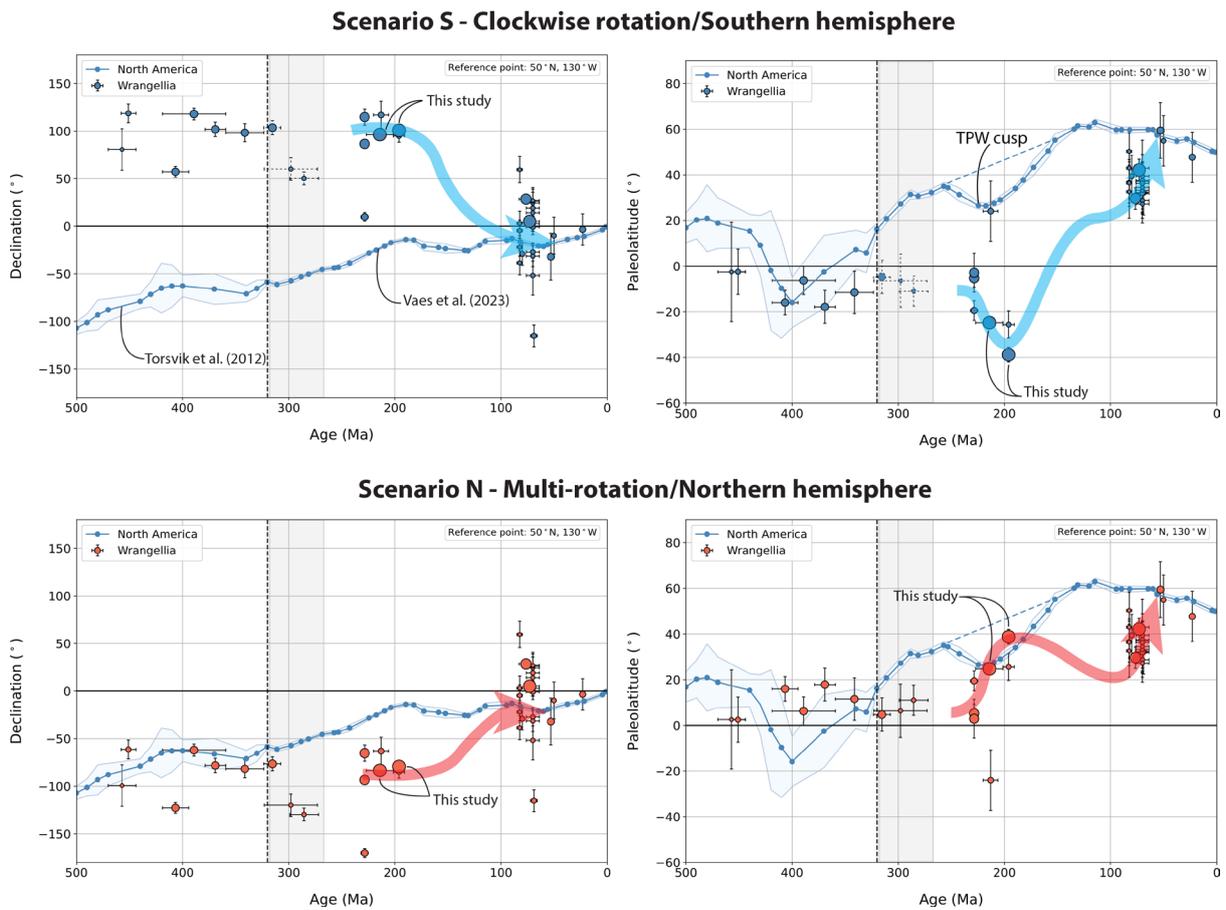


Figure 5. Declination and paleolatitude curves for the North American plate at a reference point at 50°N and 130°W, as predicted by three global APWPs. The colored bands show the 95% confidence regions for each curve. Note that the global APWP of Kent and Irving (2010) was computed for the 230 to 50 Ma time interval only.

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1651 **Figure 6.** Paleomagnetic results obtained from the Wrangellia superterrane compared to
 1652 predicted values for North America. Two possible scenarios are permitted by the hemispheric
 1653 ambiguity of the pre-100 Ma paleomagnetic data: the upper panels show Scenario S (southern
 1654 hemisphere) and the lower panels show Scenario N (northern hemisphere). Declination and
 1655 paleolatitude values are computed for a reference point at 50°N and 130°W and are shown as
 1656 blue or red dots with error bars indicating the age uncertainty and 95% confidence in the
 1657 declination or paleolatitude. Predicted declination and paleolatitude curves—based on the global
 1658 APWP of Vaes et al. (2023) for the last 320 Ma and North American APWP of Torsvik et al.
 1659 (2012) for 320–500 Ma (transition marked by vertical dashed line)—are shown as blue curves
 1660 (with 95% confidence region). The Carboniferous-Permian Kiaman superchron (~320–260
 1661 Ma) is indicated by the light grey band. The expected reversed polarity for the data from the
 1662 superchron requires those plotted for Scenario S to be remagnetized (indicated by dashed error
 1663 bars). The inferred effect of Triassic-Jurassic TPW on the paleolatitude of North America are
 1664 highlighted. Paleomagnetic data are derived from the following sources (for numerical values,
 1665 see Table 1 and Tables S1 and S2 in Data Set S3): Hillhouse (1977), Van der Voo et al. (1980),

1666 Yole and Irving (1980), Panuska and Stone (1981), Hillhouse and Grommé (1984), Panuska
1667 (1985), Hillhouse et al. (1985), Irving & Yole (1987), Irving and Brandon (1990), Haeussler et
1668 al. (1992a, 1992b), Bazard et al. (1995), Butler et al. (1997), Grommé and Hillhouse (1997),
1669 Ward et al. (1997; E/I correction from Krijgsman & Tauxe, 2006), Enkin et al. (2001),
1670 Stamatakos et al. (2001), Kim and Kodama (2004; E/I correction from Krijgsman & Tauxe,
1671 2006), and this study.

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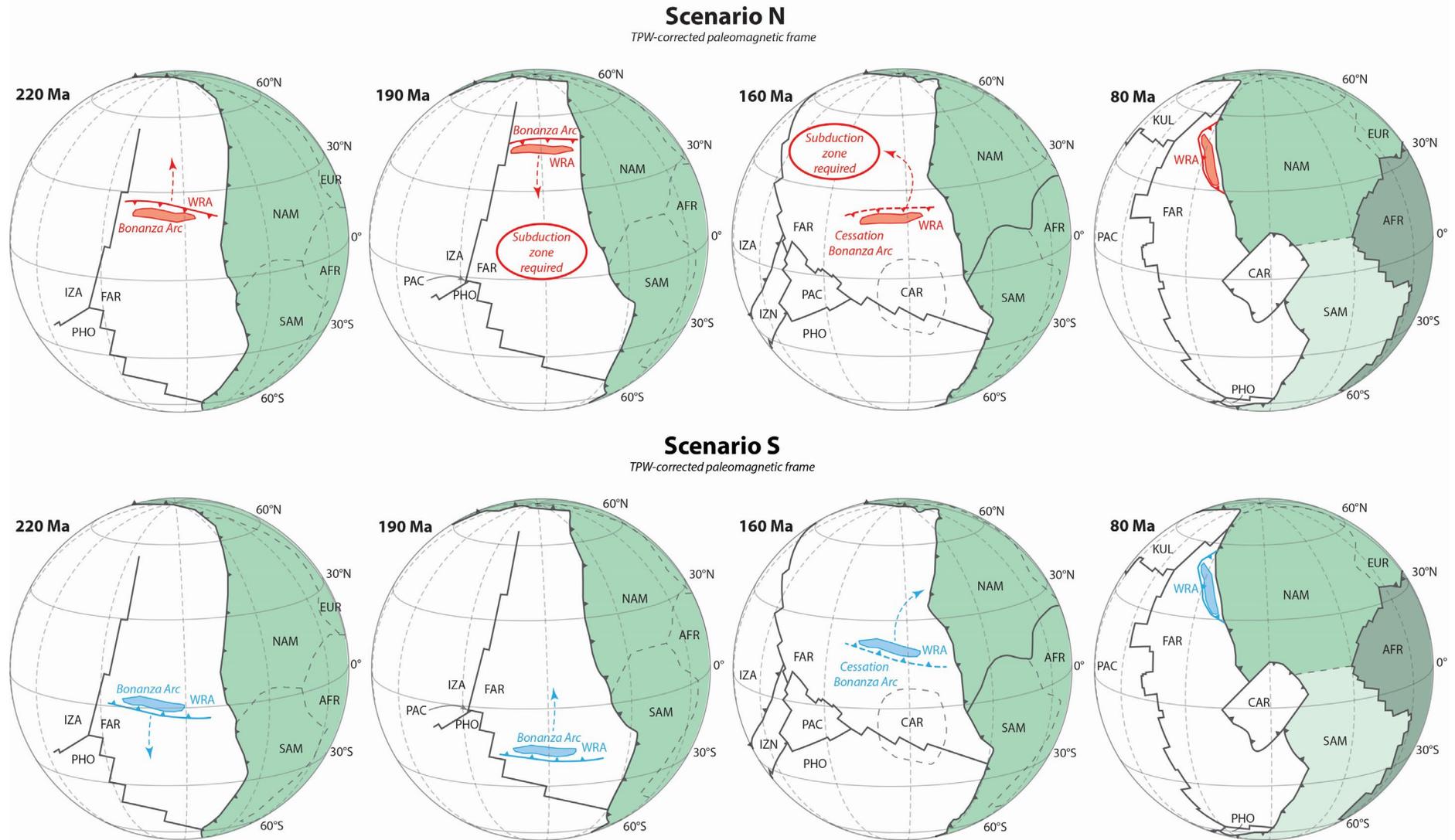


Figure 7. Schematic kinematic scenarios of the motion of the Wrangellia superterrane during the Late Triassic to Late Cretaceous (220 to 80 Ma). Scenario S shows a Late Triassic–Early Jurassic southward motion of $\sim 15^\circ$ followed by a northward motion of ~ 5000 km from the southern hemisphere accompanied by a clockwise rotation of $\sim 110^\circ$. Scenario B depicts a rapid Late Triassic–Early Jurassic northward shift of $\sim 20^\circ$ in the northern hemisphere, followed by a southward motion of lower amplitude, and then a standstill at middle latitude accompanied by counterclockwise rotation of $\sim 70^\circ$. AFR = Africa; CAR = Caribbean; EUR = Europe; FAR = Farallon; IZA = Izanagi; IZG = Izanami; KUL = Kula; NAM = North America; PAC = Pacific; PHO = Phoenix; SAM = South America; WRA = Wrangellia.

Table 1
Paleomagnetic Results of the Le Mare Lake Volcanic Unit (MD, VL, TT) and the Parson Bay Formation (YM)

Locality	slat	slon	N	In situ					Tilt-corrected					k	α_{95}	K	A_{95Min}	A_{95}	A_{95Max}	λ	plat	plong
				N45(is)	N45(tc)	D	ΔDx	I	ΔIx	D	ΔDx	I	ΔIx									
<i>Le Mare Lake Volcanic Unit</i>																						
MD	50.357	-127.439	41	n.a.	n.a.	132.2	13.0	-61.4	8.1	113.8	11.9	-63.2	6.8	13.7	6.3	8.1	2.7	8.4	7.9	-44.7	-47.3	346.7
VL	50.390	-127.392	24	n.a.	n.a.	81.5	13.5	-69.0	5.7	95.4	9.3	-56.9	7.0	32.9	5.2	17.2	3.4	7.4	11.1	-37.5	-31.1	346.3
TT	50.376	-127.546	58	n.a.	n.a.	161.3	6.3	-73.8	2.0	99.8	2.6	-48.7	2.6	98.0	1.9	70.4	2.4	2.2	6.4	-29.6	-28.7	335.7
MD-VL-TT	50.372	-127.480	123	108	117	142.9	6.8	-69.9	2.7	102.6	3.9	-54.2	3.2	29.1	2.5	17.9	1.8	3.2	4.1	-34.7	-35.0	340.8
<i>Parson Bay Formation</i>																						
YM	50.415	-127.512	138	n.a.	n.a.	347.0	3.95	66.7	1.9	278.1	1.8	40.2	2.3	48.7	1.7	52.7	1.7	1.7	3.7	22.9	23.0	151.9

Note. slat/slon = latitude and longitude of sampling location; N = total number of demagnetized samples; N45 = number of samples that fall within the 45° cutoff in in situ coordinates (is) and after tilt correction (tc); D, ΔDx = declination and associated error; I, ΔIx = Inclination and associated error; k and α_{95} = precision parameter and semiangle of the 95% cone of confidence around the computed site mean direction; K and A_{95} = precision parameter and semiangle of the 95% cone of confidence around the mean virtual geomagnetic pole; A_{95max} and A_{95min} = maximum and minimum value of A_{95} expected from paleosecular variation of the geomagnetic field; λ = paleolatitude of the locality; plat/plon = paleopole latitude and longitude; n.a. = not applicable.