Crustal structure of Sri Lanka derived from joint inversion of receiver functions and seismic ambient noise using a Bayesian approach

Jennifer Dreiling^{1,2}, Frederik Tilmann^{1,2}, Xiaohui Yuan¹, Christian Haberland¹, S.W. Mahinda Seneviratne³

¹GFZ German Research Centre for Geosciences, Telegrafenberg, 14473 Potsdam, Germany ²Freie University of Berlin, Malteserstr. 74–100, 12249 Berlin, Germany ³GSMB Geological Survey and Mines Bureau, Sri Lanka

> Please note that this is a non-peer-reviewed EarthArXiv preprint, submitted to J. Geophys. Res. Solide Earth in September 2019. https://eartharxiv.org/

11 Key Points:

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| 12 | • | Sri Lanka has mostly isostatically compensated 30–40 km thick crust |
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| 13 | • | Major mid-crustal westward dipping interface related to thrust contact between |
| 14 | | Highland Complex and Vijayan Complex |
| 15 | • | Dipping discontinuity and low velocity zone in Highland Complex support amal- |
| 16 | | gamation theory of stepwise collision |

Corresponding author: Jennifer Dreiling, dreiling@gfz-potsdam.de

17 Abstract

We study the crustal structure of Sri Lanka by analyzing data from a temporary seismic network deployed in 2016–2017 to shed light on the amalgamation process from the geophysical perspective. Rayleigh wave phase dispersion from ambient noise crosscorrelation and receiver functions were jointly inverted using a transdimensional Bayesian approach.

The Moho depths range between 30 and 40 km, with the thickest crust (38–40 km) beneath the central Highland Complex (HC). The thinnest crust (30–35 km) is found along the west coast, which experienced crustal thinning associated with the formation of the Mannar Basin. The majority of Vp/Vs ratios lies within a range of 1.66–1.8 and predominantly favor a felsic composition with intermediate-to-high silica content of the rocks.

A major intra-crustal (18–27 km), slightly westward dipping ($\sim 4.3^{\circ}$) interface with high Vs (>4 km/s) underneath is prominent in the central HC, continuing in the eastern Vijayan Complex (VC). The dipping discontinuity and a low velocity zone in the central Highlands can be related to the HC/VC contact zone and is in agreement with a wellestablished amalgamation hypothesis of a stepwise collision of the arc fragments, including deep crustal thrusting processes and a transpressional regime along the eastern suture between the HC and VC.

³⁶ 1 Introduction

Sri Lanka occupied a key region in both the assembly and the multistage breakup 37 of Gondwana. Many petrological, geochemical and geochronological studies have been 38 conducted to reconstruct the processes acting during the amalgamation. However, lit-39 tle is known about the seismic structure of the island. Until mid 2016, only three per-40 manent seismic stations existed on the island. Pathak, Ravi Kumar, and Sarkar (2006) 41 and Rai, Gaur, Rai, and Priestley (2009) analyzed receiver functions from the perma-42 nent station PALK and estimated Moho depth and Vp/Vs ratio. Prasanna, Chen, and 43 z (2013) used gravity inversion and Mishra, Vijaya Kumar, and Rajasekhar (2006) mod-44 eled gravity anomalies within Sri Lanka and other continental fragments of Gondwana 45 to determine the crustal thickness and structure beneath the island. 46

In 2016–2017 the Geological Survey and Mines Bureau (GSMB) of Sri Lanka and
 the German Research Centre for Geosciences (GFZ) installed and maintained the first

⁴⁹ broadband seismic network on the island (Fig. 1), consisting of 30 stations running for ⁵⁰ a period of 13 months. Here, we image the crustal structure of Sri Lanka using the new ⁵¹ seismic data. We performed a joint inversion of dispersion curves from seismic ambient ⁵² noise with receiver functions using a Bayesian approach, which allows us to compute a ⁵³ collection of possible models and to estimate the uncertainty of the model parameters.

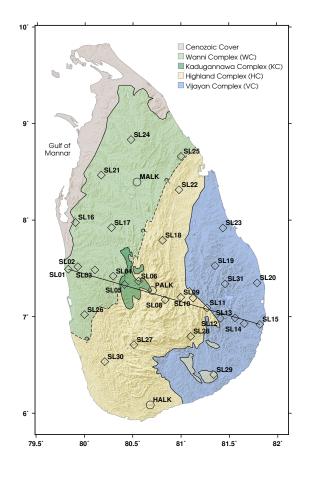


Figure 1. Map of seismic stations and major lithotectonic units (modified after Dissanayake & Chandrajith, 1999) in Sri Lanka. Diamonds represent station locations of temporary seismic array (FDSN code: 1A, 2016–2017). Circles denote three permanent stations (MALK, PALK, HALK). Black line indicates profile location for a cross section.

⁵⁴ 1.1 Geologic Background

Sri Lanka is mostly composed of Precambrian crust; only the northern and north-55 western coasts show younger Jurassic-Quaternary sedimentary deposits (Fig. 1). The Pre-56 cambrian basement consists of three major units, namely, from west to east, the Wanni 57 Complex (WC), the Highland Complex (HC), and the Vijayan Complex (VC). Some HC 58 erosion remnants (Klippen) occur around Buttala, Kataragama and Kuda Oya in the 59 southern part of the VC. The Kadugannawa Complex (KC) is a relatively small unit lo-60 cated between the WC and HC. It is contentious whether it is part of the WC, part of 61 the HC or the root zone of an island arc (?, and references therein). The WC/HC rep-62 resents as a combined unit a tilted section of the former lower-middle crust, with the 63 HC representing the lower level. The KC is at a crustal level between the WC and the 64 HC (Kriegsman, 1994; Sandiford, Powell, Martin, & Perera, 1988; ?). The WC consists 65 of metamorphic rocks of upper amphibolite- to granulite-facies, the HC predominately 66 of granulite facies and the VC of amphibolite facies (K. W. Kehelpannala, 2003, and ref-67 erences therein). 68

The contact between the WC and the HC is controversial due to absence of a struc-69 tural break between them. Stretching lineations, shear sense indicators and sheath folds 70 demonstrate that a collision has occurred in a NNW-SSE direction, i.e., the WC/KC was 71 moving on top of the HC from NNW towards SSW (K. W. Kehelpannala, 2003, and ref-72 erences therein). The boundary between the WC and the KC is less clear, while that be-73 tween the KC and the HC is well defined. The boundary between the HC and the VC 74 is considered to be a deep crustal, sub-horizontal ductile shear / thrust zone (K. W. Ke-75 helpannala, 2003; Kleinschrodt, 1994, 1996, and references therein). E-W stretching lin-76 eations in the VC and nearly N-S stretching lineations and shear sense indicators at and 77 close to the shear zone suggest a nearly E-W directed transpressional collision between 78 the combined WC/HC unit and the VC (?, and references therein). The general trend 79 of subhorizontal fold envelopes suggests the thrust to underlie large parts of the HC; Klip-80 pen south of the HC prove that the thrust plane extended nearly up to the south coast 81 (Kleinschrodt, 1994, 1996). Furthermore, Kleinschrodt (1994) suggests that the HC climbed 82 on top of the east-VC with a ramp-flat geometry or a low-angle thrust, steepening to higher 83 crustal levels. 84

The amalgamation of the Sri Lankan complexes took place within the framework of the Pan-African continental collision between West and East Gondwana. Petrolog-

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ical, geochemical and geochronological studies suggests that the WC, KC and VC have 87 been formed through arc-related events during the Early Neoproterozoic, i.e., ~ 1.0 Ga 88 (Takamura, Tsunogae, Santosh, Malaviarachchi, & Tsutsumi, 2016, and references therein). 89 A well-established theory for the amalgamation of Sri Lanka suggests a stepwise colli-90 sion of the Precambrian arcs (e.g., K. V. W. Kehelpannala, 2004) during the Pan-African 91 Orogeny, whereas Santosh et al. (2014) recently interpreted the WC, KC and VC as Early 92 to Late Neoproterozoic continental arcs, with the HC as a Neoproterozoic suture zone 93 formed by double-sided subduction and final collision of the WC and VC. 94

The hypothesis of the stepwise collision of the Precambrian arcs suggests an ini-95 tial collision of the WC and HC fragments. As a unified block the WC/HC has expe-96 rienced six phases of ductile deformation (D1–D6 in 0.61–0.55 Ga), which are not seen 97 for the VC (?). The evidence therefore suggests an early stage collision of the WC/HC 98 unit with the VC at D5 (0.58 Ga), and the WC/HC subsequently being thrust over the qq VC (K. W. Kehelpannala, 2003, and references therein). Based on post-tectonic intru-100 sion by Cambrian granites and symptotic through all three units, i.e., WC, HC and VC, 101 the fragments of Sri Lanka were united at 0.55 Ga. Most of the older structures have been 102 obliterated by strong Pan-African non-coaxial strain, which also brought all the early 103 planar and linear fabrics into parallelism with those formed during the Pan-African event 104 (?). 105

Based on petrological and geochemical data, Santosh et al. (2014) recently proposed an alternate scenario, termed divergent subduction, which involves a double-sided subduction of oceanic crust beneath the WC to the west and the VC to the east. The HC is therefore the collisional suture/ accretionary complex in between, where trench-fill sediments and ancient micro-continents or arcs are accreted and admixed during the final collision stage. Santosh et al. (2014) do not comment which larger lithospheric structures would be predicted by their model.

The Mannar Basin (west of Sri Lanka, partly onshore, Fig. 1) has been formed during Gondwana breakup, which initiated at approximately 165 Ma (?). A great amount of rifting between India and Sri Lanka together with strike slip movement and anticlockwise rotation of Sri Lanka was responsible for significant widening and rapid subsidence in the basin (Kularathna, Pitawala, Senaratne, Senevirathne, & Weerasinghe, 2015), and is associated with strong crustal thinning along the west coast.

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119 **1.2 Seismic Data**

Until May 2016 the island of Sri Lanka was equipped with only three seismic stations: PALK, MALK and HALK. PALK is an IRIS/IDA station and operates since 2000 (?). MALK and HALK are GEOFON stations and have been operating since 2010 (?). In mid 2016 a field campaign was initiated by the GSMB of Sri Lanka and executed jointly with the GFZ (?). A network of 30 three-component broadband stations has been deployed (Fig. 1), which recorded continuous data for a period of 13 months.

The temporary array was designed to perform seismic ambient noise and receiver function analyses as well as local earthquake studies. Fourteen temporary stations and the permanent station PALK form a 230 km long profile across the island, from the west to the east coast, perpendicular to the predominant geologic strike (profile in Fig. 1). Interstation distances are about 15 km. Sixteen more stations were spread out on the island at a larger spacing of about 50 km.

Three stations (SL04, SL14, SL30) were operating only for a short period of time. 132 SL14 failed after just 13 days, which was too short to record enough earthquake signals 133 or to recover stable Green's functions; hence we excluded it from both analyses. Stations 134 SL04 and SL30 were recording for 85 and 30 days, respectively. These stations were in-135 cluded for the ambient seismic noise analysis, but discarded for receiver function com-136 putation, as too few events occurred during operational time. The stations included in 137 this study and the time period considered for further analyses are summarized in Ta-138 ble 1. 139

Table 1. Seismic broadband stations and time span included in this study. SR: samplingrate; SWD: Surface wave dispersion; RF: Receiver functions. The temporary network 1A is ourprimary data source.

| Network | FDSN | Stations | SR (Hz) | SWD - Time | Span — RF |
|-----------|------|------------|---------|-----------------|-------------------|
| Temporary | 1A | SL01-SL31 | 100 | 06/2016-06/2017 | 06/2016-06/2017 |
| GEOFON | GE | MALK, HALK | 50 | 05/2016-08/2017 | 01/2015 - 12/2017 |
| IRIS/IDA | II | PALK | 40 | 05/2016-08/2017 | 01/2015 - 12/2017 |

¹⁴⁰ 2 Seismic Ambient Noise Correlation and Tomography

In order to prepare the data for calculating the cross-correlation stacks, the linear trend and the mean were subtracted from the raw data and a low pass filter was applied prior to decimation to prevent aliasing effects. The threshold was set to 85 % of the new Nyquist frequency (2.125 Hz). The data were down-sampled to a sampling rate of 5 Hz, with subsequent instrument response removal.

For ambient noise cross-correlation we applied the pre-processing procedures sug-146 gested by Bensen et al. (2007). The instrument corrected data were clipped at 3 stan-147 dard deviations and bandpass filtered between 0.01 and 1.25 Hz. Subsequently, spectral 148 whitening and 1-bit normalization were applied. The cross-correlation was performed 149 by correlating 1-hour segments of all station and component combinations and subse-150 quently rotating the full Green's tensor stream from the ENZ to the RTZ coordinate sys-151 tem. The correlograms resulting from one day were added to daily stacks. Green's func-152 tions were generated by stacking the daily stacks for the time period available. This re-153 sulted in a final correlogram stack for each of the 496 station pairs and for each of the 154 components. Here, we consider combinations of the radial and vertical component for 155 the Rayleigh surface wave, i.e., vertical-vertical (ZZ), radial-radial (RR), vertical-radial 156 (ZR), and radial-vertical (RZ). 157

Surface wave dispersion (SWD) was determined from the phase of the fundamental-158 mode Rayleigh wave, based on the zero-crossings of the real part of the correlation spec-159 trum (Ekstrm, Abers, & Webb, 2009; ?). The 2π ambiguity leads to a family of possi-160 ble period-phase velocity relations. Therefore, the average phase velocity dispersion was 161 computed for Sri Lanka and used as a guide for selection of the most likely branch for 162 each station pair. To retrieve phase velocity measurements we used the tool GSpecDisp 163 (Sadeghisorkhani, Gudmundsson, & Tryggvason, 2017). Dispersion curves were deter-164 mined for the ZZ, RR, RZ, and ZR components, separately, resulting in 478, 422, 440, 165 and 454 successful measurements, respectively. For 385 station pairs, all four components 166 could be picked. To retrieve the final Rayleigh wave dispersion curve for each station pair, 167 the four dispersion curves were averaged after interpolation. 168

The final 385 phase velocity dispersion curves for Sri Lanka are illustrated in Figure 2. The velocities increase from 2.9–3.3 km/s at the period of 1 s to 3.7–4.0 km/s at 30 s. The variations of phase velocities cover a narrow band with an average width of 0.35 km/s.

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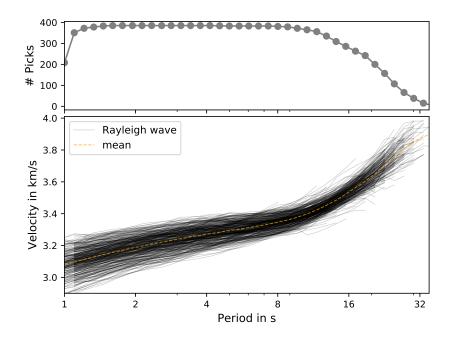


Figure 2. Rayleigh wave dispersion curves for all the individual station pairs (bottom). At most of the periods, the number of picks is >300 (top).

| 173 | For travel time tomography we used the Fast Marching Surface wave Tomography |
|-----|--|
| 174 | package (FMST; Rawlinson & Sambridge, 2005; \ref{scale}). We gridded our study area into 12 |
| 175 | x 15 cells, each having a dimension of ${\sim}25$ x 33 km. We considered five iterations, as the |
| 176 | residuals rapidly decreased and stabilized. We assumed 2.5 $\%$ outliers, which were re- |
| 177 | moved at the second iteration. As starting model each grid node was set to the period |
| 178 | dependent mean velocity. (See SI for trade-off curves, Fig. S1, and outlier pre-selection.) |
| 179 | A selection of tomography results is illustrated in Figure S2. For shorter periods |
| 180 | (1–8.5 s), velocity contours roughly follow the geological boundaries along a NNE-SSW |
| 181 | direction. The highest velocities are around the $WC/KC/HC$ contact and decrease with |
| 182 | distance towards west and east. The lowest velocities are in the SE of the island. For |
| 183 | periods longer than 10 s the pattern changes towards a north vs. south subdivision of |
| 184 | velocity regions, instead of following the geologic boundaries. |
| | Final dispersion survey more than constructed from the terror marky results, con |

Final dispersion curves were then constructed from the tomography results, corresponding to the locations of the seismic stations. These dispersion curves are smooth and stable up to a period of 30 s.

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$_{188}$ 3 Receiver Functions and H κ -Stack Analysis

We considered all earthquakes with magnitudes M>5.5 and epicentral distances 189 of $30-90^{\circ}$ (based on the USGS catalog). The temporary network recorded 246 of such 190 earthquakes, the permanent stations registered 636 events. Most of the events are located 191 NE-SE of Sri Lanka within a back azimuthal range of $40-120^{\circ}$, specifically along the West-192 Pacific and Indonesian plate boundaries. The data cover a slowness range of 4.6–8.8 s/°. 193 To ensure good quality receiver functions we selected seismograms with signal-to-194 noise ratio (SNR) > 2.5. The selection process resulted in 1979 traces from 267 events 195 (see Table S1). For receiver function (RF) computation, each trace was filtered (band 196 pass: 0.05–5 Hz), decimated to a sampling rate of 20 Hz, and trimmed to 5 s before and 197 30 s after the P-onset. Subsequently, each trace was rotated from the ZNE into the LQT 198 ray coordinate system based on the theoretical incidence angle assuming a surface Vs 199 of 3.5 km/s. The Q-component was then deconvolved with the respective L component, 200 utilizing water level stabilization (level: 0.001) and low pass filtering with a Gaussian 201 function (Gauss factor: 1.0). The receiver functions were sorted according to slowness 202 and stacked in bins of 0.2 s/° without amplitude normalization. The bin-stacked RFs 203 show a coherent signal, as can be seen in Figure S3. The Q-RFs were not move out cor-204 rected for the final stack, as the move out correction has a strong effect on the multi-205 ple timing, which would result in biased interface depths in an inversion. 206

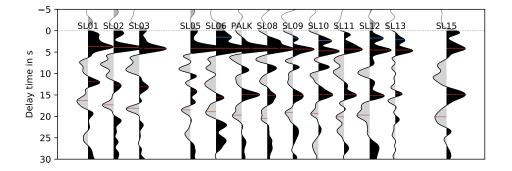


Figure 3. Receiver function stacks along main profile from SL01 to SL15. Brown markers indicate clear Ps converted phase at the Moho and corresponding multiples. Indications for a mid-crustal converter are marked in blue.

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Receiver functions along the main profile are illustrated in Figure 3. The independently derived station stacks show consistent phases throughout the profile, especially regarding the Ps phase and its multiples. An intra-crustal phase (including first multiple) is visible at the central stations SL06–SL10 and fades at SL11–SL13. This intra-crustal arrival indicates a discontinuity with a strong velocity contrast at about 15–25 km depth.

The H κ -stack grid search method following Zhu and Kanamori (2000) was applied 212 to estimate Moho depths and Vp/Vs values. The method is illustrated in Figure S4 for 213 a selection of stations. Moho depths for Sri Lanka range between 29.5–40 km with a typ-214 ical uncertainty of ~ 2 km; for Vp/Vs the range is 1.6–1.82 with an average of 1.72 and 215 a typical uncertainty of 0.06 (see Table S2). The Moho estimates strongly depend on the 216 crustal Vp, which we assumed to be 6.5 km/s for each station; this is reasonable for fel-217 sic amphibolite and granulite facies continental crust (?). A misestimation of 0.1 km/s 218 would result in a crustal thickness variation of about 1 km. The Vp/Vs ratios do not 219 depend significantly on the assumed Vp. 220

4 Bayesian Inversion of SWD and RF

As a final step, receiver functions and phase velocity dispersion curves were jointly 222 inverted with a Markov chain Monte Carlo (McMC) transdimensional Bayesian inver-223 sion tool (BayHunter; Dreiling & Tilmann, 2019), where we solve for the velocity-depth 224 structure, the number of layers, the noise parameters, and the crustal average Vp/Vs. 225 While other inversion methods often favor one best model based on the least misfit, an 226 inversion after Bayes' theorem is based on the model's likelihood and results in proba-227 bility distributions for each parameter of the model. The inversion result is represented 228 by a collection of models, the posterior distributions of which form ideally Gaussian dis-229 tributions if the chains have converged. For further details refer to ? and to the docu-230 mentation of BayHunter (Dreiling & Tilmann, 2019). 231

The model priors were set to a wide range, i.e., a depth range for the interfaces from 232 the surface to 75 km, Vs from 2 to 5 km/s, and Vp/Vs from 1.45 to 2.05. Additionally, 233 a maximum of 20 layers was imposed. The noise amplitude σ_{RF} spans from ~0 to 0.05, 234 and σ_{SWD} from ~0 to 0.05 km/s. The correlation r for the correlated noise for RFs was 235 fixed to a value of $r_{RF}=0.96$. For surface wave dispersion, the noise was assumed un-236 correlated. The model priors turn out to be sufficiently wide, relative to the values with 237 significant probability, i.e., none of the parameters inverted for have settled on a bound-238 ary. 239

The inversion was performed with 100 chains to ensure multiple independent pa-240 rameter search paths. Each chain performed 1.8 million iterations, with a 2:1 ratio for 241 the burn-in and exploration phase. The probability distributions for the proposal gen-242 eration were adjusted during the inversion to maintain an acceptance rate of ~ 40 %. Some 243 chains failed to converge, returning significantly higher misfits than most chains after 244 the burn-in phase. Such chains were declared as outlier chains. For the complete data 245 set, ~ 5 % of the chains were declared as outlier chains, which indicates that the chosen 246 number of iterations was usually sufficient enough for the chains to converge properly. 247 The final posterior distribution gathers 100,000 models from the main inversion phase 248 by sub-sampling all non-outlier chains. 249

The posterior distribution of 100,000 models was sorted according to likelihood and 250 categorized into three groups, including the best 25 %, 50 % and all models. Figure 4 251 shows an example of the McMC analysis for SL21, showing velocity-depth structures and 252 corresponding data fits from randomly selected models from each group, and the pos-253 terior distributions of likelihood, joint misfit, SWD and RF noise amplitudes, number 254 of layers and Vp/Vs for all models within a group. The grouping (colors) shows the com-255 promises the algorithm made during an inversion, e.g., increasing the number of layers 256 to reduce the noise level and the misfit. Each of the posterior parameters is unimodal. 257 The surface wave dispersion shows a good data fit. For the receiver functions, the ma-258 jority of modeled RF agrees very well in their signature, however, not all details of the 259 waveform can be matched. The first order features are modeled in nearly every chain 260 and the Vs-depth models show similar structures. The median model shows a sharp in-261 terface at 3 km depth and more gradual transitions at 13-15, 26-29 and 35-39 km; the 262 gradual transitions imply a higher uncertainty about the correct interface depth. The 263 Moho discontinuity lies between 35–39 km. 264

The quality of data fit for SL21 is representative for the other stations. (Data fits 265 and average velocity-structures are shown for all stations in the SI, Figs. S5a, S5b and 266 S6.) The posterior distributions are unimodal with the exception of the Vp/Vs of seven 267 stations, which show bimodal distributions, and the Vp/Vs of one station (SL31), which 268 did not yield plausible values (i.e., they settle on a boundary, even if extending the bound-269 ary to unrealistic values). Vp/Vs is a fine-tuning parameter, meaning, that the average 270 Vs-depth structures we derived from our data set are relatively insensitive to Vp/Vs. For 271 the seven stations showing bimodal Vp/Vs, the algorithm finds two Vp/Vs optima, and 272

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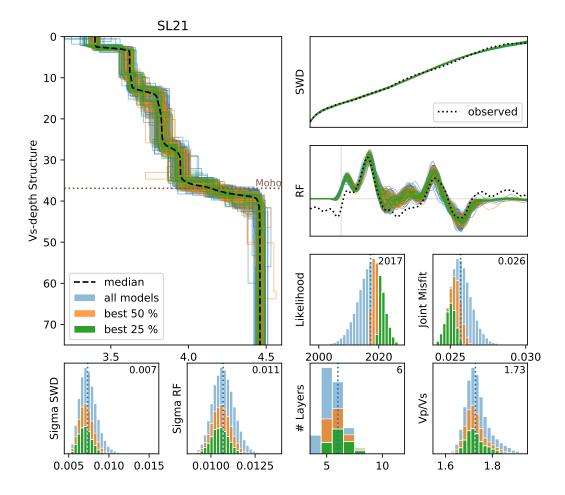


Figure 4. Selection of Vs models and corresponding data fits for station SL21, along with posterior distributions of likelihood, misfit, SWD and RF noise amplitudes, number of layers and Vp/Vs ratio. The results are color coded according to the likelihood, i.e., three groups showing 25 %, 50 % and 100 % of the best models. Dotted vertical lines illustrate the median, whose value is displayed in the upper right corner of each panel.

therefore compromises by slight modifications of the other parameters, but still leading 273 to Gaussian distributions for Vs. The most probable Vs-depth models corresponding to 274 either of the Vp/Vs optima, show equal major structures. For the station not converg-275 ing in Vp/Vs, we compared the results with those from an inversion assuming a fixed 276 Vp/Vs (=1.73), resulting in models that are very similar in their Vs-depth structure. 277 Figure 5a shows the posterior distribution for SL10 for the velocity-depth struc-278 tures, including mean and mode model, and interface depth probabilities. The surface 279 velocity is ~ 3.4 km/s, the interfaces are well defined at 3, 12, 20, and 38 km. The Moho 280

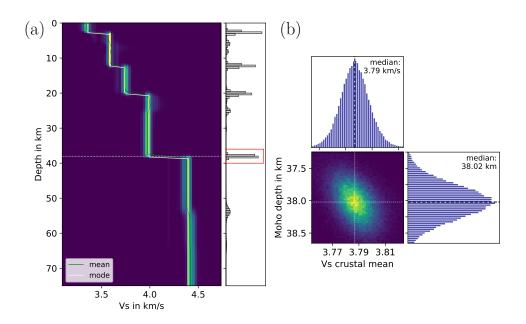


Figure 5. (a) Posterior distributions of shear wave velocities for station SL10 and interface depth probabilities. The red box marks the interface probability of the Moho. (b) Posterior distributions of the Moho depth and crustal average Vs. The median Moho depth is 38 ± 0.2 km and the average crustal Vs is $\sim 3.79\pm0.01$ km/s.

depth is between 37 and 39 km, emphasized by the red box in Figure 5a. To retrieve a 281 robust estimate for the Moho depth, the Vs model of each station was inspected to give 282 a pre-selected depth range (e.g., 37-39 km for SL10, 35-39 km for SL21). Each of the 283 100,000 models was then analyzed to find the interfaces within the pre-selected range, 284 with the last crustal layer having a Vs <4.2 km/s. (See Fig. 5b for the distribution of 285 Moho depths at SL10.) Additional parameters, e.g., Vs in the last crustal layer, aver-286 age crustal Vs, Vs increase across the Moho and upper mantle Vs were extracted. Those 287 values show a moderate trade-off between crustal thickness and velocity, as is illustrated 288 in Figure 5b; although subtle, a deeper Moho estimate is accompanied by a larger av-289 erage crustal Vs. This trade-off is well known for receiver functions, but reduced in its 290 impact by the inclusion of surface wave dispersion. The Moho interface for SL10 is at 291 38 ± 0.2 km and average crustal Vs is 3.79 ± 0.01 km/s. Maps of median Moho depths and 292 average crustal Vs are shown in Figure 6 with values as summarized in Table S2. 293

²⁹⁴ 5 Crustal Velocity Structure

Figure 6 shows Moho depths and Vp/Vs derived by joint Bayesian inversions (a, b) and H κ -stack analysis (d, e). Figure S7 illustrates a Vs cross section along the main profile (see Fig. 1 for profile location), comparing the differently derived Moho depths.

The Moho depths derived from joint inversion and $H\kappa$ -stacking generally agree well 298 with each other. Absolute differences between both methods are between 0.1 and 3.2 km 299 with a median difference of 0.7 km. The largest Moho depths (38–40 km) are found be-300 low the topographic high in the HC. The three northernmost stations in Sri Lanka (SL24, 301 SL25, MALK) also have a deep Moho interface at $>\sim 38$ km depth. The west coastal sta-302 tions SL01–SL03 show the thinnest crust (30–35 km). SL20 at the east coast also shows 303 a shallow Moho depth in the H κ -stack (\sim 33±2 km), but not in the joint inversion (\sim 36 km). 304 We note that there is a strong interface at ~ 31 km depth that might have been inter-305 preted as the Moho in the H κ -stacking (Fig. 7b, right); this station has the largest Moho 306 deviation of 3.2 km. 307

The Moho interface generally mirrors the topography, i.e., higher crustal elevations correspond to larger Moho depths (Fig. S7). The crustal thickness is continuously increasing from SL01 to SL05 (30–36 km), with a sudden increase of 3 km to 39 km at SL06, which corresponds to the topographic trend with an elevation change by a factor of 4 from SL05 to SL06. SL08 shows the deepest Moho interface, which is thus getting slightly shallower again towards the east coast.

The median Vp/Vs values from joint inversion are between 1.5–1.93, with the ma-314 jority between 1.68–1.8. The H κ -stack results range between 1.6–1.82, with the major-315 ity of the stations between 1.66-1.73. The differences of Vp/Vs from both techniques 316 are up to 0.28 with a median difference of 0.03 km. Vp/Vs results from both methods 317 agree in their general range for the study region, but do not show a common pattern. 318 $H\kappa$ -stack results for Vp/Vs are more reliable, as they include a range of RFs and con-319 sider the arrival times of Ps conversion and multiples directly associated with the slow-320 ness. For joint inversion, we considered the RF stack with its median slowness. 321

The average crustal Vs (Fig. 6c) ranges from 3.7–3.9 km/s, with increased velocities in the central HC (>3.83 km/s), decreasing with distance towards the coastal regions. The southern and westernmost coastal stations (SL01, SL02, HALK) have the lowest crustal Vs. Moho depths, crustal average Vs and Vp/Vs do not show a clear correlation, neither with each other, nor with the surface geologic units.

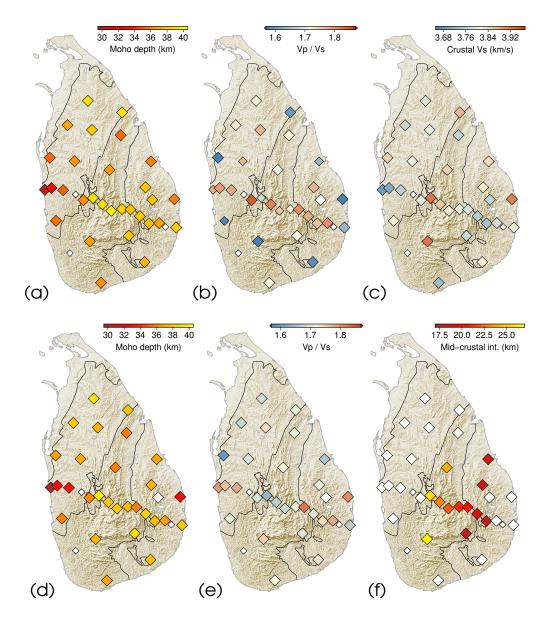


Figure 6. Spatial distribution of (a) Moho depths, (b) Vp/Vs ratio, (c) crustal average Vs and (f) mid-crustal interface depths derived from the McMC Bayesian inversion. (d) and (e) are the Moho depths and Vp/Vs ratios derived from the H κ -stack grid search, respectively. White symbols denote stations with no inversion performed (small diamonds) or no results gained (large diamonds). The smaller colored symbols in (b) indicate that the distributions are bimodal.

The cross section in Figure 7a (and Fig. S7) shows a prominent westward dipping 327 mid-crustal interface with an apparent angle of $\sim 4.3^{\circ}$ between SL06–SL12, and an av-328 erage velocity increase from 3.75 to 4 km/s. These lower crustal high velocities are ab-329 sent at the stations adjacent to the west (SL01-SL05), while the stations to the east (SL11-330 SL13) show a thinner or inter-layered section of the higher velocities. The probability 331 of interfaces (Fig. S7) furthermore suggest the dipping interface to be traceable across 332 the entire profile (SL03–SL15). Figure 7a shows our interpretation of the intra-crustal 333 interface; it is also evident on five other stations across Sri Lanka (Fig. 7b, left). Figure 334 6f shows the spatial extent of the mid-crustal discontinuity; values are summarized in 335 Table S2. 336

The mid-crustal interface is observable on the central stations in the HC, and on three additional stations in the VC. For the stations in the HC the interface depth lies between $\sim 18-27$ km; the interface in the VC is at ~ 18 km depth. The strike has an orientation similar to the geologic strike with a dip towards WNW.

A well constrained low velocity zone is observed along SL05–SL09, and SL13 at depths of 10 km, with Vs between 3.4–3.6 km/s. Stations SL15, SL20, SL23, SL25 and SL26 show low velocity zones at mid- to lower crustal depths (20/30 km); they are located at the western and eastern coastlines.

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6.1 Comparison with other geophysical studies

6 Interpretation and Discussion

The H κ -stack analysis from Pathak et al. (2006) reveals a Moho depth of 34 ± 1 km 347 beneath PALK, much shallower compared to our results from two independent analy-348 ses, which are 38.25 ± 1.9 km and 39 ± 0.3 km from H κ -stack and Bayesian inversion, re-349 spectively. This discrepancy can (partially) be explained by the average Vp they assumed 350 for the crust (6.0 or 6.1 km/s), which is lower than our Vp assumption of 6.5 km/s. Rai 351 et al. (2009) obtained a crustal thickness of 37.5 ± 1 km and Vp/Vs value of 1.721 ± 0.02 352 for PALK. Their estimates of Moho depth and Vp/Vs agree borderline with ours. By 353 forward and inverse modeling of RF and surface wave data they inferred a velocity-depth 354 structure with a low velocity layer in the upper crust and a mid-crustal discontinuity at 355 a depth of 22.5 km. We also observe a shallow low velocity zone (~ 10 km) and an intra-356 crustal discontinuity at 23 ± 0.4 km. 357

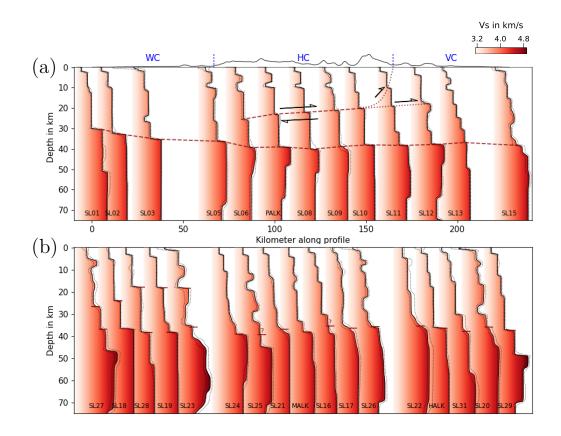


Figure 7. Vs models at stations along the main profile (a) and at other stations away from the profile (b). The shallow and deeper red markers indicate the mid-crustal and Moho interface, respectively. The dotted lines mark an interpretation of the outcropping mid-crustal interface, and a possible continuation within the crust. Shear sense indicators denote an ancient thrust. The Vs models in (b) are divided into 3 groups, from left to right, with stations that include the mid-crustal interface, stations in the WC, and stations located in the HC and VC, respectively. Moho depths labeled with a question mark indicate an interpretation leaned on surrounding stations (see also SL13).

| 358 | Mishra et al. (2006) used gravity data and modeled the anomalies along an E-W $$ |
|-----|--|
| 359 | profile through PALK. They modeled crustal thicknesses of up to $40-41$ km under the |
| 360 | eastern part of the HC, close to our observations, and explain the central gravity high |
| 361 | with a higher density crustal section protruding in the upper crust (10–15 km). We also |
| 362 | observe a central anomalous higher velocity section (which can correlate to higher den- |
| 363 | sities), however, situated in the lowermost crust. A fresh gravity modeling based on our |
| 364 | results might be of interest, but would exceed the scope of this study. |

365

6.2 Average crustal Vp/Vs

Vp/Vs can be helpful to distinguish between felsic and mafic rocks as a matter of the relative proportions of quartz (Vp/Vs~1.49) and plagioclase (Vp/Vs~1.87) (Christensen, 1996). Musacchio, Mooney, Luetgert, and Christensen (1997) grouped crustal rocks based on Vp/Vs and Vp into three categories: felsic, anorthositic and mafic rocks.

Classifying the results from H κ -stacking, none of the Vp/Vs lies beyond 1.82, which 370 would exclude an anorthosite rock composition. As our Vp/Vs for both analyses are gen-371 erally more on the lower end (1.66-1.73 and 1.68-1.8), most of Sri Lanka is represented 372 by felsic rocks with intermediate-to-high silica content. However, our Vp/Vs estimates 373 are crustal averages; it is possible that sections of the crust are dominated by different 374 compositions. The joint inversion shows average crustal Vs between 3.7–3.9 km/s (Vp: 375 5.9-7.3 km/s), which would predominantly still favor a felsic composition over a matic 376 one. 377

378

6.3 Moho depths and intra-crustal features

The Moho depths are not obviously correlated with the geologic units, which sug-379 gests that the crustal fragments have been unified through reworking and deformation 380 through the Pan-African collision and possibly later erosive processes. The Moho inter-381 face generally mirrors the topography, except for the thicker crust in the northernmost 382 part of the island, which might be caused by density differences through crustal com-383 position. The thinner crust along the west coast (<36 km), including the thinnest crust 384 at SL01-SL03 (30–35 km), can be explained by the formation of the adjacent Mannar 385 Basin, including rifting and crustal thinning. 386

Our study reveals a major WNW-dipping mid-crustal interface in the central HC 387 with an apparent dip of $\sim 4.3^{\circ}$ along the profile. Stations in the VC and close to the HC/VC 388 border (i.e., SL12, SL19, SL23, SL28) also show a discontinuity at a depth of ~ 18 km. 389 It is unclear whether the four stations see the same structure as the stations in the HC 390 or image a separate feature within the eastern VC. The mid-crustal interface might be 391 a feature from before the Pan-African collision; however, as the extent of the disconti-392 nuity from the HC into the eastern VC continues at the same depth and shows coher-393 ent Vs contrast, it is likely that the interface is the result of a shared event. Therefore, 394

we are inclined to interpret this mid-crustal feature as being related to the HC/VC thrust contact.

Kleinschrodt (1994, 1996) suggested that the HC was thrust onto the eastern VC 397 along a deep crustal, sub-horizontal to gently west-dipping thrust surface, which under-398 lies large parts of the HC, with a thrust geometry of a ramp-flat geometry or a low-angle 399 thrust that steepens to higher crustal levels. Our results are in agreement with this hy-400 pothesis, which is supported by several other studies (see introduction). The interface 401 might be interpreted as the HC/VC thrust contact that steepens to shallower crustal lev-402 els and the surface (Fig. 7a). Stations SL11-SL13 show a slightly different Vs structure 403 below the discontinuity (Fig. S7), i.e., high Vs inter-layered with lower Vs, which could 404 reflect the complicated contact zone between the HC and VC and might even image a 405 buried continuation of the thrust contact within the VC, i.e., a blind thrust. 406

A low velocity layer as we observe at the central stations within the HC, was also 407 observed by Rai et al. (2009) in the upper crust of other Pan-African terranes. Such intra-408 crustal structure is assumed to be a relic of deformation and magmatism caused by up-409 welling of lower crust or subcrustal melts. Low velocity zones are thought to be the con-410 sequence of an influx of CO_2 -rich fluids, that are trapped at these depths or originated 411 from retrograde metamorphism to amphibolite and greenschist facies, and were brought 412 there through deep-seated thrusting and lateral shearing during a transpressive regime 413 (Rai et al., 2009, and references therein). 414

415

6.4 The amalgamation of Sri Lanka

The hypothesis of the stepwise collision predicts westwards dipping thrust contacts 416 between the WC/KC, the HC and the VC island arcs (e.g., K. V. W. Kehelpannala, 2004). 417 We observe a gently westward dipping mid-crustal interface beneath the HC which shows 418 a strong Vs increase and thus indicates a change of rock material. Our observation matches 419 the proposed position and orientation of the HC/VC thrust contact. The velocity change 420 is also seen within the eastern VC, which suggests that the structure might has been part 421 of the VC crust before thrusting, or evolved alongside. We assume a steepening of the 422 thrust contact to the surface; as the signature of the mid-crustal interface in the central 423 HC does not disappear, but fades towards the east, we propose a buried continuation 424 within the VC. The low velocity layer in the HC along the main profile (~ 10 km depth) 425 might be caused by influx of CO_2 -rich fluids of retrograde metamorphism to amphibo-426

lite facies brought about by deep-seated thrusting and lateral shearing during a transpressive regime. The dipping mid-crustal interface and the low velocity zone, both relate to deeper thrusting and a transpressive regime, which clearly favors the stepwise collision theory as described with its details (see section 1.1).

Does this exclude the possibility of an amalgamation through divergent double subduction? Santosh et al. (2014) sketch the amalgamation with processes such as slab melting and arc magmatism, basaltic underplating, astenospheric upwelling and slab breakoff; features, that are also included in the stepwise collision and we cannot asses with our data. In their study, Santosh et al. (2014) did not focus on crustal structures, which makes it impossible for us to discuss the matter based on their information.

Divergent double subduction occurs rather rarely (e.g., the Lachlan fold belt in southern Australia, the Molucca Sea collision zone in Indonesia). Also, a large scale dipping structure in the accretionary zone is not a feature seen for this type of subduction (e.g., ??); however, the crustal structure is mostly not the focus of these studies. As argued above, we prefer the stepwise collision hypothesis, although we cannot completely exclude the theory of the double-sided subduction.

443 7 Conclusions

Rayleigh wave dispersion curves from ambient noise correlation and receiver functions were computed and jointly inverted using McMC transdimensional Bayesian inversion. Based on the median of the posterior distributions received from each station, we evaluated the crustal velocity structure of Sri Lanka.

Our results show a Moho interface at 30–40 km depth with a distinct velocity increase. The Moho depths show no correlation to the geologic units and largely mirror the topography, which suggests Airy isostacy for most of the Sri Lankan continental crust. A thicker crust in the northernmost part of the island might be caused by compositional effects on density. The lower Moho depths along the west coast emerged presumably through rifting and crustal thinning processes through the formation of the adjacent Mannar Basin.

We identify a prominent intra-crustal interface beneath the HC (18–27 km), and the eastern VC (~18 km). We relate this westward dipping interface to the HC/VC thrust contact, which steepens to shallower crustal levels. The interface within the VC might have been part of the VC unit before the thrusting event, or evolved alongside. A low velocity zone in the central HC supports deep-seated thrusting and lateral shearing dur-

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⁴⁵⁹ ing a transpressive regime. Our results clearly favor the amalgamation theory of a step-

460 wise collision of arc fragments to form Sri Lanka.

461 Acknowledgments

462 Many thanks to Dr. Robert Trumbull for the vivid discussions about geologic interre-

463 lationships. We thank the GSMB of Sri Lanka, which made it possible to deploy a tem-

464 porary seismic network in Sri Lanka, and the employees for their local support and en-

465 gagement during the field work. We acknowledge the GFZ for funding the field exper-

- ⁴⁶⁶ iment from expedition funds, the Geophysical Instrument Pool Potsdam (GIPP) for sup-
- ₄₆₇ plying the instruments, the GEOFON data center for hosting the project data (https://

468 geofon.gfz-potsdam.de/waveform/archive), and DFG research support through grant

469 TI316/4-1.

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