Deep and rapid thermo-mechanical erosion by a small-volume lava flow

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Abstract

We document remarkably efficient thermo-mechanical erosion by a smallvolume lava flow. Downcutting by a basaltic-andesite lava flow on the steep-sided Momotombo volcano, Nicaragua, occurred at 100 times the rate commonly reported for thermal erosion in lava flow fields, even though this flow was small-volume $(0.02 \,\mathrm{km}^3)$ and effused at a low rate for <1week. The erosion depth, up to 30 m incision, is explained by reduction of hardness, H, of the pyroclastic substrate into which the lava flow incised. We show that incision depth decreases, approximately exponentially, with distance along the flow path, until erosion stopped and the flow became constructional. This transition occurs 650 m from the vent on a slope averaging a 32° incline. Results indicate that syn-eruptive erosion is an important morphological process on some steep-sided volcanoes that are predominantly composed of layered pyroclasts. Rapid erosion and incision increased flow run-out for the 1905 flow, which in turn directed the flow and run-out of the 2015 lava flow. Mapping and understanding these features is critical for improving lava flow hazard assessments and provides insight into the construction and growth of composite cones.

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Keywords: lava flow physics; thermal and mechanical erosion; lava inundation; volcano morphology; remote sensing; thermal dynamics of lava flows

1 Introduction

Lava flows are responsible for altering landscapes on geologically short timescales. The overwhelming majority of lava flows construct topography by building raised channels and/or compound flow fields, both of which evolve through time and along-flow (Kerr et al., 2006; Dietterich and Cashman, 2014). The 31 morphologies of flow features are mainly determined by the composition and effusion rate of the flow, as well as the pre-existing and syn-eruptive topography (Richardson and Karlstrom, 2019; Bilotta et al., 2019). These factors also control the thickness of lava flows, which in turn influence a flow's run-out distance 35 and inundation hazard potential (Kilburn and Lopes, 1988; Dietterich et al., 2017). A small fraction of channelized flows and lava tubes erode into older surfaces during emplacement via thermal, mechanical, or thermo-mechanical processes (Greeley et al., 1998; Fagents and Greeley, 2001; Kerr, 2001; Siewert and Ferlito, 2008; Hurwitz et al., 2010, 2013). 40 The 1905 eruption of Momotombo volcano, Nicaragua, provides an example of thermo-mechanical erosion by a small volume ($<0.02 \,\mathrm{km}^3$) lava flow on 42 a steep-sided edifice (Figs. 1 and 2). We first document the morphology of the channel using a combination of satellite and terrestrial radar generated dig-44 ital elevation models (DEMs) from 2012-2017. Erosion depths from the 1905 flow are then calculated by reconstructing pre-channel topography, extracting cross-sectional profiles, and calculating the maximum difference between the measured and modeled surfaces normal to the channel. We use these results to test thermal and thermo-mechanical models of erosion. Model inputs are in-49 formed by observations from Momotombo's most recent eruption in 2015, which we capture with the range of our satellite observations. We find the channel was 51 thermo-mechanically eroded by a lava flow that erupted in 1905. Additionally, we assert that thermo-mechanical erosion is an important morphological process

- on some steep-sided volcanoes composed predominantly of layered pyroclasts.
- This study is the first to look at lava flow erosion on steep sided slopes, expands
- our knowledge of the rate at which syn-eruptive erosion occurs, and mathemat-
- 57 ically couples thermal and mechanical models of erosion.

58 2 Background

59 2.1 Erosion by Lavas

- Erosion by lava has been hypothesized for the formation of rilles on both Mars
- 61 (Dundas and Keszthelyi, 2014) and the Moon (Head and Wilson, 2017; Wilson
- and Head, 2017), canali on Venus (Baker et al., 1992; Williams-Jones et al.,
- 1998), and channels on Io (Schenk and Williams, 2004). Studies of active erosion
- by flowing lava have occurred on the island of Hawai'i during the 1972–1974
- Mauna Ulu eruption and the initial stages of the 1983-2018 Pu'u O'o eruption,
- where erosion rates of 4 cm depth/day and 10 cm depth/day were observed in
- lava tubes via skylights, respectively (Peterson et al., 1994; Kauahikaua et al.,
- 68 1998). Erosion by turbulent komattite flows during the Archean, responsible
- for large Ni-sulphide ore deposits, is also widely noted (Williams et al., 1998;
- ⁷⁰ Beresford et al., 2002; Staude et al., 2017).

71 Thermal Erosion

- Thermal erosion occurs when lava moves with sufficient flux and temperature to
- ₇₃ melt and incise the underlying terrain (Kerr, 2001). Thermal erosion by flowing
- 14 lava requires the complete or partial melting and assimilation of a substrate
- into the overriding flow. A lava flow's total available thermal energy $(E_{thermal})$,
- ⁷⁶ sourced from advection and crystallization, is modeled as:

$$E_{thermal} = m_l[c_{pl}(T_l - T_a) + \phi F] \tag{1}$$

where m_l is the mass of the erupted lava, c_{pl} is the specific heat capacity of the lava, T_l is the erupted temperature of the lava, T_a is the temperature of the environment into which heat is being transferred (the substrate in this case), ϕ 79 is the mass fraction crystallization, and F is the latent heat of fusion (Wooster et al., 1997). Lava flows are an open system where available thermal energy 81 is eventually balanced out by heat loss through conduction, convection, and radiative heat transfer. Studies of the thermal energy balance of lava flows on 83 Mt Etna show that upwards of 85% of energy was retained during the initial phases of the eruption, which can be used to further bound the amount of energy available to melt and erode the substrate (Wooster et al., 1997; Patrick et al., 2004). The presence of multiple heat sinks also highlights the fact that not all 87 available thermal energy can be partitioned into eroding the substrate, so we 88 need to quantify heat transfer between the base of the lava flow and substrate. The rate of conductive heat transfer into the substrate (i.e., the growth of a thermal boundary layer) can be modeled as:

$$\frac{dy}{dt} = \frac{\eta_T \sqrt{\kappa}}{t} \tag{2}$$

where y is the depth into the substrate, t is the duration of the flow, and η_T is a dimensionless similarity variable (Turcotte et al., 2002; Fagents and Greeley, 2001). The growth of this boundary layer is illustrated in Figure 3.

5 Mechanical Erosion

Mechanical erosion occurs when the wearing material (the lava flow) is harder than the substrate (the edifice) (Sklar and Dietrich, 1998; Siewert and Ferlito, 2008; Hurwitz et al., 2010). The early stages of an eruption are most conducive

to erosion because flow velocity is often highest and basal friction is also high due to the vertical load (Siewert and Ferlito, 2008; Hurwitz et al., 2010). Mechanical 100 erosion as a function of substrate hardness can be modeled as: 101

$$H = \frac{k\rho ghvtsin\theta}{d_{channel}} \tag{3}$$

where H is the hardness of the substrate, k is a dimensionless proportionality constant that captures the material hardness of the erosive layer (the lava flow), 103 ρ is the density of the lava flow, h is the thickness of the lava flow, v is the velocity of the lava flow, θ is the slope of the edifice, and $d_{channel}$ is the depth 105 of erosion. This equation implies that the depth of an eroded channel will be 106 constant, so long as the velocity of the flow is constant. If the depth of the 107 channel changes with distance, then either the velocity of the flow is changing, 108 or the hardness, H, is changing, or both. 109 The yield strength is an exponential function of temperature, which means 110 111

that H can also be modeled as a function of temperature:

$$H \approx \frac{ae^{(-bT_a)}}{3} \tag{4}$$

where a and b are flow-dependent variables that vary with magma composition 112 and e is Euler's number. This relationship describes softening of the substrate 113 as temperature increases or for a longer duration of heating, which means that 114 mechanical lava erosion models presented in Siewert and Ferlito (2008) can be 115 reworked to become thermo-mechanical models. Values of 1.18 \times $10^5\,\mathrm{Pa}$ and 116 $0.12\,\mathrm{K^{-1}}$ correspond to a and b for Momotombo. We note that experimental data have indicated that the Arrhenius relationships shown in Equation 4 break 118 down around the glass transition point (Miller, 1963; Gottsmann and Dingwell, 119 2002). 120

2.2 Geology of Momotombo and Recent Activity

Momotombo (1,297 m) is located at the southern end of the Cordillera de Los 122 Maribios in central Nicaragua (Fig. 1). The edifice is composed primarily of basaltic to basaltic andesite lavas, cinders, and other tephra that erupted dur-124 ing the last 4,500 years (Kirainov et al., 1988). Sixteen historical eruptions 125 have been documented, the majority of which have been strombolian to violent 126 strombolian (VEI 1-2), with several plinian events (up to VEI 4) (Global Vol-127 canism Program, 2017). A VEI 4 eruption in 1605-1606 and large earthquake in 128 1610 lead to the abandonment of city of León (Viejo), the capitol of the region 129 at that time (Sapper, 1925). Though the specific morphological changes to the 130 edifice from the 1605-1606 plinian event are not well documented, it's possible 131 that serious damage to the structural integrity of the summit occurred given the impact on surrounding municipalities (Sapper, 1925). The subsequent steady 133 activity throughout the 1800's rebuilt the summit from cinders, agglutinate, and channelized lava flows, as shown in photographs from the late 1800's and 135 early 1900's (Vincent, 1890; Intercontinental Railway Commission, 1898; Sap-136 per, 1925). Analogous volcanoes, such as Ngauruhoe (New Zealand) (Hobden 137 et al., 2002), Izalco (El Salvador) (Carr and Pontier, 1981), and Cerro Negro 138 (Nicaragua) (Hill et al., 1998; Courtland et al., 2012), have built moderately 139 sized pyroclastic cones in only a few hundred years and can provide insight into 140 the constructional history of Momotombo. 141 The 1905 eruption (VEI 2) occurred between January 16–21 (Sapper, 1925). 142

The 1905 eruption (VEI 2) occurred between January 16–21 (Sapper, 1925).

The basaltic andesite lava flow was accompanied by an eruptive column of sufficient height to deposit ash 15 km to the west on the city of León. The effusive
component consisted of an eruptive volume of <0.02 km³ of basaltic andesite
(Fig. 1). Intermittent explosions occurred that sent incandescent blocks and
bombs "a great distance" from the crater (Sapper, 1925). First person accounts

also describe destruction to the summit during the eruption, which may have resulted in the drainage of a small summit lava lake (Sapper, 1925) (Fig. 2). The 1905 eruption was followed by 110 years of quiescence, which ended on 30 November, 2015. A small volume basaltic andesite lava flow was emplaced between 1 December and 7 December, 2015, and was followed by several months of intermittent explosions (Global Volcanism Program, 2017).

54 3 Methods and Results

3.1 Digital Elevation Model Generation

156 3.1.1 TanDEM-X Satellites

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(TDX) and collected on 24 October, 2012 and 18 March, 2017. These DEMs, 158 which also capture the change in topography due to the eruption in 2015–2016, 159 allow us to obtain baseline measurements for the 1905 channel (Fig. 4) and 160 determine if any erosion occurred during the most recent eruption. The bistatic mode of TDX allows these two satellites to fly in tandem formation and observe 162 the same ground point simultaneously (Krieger et al., 2007). We note that the 163 flight paths for these acquisitions were not the same, which resulted in an offset 164 due to a heading difference of $\sim 21^{\circ}$. 165 GAMMA software was used to process the TDX SAR images to generate 166 DEMs with the InSAR (Interferometric Synthetic Aperture Radar) technique 167 (e.g., Deng et al., 2019). A 30-m SRTM (Shuttle Radar Topography Mission) 168 DEM provided independent ground control points. Two (range) by two (az-169 imuth) pixel multilooking was used to reduce speckle noise. The final DEMs have a spatial resolution of 5×5 meters with a vertical precision of < 2 m.

Digital elevation models (DEMs) were generated from TanDEM-X Satellites

3.1.2 Terrestrial Radar

We employed terrestrial radar interferometry (TRI) to assess the level of noise 173 in our topographic profiles (described below) from the 2017 TDX acquisition. Although this comparison does not give us a direct assessment of noise for the 175 2012 DEM (the model from which we are measuring channel incision depths), it gives a relative level of confidence in our TDX DEM processing methods. 177 TRI is a ground-based scanning radar that measures the amplitude and phase 178 of a backscattered microwave signal. A GAMMA real aperture radar operating 179 at Ku-band (1.74cm wavelength) was used for this study. The TRI has one 180 transmitting antenna and two receiving antennas, which allows for topographic 181 mapping with a single scan (e.g. Dixon et al., 2012; Caduff et al., 2015; Voytenko 182 et al., 2015; Xie et al., 2018; Deng et al., 2019). The resolution of the range measurements is ~ 1 m, and the azimuth resolution varies linearly with distance 184 (e.g., 1.8 m at 1 km distance, 7 m at 4 km). The spatial coverage of TRI is much smaller compared to satellite imaging. Details of TRI data processing for DEM 186 generation are given in Strozzi et al. (2012) and Xie et al. (2018). TRI surveys 187 were conducted in December 2015 and April 2016. We use results from the 188 2016 campaign because it occurred towards the end of the eruption period and 189 is temporally closer to the 2017 TDX acquisition. The TRI DEM has a 5 \times 5 m 190 resolution with an accuracy <5 m and covers the incised portion of the channel 191 (Fig. 5). 192

3.2 Channel Profiles and Depth Calculation

Previous GIS-based methods used to determine paleotopography of volcanic terrains (e.g., Germa et al., 2015) interpolate missing topography based on connecting high points in elevation. Studies of fluvial channel erosion in steep terrain generally do not deal with incision into conical edifices (Robl et al., 2008;

Fox, 2019). Additionally, these approaches model down-section and not crosssection morphology, which we require to accurately measure incision depth and 199 extract cross-section profiles. We developed an elliptical least-square best-fit 200 contour method to obtain incision depths along the channel on Momotombo's 201 steep slopes to fill this application gap. We use this method to obtain channel 202 depths and cross-channel profiles from the 2012 and 2017 TDX DEMs and the 203 2016 TRI DEM. Our depth measurements are minimum values, as it is likely 204 that the eroding flow would have emplaced some volume of lava within the 205 channel. We also use this method to determine if any incision occurred during the 2015 lava flow. Additionally, comparing the post-2015 TDX and TRI DEMs 207 provides an indication of noise within our channel measurements.

This approach measures incision depth against modeled paleotopography 209 created from optimized elliptical contours. A path down the channel's center was defined with a sampling density set to the resolution of the DEM (5 m 211 for this study) (Fig. S1a). The widths of the channel and levees, determined 212 visually, were masked out in order to separate their signal from the overall signal 213 of the cone. A refined elliptical fit for the uppermost contour of the channel path 214 was calculated by minimizing the mean-squared difference between the actual 215 elevation contour and an elliptical contour. This optimized ellipse was then used 216 to calculate a fit to the elevation contour below. This second recalculated ellipse 217 was then fit to the next elevation contour, and so forth, until the end of the 218 designated channel path was reached. The output of this process is a modeled 219 paleotopography with no channel or levee structures (Fig. S1b). The incision 220 depth was calculated against the modeled paleotopography. The normal vector to the paleotopographic model was calculated at each point, spaced 5 m apart 222 along each elliptical contour. As before, the normal vector was calculated by 223 fitting a plane to 8 adjacent points (three from the contour above, three from 224

the contour below, and two adjacent points from the same contour) (Fig. S1c). The DEM was then re-orientated such that the z-axis was coincident with the 226 calculated normal vector. The incision depth was returned as a weighted average 227 of the constituent points, with weighting criteria based on the distance of the 228 points to the center of the plane (i.e., the point closest to the center had the 229 greatest weight) (Fig. S1d). This process was repeated for each point along the 230 contour, and then for every contour. This process allowed us to measure the 231 incision depth throughout the channel while removing the conic signal of the 232 edifice (Fig. 6). Incision depth varied from 35 m at the summit rim and tapers 233 off to 0 at ~ 600 m elevation (Fig. 7). We calculate the eroded volume of the 234 channel to be 4×10^5 m³. A profile of each cross-channel contour was calculated for the 2012 and 2017 TDX and 2016 TRI DEMs (Fig. 8). 236 Results show the pre-2015 eruption channel extended down the northeast side of the edifice from the summit and incised into the summit (Fig. 6). The 2015 lava flow follows the same path as the pre-eruption channel (Fig 8). The

Results show the pre-2015 eruption channel extended down the northeast side of the edifice from the summit and incised into the summit (Fig. 6). The 2015 lava flow follows the same path as the pre-eruption channel (Fig 8). The elliptical contour fit method was also applied to the 2016 TRI DEM. Data gaps within the DEM were filled using a regularized spline with tension interpolation method in QGIS and the same process described above was utilized. We provide the code for this method (S1) and an additional code for a circular fit method in the Supplemental Documentation (S2).

$_{\scriptscriptstyle 245}$ 4 Discussion

246 4.1 Channel Origins

The difference between modeled paleotopography and the 2012 TDX DEM shows that a channel has cut into the edifice. The distribution of lava flows beneath the 1905 units, shown in Figure 1, implies that no structure existed in

this location prior to 1905 to consistently direct the path of subsequent lava flows (as was the case for the 2015 eruption). No historical reports note a leveed lava 251 channel on the NE flank of Momotombo prior to the 1905 eruption (Vincent, 252 1890; Intercontinental Railway Commission, 1898; Sapper, 1925). Examination 253 of the area surrounding Momotombo shows no down-slope deposition of suffi-254 cient volume to support a channel having carved into the NE edifice prior to the 255 1905 eruption by environmental erosion (e.g. hydrologic erosion) and then infilled by subsequent lava flows. The rest of the edifice is similarly devoid of any 257 large-scale drainage features (e.g. barrancas, rilles, or gullies). DEM analyses show the channel width is uniform from summit to low on the slopes (Fig. 2a). 259 Incised channels on other volcanoes (e.g., Merapi, Nevado del Ruiz) are generally much wider at the mouth, less consistent in width, and are less linear than 261 Momotombo's channel because they are related to more violent hazards (e.g., pyroclastic density currents and lahars). Results also show it was unlikely that 263 erosion occurred during the 2015 lava flow, given that it flowed over the armored 264 channel and not a variably consolidated slope of cinders and spatter/agglutinate (Fig. 7). In the absence of evidence that suggests otherwise, we conclude that 266 channel most likely formed during the emplacement of the 1905 lava flow. 267

268 4.1.1 Thermal Erosion

The total energy emitted by the 1905 lava flow, calculated using Equation 1, is about $7 \times 10^{16} \,\mathrm{J}$ for an eruptive volume of $2 \times 10^7 \,\mathrm{m}^3$. Taking into account the energy balance between heat sources (85% retained) and sinks (15% lost) described in Wooster et al. (1997), we calculate that $6 \times 10^{16} \,\mathrm{J}$ is the minimum amount of the original energy that remained within the flow. Although the length of time used to calculate this ratio for the Mount Etna eruption was greater than the duration of the Momotombo eruption, it is helpful to place a first order constraint on how much thermal energy is lost into the environment.

We calculate the total energy needed to fully erode the channel is about 2 \times 10¹⁴ J, by substituting the mass of the eroded section of the channel into Equation 1. Although we find that sufficient energy exists within the system 279 to erode the substrate, we note that not all of the energy present is available for this purpose. Had all of this energy been used to erode, the thermal loss 281 would have been sufficiently great that the emplaced flow would be much thicker 282 and shorter than what is observed. Modeling the depth of heat transfer into 283 the substrate as a function of time using Equation 2 estimates the growth of 284 the substrate thermal boundary layer at 8.8 cm per day. Thermal erosion rates in Hawaiian tubes and channels can reach ~10 cm per day (Kauahikaua et al., 286 1998); given the lower temperature of lavas erupted on Momotombo, a slower rate of thermal boundary layer growth makes sense. Given the short duration of 288 the 1905 eruption (<1 week; Sapper, 1925), thermal erosion by itself is unlikely to have formed the observed morphology. 290

1 4.1.2 Thermo-mechanical Erosion

Conceptually, once the near-subsurface reaches the threshold temperature, which 292 we assert is the glass transition (1013 K, Giordano et al., 2005), the hotter sub-293 surface material is transported downhill by the lava flow. For this assumption 294 to be correct, the slope has to be steep and the lava flow velocity at its base has to be fast. This implies that the thermal boundary layer reaches some critical 296 thickness and is then eroded away mechanically by the lava flow. While this process surely contributes to incision of the channel, it cannot fully account for 298 the observed depths. It is possible H varies with distance along the channel because the underlying substrate changes from very soft material at the top of 300 the cone to harder material lower on the flanks. We reject this because the sur-301 face of the edifice is a relatively uniform slope without major lithologic changes. 302 These results provide insight into how heat is transferred into the substrate over

time, which helps us understand how the material hardness of the substrate, H, is reduced over time as the lava flow moves downslope.

We assume that the growth and removal of the thermal boundary layer over 306 time can be approximated as a steady state process, captured by the depth 307 of incision at each point along the channel. An exponential was fit to the 308 incision depth data to extract the quadradic function of $d_{channel}$ in terms of 309 incision depth and elevation (Fig. 10). We solve for H and find that the lowest 310 value (corresponding to softest substrate) was $4.8 \times 10^5 \,\mathrm{Pa}$ corresponding to 311 an incision depth of 30 m (Fig. 10). These results show the hardness of the 312 substrate increases as a function of time, but note that we use depth as proxy 313 for time. Realistically, the deepest incision points are closest to the vent because they have been exposed to erosive work for a greater amount of time. The 315 function modeled to from Equation 3 tackles that problem from the opposite perspective (i.e., it models the decrease in erosive depth, which suggests increase 317 in hardness). We know that temperature is increasing in the substrate based on 318 the relationships described in Equation 2, which decreases the hardness, so we 319 revise these results and report them as a decrease in hardness over time to more 320 accurately reflect the physical processes controlling the incision depth. Detailed 321 historical observations list the eruption duration as six to seven days in length, 322 which allows us to place constraints on the rate of change of hardness (Fig. 9) 323 (Sapper, 1925). 324

325 4.2 Channel Growth

We calculate that the eroded volume is equal to roughly 2% of the total volume of the 1905 lava flow. Given the low eruptive temperature of basaltic andesites, it is unlikely that this material was fully assimilated into the flow. As the material heats and softens it is likely dragged downslope along the base of the flow for

a short distance and then re-deposits, which creates a scalloped type signal for the measure of incision depth against elevation (Fig. 7). The thicknesses of 331 subsequent lava flows will respond to this subtle topographic variability (i.e., 332 more lava will be deposited in the troughs, less on the crest), which we find to 333 be true for the 2015 flow (Fig. 7). We also note a strong correlation between 334 incision and lack of developed levees in the 2012 TDX DEM, which suggests that 335 erosion began early on during the eruption (Fig. 8, conceptualized in Fig. 12). 336 This implies that levee bounded incised channels are not thermo-mechanically 337 eroded, which is an important consideration for studies of incised channels on other planetary bodies. 339

340 4.3 Morphologic Implications

The depth of a thermally eroded channel is limited by the efficiency of heat 341 transfer across the boundary between the lava flow and the substrate, the rate of heat transfer in the subsurface, and the duration of an eruption (Kerr, 2001); 343 when combined with morphologic studies of emplacement conditions (especially time), a constraint can be placed on the maximum depth of erosion. Channels 345 whose incision depth exceeds this threshold indicate a preferential hardness 346 ratio of the substrate to the flow (i.e., it is softer than the overriding flow) and can therefore be used to determine the presence of pyroclastic rocks and other 348 easily eroded terrain (e.g., unconsolidated regolith, alluvial deposits, etc.) on 349 planetary surfaces that can only be observed remotely. Large flows on the Moon, 350 in particular, may be worth revisiting in light of these findings (Hurwitz et al., 2013). 352

A 1528 drawing of Momotombo by Oviedo (Fernandez de Oviedo y Valdes, 1528) shows a similar channel on the volcano's west side. A channel on the northern flank (likely emplaced during an eruption in the second half of the 1800's Sapper (1925)) is also incised into the cone, and has been infilled by cinders from subsequent eruptions, which suggests the processes of slope incision
and subsequent infill occurs with relative frequency on Momotombo. The preferential diversion of lava flows into incised channels for future events suggest
that lava flow hazards on some steep-sided volcanoes are influenced by the creation, infill, and eventual abandonment of these structures. The channel may
limit the lava flow hazard for the western flank if the next eruption is similar in
size. Understanding the evolution of these features has important implications
for lava flow hazards and growth patterns and erosion of composite volcanoes.

5 5 Conclusions

We use satellite (TDX) and terrestrial radar (TRI) DEMs to obtain a detailed record of recent changes to the edifice of Momotombo Volcano from 2012–2017, 367 during which a VEI-2 eruption occurred. We describe a unique lava channel 368 that incised 25–35 m into the northeast sector of the volcano near the summit 369 and transitions into a constructional channel roughly halfway down the edifice. 370 We assert that this feature formed erosively during the emplacement of a lava 371 flow in 1905 and note a direct correlation between a lack of levees and incision 372 depth. Thermal erosion alone was unable to account for the full depth of incision and we suggest that thermo-mechanical erosion is the likely cause. We examine 374 inputs from mechanical models of erosion and determine that, based on the relationship between material hardness and shear strength, these models should 376 be re-classified as thermo-mechanical. We propose that the transfer of heat into the substrate decreases the hardness of the material, which encourages it to 378 flow more readily and excavate. We establish that the critical temperature at 379 which this occurs is lower than previously thought, likely at the glass transition 380 temperature (1013K), instead of the liquidus of a given lava. We calculate the 381

- total eroded volume to be $4 \times 10^5 \, \mathrm{m}^3$ and determine a minimum hardness of 4.8
- $\times 10^5 \,\mathrm{Pa}$ at the deepest point of incision for the 1905 eruption. Deeply incised
- channels control the distribution of future flows and can also be used to infer
- the material properties of the substrate into which they are excavated.

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References

- Baker, V. R., Komatsu, G., Parker, T. J., Gulick, V. C., Kargel, J. S., and
- Lewis, J. S. (1992). Channels and valleys on Venus: Preliminary analysis of
- Magellan data. Journal of Geophysical Research. Planets, 97(E8):13421.
- Beresford, S., Cas, R., Lahaye, Y., and Jane, M. (2002). Facies architecture of an
- archean komatiite-hosted ni-sulphide ore deposit, victor, kambalda, western
- australia: implications for komatiite lava emplacement. Journal of Volcanol-
- $ogy \ and \ Geothermal \ Research, \ 118(1):57-75.$
- Bilotta, G., Cappello, A.and Herault, A., and Del Negro, C. (2019). Influ-
- ence of topographic data uncertainties and model resolution on the numerical
- simulation of lava flows. Environmental Modelling and Software, 112:1–15.
- ⁴⁰⁴ Caduff, R., Schlunegger, F., Kos, A., and Wiesmann, A. (2015). A review of ter-
- restrial radar interferometry for measuring surface change in the geosciences.
- Earth Surface Processes and Landforms, (2):208.
- 407 Carr, M. J. and Pontier, N. K. (1981). Research paper: Evolution of a young
- parasitic cone towards a mature central vent; Izalco and Santa Ana volcanoes
- in El Salvador, Central America. Journal of Volcanology and Geothermal
- 410 Research, 11:277 292.
- Courtland, L. M., Kruse, S. E., Connor, C. B., Connor, L. J., Savov, I. P.,
- and Martin, K. T. (2012). Gpr investigation of tephra fallout, cerro negro
- volcano, nicaragua: a method for constraining parameters used in tephra
- sedimentation models. Bulletin of Volcanology, (6):1409.

- Deng, F., Rodgers, M., Xie, S., Dixon, T. H., Charbonnier, S., Gallant, E. A.,
- Vlez, C. M. L., Ordoez, M., Malservisi, R., Voss, N. K., and Richardson,
- J. A. (2019). High-resolution dem generation from spaceborne and terrestrial
- remote sensing data for improved volcano hazard assessment a case study at
- nevado del ruiz, colombia. Remote Sensing of Environment, 233:111348.
- Dietterich, H., Lev, E., Chen, J., Richardson, J., and Cashman, K. (2017).
- Benchmarking computational fluid dynamics models of lava flow simulation
- for hazard assessment, forecasting, and risk management. Journal of Applied
- Volcanology, 6(1):1.
- Dietterich, H. R. and Cashman, K. V. (2014). Channel networks within lava
- flows: Formation, evolution, and implications for flow behavior. Journal of
- Geophysical Research. Earth Surface, 119(8):1704.
- Dixon, T. H., Voytenko, D., Lembke, C., Pea, S., Howat, I., Gourmelen, N.,
- Werner, C., and Oddsson, B. (2012). Emerging technology monitors ice-sea
- interface at outlet glaciers. Eos (0096-3941), 93(48):497.
- Dundas, C. M. and Keszthelyi, L. P. (2014). Emplacement and erosive effects
- of lava in south Kasei Valles, Mars. Journal of Volcanology and Geothermal
- 432 Research, 282:92 102.
- 433 Fagents, S. A. and Greeley, R. (2001). Factors influencing lava-substrate heat
- transfer and implications for thermomechanical erosion. Bulletin of Volcanol-
- ogy, (8):519.
- Fernandez de Oviedo y Valdes, G. (1528). Infierno de mamea.
- Fox, M. (2019). A linear inverse method to reconstruct paleo-topography. Ge
- omorphology, 337:151 164.

- Germa, A., Lahitte, P., and Quidelleur, X. (2015). Construction and destruc-
- tion of Mont Pele volcano: Volumes and rates constrained from a geomorpho-
- logical model of evolution. Journal of Geophysical Research. Earth Surface,
- 442 120(7):1206.
- Giordano, D., Nichols, A. R., and Dingwell, D. B. (2005). Glass transition
- temperatures of natural hydrous melts: a relationship with shear viscosity
- and implications for the welding process. JOURNAL OF VOLCANOLOGY
- 446 AND GEOTHERMAL RESEARCH, (1-2):105.
- Global Volcanism Program (2017). Report on Momotombo (Nicaragua). In Ven-
- zke, E., editor, Bulletin of the Global Volcanism Network, volume 42:1, Smith-
- sonian Institution. https://doi.org/10.5479/si.GVP.BGVN201701-344090.
- 450 Gottsmann, J. and Dingwell, D. B. (2002). The thermal history of a spatter-fed
- lava flow: the 8-ka pantellerite flow of mayor island, new zealand. Bulletin of
- 452 Volcanology, 64(6):410–422.
- Greeley, R., Fagents, S. A., Scott Harris, R., Kadel, S. D., Williams, D. A.,
- and Guest, J. E. (1998). Erosion by flowing lava: Field evidence. Journal of
- 455 Geophysical Research, (B11):27.
- 456 Head, J. W. and Wilson, L. (2017). Generation, ascent and eruption of magma
- on the Moon: New insights into source depths, magma supply, intrusions
- and effusive/explosive eruptions (Part 2: Predicted emplacement processes
- and observations). *Icarus*, 283(Lunar Reconnaissance Orbiter Part II):176
- -223.
- 461 Hill, B. E., Connor, C. B., Jarzemba, M. S., La Femina, P. C., Navarro, M., and
- Strauch, W. (1998). 1995 eruptions of cerro negro volcano, nicaragua, and
- risk assessment for future eruptions. Geological Society of America Bulletin,
- 464 (10):1231.

- Hobden, B. J., Houghton, B. F., and Nairn, I. A. (2002). Growth of a young,
- frequently active composite cone: Ngauruhoe volcano, New Zealand. Bulletin
- of Volcanology, (6):392.
- Hurwitz, D. M., Fassett, C. I., Head, J. W., and Wilson, L. (2010). Formation
- of an eroded lava channel within an Elysium Planitia impact crater: Dis-
- tinguishing between a mechanical and thermal origin. Icarus, 210(2):626 –
- 471 634.
- Hurwitz, D. M., Head, J. W., and Hiesinger, H. (2013). Lunar sinuous rilles:
- Distribution, characteristics, and implications for their origin. Planetary and
- Space Science, 79-80:1-38.
- Intercontinental Railway Commission (1898). A Condensed Report of the Trans-
- actions of the Commission and of the Surverys and Explorations of its Engi-
- neers in Central and South America.
- Kauahikaua, J., Cashman, K. V., Mattox, T. N., Heliker, C. C., Hon, K. A.,
- Mangan, M. T., and Thornber, C. R. (1998). Observations on basaltic lava
- streams in tubes from Klauea Volcano, island of Hawai'i. Journal of Geophys-
- ical Research. Solid Earth, 103(B11):27303.
- ⁴⁸² Kerr, R. C. (2001). Thermal erosion by laminar lava flows. Journal of Geophys-
- ical Research, (11):26.
- 484 Kerr, R. C., Griffiths, R. W., and Cashman, K. V. (2006). Formation of chan-
- nelized lava flows on an unconfined slope. Journal of Geophysical Research:
- Solid Earth (19782012), 111(B10).
- 487 Kilburn, C. R. J. and Lopes, R. M. C. (1988). The growth of AA lava flow
- fields on Mount Etna, Sicily. Journal of Geophysical Research: Solid Earth,
- 93(B12):14759-14772.

- ⁴⁹⁰ Kirainov, V., Melekestev, I., Ovsyannikov, A., and Andreev, V. (1988). Re-
- construction of the eruptive activity of momotombo volcano (nicaragua) to
- assess volcanic hazards. pages 495 498.
- Krieger, G., Moreira, A., Fiedler, H., Hajnsek, I., Werner, M., Younis, M.,
- and Zink, M. (2007). TanDEM-X: A Satellite Formation for High-Resolution
- SAR Interferometry. IEEE Transactions on Geoscience and Remote Sensing,
- 496 (11):3317.
- ⁴⁹⁷ Miller, A. A. (1963). Free volume and viscosity of liquids: Effects of temperature.
- The Journal of Physical Chemistry, 67(5):1031-1035.
- Patrick, M. R., Dehn, J., and Dean, K. (2004). Numerical modeling of lava
- flow cooling applied to the 1997 okmok eruption: Approach and analysis (doi
- 10.1029/2003jb002537). JOURNAL OF GEOPHYSICAL RESEARCH -ALL
- SERIES-, (3):B03202.
- Peterson, D. W., Holcomb, R. T., Tilling, R. I., and Christiansen, R. L. (1994).
- 504 Development of lava tubes in the light of observations at Mauna Ulu, Kilauea
- Volcano, Hawaii. Bulletin of Volcanology, (5):343.
- Richardson, P. and Karlstrom, L. (2019). The multi-scale influence of topogra-
- phy on lava flow morphology. Bulletin of Volcanology, (4):1.
- Robl, J., Stwe, K., and Hergarten, S. (2008). Channel profiles around himalayan
- river anticlines: Constraints on their formation from digital elevation model
- analysis. Tectonics, 27(3).
- Sapper, K. (1925). Los volcanes de la America Central. Halle: Verland von Max
- Niemeyer.
- 513 Schenk, P. M. and Williams, D. A. (2004). A potential thermal erosion lava
- channel on Io. Geophysical Research Letters, 31(23).

- Siewert, J. and Ferlito, C. (2008). Mechanical erosion by flowing lava. Contem porary Physics, (1):43.
- Sklar, L. and Dietrich, W. E. (1998). River Longitudinal Profiles and Bedrock
- Incision Models: Stream Power and the Influence of Sediment Supply. Geo-
- physical Monograph American Geophysical Union, page 237.
- Staude, S., Barnes, S. J., and Le Vaillant, M. (2017). Thermomechanical ero-
- sion of ore-hosting embayments beneath komatiite lava channels: Textural
- evidence from kambalda, western australia. Ore Geology Reviews, 90:446 –
- ₅₂₃ 464.
- 524 Strozzi, T., Werner, C., Wiesmann, A., and Wegmuller, U. (2012). Topogra-
- phy Mapping With a Portable Real-Aperture Radar Interferometer. IEEE
- Geoscience and Remote Sensing Letters, (2):277.
- Turcotte, D. L., Turcotte, D. L., and Schubert, G. (2002). Geodynamics. Cam-
- bridge; New York: Cambridge University Press.
- Vincent, F. (1890). In and out of Central America.
- Voytenko, D., Dixon, T. H., Luther, M. E., Lembke, C., Howat, I. M., and de la
- Pea, S. (2015). Observations of inertial currents in a lagoon in southeastern
- Iceland using terrestrial radar interferometry and automated iceberg tracking.
- Computers and Geosciences, 82:23 30.
- Williams, D. A., Kerr, R. C., and Lesher, C. M. (1998). Emplacement and
- erosion by archean komatiite lava flows at kambalda: Revisited. Journal of
- Geophysical Research, (B11):27.
- Williams-Jones, G., Williams-Jones, A. E., and Stix, J. (1998). The nature
- and origin of Venusian canali. Journal of Geophysical Research. Planets,
- 103(E4):8545.

- ⁵⁴⁰ Wilson, L. and Head, J. W. (2017). Generation, ascent and eruption of magma
- on the Moon: New insights into source depths, magma supply, intrusions and
- effusive/explosive eruptions (Part 1: Theory). Icarus, 283(Lunar Reconnais-
- sance Orbiter Part II):146 175.
- Wooster, M. J., Wright, R., Blake, S., and Rothery, D. A. (1997). Cooling
- mechanisms and an approximate thermal budget for the 1991-1993 Mount
- Etna lava flow. Geophysical Research Letters, (24):3277.
- Xie, S., Dixon, T. H., Voytenko, D., Fanghui, D., and Holland, D. M.
- 548 (2018). Grounding line migration through the calving season at Jakobshavn
- Isbr, Greenland, observed with terrestrial radar interferometry. Cryosphere,
- 12(4):1387 1400.

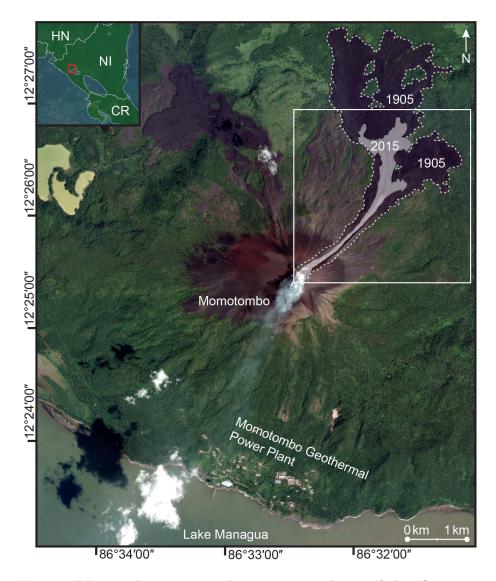


Figure 1: Momotombo area map. The most recent volcanism (a lava flow emitted during the 2015-2016 eruption) is noted by a light-grey overlay. The 1905 eruption is noted by a dotted line. The area shown in Fig. 4 is noted by the white box. Note the widely dispersed flows that underlie the 2015 and 1905 flows. Their distribution suggests that no incised channel existed at the time of their emplacement to direct flow paths. Background image from GoogleEarth.

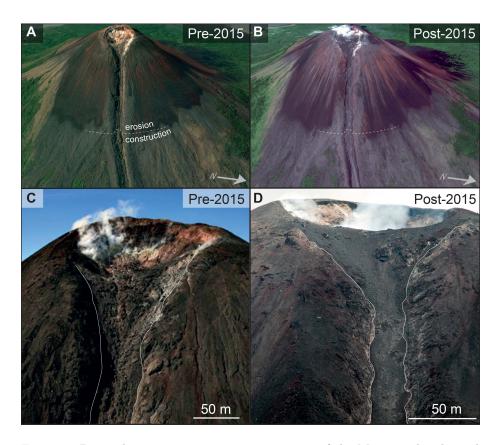


Figure 2: Pre and post 2015-2016 eruption images of the Momotombo channel and summit. A) A pre-eruption image of the channel, where the dashed line shows the approximate transition between erosional and constructional behavior. B) Major changes in morphology can be seen in and around the summit crater, where the 2015 lava flow first filled and was then partially excavated during subsequent explosions in 2016. The pre-existing floor of the channel has been paved over by a lava flow and appears less 'rough' than the pre-eruption channel floor. Images A and B from Google Earth. C) The summit crater prior to the 2015-2016 eruption, with white lines bounding the channel. Textures within the channel indicate downslope flow. Image from INETER. D) The summit area on 6 April, 2016. Several hundred small explosions have partially excavated a small dome from December, 2015. Blocks have been deposited atop the recent lava flow and a fine, grey layer of ash from repeated pyroclastic density currents coats the channel. Image from E. Gallant.

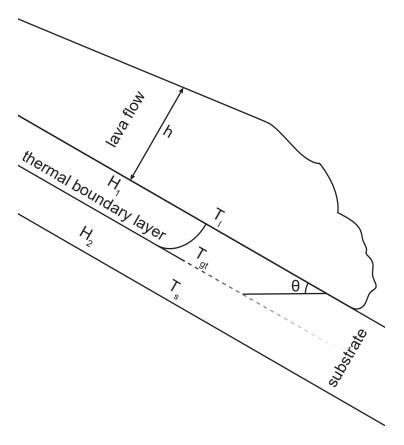


Figure 3: Thermal boundary formation illustration. The energy required to melt a pyroclast-rich substrate is less than that of a lava flow due to a lower density; we can therefore substitute the peak glass transition temperature (1013 K) as the minimum temperature required to initiate melting for such substrates (Giordano et al., 2005). The formation of the thermal boundary layer shows the transition between the temperature of the lava (T_l) and the substrate (T_s) . This layer defines the boundary between a thermally softened substrate (H_1) and the unaffected substrate (H_2) . The height of the lava flow is noted by h and the slope of the edifice by θ .

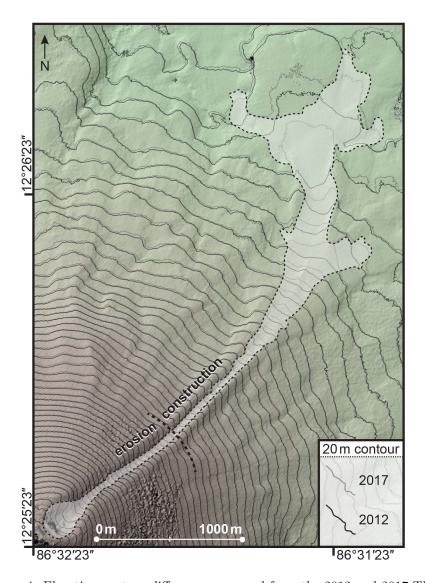


Figure 4: Elevation contour difference measured from the 2012 and 2017 TDX DEMs. Elevation contours at 20 m intervals were mapped from the TDX datasets; the grey contours represent the 2017 elevations (post 2015-2016 eruption), and the black represent the 2012 elevations. Slight contour variations exist in areas not impacted by the 2015-2016 eruption due to the different look angles of the TDX data pairs and georeferencing uncertainties. The dashed line and white infill show the area covered by the 2015 lava flow.

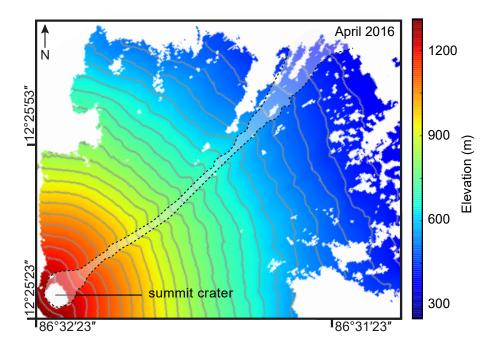


Figure 5: 2016 Terrestrial Radar DEM. The white infill and black dashed line shows the area covered by the 2015 lava flow. The absence of signal near the summit shows the $120\times100\,\mathrm{m}$ crater, which was excavated by several hundred small explosions between December 2015 and April 2016. The grey lines indicate 20 m contour intervals.

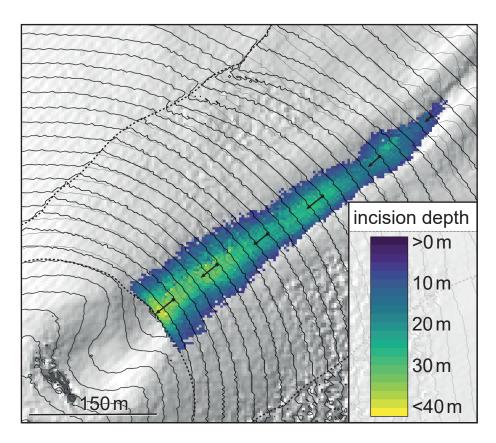


Figure 6: Elliptical contour fit incision depth process (from 2012 DEM). The original elevation contours at 20 m intervals are noted in grey, the modeled fit in black. Incision depths, with arrows that indicate the horizontal distance between the modeled paleotopography and the current point of corresponding incision. Incision is deepest at the summit and decreases downslope.

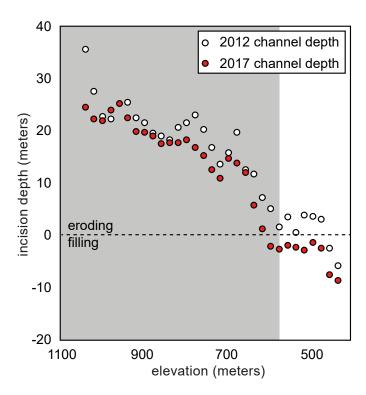


Figure 7: Calculated channel depths at 20-m elevation intervals before and after the eruption, using the elliptical fit method described in Figure 6. Calculated channel depths and slopes at 20-m elevation intervals before and after the eruption. White and red dots are channel depths of the pre- and post-eruption channel, respectively. The difference of the two DEMs gives us the thickness of the 2015 lava flow. The grey shading indicates the elevation range where erosion dominates $(>650\,\mathrm{m})$.

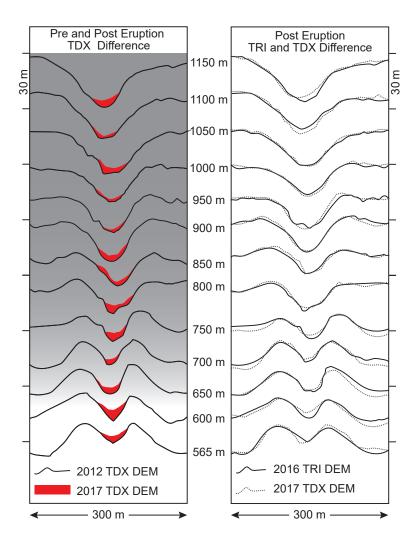


Figure 8: Momotombo lava channel profiles. The left figure shows channel profiles from the 2012 TDX DEM and an overlay of the lava flow from 2015. The 2012 profile is noted by the solid black line, with the 2015 lava flow (imaged by the 2017 TDX acquisition) by the red polygon. The grey shading illustrates the transition of channel into a constructional feature. The profiles have been visually adjusted to match up topography in order to account for the 21° difference in acquisition angles. The right figure shows the difference between the 2016 TRI DEM and the 2017 TDX DEM. Comparison of the post-eruption profiles estimates the relative noise of the DEMs.

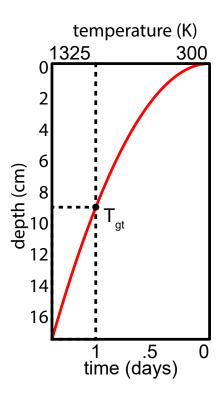


Figure 9: Substrate thermal boundary layer. At one day, we calculate the thermal boundary layer for a flow on Momotombo to grow to $8.8\,\mathrm{cm}$. Thermal erosion rates in Hawaiian tubes and channels can reach $\sim\!10\,\mathrm{cm}$ per day (Kauahikaua et al., 1998); given the lower temperature of lavas erupted on Momotombo, it is reasonable to assume a slower rate of thermal boundary growth.

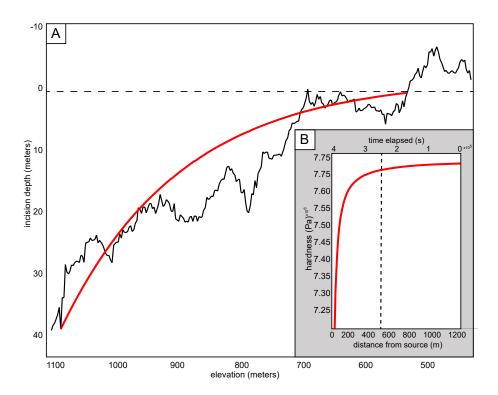


Figure 10: Incision depth fit models. A: Depth is used as a proxy for time in order to constrain the hardness (H) of the eroded substrate. The red line indicates the exponential fit to the eroded depth as a function of elevation. The dotted line indicates the transition from erosion (below) to construction (above). B: Substrate hardness as a function of time and distance. The dotted line notes the transition from erosion (left) to construction (right). Time elapsed is the total time the substrate is in contact with flowing lava.

551 Supplemental Material

Paleotopography Modeling Code

- The MatLab code and associated functions used to model the paleotopogra-
- phy of the erosive channel and determine excavation depths is available at
- https://github.com/elisabeth-gallant/USF_dissertation.

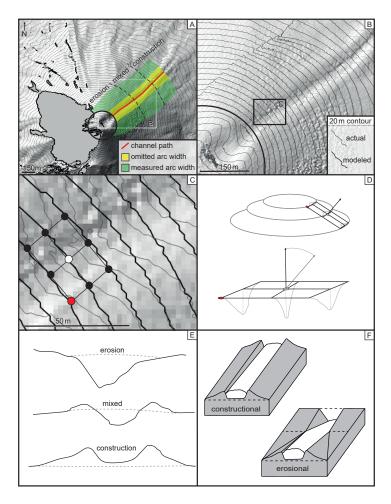


Figure 11: Elliptical contour fit incision depth process (from 2012 DEM). A: The center of the channel is noted in red, area used to calculate the best fit contours in green, and the area omitted from the contour calculation in yellow. The dashed line shows the approximate boundary between erosional, mixed, and constructional regimes (described in profile in panel F). B: The original elevation contours at 20 m intervals are noted in grey, the modeled fit in black. C: The orientation of vectors normal to the edifice are calculated using a 9 \times 9 matrix. The distance between sampling points is enlarged to illustrate the concept. D: The matrix is transposed so that the only magnitude of the vector is in the z-direction. The elevation value is returned as a weighted-average of the constituent points (with distance from the center as the weighting criteria). E: The simplified profiles for erosional, mixed, and constructional regimes, with a dashed line denoting the paleosurface. F: Constructional (left) vs erosional (right) behaviour in lava channels on an incline. We note the change in slope associated with the original orientation of the incline and the eroded portion. The dashed line notes the paleotopography surface. Additionally, levees are absent from the incised section of the channel. Profiles of these different regimes are presented in Figure 8.

556 Circular Contour Fit

Channel incision depths were measured by interpolating best-fit 20-m interval 557 contour lines to the overall shape of the edifice using the 2012 TDX DEM. Each interpolated contour line was calculated by fitting the original contour line with 559 an arc using least-squares. The section proximal to the channel was not included in the fitting process due to its wide deviation from the overall shape of the cone. 561 Depth at each sampled elevation was measured by determining the minimum 562 distance between the interpolated contour line and the deepest point of the 563 channel, measured normal to the slope of the edifice (Fig. 12b). All points on 564 the contour line (not including the flow levees and the channel) were averaged to calculate the slope for each measured depth. 566 Results show the pre-2015 eruption channel extended down the northeast side of the edifice from the summit and continued for 2 km onto the surrounding

Results show the pre-2015 eruption channel extended down the northeast side of the edifice from the summit and continued for 2 km onto the surrounding plain (Fig. 13). A consistent width of \sim 30 m was maintained throughout the channels length, while incision depth varied from 35 m at the summits rim and tapered off to 0 at \sim 700 m elevation (Fig. 12a). The 1905 channel has been infilled with a lava flow from 2015 with a thickness <3 m (Fig. 13).

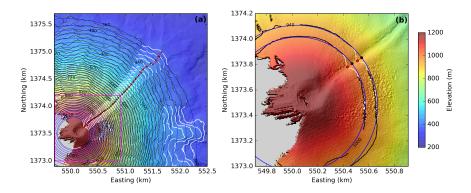


Figure 12: Channel depth calculation using the pre-eruption DEM as an example (a) 2012 DEM. Black and white lines are contours with 20 m interval. The blue circles are the best-fit arcs for each contour line. Contour lines in white were not used in the fitting processing because they include the topographic influence of the levees and channel, which greatly deviates from the circular fit. The red lines indicate the depth of the channel thalweg. (b) A zoomed in view of the fitting process. Red dots indicate the deepest points in the channel.

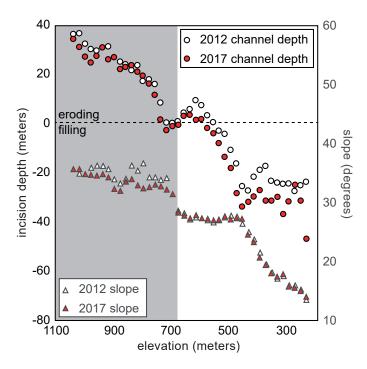


Figure 13: Calculated channel depths, slopes, and differences using the circular fit method. Calculated channel depths and slopes at 20-m elevation intervals before and after the eruption. White and red dots are channel depths of the preand post-eruption channel, respectively. The difference of the two DEMs gives us the thickness of the 2015 lava flow. White and red triangles are averaged slops at 20-m elevation intervals of the pre- and post-eruption edifice.

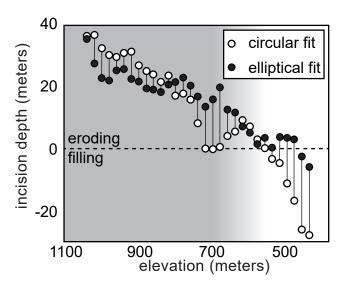


Figure 14: Comparison of the circular and elliptical method incision depths for the 2012 TDX DEM. The grey area indicates the transition between incision and construction, with a gradient to indicate the uncertainty between the two models as to where this occurs.