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1 Multi-satellite data depicts record-breaking methane leak from a well 2 blowout

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15 **Accidental blowouts in oil and gas wells can result in large and prolonged methane**
16 **emissions, which are often unreported when happening in remote places. The**
17 **rapid advancement of space-based methods for detecting and quantifying methane**
18 **plumes provides an essential tool for uncovering these super-emission events. We**
19 **use a range of methane-sensitive satellites to document a methane leak from a**
20 **well blowout in Kazakhstan’s Karaturun East oil field in 2023. A dense time series**
21 **of observations from multiple satellites shows that the leak was active during 205**
22 **days. Using 48 high-quality plume observations from this time series, we estimate**
23 **that a total of 128 ± 36 kt of methane were released to the atmosphere during this**
24 **leak. Our results reveal that the total methane emissions from the Karaturun 2023**
25 **event exceeded those from all previously documented accidents, and highlight the**
26 **pivotal role of satellites in detecting and quantifying large methane plumes around**
27 **the planet.**

28 Introduction

29 Human-induced methane emissions are responsible for about 30% of the global warm-
30 ing since the pre-industrial period¹. The oil and gas industry accounts for a large share
31 of those emissions², although the mitigation of oil and gas emissions has actually been
32 found to be technically feasible and cost-effective³. This holds particularly true for high-
33 emitting point sources, also known as super-emitters⁴. Methane super-emitters in the
34 oil and gas industry are usually linked to unexpected infrastructure failures, including
35 blowouts during drilling, completion, or production activities in oil and gas wells. Well

36 blowouts result in uncontrolled releases of substantial amounts of natural gas, consisting
37 primarily of methane. However, it is difficult to achieve a global perspective of the methane
38 emissions originated by blowouts in oil and gas wells, as these accidents often occur in
39 remote areas, which complicates the acquisition of surface and airborne measurements.

40 A growing constellation of methane-sensitive satellites is now improving our ability
41 to detect and monitor large methane leaks around the planet. This constellation com-
42 prises the Sentinel-5P/TROPOMI mission and a number of high spatial resolution mis-
43 sions, which can be separated in hyperspectral and multispectral missions. TROPOMI
44 provides a systematic daily global surveillance of the largest methane emissions since
45 2018^{5;6}. In contrast, the hyperspectral high-resolution missions have a sparse spatio-
46 temporal sampling as compared with TROPOMI, but scan the Earth at a much higher
47 spatial resolution, which enables the detection of smaller plumes and the attribution of
48 those to facility-level sources. Among these hyperspectral high-resolution missions we
49 find the GHGSat private constellation^{7;8}, specifically designed for methane and carbon
50 dioxide mapping at 25–50 m resolution, and the EnMAP, PRISMA and EMIT scientific mis-
51 sions, which also have a relatively high sensitivity to methane^{9–11}. Finally, the Sentinel-2
52 and Landsat multispectral radiometers offer a lower sensitivity to methane than the hyper-
53 spectral missions, but have a 20-30 spatial sampling and a systematic global coverage
54 every few days^{12–14}, which compensates for the sporadic and under-demand sampling of
55 the hyperspectral satellites.

56 There are examples of the potential of the previous satellites to document methane
57 emissions from well blowouts in the last years, although only a few satellite observations
58 were available for each of those studies: Thompson et al. conducted the first detection
59 of an individual methane plume with observations of the 2015 Aliso Canyon event (Los
60 Angeles, USA) by the Hyperion high-resolution spectroscopy demonstration mission¹⁵;
61 Pandey et al. used one TROPOMI overpass to estimate the emissions caused by a shale
62 gas well blowout in Ohio (USA)¹⁶; Maasackers et al. used six TROPOMI overpasses
63 and gas flaring data from the VIIRS satellite instrument to estimate emissions from a
64 natural gas well blowout in Louisiana (USA)¹⁷; Cusworth et al. combined observations
65 from TROPOMI, GHGSat and PRISMA (four observations in total) to characterise the
66 emissions from a 20-day leak event due to a gas well blowout in the Eagle Ford Shale
67 (USA)¹⁸.

68 All of those blowouts happened at US sites accessible to well operators and miti-
69 gation teams, so the associated emissions could therefore be well documented despite
70 the limited availability of satellite observations. This is not the case of the well blowout at
71 the Karaturun East oil field in Kazakhstan's Mangistau region that we report in this study
72 (Fig. 1), which happened at a remote site. According to media reports, the blowout and
73 subsequent fire happened on the morning of 9 June 2023 during exploration works at well
74 303¹⁹ (Fig. 1a). The fire destroyed different pieces of safety equipment, leading to a loss

75 of well control and a 10-m high fire blaze. Days later, a 15-m wide crater was formed
76 by the collapse of rocks around the wellhead, which prevented an early seal of the well.
77 The first attempt to halt the flow of gas consisted of pumping thousands of tons of water
78 through two injection holes between 13 October and 20 November. This action mitigated
79 the gas leak, but did not completely resolve it. The flow of gas and the fire could finally
80 be stopped on 25 December 2023 by injecting heavy drilling mud via a special-purpose
81 probe, which connected with the wellbore of the accident well at a depth of about 1000
82 m²⁰.

83 [Figure 1 about here.]

84 In this work, we use satellites to document the methane emissions originated by
85 the blowout in the Karaturun East oil field. We have generated a dense time series with
86 more than hundred satellite observations from TROPOMI and high-resolution satellites to
87 monitor and quantify the massive methane plumes following this blowout. The satellite
88 data were processed with state-of-the-art algorithms for the detection and quantification
89 of methane plumes from space, optimised for the particular emission intensity and site
90 characteristics of this event (Methods).

91 **Results and discussion**

92 The leak was first detected in global methane concentration data from the Sentinel-
93 5P/TROPOMI mission¹⁹ (Fig. 1b, Fig. S1). The exact location and date of the blowout
94 could be confirmed retrospectively with data from the Sentinel-2 multispectral radiometer,
95 which flew over the site hours after the accident. The evolution of the leak was then moni-
96 tored with TROPOMI and a range of high-resolution missions (including PRISMA, EnMAP,
97 EMIT, GHGSat and the Sentinel-2 and Landsat-8/9 multispectral radiometers), some of
98 which were specifically tasked to acquire data over the site. Extremely large methane
99 plumes were detected during the entire time series (Fig. 2).

100 [Figure 2 about here.]

101 In particular, we detected methane plumes from the site 115 times between 9 June
102 and 25 December 2023. After quality screening of all the detected plumes, we retained
103 48 for the quantification of emission rates (Methods). We obtained flux rates between
104 3.6 ± 1.3 and 63 ± 42 t/h, with typical values between 20 and 50 t/h (Fig. 3a, Supplemen-
105 tary Fig. S2, and Supplementary Tables S1 and S2). The most substantial emission rates
106 occurred in the weeks following the blowout (Fig. 3a). Plume intensity gradually decreased

107 over time until the leak repair on 25 December. Notably, a relatively large plume of
108 12 ± 3 t/h was detected by GHGSat on 25 December, which suggests that the satellite
109 flew over the site shortly before the final repair action on the same day. Three subsequent
110 observations on 1, 12 and 14 January 2024 confirmed the definitive cessation of the leak.

111 [Figure 3 about here.]

112 We utilized a time series of fire radiative power derived from the VIIRS FIRMS satel-
113 lite product to track the fire intensity during the event (Methods). Additionally, we used
114 observations from high resolution satellites to detect (albeit not quantify) fire at the site.
115 We found that the strongest fire occurred immediately after the well blowout, and kept a
116 relatively high intensity during about 20 days (Fig. 3b). Subsequently, the fire remained
117 active as indicated by the fire detections from the high resolution satellites, but its intensity
118 decreased. The fire intensity in this event is actually substantially lower than that of strong
119 flares in oil and gas installations, and also than the intensity of the fires measured during
120 the Louisiana 2019 blowout event¹⁷ (Supplementary Fig. S3). The relatively low intensity
121 of the fire at the Karaturun East site would indicate that only a small fraction of the gas
122 outflow was flared. It is unclear whether the accident well 303 was an oil or a gas well,
123 but in the first case it could be hypothesized that the 2-phase oil and gas mixture led to a
124 low combustion efficiency during the event²¹.

125 To give context to the magnitude of the methane plumes detected during the entire
126 Karaturun East leak, the majority of plumes identified in this event exceeded 10 t/h and
127 are comparable to the largest individual plumes detected using global TROPOMI data
128 worldwide^{5,6}. We obtain an estimate of 128 ± 36 kt for the total amount of methane re-
129 leased to the atmosphere during the event (Methods). This is substantially larger than
130 the total emission of 97 kt (no uncertainty available) reported for the outstanding 2015 Al-
131 iso Canyon blowout²², which is considered the largest methane leak from regular oil and
132 gas operations documented to date. The total emission from the Karaturun event also
133 exceeds the estimates for the massive releases from the Ohio 2018¹⁶ and the Louisiana
134 2019¹⁷ blowout events, for which 60 ± 15 kt and 21–63 kt (95% confidence interval) were
135 estimated, respectively (Fig. 3c). Only the sabotage of the Nord Stream 1 and 2 subsea
136 twin pipelines in the Baltic Sea on 26 September 2022, for which a total of 420-490 kt
137 (95% confidence interval) has been estimated²³, may have led to greater emissions than
138 the Karaturun 2023 blowout event.

139 Our results show that the 2023 well blowout in Kazakhstan's Karaturun East oil field
140 has likely caused the largest methane emission from an infrastructure accident ever doc-
141 umented. The detection and quantification of this leak has only been possible because
142 of the recent availability of methane-sensitive satellites. It is unknown how many of such

143 large methane leaks from oil and gas infrastructure failures may have occurred in the last
144 decades around the world, and how this may have led to underestimated emission inven-
145 tories. The new era of methane monitoring from space, boosted by international initiatives
146 such as the Methane Alert and Response System (MARS)²⁴ of the United Nations En-
147 vironment Programme, and new methane satellite missions such as MethaneSAT²⁵ and
148 Carbon Mapper²⁶, will be crucial for the detection and quantification of large methane
149 leaks around the world.

150 **Methods**

151 *Satellite data*

152 We generated a dense time series of methane observations over the Karaturun
153 East site using TROPOMI and high-resolution satellite missions. The latter included
154 hyperspectral observations from the GHGSat private satellite constellation, which offers
155 the highest sensitivity to methane for this type of high-emitting point source, as well as
156 from the public hyperspectral missions EnMAP (German Aerospace Agency, Germany),
157 PRISMA (Italian Space Agency, Italy) and EMIT (NASA Jet Propulsion Laboratory, USA),
158 which share a relatively similar configuration, a medium-high sensitivity to methane, and
159 an open data policy. Acquisitions from the less sensitive Sentinel-2 and Landsat multi-
160 spectral radiometers were mostly used for plume detection along the entire time period.
161 A total of 115 plumes were detected between 9 June and 25 December 2023 (26 from
162 TROPOMI and 89 from the high resolution satellites). After quality control, 48 of those
163 plumes (15 from TROPOMI and 33 from the high resolution satellites) were retained for
164 the quantification of emissions (Supplementary Figure S2).

165 Satellite data were also used to monitor fire activity at the site. Fire radiative power
166 data from the Fire Information for Resource Management System (FIRMS) based on VI-
167 IRS data were used to assess the evolution of fire intensity (Fig. 3b and Supplementary
168 Fig. S3). In addition, the observations from the high resolution satellites from which we
169 derived methane plumes were also utilised to detect (but not quantify) smaller active fires
170 at the site.

171 *Quantification of methane plumes with high spatial resolution satellites*

172 For the retrieval of methane plume information from EnMAP, PRISMA and EMIT
173 hyperspectral data, we optimised the widely-used matched-filter approach to deal with
174 the large plumes and high methane concentration values found during the Karaturun East
175 leak. This included the implementation of a log-normal version of the matched filter and
176 the removal of plume pixels when calculating the statistics needed for the matched-filter,
177 as described in Pei et al.²⁷. The ΔXCH_4 maps obtained with this retrieval were screened

178 for quality (cloud-free observations, no retrieval artifacts, no substantial fractions of the
179 plume lying outside the image area).

180 The selected high-quality observations were used for the subsequent flux rate es-
181 timation. We manually delineated the plumes and calculated the integrated methane
182 enhancement (IME), which is the total mass of methane contained in the plume. The
183 IME values were converted into flux rate estimates using the IME-based model²⁸, which
184 relates the IME, the plume length, and an effective wind speed parameter (U_{eff}) to the flux
185 rate. For U_{eff} , we used empirical linear models linking U_{eff} with 10-m wind speed data
186 (U_{10}). These models were specifically derived for the typical ΔXCH_4 retrieval precision
187 and plume length estimated for this event (simulated plumes are 5-10 km long). One
188 common $U_{\text{eff}} - U_{10}$ model was used for EnMAP and PRISMA, which share the same 30-m
189 resolution and retrieval precision, and a second one was derived for EMIT in order to ac-
190 count for its 60 m pixels (Supplementary Fig. S4). Uncertainties in flux rate estimates were
191 derived assuming a 50% uncertainty in the input wind speed data, which is consistent with
192 previous studies.

193 In the case of the multispectral missions (Sentinel-2 and Landsat), we have followed
194 the approach described in Gorroño et al.¹³. Same as for the hyperspectral data, this ap-
195 proach consists of a two-step processing scheme, in which ΔXCH_4 maps are first derived
196 with a multi-band and multi-pass retrieval, and the flux rates are subsequently estimated
197 using the IME-based method. Since the methane sensitivity of the multispectral missions
198 is lower than that of GHGSat and the hyperspectral missions, data from Sentinel-2 and
199 Landsat were mostly used to detect methane plumes during the event. Thanks to their
200 systematic acquisitions and combined revisit time of 2-3 days, more than 70 methane
201 plumes could be detected with Sentinel-2 and Landsat. In addition, 5 of them acquired
202 under optimal observation conditions were included in the list of 48 plumes used for quan-
203 tification.

204 Finally, in the case of GHGSat, the physically-based methane concentration retrieval
205 described in Jervis et al.⁸ was applied. The performance of GHGSat for the detection and
206 quantification of methane emissions from oil and gas infrastructure has been extensively
207 tested during operations in the last years⁷.

208 *Verification of the detection and quantification of methane plumes with high spatial reso-* 209 *lution satellites*

210 Our processing chain to convert spectral radiance data cubes to flux rates, which
211 involves ΔXCH_4 retrieval, plume segmentation, and IME-based flux rate estimation, has
212 been validated with controlled methane release tests for all the previous high-resolution
213 instruments²⁹.

214 However, since those controlled-release tests were made for plumes weaker than
215 the ones detected in this event, we also used end-to-end simulations to test our ability
216 to quantify flux rates for the large plumes and particular conditions of the Karaturun East
217 site. Real top-of-atmosphere radiance data acquired during the event were used as input
218 for the simulations, so the particular acquisition conditions at the site (atmospheric state,
219 illumination angles, potential water vapour and smoke co-emissions ...) were properly rep-
220 resented. This end-to-end simulation approach has already been applied to hyperspectral
221 and multispectral data^{9;13}. In this study, we did the simulations for PRISMA observations
222 of methane plumes over the Karaturun East site within a 5–50 t/h emission range. The re-
223 sults show that our processing is able to produce reliable flux estimates for that entire flux
224 rate range (Supplementary Fig. S5). Even if this end-to-end verification exercise was only
225 done for PRISMA, the overall conclusions can be extended to the other high resolution
226 missions, as they share a relatively similar processing chain.

227 Those simulations confirm the robustness of our methane retrieval and quantification
228 methods for the particular conditions of the Karaturun East event. In addition, it must be
229 remarked that we do not find any distortion of our ΔXCH_4 maps with the water vapour and
230 smoke being potentially co-emitted by the source (Supplementary Fig. S6). We verified
231 this by generating water vapour anomaly maps using a similar retrieval method as the
232 one we use for methane. The resulting water vapour maps show the expected turbulence
233 patterns, but no water vapour plume superposed to the methane plumes. Also, we do not
234 detect any smoke signal in the 2300 nm spectral window used for the methane retrievals,
235 indicating that the smoke plumes may not have a relevant optical activity in this spectral
236 range for this event, which has also been found in previous studies¹⁰.

237 *Quantification of methane plumes with TROPOMI*

238 TROPOMI (aboard Sentinel-5P)³⁰ observes methane with high precision at a res-
239 olution of $5.5 \times 7 \text{ km}^2$, allowing detection of the plume further downwind. We used the
240 Weather Research and Forecast (WRF) version 4.1³¹ to simulate the enhanced methane
241 concentrations associated with the blowout at a resolution of $3 \times 3 \text{ km}^2$ from June to De-
242 cember 2023. We then compared these modelled concentrations to TROPOMI data in a
243 Bayesian inversion framework³² to infer daily emissions rates. To obtain simulated plumes
244 that best match TROPOMI, we ran WRF using two meteorological boundary conditions
245 products and four planetary boundary layer physics schemes, and sampled the model
246 at several timesteps around the TROPOMI overpass. Based on daily inversions with all
247 model setups, we selected the simulations that gave the lowest posterior observation cost
248 for each day to be used. We only report quantifications for days with clear plumes and
249 good matches with simulated plumes based on visual inspection. To estimate uncertainty,
250 we built an ensemble of inversions by varying critical inversion parameters such as data
251 filtering and the selected simulation. We conservatively report the 2-standard deviation
252 range from the ensemble as uncertainty. Details on the TROPOMI quantification approach

253 are given in the Supplementary Text S1.

254 *Quantification of total methane emission*

255 We estimate a total of 128 ± 36 kt of methane being released to the atmosphere
256 between 9 June and 25 December. This amount was calculated through the integration
257 of a polynomial fitted to the time series of 48 flux rate estimates from the plume data
258 passing the quality screening (Fig. 3a). The estimation of the uncertainty associated to
259 the total emission is the result of propagating the uncertainty from each flux rate estimate
260 through the flux rate interpolation and curve integration using multivariate Monte Carlo
261 simulations. We assume a 50% correlation between flux rate errors in order to account
262 for both uncorrelated error components (e.g. plume shape changes) and correlated error
263 components (e.g. same source area) (Supplementary Fig. S7).

264 We tested an alternative option for the quantification of the total emission. This
265 consisted in using only the most accurate flux rate estimates from GHGSat and the hy-
266 perspectral missions, which are a priori the best suited satellites for the quantification
267 of single plumes over a relatively complex surface. This alternative approach prioritises
268 the quality of observations over quantity. For this alternative approach, we found similar
269 total emission numbers than for the default all-satellite configuration, but the uncertainty
270 of the total emission was substantially lower for the latter case with all 48 plumes (36
271 versus 58 t/h, Supplementary Fig. S7), so we opted for using all the available high-quality
272 observations for the calculation of the total emission and its uncertainty.

273 **List of Supplementary Materials**

- 274 • Text S1
- 275 • Tables S1 and S2
- 276 • Figures S1–S7

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369 be addressed to L.G.

370 **Author contributions**

371 Conceptualization and methodology: L.G., J.D.M., and I.I.-L. Formal analysis: L.G., J.R.,
372 S.S., A.V., I.I.-L., J.G., X. Z., B.J.S., J.D.M., A.G. Investigation: All authors. Supervision:
373 L.G., J.D.M., I.A., Q.P., A.B., D.Z. Writing-original draft: L.G. Writing-review and editing:
374 All authors

375 **Competing interests**

376 The authors declare no competing interests.

377 **List of Figures**

378 1 Blowout and first methane plumes from the 2023 Karaturun East leak. **a**,
379 Location of the well No. 303 (45.3324°N, 52.3730°E) in the Karaturun East
380 oil field where the blowout happened on 9 June 2023, including a view of
381 the active fire at the site (photo from Mangistau Regional Administration).
382 **b**, Synoptic map of total methane concentration (in parts-per-billion, ppb)
383 from a Sentinel-5P/TROPOMI overpass 17 days after the blowout. A large
384 methane plume emitted from the accident well is observable. 13

385 2 Sample of methane plumes detected with the PRISMA, EMIT, EnMAP and
386 GHGSat satellite sensors on different days. The color scale in the maps
387 represent methane concentration enhancements above background methane
388 levels (ΔXCH_4), expressed in parts-per-billion (ppb). The emission rate (Q)
389 estimated for each plume is provided on the top right side of each map panel. 14

390 3 Quantification of methane emissions and fire intensity from the Karaturun
391 East 2023 blowout. **a**, Time series of methane emission rates (in metric
392 tonnes per hour, t/h) derived from the satellite observations which passed
393 the quality screening (see also Supplementary Fig. S2). The black line and
394 the shaded area represent the polynomial fit that has been integrated for
395 the calculation of the total amount of methane released during the event.
396 The black bars depict all satellite observations from which a plume could
397 be detected, including those that could not be quantified. **b**, Time series of
398 fire radiative power derived from the VIIRS FIRMS data product. The black
399 bars depict all satellite detections of active fire at the site. **c**, Comparison
400 of the total amount of methane released during the Karaturun East 2023
401 event with the Aliso Canyon 2015²², Ohio 2018¹⁶, and Louisiana 2019¹⁷
402 blowout events also leading to massive methane emissions. 15

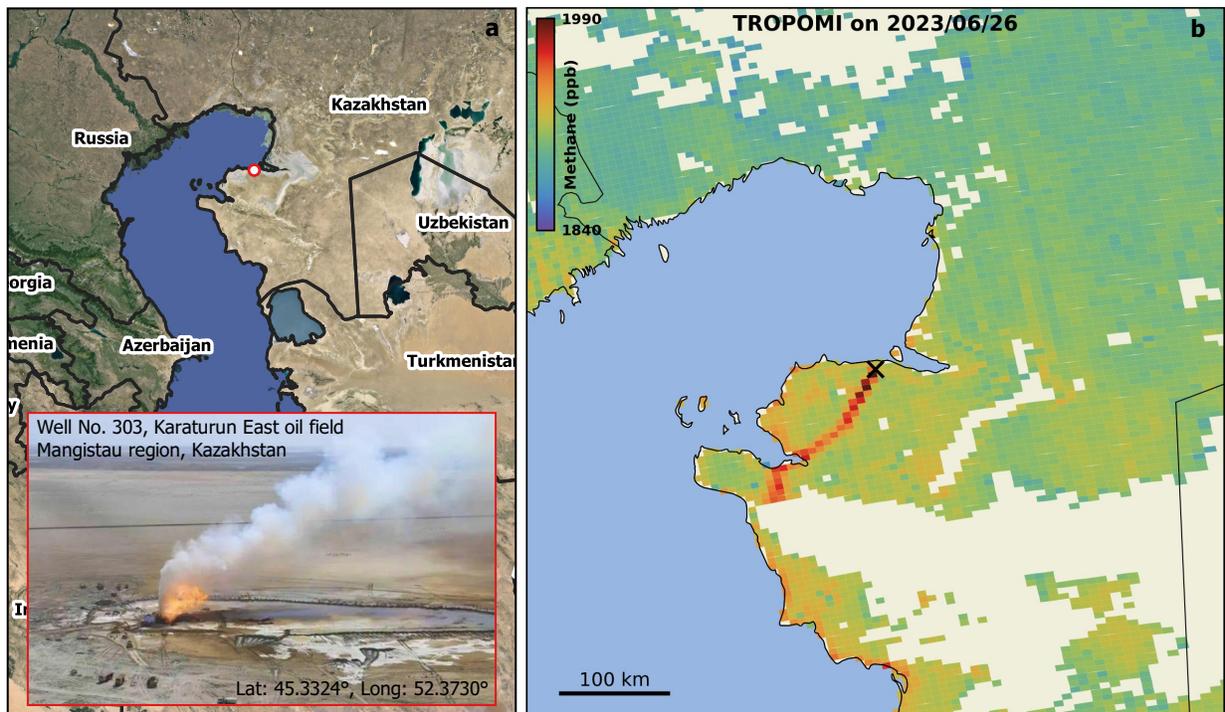


Figure 1: Blowout and first methane plumes from the 2023 Karaturun East leak. **a**, Location of the well No. 303 (45.3324°N , 52.3730°E) in the Karaturun East oil field where the blowout happened on 9 June 2023, including a view of the active fire at the site (photo from Mangistau Regional Administration). **b**, Synoptic map of total methane concentration (in parts-per-billion, ppb) from a Sentinel-5P/TROPOMI overpass 17 days after the blowout. A large methane plume emitted from the accident well is observable.

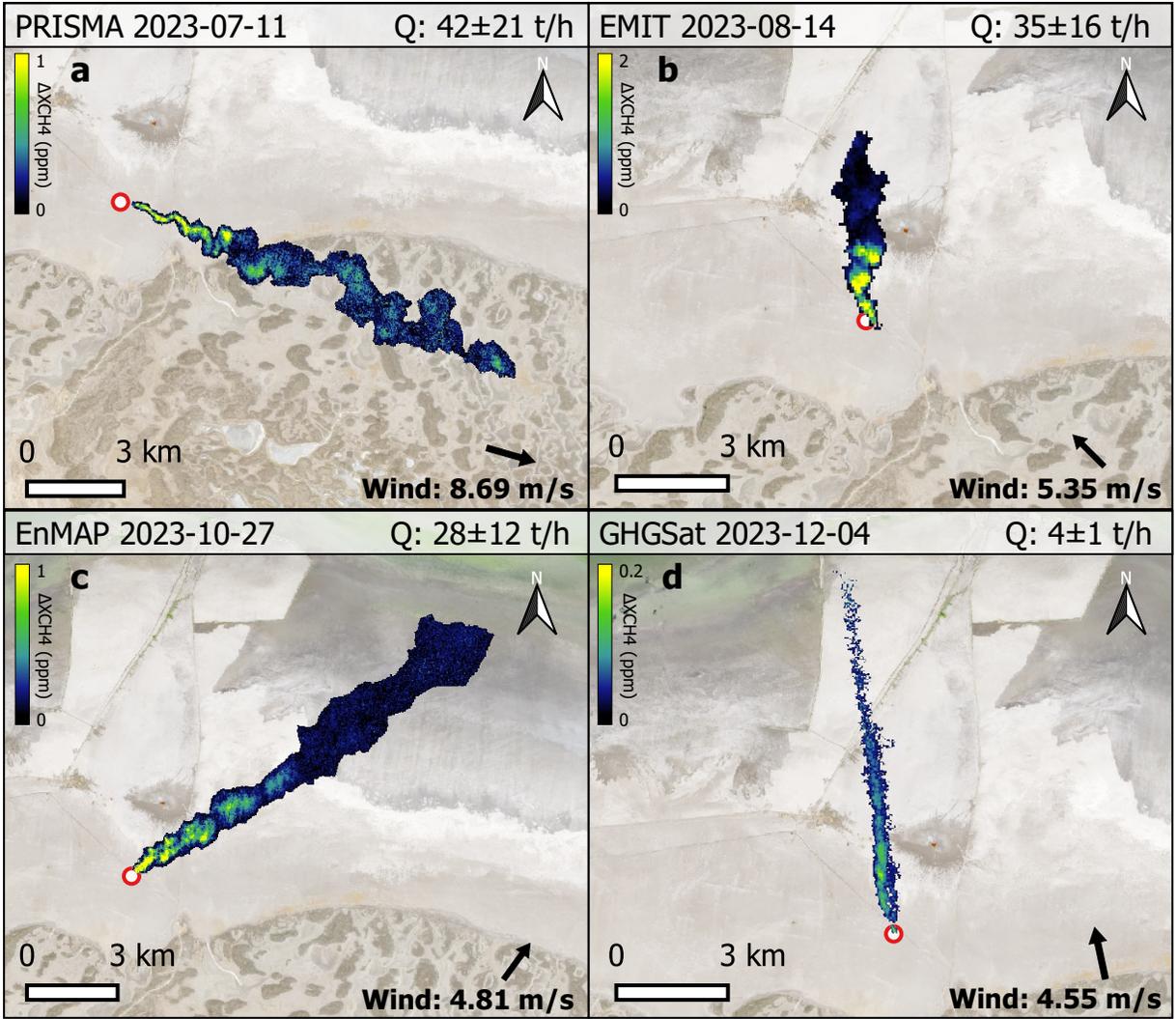


Figure 2: Sample of methane plumes detected with the PRISMA, EMIT, EnMAP and GHGSat satellite sensors on different days. The color scale in the maps represent methane concentration enhancements above background methane levels (ΔXCH_4), expressed in parts-per-billion (ppb). The emission rate (Q) estimated for each plume is provided on the top right side of each map panel.

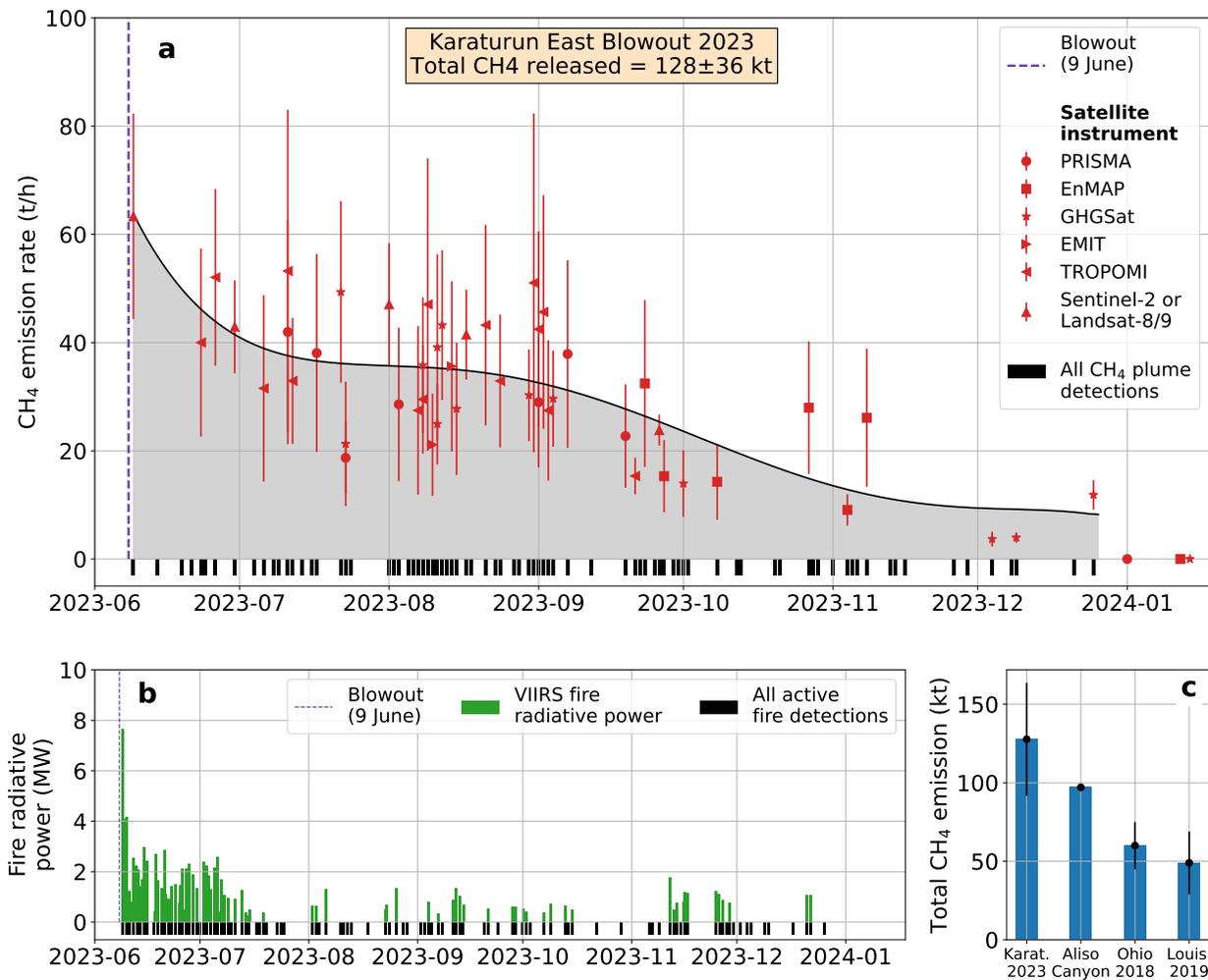


Figure 3: Quantification of methane emissions and fire intensity from the Karaturun East 2023 blowout. **a**, Time series of methane emission rates (in metric tonnes per hour, t/h) derived from the satellite observations which passed the quality screening (see also Supplementary Fig. S2). The black line and the shaded area represent the polynomial fit that has been integrated for the calculation of the total amount of methane released during the event. The black bars depict all satellite observations from which a plume could be detected, including those that could not be quantified. **b**, Time series of fire radiative power derived from the VIIRS FIRMS data product. The black bars depict all satellite detections of active fire at the site. **c**, Comparison of the total amount of methane released during the Karaturun East 2023 event with the Aliso Canyon 2015²², Ohio 2018¹⁶, and Louisiana 2019¹⁷ blowout events also leading to massive methane emissions.

Supplementary Materials

Multi-satellite data depicts record-breaking methane leak from a well blowout

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Text S1. Quantification of methane plumes with TROPOMI

With daily global coverage, TROPOMI, onboard Sentinel-5P, enables global mapping of methane concentrations at $5.5 \times 7 \text{ km}^2$ resolution using the shortwave infrared (SWIR) spectrum at $2.3 \mu\text{m}$. For this analysis, we use version 02.05.00 of the TROPOMI-CH₄ operational product corrected for stripes¹ with a custom quality filter (qa value ≥ 0.4 , SWIR aerosol optical thickness < 0.1 , SWIR surface albedo > 0.05 , surface classification $\neq 3$, and SWIR cloud fraction < 0.015).

Methane concentrations are simulated using the Weather Research and Forecast (WRF) version 4.1² around the blowout location at a $3 \times 3 \text{ km}^2$ resolution for an area of $800 \times 800 \text{ km}^2$ from June 2023 to December 2023. We perform simulations with both 6-hourly National Centre for Environmental Prediction (NCEP)³ and hourly ERA5⁴ meteorological fields. The 6-hourly Copernicus Atmosphere Monitoring Service (CAMS) atmospheric composition forecast data⁵ is used to provide the initial and boundary conditions. To infer emissions, the Bayesian cost function J is minimized to optimize the state vector $\hat{\mathbf{x}}$ ⁶:

$$\hat{\mathbf{x}} = \mathbf{x}_A + \mathbf{S}_A \mathbf{K}^T (\mathbf{K} \mathbf{S}_A \mathbf{K}^T + \mathbf{S}_0^{-1} (\mathbf{y} - \mathbf{K} \mathbf{x}_A)) \quad (1)$$

where \mathbf{x}_A is the prior state vector, considered 30 t/hr, \mathbf{S}_A is the prior error covariance matrix assuming 10% uncertainty for the CAMS boundary conditions and 100% uncertainty for the blowout emissions, \mathbf{K} is the Jacobian matrix, and \mathbf{y} is the observational vector containing TROPOMI observations. The model output is resampled to match the TROPOMI pixel spatial footprint using TROPOMI averaging kernels. The observations and model are then aggregated to $0.2^\circ \times 0.2^\circ$ grids to negate model errors. \mathbf{S}_0 , the observational error covariance matrix, is constructed as a diagonal matrix using the standard deviation of the difference between the prior modeled concentrations and TROPOMI observations.

To obtain a simulated plume that best matches the TROPOMI observed plume, we perform an ensemble of WRF simulations using: meteorological fields from either NCEP or ERA5, using four different planetary boundary layer physics options, and sampling the WRF outputs at the TROPOMI overpass time as well as up to 3 hours before and after the overpass time. Preliminary inversions are performed daily using the 56 simulated plumes, and the plumes with the lowest posterior observation cost function are selected. For the final inversion, we optimize the CAMS boundary conditions and the blowout emissions for each day. We only report quantifications for days with clear plumes, no potentially interfering downwind coastal artifacts, and good matches with simulated plumes based on visual inspection. We find that for all reported daily plume quantifications, the averaging kernel value for the blowout is above 0.5, showing the prior value has no significant impact on the estimated emissions. For estimating the uncertainty associated with the quantified emissions, an ensemble of inversions is computed by varying inputs and the assumptions used for the inversion⁷. This includes: increasing and decreasing the prior emissions for the blowout by a factor of 50%; using model outputs

sampled every one hour before and after the overpass hour; using the second best plume match based on observation cost; performing the inversions using different aggregation resolutions ($0.15^\circ, 0.25^\circ$); using TROPOMI data with highest quality flag of 1 (qa value = 1); using TROPOMI data without filtering for albedo; and following the central limit theorem instead of using mean observational error when aggregating observations, giving us a total of 1728 ensemble members for each day.

	Sensor	Acquisition date	Wind (m/s)	Q (t/h)	Q_err (t/h)
1	Sentinel-2	09/06/2023	3.0	63.3	42.0
2	Landsat 8-9	30/06/2023	2.9	42.9	14.6
3	PRISMA	11/07/2023	8.6	41.9	20.9
4	PRISMA	17/07/2023	7.3	38.0	18.2
5	GHGSAT	22/07/2023	4.6	49.3	16.8
6	GHGSAT	23/07/2023	1.5	21.2	11.4
7	PRISMA	23/07/2023	2.5	18.7	6.7
8	Landsat 8-9	01/08/2023	5.8	47.0	19.1
9	PRISMA	03/08/2023	8.9	28.5	14.1
10	GHGSAT	08/08/2023	4.5	35.8	12.5
11	EMIT	10/08/2023	5.6	21.1	9.4
12	GHGSAT	11/08/2023	2.8	39.1	17.2
13	GHGSAT	11/08/2023	5.6	24.9	7.5
14	GHGSAT	12/08/2023	5.1	43.2	13.8
15	EMIT	14/08/2023	5.3	35.6	15.7
16	GHGSAT	15/08/2023	2.8	27.7	12.2
17	Landsat 8-9	17/08/2023	3.9	41.4	15.4
18	GHGSAT	30/08/2023	6.3	30.2	8.5
19	PRISMA	01/09/2023	3.9	28.9	12.0
20	GHGSAT	04/09/2023	5.6	29.6	8.9
21	PRISMA	07/09/2023	5.7	37.8	17.3
22	PRISMA	19/09/2023	4.0	22.7	9.5
23	ENMAP	23/09/2023	6.9	32.4	15.4
24	Landsat 8-9	26/09/2023	4.9	23.8	9.3
25	ENMAP	27/09/2023	4.7	15.3	6.7
26	GHGSAT	01/10/2023	2.6	13.9	6.1
27	ENMAP	08/10/2023	8.5	14.2	7.0
28	ENMAP	27/10/2023	4.8	27.9	12.2
29	ENMAP	04/11/2023	1.9	9.0	2.9
30	ENMAP	08/11/2023	8.1	26.1	12.8
31	GHGSAT	04/12/2023	4.5	3.6	1.3
32	GHGSAT	09/12/2023	9.7	3.9	0.9
33	GHGSAT	25/12/2023	8.2	11.8	2.7

Table S1: Summary of the emission rates derived from the high-resolution satellite observations passing the quality screening process. This quality screening has removed the observations with cloud contamination, retrieval artifacts and plumes for which a substantial part of the tail lied outside the imaged area. Q refers to the flux rate estimated for each plume, and Q_err to the associated 1- σ uncertainty.

	Sensor	Acquisition date	Q (t/h)	Q_err (t/h)
1	TROPOMI	23/06/2023	40.0	17.3
2	TROPOMI	26/06/2023	52.0	16.3
3	TROPOMI	06/07/2023	31.5	18.4
4	TROPOMI	11/07/2023	53.2	29.9
5	TROPOMI	12/07/2023	32.9	14.3
6	TROPOMI	07/08/2023	27.4	12.6
7	TROPOMI	08/08/2023	29.5	12.6
8	TROPOMI	09/08/2023	47.0	23.1
9	TROPOMI	21/08/2023	43.2	20.8
10	TROPOMI	24/08/2023	32.9	17.1
11	TROPOMI	31/08/2023	51.0	31.3
12	TROPOMI	01/09/2023	42.4	29.7
13	TROPOMI	02/09/2023	45.6	21.5
14	TROPOMI	03/09/2023	27.4	17.1
15	TROPOMI	21/09/2023	15.3	3.3

Table S2: Summary of the emission rates derived from the TROPOMI satellite observations passing the quality screening process. This quality screening has consisted in the selection of only clear plumes and of good matches with simulated plumes based on visual inspection. Q refers to the flux rate estimated for each plume, and Q_err to the associated 1- σ uncertainty.

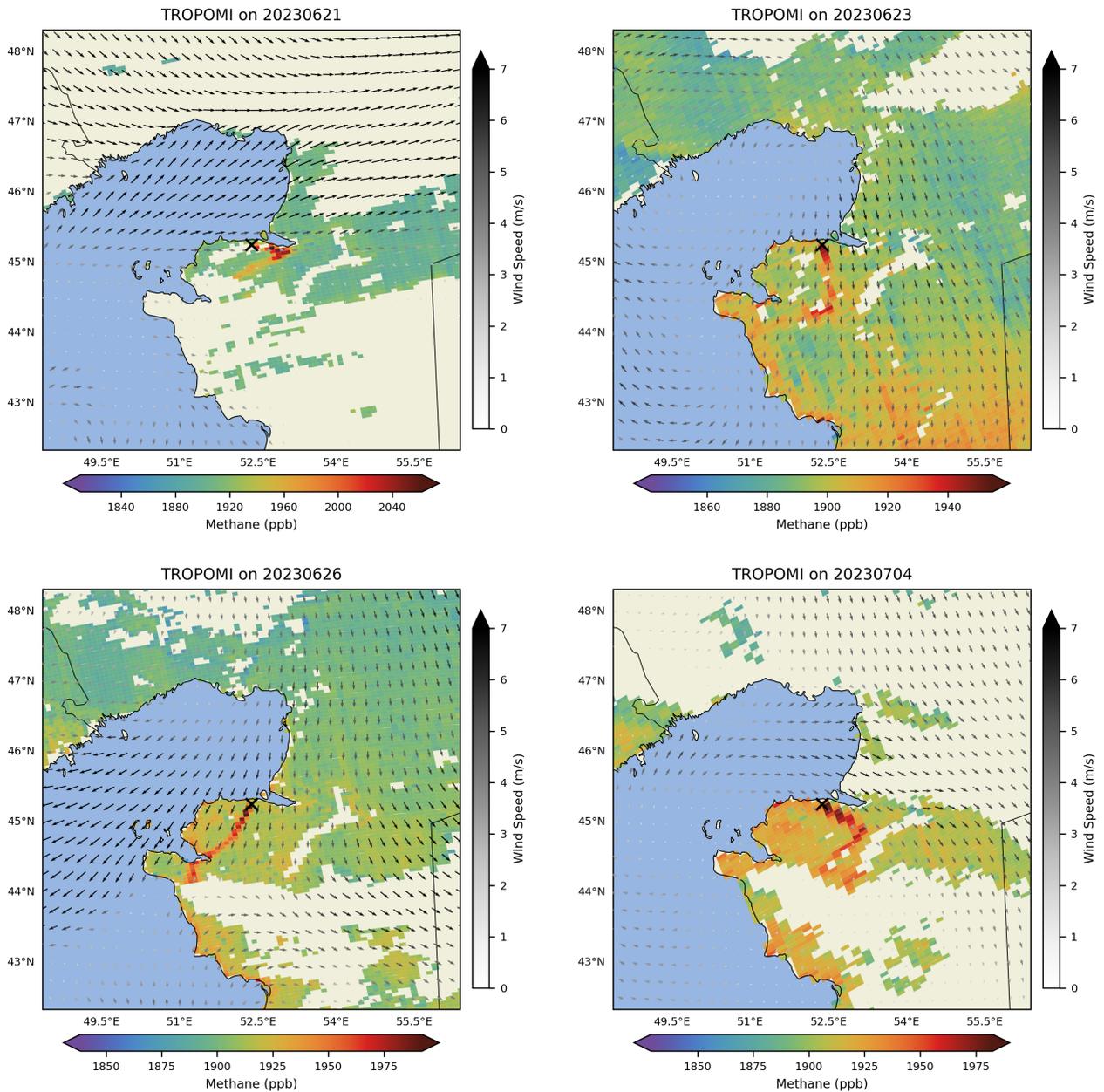


Figure S1: Methane column concentration maps for the first observations of the Karaturun 2023 leak by Sentinel-5P/TROPOMI. Data are from the TROPOMI-CH₄ operational product (02.05.00 version). The color scale indicates the total methane column concentration estimated from TROPOMI (as opposed to the concentration enhancement derived from the high resolution data). Arrows indicate wind intensity and direction. TROPOMI data on 21 June and 4 July were not used to quantify emissions as we could not obtain a good match with a modeled plume.

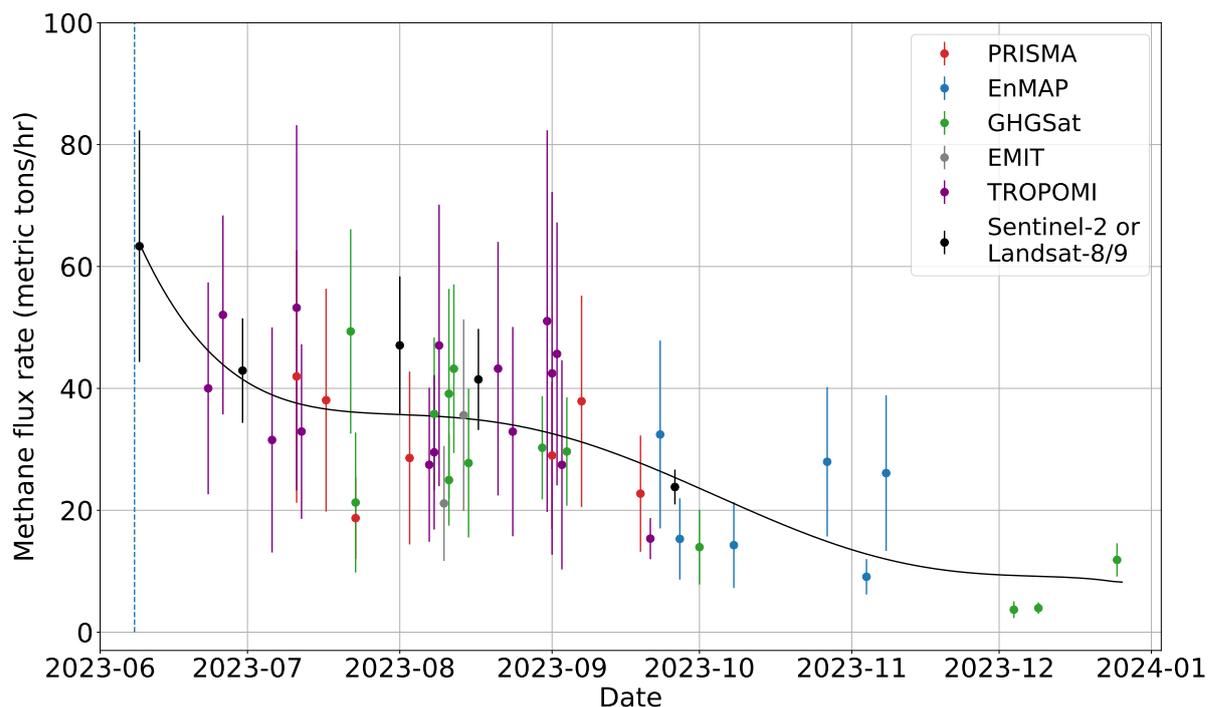
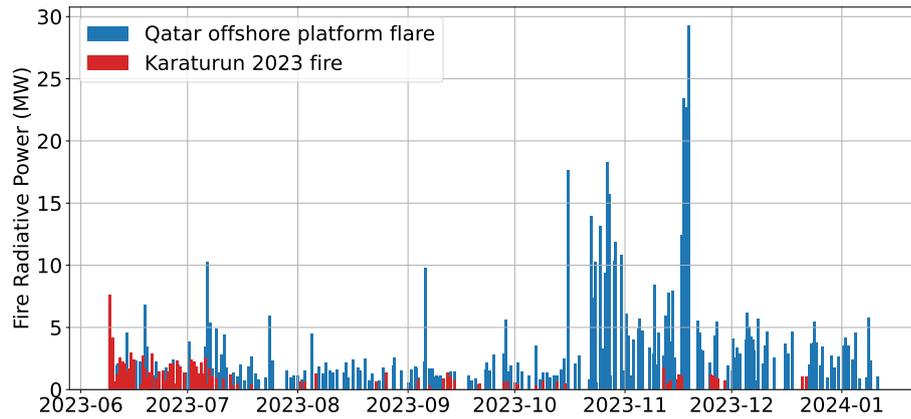
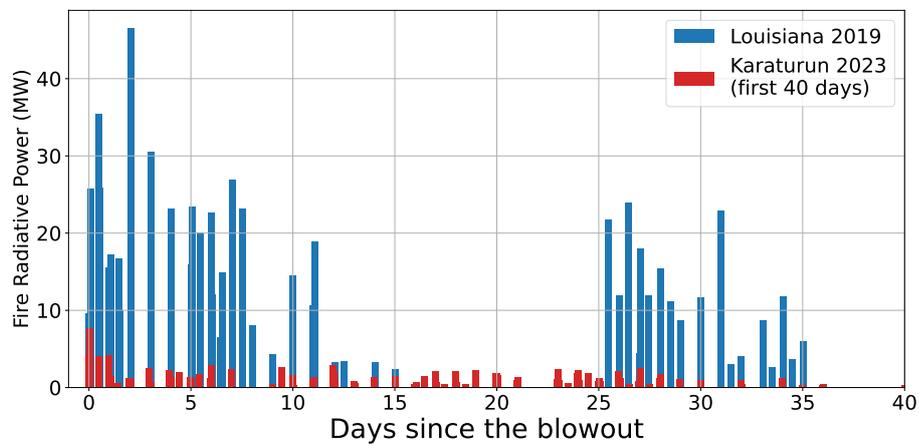


Figure S2: Time series of flux rate estimates obtained from the different satellites used in this work. The points represent the flux rate estimates for the 48 plumes retained for quantification after quality screening. This figure offers a more clear representation of the data derived from each satellite as compared with Fig. 3a of the Main Text.



(a)



(b)

Figure S3: Comparison of the fire intensity at the Karaturun 2023 blowout site with that of other gas flaring events. (a) Regular flare in an offshore platform in Qatar (26.59°N , 52.00°E). (b) Fire intensity during the Louisiana 2019 event ([Maasackers et al., 2022](#)), where gas burned first at the wellheads for two weeks, and then at a flare pit for 10 days, with a 10-day period in between during which the gas was vented. In both cases, fire intensity is proxied by the fire radiative power variable provided in the VIIRS Fire Information for Resource Management System (FIRMS) product.

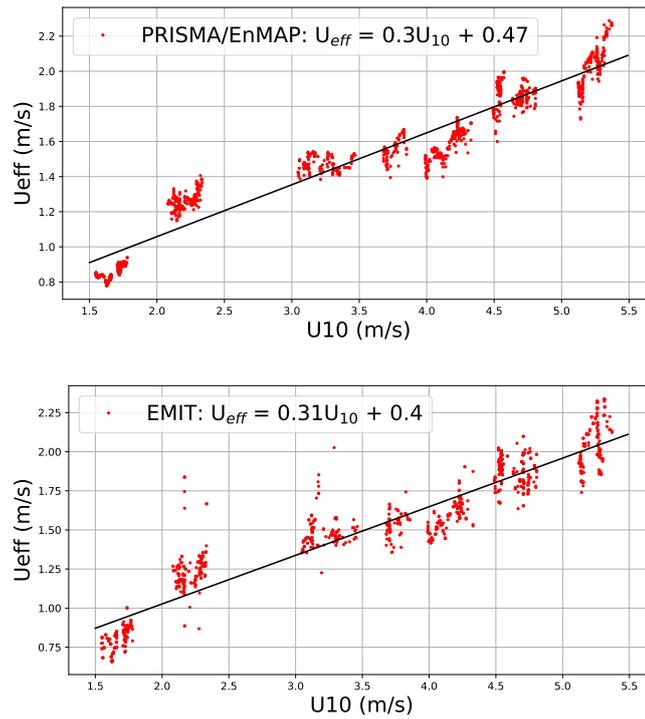


Figure S4: Empirical $U_{\text{eff}} - U_{10}$ models for IME-based flux rate estimates from EnMAP, PRISMA and EMIT ΔX_{CH_4} retrievals. The linear models have been generated using a database of plumes simulated with a WRF-LES modeling approach. The plume simulations cover the flux rate range of 10–90 t/h, and are done for the ΔX_{CH_4} retrieval noise found in the Karaturun East site for those missions. Separate models have been generated for EnMAP-PRISMA and EMIT because of the different spatial sampling (30 m for EnMAP-PRISMA and 60 m for EMIT). Retrieval precision is assumed to be similar for EnMAP and PRISMA.

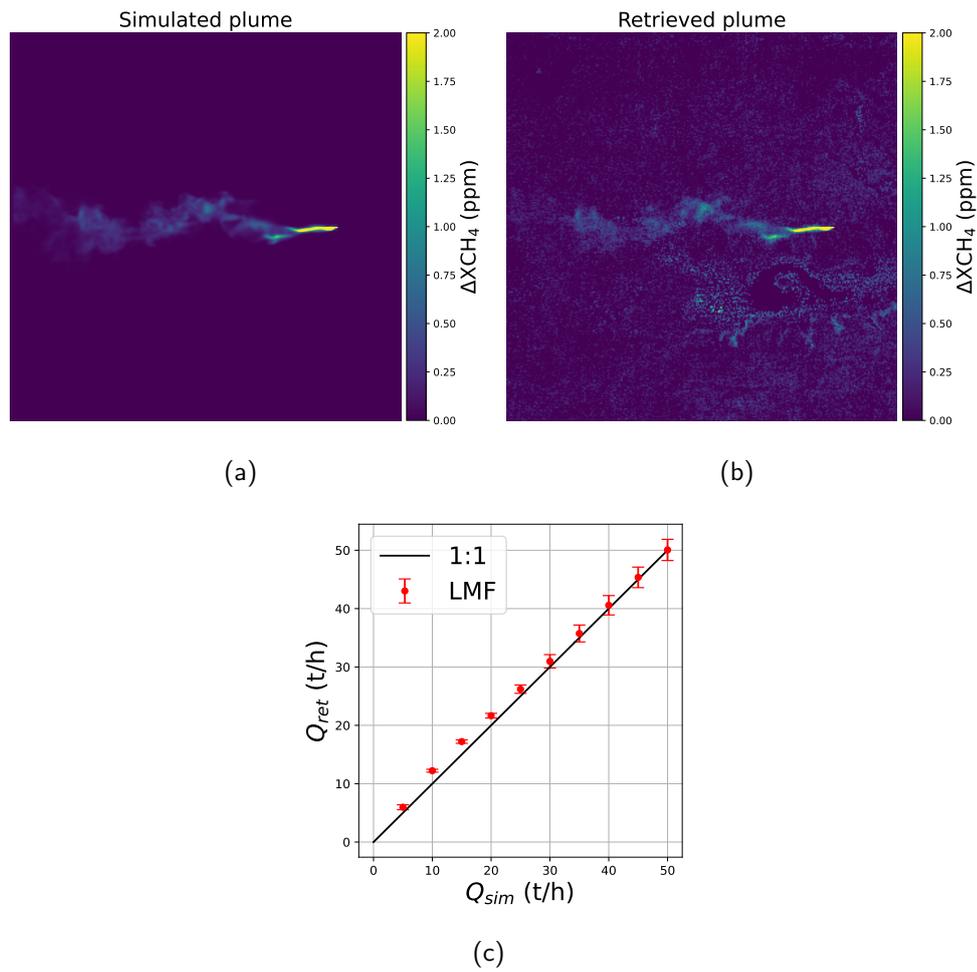


Figure S5: Verification of ΔXCH_4 retrievals and flux rate estimates from hyperspectral data with end-to-end simulations. (a), Simulated methane plume (25 t/h) added to a real PRISMA radiance dataset. (b), Methane plume retrieved from the processing of the resulting PRISMA radiance dataset using the ΔXCH_4 retrieval scheme implemented for this study. (c), Comparison of the input and estimated flux rates (Q_{sim} and Q_{ret} , respectively) from the entire end-to-end simulation process.

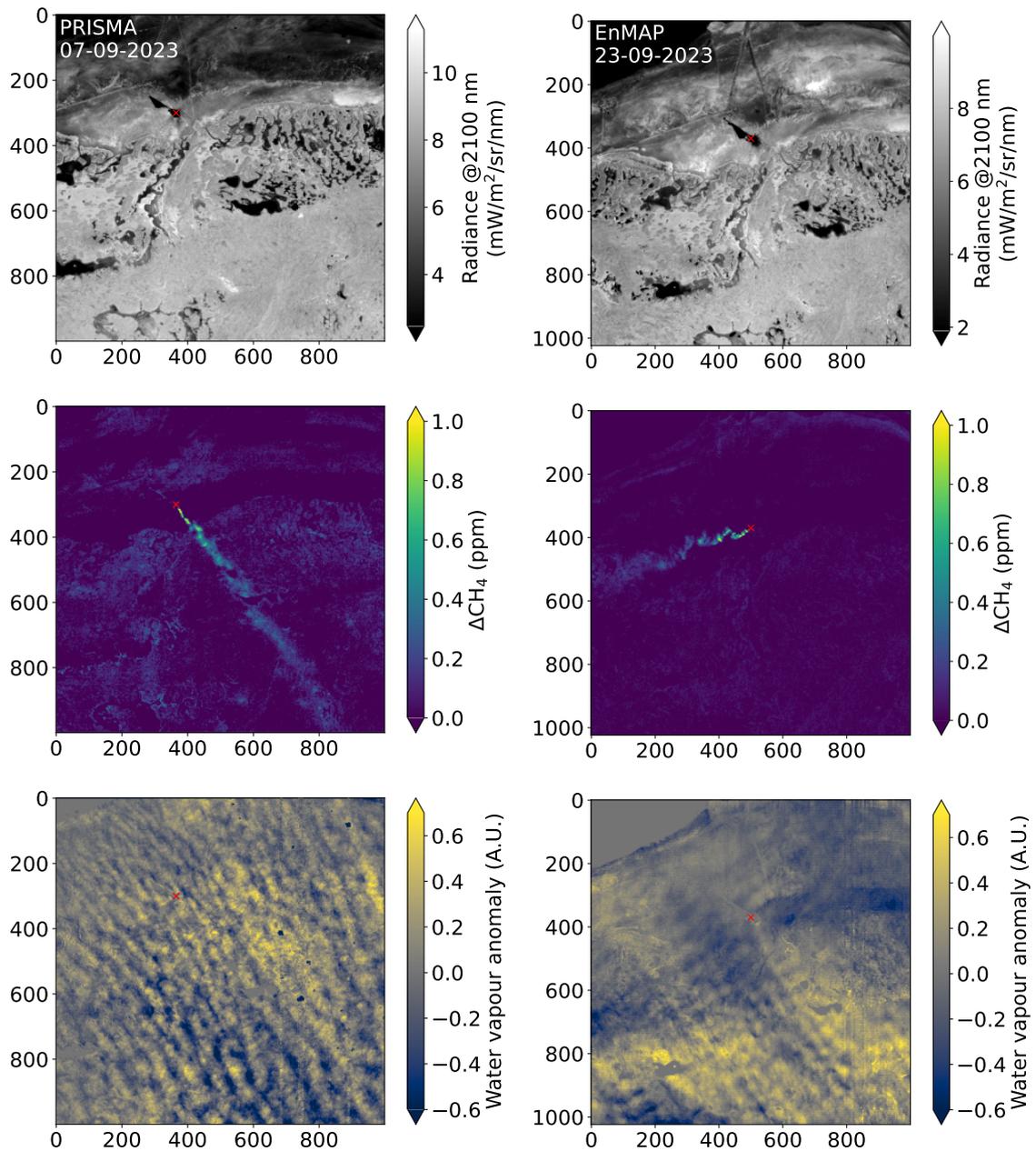
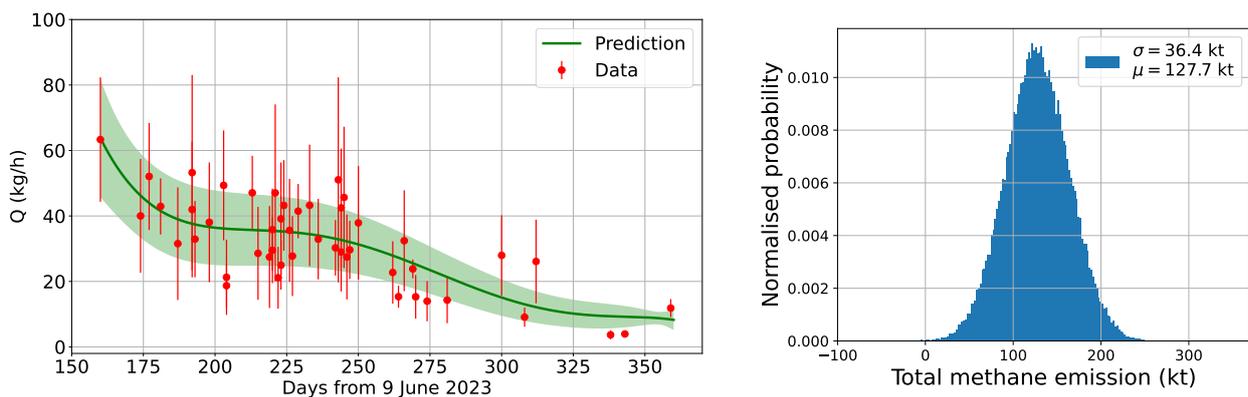
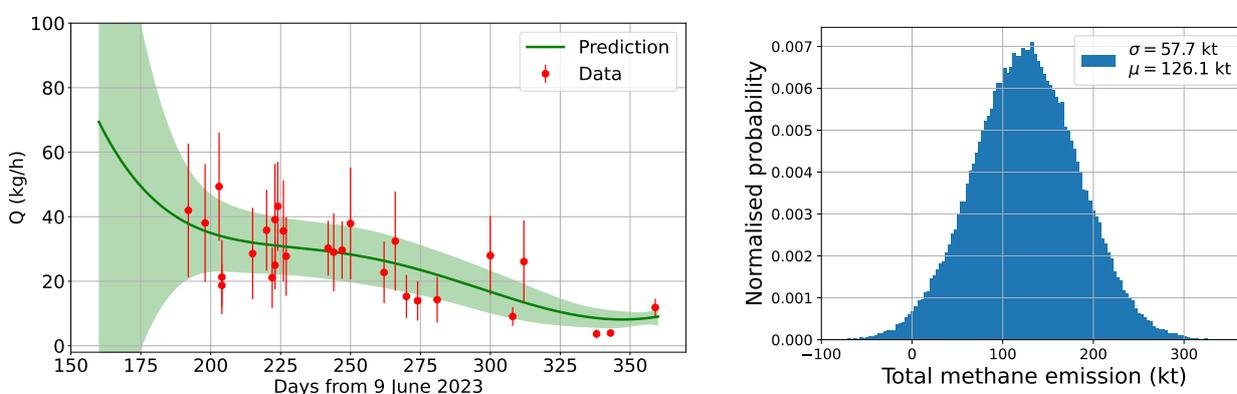


Figure S6: Assessment of the potential distortion of ΔXCH_4 retrievals by water vapour and smoke. Maps of at-sensor radiance at 2100 nm (top row), ΔXCH_4 (center row), and water vapour concentration anomaly (bottom row) derived from a PRISMA acquisition from 7 September 2023 (left column), and an EnMAP acquisition from 23 September 2023 (right column). No water vapour plume superposed to the methane plume can be observed from the comparison of ΔXCH_4 and water vapour anomaly maps. Also, no smoke plume can be observed in the 2100 nm radiance maps, which suggests that, in this particular event, smoke has a negligible optical activity on the shortwave infrared window from which ΔXCH_4 maps are retrieved.



(a) All 48 high-quality plumes used for the total emission quantification (see Fig. 3a and Fig. S2)



(b) Hyperspectral-only plume detections (PRISMA, EnMAP, EMIT, GHGSat)

Figure S7: Quantification of the total amount of methane released by the leak. Top row, results from the dataset consisting of the 48 high quality plumes (including TROPOMI as well as hyperspectral and multispectral high-resolution observations) used in this study for the quantification of the leak in this study. Bottom row, results from an alternative hyperspectral-only configuration. The time series on the left hand side depict a polynomial fit of the satellite-based flux rate estimates (Q) selected after quality screening. The fitted model is integrated to obtain an estimate of the total leak. The shaded green area corresponds to the uncertainty ($k=1$) of the flux rate fitting. The probability distribution functions on the right hand side show the result of propagating the temporal flux rate together with the uncertainty using multivariate Monte Carlo simulations. An error correlation of 0.5 is assumed for the individual satellite observations. The comparison between the top and bottom row illustrates the fact that the extra observations from TROPOMI and the multispectral missions contribute to decrease the uncertainty range but have a very low impact on the total emission estimate.

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