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Sink of eddy energy by submesoscale sea surface temperature variability in a coupled regional model

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8 ABSTRACT: Air-sea interaction impacts ocean energetics via modifications to the exchange of
9 momentum and buoyancy. Prior work at the submesoscale has largely focused on mechanisms
10 related to the eddy kinetic energy (EKE), such as the current feedback on stress, which generates
11 negative wind work, or variations in sea surface temperature (SST) that modify surface winds.
12 However, less is known about the influence of submesoscale SST variability on ocean energetics
13 through its direct effect on the surface flux of eddy potential energy (EPE). Here the role of
14 EPE flux on submesoscale ocean energetics is investigated using a fully-coupled model of the
15 California Current region, including a numerical experiment that suppresses the thermal response
16 in the computation of air-sea fluxes at the submesoscale. Correlations between surface buoyancy
17 anomalies and surface buoyancy fluxes lead to an approximately 10–25% loss of submesoscale
18 EPE, which results in similar magnitude reductions of the vertical buoyancy production, EKE,
19 and eddy wind work. The changes induced by this mechanism in the energy reservoirs and
20 dissipation/conversion pathways are on the same order of magnitude as the negative wind work
21 induced by the current feedback. An approximate form of the EPE flux shows that it is a function
22 of the density ratio and proportional to the surface EPE reservoir of the system. These findings
23 indicate the importance of the submesoscale SST variability, and small-scale variability in surface
24 heat fluxes, in modifying energy reservoirs and conversion pathways of the ocean via the direct
25 flux of EPE at the air-sea interface.

26 SIGNIFICANCE STATEMENT: This work investigates the impact of small oceanic frontal
27 features in the ocean, classified as submesoscale, in the exchange of energy at the air-sea boundary.
28 Submesoscale fronts and filaments range from approximately 0.1 to 10 km and are characterized
29 by strong horizontal density changes and fast-evolving flow. The associated density anomalies at
30 the surface may be important in the overall energy budget of the surface ocean since they can affect
31 the energy fluxes at the air-sea boundary. Two numerical experiments were set up for a comparative
32 analysis of the energy transfer, conversion, and storage in the upper layer of the California Current
33 region. In one experiment, the role of sea surface temperature anomalies in generating air-sea
34 fluxes is suppressed. A comparison between the two experiments shows a difference of 10-25%
35 in the energy storage and conversion. Sea surface temperature variability may induce a reduction
36 of energy via air-sea fluxes similar to energy dissipation driven by wind-current interactions in the
37 same scale of phenomena.

38 **1. Introduction**

39 The turbulent heat and momentum exchanges across the ocean-atmosphere interface are intrinsi-
40 cally dependent on the scale of the ocean features (Seo et al. 2023). SST variability at the mesoscale
41 plays an essential role in modifying the overlaying atmosphere dynamics which in turn leads to
42 substantial coupled responses of the ocean (Bishop et al. 2017; Chelton and Xie 2010; O'Neill
43 et al. 2012; Small et al. 2008). However, much of our understanding of how ocean variability leads
44 to coupled interactions is constrained to mesoscale resolution (10-100 km). At smaller scales in
45 the ocean, frontal and filamentous features of the order of 0.1-10 km – denoted submesoscale –
46 are characterized by sharper temperature gradients and ageostrophic flows. Submesoscale currents
47 are common oceanic features driven by the downscale eddy cascade of mesoscale flows and are
48 important to global ocean dynamics (McWilliams 2016; Wenegrat et al. 2018). As the dynamics
49 of submesoscale currents are strongly ageostrophic, strong vertical velocities are characteristic
50 in the flow which allow for strong transport of properties such as dissolved gases, nutrients, and
51 heat (Mahadevan et al. 2012; Renault et al. 2016; Balwada et al. 2021). The vertical flux of heat
52 (buoyancy) affects both the timing and strength of ocean stratification (Mahadevan et al. 2012;
53 Johnson et al. 2016) and the surface flux of heat between the ocean and atmosphere (Su et al. 2018,
54 2020; Iyer et al. 2022).

55 Air-sea interaction at the submesoscale is somewhat less well understood since numerical simu-
56 lations are computationally costly and observations are challenging. Observation works that have
57 investigated the air-sea fluxes at the submesoscale, although scarce, have shown larger fluxes of
58 heat, moisture, and momentum at fronts (Shao et al. 2019; Iyer et al. 2022), also consistent with
59 submesoscale-permitting global ocean models analysis that used uncoupled air-sea bulk formulae
60 (e.g., Su et al. 2018, 2020). Coupled numerical simulations have shown an active EKE transfer at
61 the air-sea interface by submesoscale variations in surface wind stress (e.g., Renault et al. 2018; Bai
62 et al. 2023; Conejero et al. 2024). For example, coupled modeling experiments of the California
63 Current System indicate that modifications to the wind stress by small-scale currents (the current
64 feedback on stress, CFB) lead to a 17% reduction in submesoscale EKE (Renault et al. 2018).
65 These changes to the surface stress also modify the Ekman transport of buoyancy at fronts, and
66 consequently the PV budget of the surface mixed-layer (Wenegrat 2023). In addition, modulations
67 of the marine atmospheric boundary layer and changes in atmospheric kinetic energy by SST vari-
68 ability, namely, the thermal feedback mechanism (TFB), were also explored in idealized models
69 (Wenegrat and Arthur 2018; Sullivan et al. 2020, 2021). These results indicate that sharp fronts
70 at the submesoscale impact the response of the marine atmospheric boundary layer by driving
71 secondary circulations in the atmosphere which in turn modify the surface wind stress and wind
72 work (Skylvingstad et al. 2007; Wenegrat and Arthur 2018; Sullivan et al. 2021). Recent studies
73 also show the combined effect of CFB and TFB in the wind stress (Bai et al. 2023; Conejero et al.
74 2024), which indicates that submesoscale SST variability shows a direct influence on the transfer
75 of momentum between the atmosphere and ocean, modifying the surface flux of EKE.

76 The influence of submesoscale SST variability on ocean energetics through its direct effect on
77 the surface flux of EPE, however, is less explored. Observations show strong covariability between
78 surface heat fluxes and surface buoyancy anomalies at the submesoscale (Shao et al. 2019; Iyer
79 et al. 2022; Yang et al. 2024), suggesting there will also be a direct surface flux of EPE. This has
80 been shown to be an important sink of mesoscale EPE (Bishop et al. 2020; Guo et al. 2022), which
81 can impact the baroclinic conversion rate in boundary currents in the first 100 m of the upper ocean
82 (Ma et al. 2016; Renault et al. 2023), but has not yet been explored at the submesoscale. Here we
83 investigate the impact of SST anomalies on submesoscale EPE flux using a fully-coupled regional
84 model of a portion of the California Current system, a region where submesoscale features have

85 been indicated as important drivers of air-sea fluxes as momentum and heat (Capet et al. 2008b;
 86 Renault et al. 2018).

87 Two coupled ocean-atmosphere simulation setups are used to assess the effect of submesoscale
 88 SST variability on the EPE flux, including both a fully coupled simulation and one where sub-
 89 mesoscale SST anomalies are not included for air-sea flux calculations. The flux, conversion,
 90 and storage components of eddy energy in the mixed layer for both simulations are compared,
 91 highlighting an increase of eddy energy when SST anomalies do not affect surface fluxes. The
 92 impact of the EPE flux is not limited to the EPE but also propagates to changes in EKE through
 93 modification of the vertical buoyancy production and changes to the surface wind work. This
 94 analysis shows that the flux of EPE driven by SST at the submesoscale is comparable to analogous
 95 transfers of EKE by surface momentum transfer (wind work) at the submesoscale.

96 The work is organized as follows. Section 2 introduces the theoretical background for the eddy
 97 energy equations in spectral space. The numerical experiments are described in section 3. In
 98 section 4, the submesoscale dynamics of the numerical simulations are described, and the impact
 99 of the surface EPE flux is estimated. An approximated form of the EPE flux is obtained and
 100 compared with the flux of EKE by the surface wind work in section 5. Finally, the results are
 101 summarized in section 6.

102 2. Energy equations and spectral analysis

103 *a. Eddy potential energy and eddy kinetic energy equations*

104 This work compares the eddy energy pathways and reservoirs of the upper ocean, assessing
 105 the influence of submesoscale SST anomalies in the air-sea fluxes variability. We consider the
 106 reservoirs, conversion rates, and flux terms of eddy energy in horizontal wavenumber space. The
 107 mixed-layer integrated EPE and EKE equations for a Boussinesq fluid in spectral space are written
 108 as

$$\int_{z_m}^0 \frac{\partial}{\partial t} EPE \, dz = \int_{z_m}^0 \left(-A_b - C_{(EKE,EPE)} + V_b \right) dz, \quad (1)$$

$$\int_{z_m}^0 \frac{\partial}{\partial t} EKE \, dz = \int_{z_m}^0 \left(-A_m - PH + C_{(EKE,EPE)} + V_m \right) dz, \quad (2)$$

109 where the left-hand side terms represent the rate of change of EPE and EKE integrated from the
 110 surface to the mixed layer depth (z_m). The eddy terms analyzed in this work refer to the variability
 111 encompassed at the small mesoscale and submesoscale horizontal wavenumbers (see section 3c).
 112 The reservoirs are described as follows

$$EPE = \mathbb{R} \left[\frac{\widehat{b} \widehat{b}^*}{2N_r^2} \right], \quad (3)$$

$$EKE = \mathbb{R} \left[\frac{\widehat{u} \widehat{u}^* + \widehat{v} \widehat{v}^*}{2} \right], \quad (4)$$

113 where b is buoyancy, $N_r^2 = \frac{\partial \langle b \rangle}{\partial z}$ is the reference squared Brunt Vaissala frequency (i.e., $\langle b \rangle$ is
 114 the horizontally and temporally averaged buoyancy), and u and v are the zonal and meridional
 115 velocity components. The caret ($\widehat{\cdot}$) denotes the two-dimensional Fourier transform. The symbol \mathbb{R}
 116 represents the real component of the spectra and the asterisk ($*$) indicates the complex conjugate
 117 operator.

118 The A_b term in (1) and the A_m term in (2) include the horizontal advection of EPE and the 3D
 119 advection of EKE, respectively. Those terms also include the cross-scale fluxes of energy which
 120 are not analyzed here due to domain size constraints (briefly discussed in section 6). The PH term
 121 in (2) represents the pressure work, also not described in this work.

122 The rate of conversion between EPE and EKE, $C_{(EKE,EPE)}$, is

$$C_{(EKE,EPE)} = \mathbb{R} \{ \widehat{w} \widehat{b}^* \} \quad (5)$$

123 where w represents the vertical velocity component. Conversion of EPE at the submesoscale is
 124 generated by baroclinic mixed-layer instabilities and other ageostrophic secondary circulations that
 125 extract available potential energy from fronts (Fox-Kemper et al. 2008).

126 The V_b term in (1) arises from the diffusive term in the buoyancy budget, and hence once
 127 integrated over the mixed-layer contains contributions from both the dissipation of EPE and the
 128 diffusive fluxes of EPE at the surface and mixed-layer base (detailed derivation in Appendix A).
 129 The surface EPE flux is the focus of this work, which can be determined using the surface boundary
 130 condition (Cronin and Sprintall 2001; Storch et al. 2012)

$$\kappa \frac{\partial}{\partial z} b \Big|_{z=0} = B_o, \quad (6)$$

131 where κ is the vertical diffusivity and B_o is the surface buoyancy flux. The EPE flux, hereafter
 132 referred to as G_{EPE} , is then

$$G_{EPE} = \mathbb{R} \left[\frac{\widehat{b_o^* B_o}}{N_r^2} \right] \quad (7)$$

133 where b_o is the surface buoyancy. Negative G_{EPE} values indicate EPE loss from the ocean, whereas
 134 positive G_{EPE} indicates a gain of EPE.

135 Analogously, the diffusive flux of EKE at the surface (or wind work) may be calculated by
 136 vertically integrating V_m and using the surface boundary conditions for momentum (more details
 137 of the derivation in Capet et al. 2008b; Storch et al. 2012). The wind work, G_{EKE} , is thus defined
 138 as

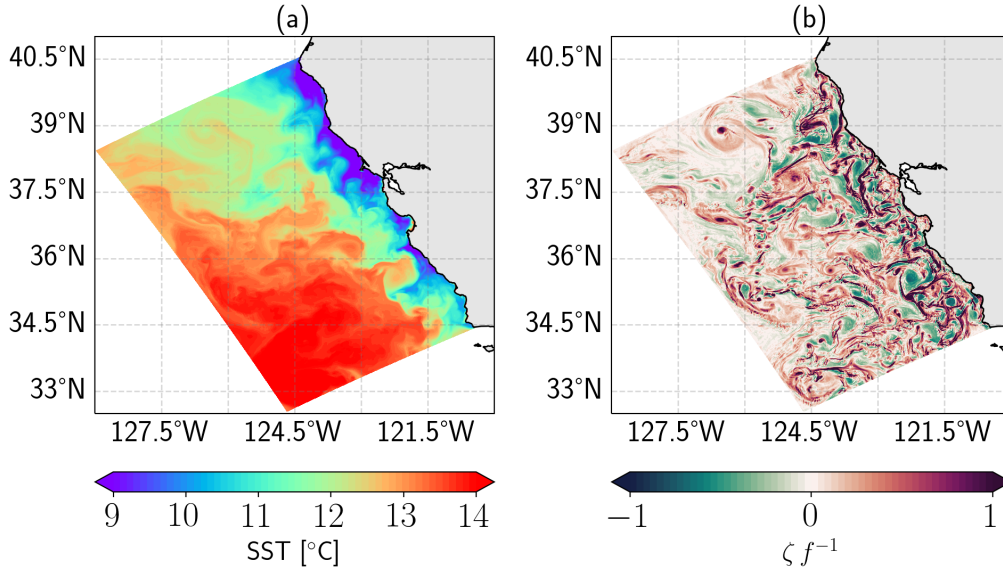
$$G_{EKE} = \frac{1}{\rho_o} \left(\mathbb{R} \left[\widehat{\tau_x u_o^*} + \widehat{\tau_y v_o^*} \right] \right), \quad (8)$$

139 where ρ_o is the background surface density, $\boldsymbol{\tau} = (\tau_x, \tau_y)$ is the surface momentum flux, and $\mathbf{u}_o =$
 140 (u_o, v_o) is the surface velocity, with the zonal and meridional velocity components, respectively.

141 3. Numerical simulation

142 a. Model description

143 The ocean components of the coupled model in the California Current System region use
 144 the Regional Oceanic Modeling System (ROMS) in its CROCO version (Coastal and Regional
 145 Oceanic COmmunity) (Shchepetkin and McWilliams 2005; Debreu et al. 2012; Shchepetkin 2015).
 146 CROCO is a free-surface, terrain-following coordinate model with split-explicit time stepping.
 147 The equations solved in this model's configurations were set for Boussinesq and hydrostatic
 148 approximations. The numerical experiments used in this work are the highest resolution products
 149 from a four nest configuration described in Renault et al. (2018). The domain for the simulations
 150 covers 119.9° W to 128.98° W and from 32.54° N to 40.73° N (Fig. 1). The simulations were spun
 151 up from the same initial state from June to November 2011, after which they were run separately



160 FIG. 1. Snapshots of (a) sea surface temperature and (b) surface vorticity normalized by the Coriolis frequency
 161 (*f*) from the fully coupled simulation illustrating the model domain.

152 from November 2011 to June 2012 (more detailed description in section 3b). The boundary and
 153 initial conditions are taken from a coarser 4-km nested grid. For the horizontal grid, 1000×1520
 154 points with a grid spacing of $(\Delta x, \Delta y) = 0.5$ km were set with 80 terrain-and-surface-following
 155 sigma levels in the vertical with stretching parameters $h_{\text{cline}} = 200$ m, $\theta_b = 3.0$, and $\theta_s = 6$. The
 156 turbulence closure used is the K-Profile Parameterization (KPP, Large et al. 1994). The outputs
 157 analyzed in this work have a 6-month time span (January to June 2012) with a 6-hour temporal
 158 resolution. More information about the settings and spin-up of the model can be found in Renault
 159 et al. (2018).

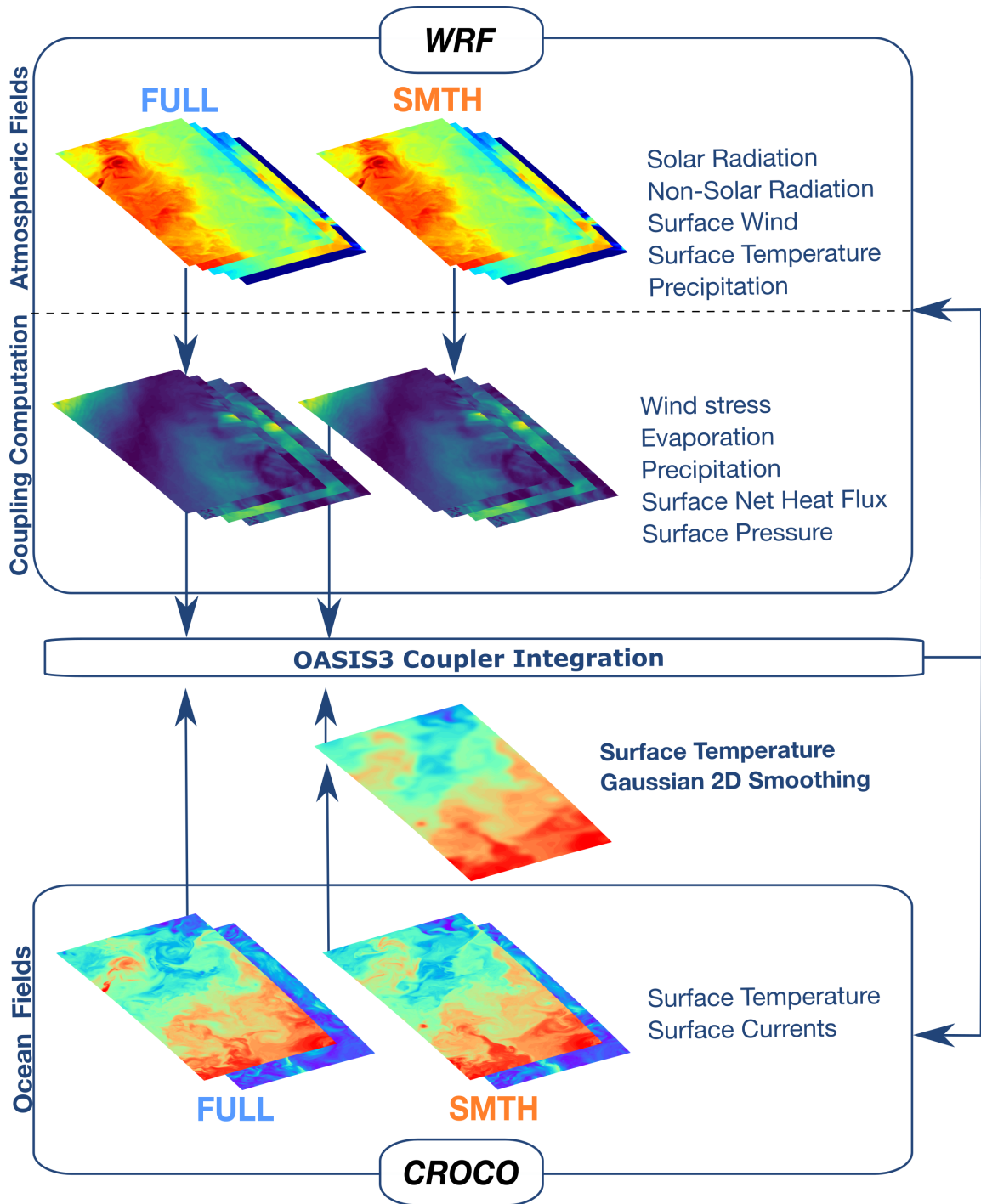
162 For the atmospheric component of the fully-coupled system, the Weather Research and Forecast
 163 Model (WRF, version 4.1) was used (Skamarock et al. 2019). An implementation of a nesting grid
 164 is also used in this model as in Renault et al. (2018). The atmospheric component used in this work
 165 has a spatial resolution of 2 km with initial and boundary conditions provided by the simulation
 166 from the previous nesting with a 6 km horizontal resolution. The domain for the simulations covers
 167 118.98° W to 129.14° W and from 32.44° N to 41.20° N, which is slightly larger than the ocean
 168 domain to avoid the WRF sponge boundaries. For the horizontal grid, 300×390 points with a
 169 grid spacing of $(\Delta x, \Delta y) = 2$ km were set with 50 vertical levels. In the boundary layer model,

170 bulk formulae (COARE formulation, Edson et al. 2013) are used to compute the surface turbulent
171 heat, freshwater, and momentum fluxes, which are subsequently provided to CROCO. Note that,
172 in this work, the implementation used of a surface-layer vertical mixing parameterization for the
173 planetary boundary layer (MYNN, Nakanishi and Niino 2006) and a tri-diagonal matrix for vertical
174 turbulent diffusion is necessary for the implementation of relative winds in the atmospheric model
175 and to reproduce the CFB mechanism appropriately (Renault et al. 2019).

176 The OASIS3 software was used for the surface data exchange between the two models (Valcke
177 2013) to couple CROCO and WRF. This procedure supports the communication of two-dimensional
178 fields between the two numerical codes for the integration of the coupled system. The diagram
179 in Fig. 2 illustrates the surface fluxes computation using this software. In these experiments,
180 WRF provides the hourly averages of freshwater, heat, and momentum fluxes to CROCO whereas
181 CROCO feeds the hourly SST and surface currents to WRF for the calculation of fluxes. OASIS3
182 is implemented in the 4 km and 6 km grids for CROCO and WRF, respectively, and nested into the
183 higher resolution grids.

184 *b. Experiment setup*

185 To observe the impacts of SST variability at the submesoscale on the upper ocean dynamics,
186 two fully-coupled numerical simulations were implemented using two different air-sea coupling
187 configurations. A schematic of the two experimental setups is illustrated in Fig. 2. The first
188 experiment consists of a fully-coupled model system, hereafter referred to as the FULL experiment.
189 In the second experiment, SST anomalies are low-pass filtered before being passed to WRF for
190 the calculation of surface fluxes, suppressing the role of submesoscale SST variability in air-
191 sea interaction as illustrated in Fig. 2. The latter experiment will be referred to as the SMTH
192 experiment. This comparison between experiments assesses the impact of the ocean submesoscale
193 SST variability on the exchange of heat and momentum at the air-sea interface. This analysis is
194 similar to previous studies performed in mesoscale-resolving simulations (Zhai and Greatbatch
195 2006; Seo et al. 2016; Renault et al. 2023).



196 FIG. 2. Schematics of the different coupling computations for the FULL and SMTH experiments using WRF
 197 (The Weather Research and Forecast Model) and CROCO (Coastal and Regional Ocean Community Model). The
 198 examples illustrate the computation of sensible heat flux. The filtering of submesoscale sea surface temperature
 199 variability for the coupling computation is illustrated for the SMTH experiment.

200 *c. Spatial filtering and spectral analysis*

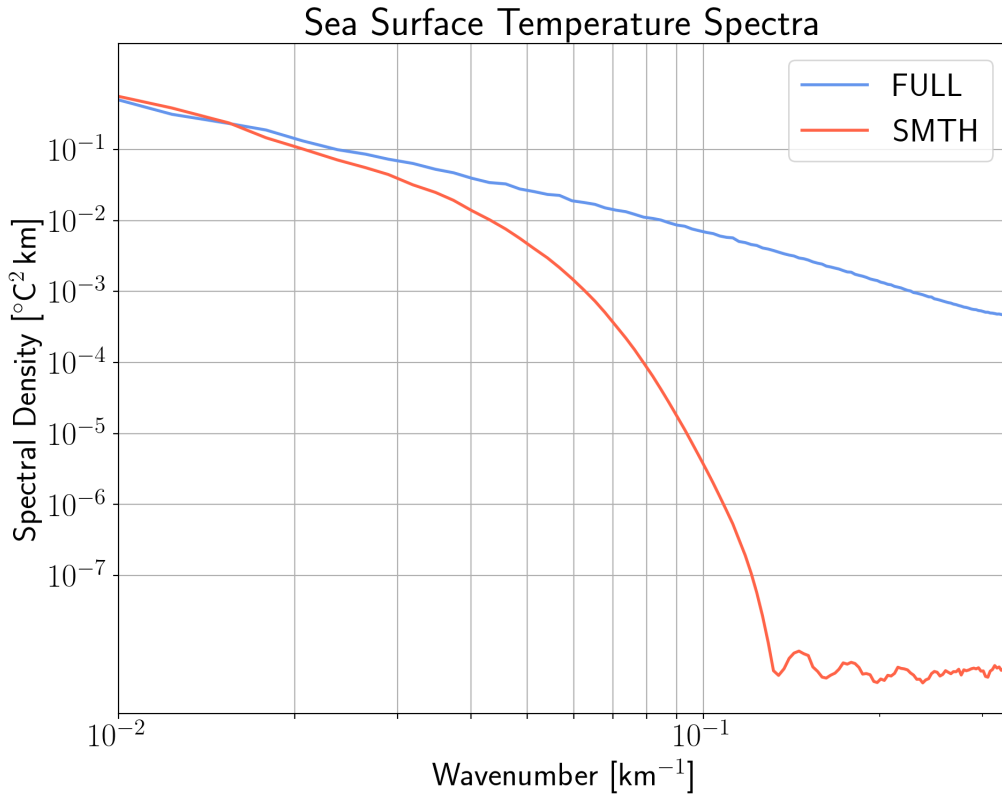
201 A two-dimensional spatial Gaussian filter is used to isolate submesoscale anomalies from the
202 mesoscale and large-scale signals. The filter applies a $(6\sigma + 1)$ window in both horizontal di-
203 mensions and has a $\sigma=3$ and a cutoff value of 0.5 as performed in Renault et al. (2023). This
204 configuration allows for an assessment of the impact of SST submesoscale anomalies in the energy
205 fluxes, reservoirs, and conversion rates. In Fig. 3 an example of the differences in the SST field
206 used in the air-sea coupling between simulations is shown. The filter reduces variability from
207 approximately 50 km wavelength (0.02 km^{-1} wavenumber) to smaller scales, such that at 20 km
208 wavelength (0.05 km^{-1} wavenumber) SST variability is reduced by an order of magnitude. Here we
209 refer to the range of scales smaller than this filter scale as ‘submesoscale’ however we note that the
210 submesoscale is more accurately defined as a dynamical regime, and hence the definition employed
211 here is only approximate. The Fourier transform calculation in this work includes subtraction of
212 the spatial mean and tapering using a Hanning window. A temporal average of the period of the
213 simulations (i.e., 6 months) is also applied in all spectra.

216 **4. Results**

217 *a. Model characterization*

218 The submesoscale dynamics of the California Current are depicted in Fig. 1 where SST and
219 normalized relative vorticity fields (i.e., Rossby number) show strong variability in the region.
220 Smaller-scale vortices and their associated high normalized relative vorticity, $Ro \sim \mathcal{O}(1)$, indicate
221 the presence of flows that are dynamically submesoscale, a consequence of mesoscale strain and
222 frontal instabilities of the California Current (Capet et al. 2008a).

223 The surface eddy energy reservoirs, vorticity, and divergence are impacted by air-sea fluxes
224 driven by SST anomalies at the submesoscale as shown in Fig. 4. The SMTH experiment has more
225 eddy energy (in both EPE and EKE) than the FULL simulation that accounts for submesoscale
226 SST variability in air-sea fluxes (Fig. 4a). This surplus of energy indicates more variability
227 in velocity and buoyancy at the submesoscale when air-sea fluxes driven by SST anomalies are
228 suppressed. Both EPE and EKE surface spectra present a slope of $\sim k_h^{-2}$ which is associated with
229 flows with energetic submesoscale currents (Capet et al. 2008a) and a white horizontal gradient
230 spectra. The EKE spectral slope found is similar to observations in adjacent regions such as the

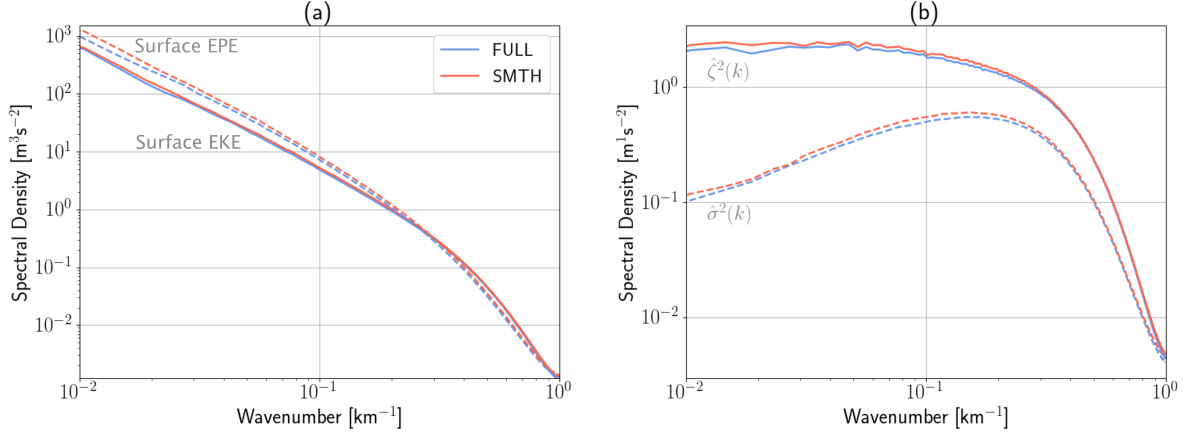


214 FIG. 3. Isotropic wavenumber spectra comparing the sea surface temperature fields in the coupling computation
 215 for the model simulation setups FULL and SMTH.

231 Southern California Current (Chereskin et al. 2019), which attributes the behavior to an energetic
 232 submesoscale and relatively weaker mesoscale than in Western Boundary Currents. Vorticity (ζ)
 233 and divergence (δ) spectra are proportional to the horizontal velocity gradient, which indicate sharp
 234 velocity gradients commonly observed in submesoscale fronts and filaments (Barkan et al. 2019).
 235 Fig. 4b indicates weaker velocity gradients in the FULL experiment compared to the SMTH case.

239 *b. Eddy potential energy flux at the submesoscale*

240 SST anomalies at the submesoscale enhance the loss of EPE via correlations between the thermal
 241 components of buoyancy and buoyancy flux. A schematic representation of the mechanism above
 242 is shown in Fig. 5 where spatial anomalies of buoyancy (b') and buoyancy flux (B'_o) are correlated.
 243 The heat flux anomalies respond to SST anomalies at the front to diminish the differences in
 244 temperature between the surface ocean and the atmosphere. This mechanism decreases the absolute



236 FIG. 4. Surface dynamics and energetics are influenced by air-sea fluxes at the submesoscale. Two-dimensional
 237 spectra of surface (a) eddy kinetic energy (EKE) and eddy potential energy (EPE), and (b) vorticity (ζ) and
 238 divergence (σ) for the FULL (blue line) and SMTH (orange line) simulation outputs.

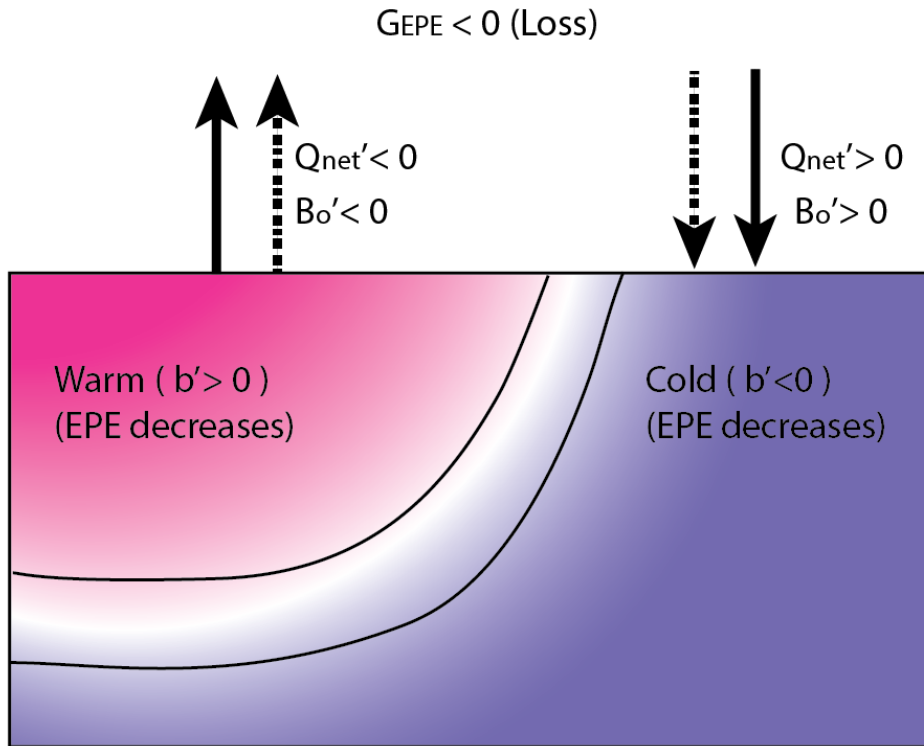
245 values of b' and hence the mixed-layer EPE (assuming temperature anomalies and buoyancy
 246 anomalies are of the same sign, discussed further in section 5).

251 The air-sea buoyancy flux, B_o , may be parameterized as proportional to heat and freshwater
 252 fluxes (Cronin and Sprintall 2001):

$$B_o = \frac{\alpha_\theta g}{\rho_o C_p} Q_{net} - \beta_s g S_o (E - P), \quad (9)$$

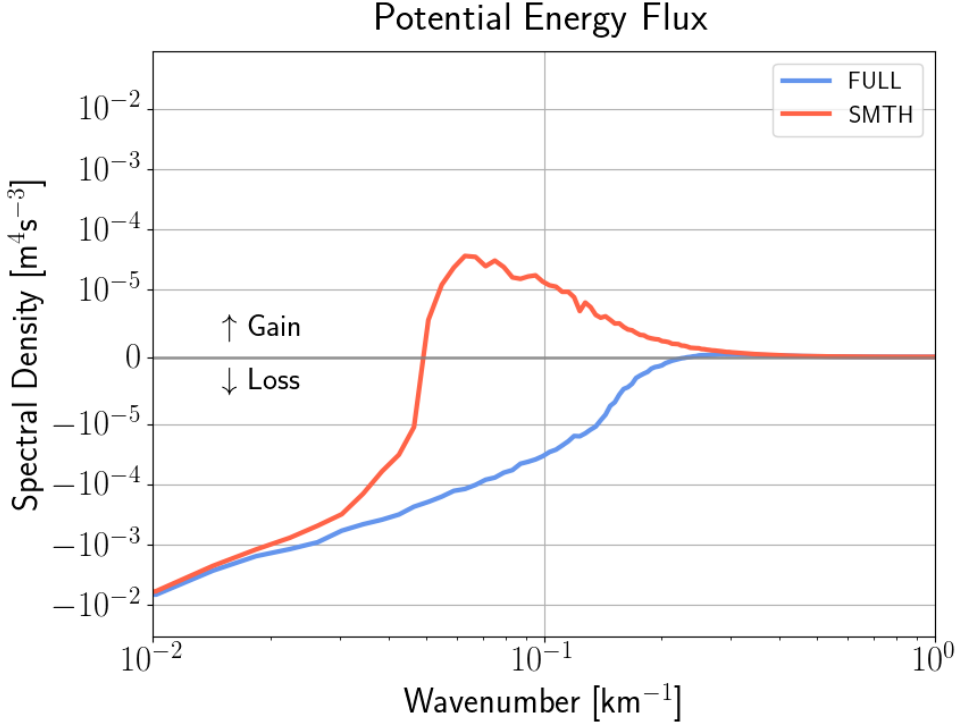
253 where g is gravity, S_o is the surface salinity, C_p is the specific heat of water, Q_{net} is the net surface
 254 heat flux, E is evaporation and P is precipitation. α_θ and β_s represent the thermal expansion and
 255 salinity contraction coefficients calculated at each point. This parameterization allows for the
 256 computation of G_{EPE} .

257 The spectra of G_{EPE} are shown in Fig. 6. The loss or gain of EPE in the ocean is represented as
 258 negative and positive spectral density values, respectively. The co-spectra of the surface buoyancy
 259 and buoyancy flux (G_{EPE}) show a difference in EPE flux exceeding an order of magnitude between
 260 the SMTH and FULL experiments. The FULL experiment spectrum shows loss of EPE in the
 261 submesoscale and lower mesoscale spatial range, which indicates that submesoscale EPE flux is
 262 working as a sink of energy to the atmosphere, similar to mesoscale SST anomalies (Storch et al.
 263 2012; Bishop et al. 2020; Guo et al. 2022; Renault et al. 2023). Conversely, the SMTH experiment



247 FIG. 5. Schematic representation of the surface flux of EPE driven by SST anomalies and heat flux in a
 248 submesoscale front. Heat flux counteracts the SST anomalies, resulting in a decrease in buoyancy anomaly on
 249 both sides of the front and an overall loss of EPE. The prime symbol represents the spatial anomalies due to the
 250 front.

264 spectrum indicates a gain of EPE flux at the submesoscale and a smaller loss at the small mesoscale
 265 range compared to the FULL experiment. The gain of EPE, $G_{EPE} > 0$, likely occurs in the SMTH
 266 simulation because the heat flux is not necessarily driven by SST anomalies and thus may act to
 267 enhance the buoyancy anomalies' magnitude, increasing the EPE. These differences in EPE flux
 268 between the numerical experiments indicate that submesoscale air-sea fluxes are acting as a sink
 269 of submesoscale EPE.



270 FIG. 6. Submesoscale buoyancy anomalies are correlated with buoyancy flux anomalies, driving a loss of
 271 EPE to the atmosphere. Two-dimensional spectra of available potential energy flux for the FULL (blue line)
 272 and SMTH (orange line) experiments. The spectra are averaged over time period of the simulations. Positive
 273 (negative) values represent the gain (loss) of EPE to the atmosphere.

274 *c. Decomposition of eddy potential energy flux and approximations*

275 The G_{EPE} term is further expanded to assess the importance of each component contributing
 276 to b and B_o anomalies. It is possible to approximate surface buoyancy into a linear equation that
 277 takes into account SST and salinity anomalies and surface values of α_θ and β_S . The linearized
 278 surface buoyancy in spectral space is:

$$\widehat{b}_o \approx g \left[\widehat{\alpha_\theta T_o} - \widehat{\beta_S S_o} \right]. \quad (10)$$

279 where T_o is the surface temperature.

280 Using (10) and (9), the EPE flux can be divided into components driven by thermal and salinity
 281 anomalies and fluxes. This decomposition allows for the assessment of the relative contributions

282 of surface temperature and salinity anomalies and fluxes of heat and freshwater in G_{EPE} . The
 283 expansion can be written as

$$G_{EPE} \approx \frac{1}{N_f^2} \mathbb{R} \left[\widehat{b_T^* B_{oT}} + \widehat{b_T^* B_{oS}} + \widehat{b_S^* B_{oS}} + \widehat{b_S^* B_{oT}} \right], \quad (11)$$

284 where the components of buoyancy and buoyancy flux are defined as follows:

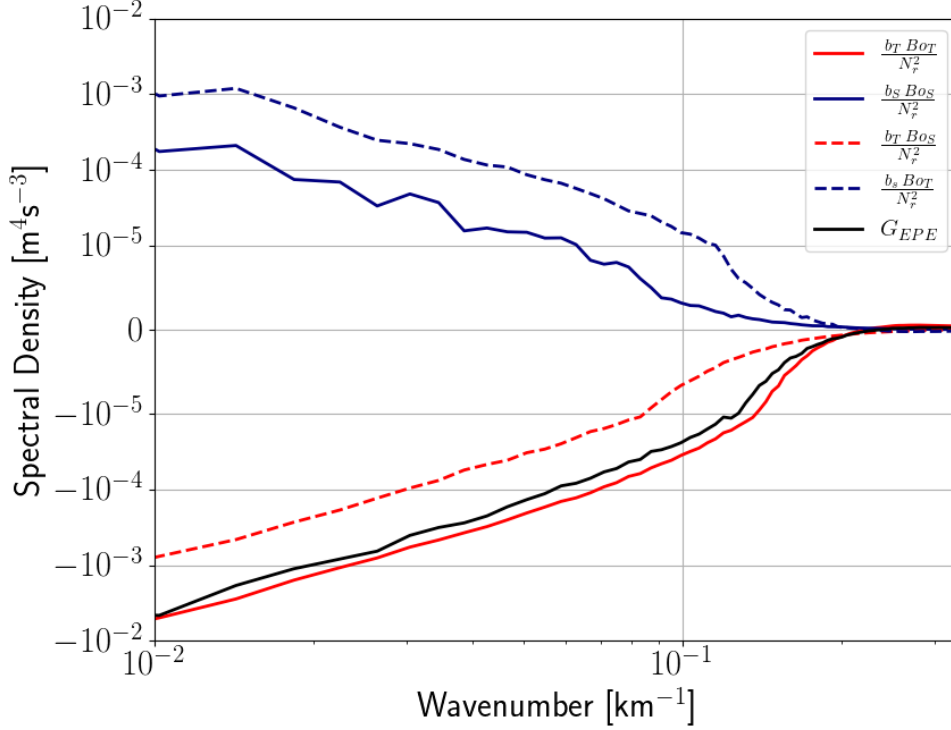
$$\widehat{b_T} = g \widehat{\alpha_\theta T_o}, \quad (12)$$

$$\widehat{b_S} = -g \widehat{\beta_S S_o}, \quad (13)$$

$$\widehat{B_{oT}} = \frac{g}{\rho_o C_p} \widehat{\alpha_\theta Q_{net}}, \quad (14)$$

$$\widehat{B_{oS}} = -g \widehat{\beta_S (E - P) S_o}, \quad (15)$$

288 The total EPE flux thus consists of components from correlations between surface temperature
 289 anomalies and heat fluxes, surface salinity anomalies and freshwater fluxes, and cross-terms
 290 (eg. correlations between surface heat fluxes and salinity anomalies). The spectra for the four
 291 components for the FULL experiments are shown in Fig. 7. G_{EPE} components correlated with
 292 temperature anomalies (i.e., b_T , Fig. 7 - red lines) indicate a net loss of EPE to the atmosphere,
 293 whereas the components generated by salinity anomalies (i.e., b_S , Fig. 7 - blue lines) show a
 294 net gain of EPE. The product of the thermal components of buoyancy and buoyancy flux (i.e.,
 295 $b_T B_{oT}$, Fig. 7 - red solid line) is the dominant component of EPE flux to the atmosphere at the
 296 submesoscale and is responsible for the net loss of EPE shown in Fig. 6. The term that correlates
 297 the salinity component of buoyancy and buoyancy fluxes (i.e., $b_S B_{oS}$, Fig. 7 - blue solid line) has
 298 the smallest magnitude at the submesoscale, indicating that B_{oS} (proportional to freshwater fluxes)
 299 is not as efficient as B_{oT} (proportional to heat fluxes) in fluxing EPE in this region. Instead, the
 300 component that contributes to the largest gain of EPE in the analysis is the cross-term $b_S B_{oT}$ (Fig.
 301 7 - dashed blue line). Temperature and salinity anomalies drive inverse changes in the EPE of
 302 the upper ocean, which as shown below results from the partial density compensation of fronts
 303 in the California Current Region (Rudnick and Ferrari 1999; Mauzole et al. 2020). B_{oT} may also
 304 be further approximated to the latent and sensible components of heat flux anomalies since those



307 FIG. 7. Correlation between heat flux and surface buoyancy anomalies has the greatest contribution to the
 308 EPE flux. Decomposition of the total EPE flux (FULL) in terms of the contributions of temperature and salinity
 309 components (see section 4c). Blue lines represent the components of G_{EPE} proportional to salinity anomalies
 310 (b_S). Red lines represent the components proportional to temperature anomalies (b_T). The solid blue and red
 311 lines represent the component proportional to temperature and heat flux anomalies (B_{oT}) and to salinity and
 312 freshwater flux anomalies (B_{oS}). Dashed blue and red lines represent the cross-term components proportional to
 313 temperature and freshwater flux anomalies and to salinity and heat flux anomalies. The black solid line represents
 314 the total EPE flux.

305 are the components correlated to surface buoyancy anomalies. This approximation is useful for
 306 scaling the EPE flux mechanism and is explored in the next section.

315 This analysis suggests the EPE flux in these simulations is well approximated by

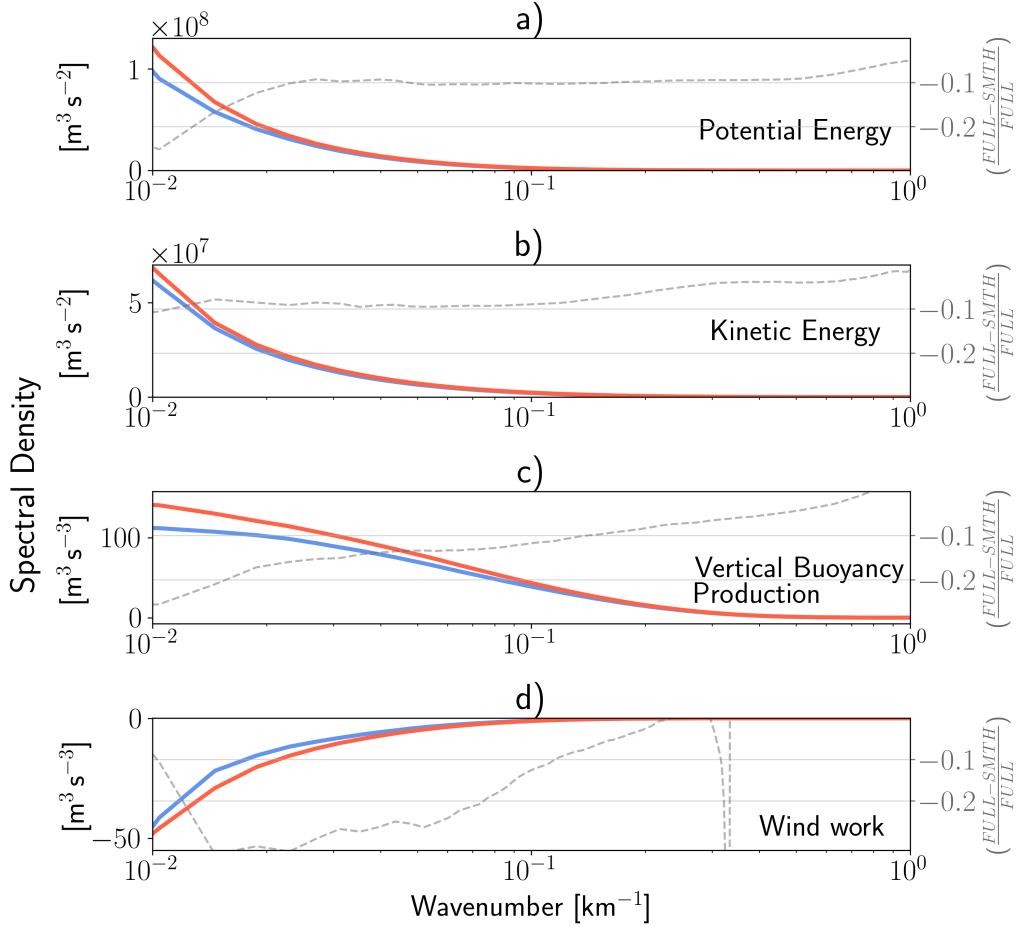
$$G_{EPE} \approx \frac{1}{N_f^2} \mathbb{R} \left[\widehat{b_o}^* \widehat{B_{oT}} \right]. \quad (16)$$

316 In the California Current system, the partial T/S compensation means that while the thermal
317 component of the buoyancy flux drives a loss of EPE through the temperature anomalies (b_T),
318 there is also a partially compensating gain of EPE through the correlation of heat flux anomalies
319 and salinity anomalies (b_S). How the EPE flux depends on density compensation more generally
320 is discussed further in section 5 below.

321 *d. Eddy energy reservoirs, conversion rates and fluxes*

322 Changes in the EPE flux have impacts on the EPE reservoir, but also can affect the EKE through
323 the energy conversion terms. The spectra shown in Fig. 4 suggest that, at the surface, the reservoirs
324 of EPE and EKE are both impacted by the response of SST variability in the air-sea energy transfer
325 via EPE flux. Cumulative spectra (or ogives) of the vertically-integrated reservoirs EKE and EPE
326 shown in Fig. 8a,b confirm the same pattern in the mixed-layer integral. The EPE flux drives a
327 sink of EPE to the atmosphere due to SST-induced heat flux anomalies (Fig. 6), which generates a
328 reduction of submesoscale EPE in the mixed-layer of 10 – 25% (Fig. 8a). At the same scales, EKE
329 is also reduced by approximately 10% as seen in Fig. 8b. The EKE reservoir is likely reduced by
330 the smaller rate of eddy energy conversion, namely, vertical buoyancy production ($C_{(EPE,EKE)}$),
331 which decreases significantly (10 – 25%) in the FULL experiment. As mentioned in section
332 2, $C_{(EPE,EKE)}$ may be attributed to mixed-layer instabilities (and other ageostrophic secondary
333 circulations) where available potential energy stored in thermal-wind balanced fronts is extracted
334 and converted into perturbation flows such as eddies (Capet et al. 2008a; Fox-Kemper et al. 2008)
335 again reflecting the weaker submesoscale in FULL vs SMTH (Fig. 4). This comparative analysis
336 indicates that at the submesoscale, G_{EPE} directly reduces the EPE which induces a lower baroclinic
337 conversion rate ($C_{EKE,EPE}$) and, consequently, a decrease in the EKE reservoir in the mixed layer.

344 Fig. 8d also depicts the cumulative difference in surface EKE flux (G_{EKE}) between the two
345 models. Loss of EKE is present in both experiments at the submesoscale since the CFB effect
346 is accounted for in the wind stress parameterizations. At the submesoscale, there is a relative
347 decrease in wind work in the FULL experiment of 15-30%, a reduction of the EKE flux driven
348 by SST variability. The ratio between the two wind work spectra shown is approximately one or
349 greater than one in scales smaller than the effective resolution of the simulation and hence not
350 considered in this analysis. Scalings of the CFB mechanism on the wind work indicate a direct



338 FIG. 8. Cumulative spectra of vertically integrated parameters depict the impact of SST variability in air-sea
 339 coupling. Blue (Orange) lines represent the FULL (SMTH) experiment spectra. Ogive graphs are integrated
 340 from the larger to smaller horizontal wavenumber. The panels represent (a) Potential Energy, (b) Kinetic Energy,
 341 (c) vertical buoyancy production, and (d) wind work. Grey dashed lines indicate the relative difference between
 342 the spectra for both experiments. EKE and EPE, and vertical buoyancy production were integrated from 50 m
 343 depth to surface, the averaged mixed-layer depth for the region.

351 relationship between EKE flux and EKE reservoir in the upper ocean (Renault et al. 2017), which is
 352 consistent with the decrease of wind work observed in the less energetic FULL experiment (section
 353 5). Concurrently, the TFB mechanism may induce wind anomalies that are partly correlated with
 354 surface currents and hence decrease the net loss of EKE by wind work at the submesoscale (Renault
 355 et al. 2018; Bai et al. 2023; Conejero et al. 2024). This suggests that the more negative wind work

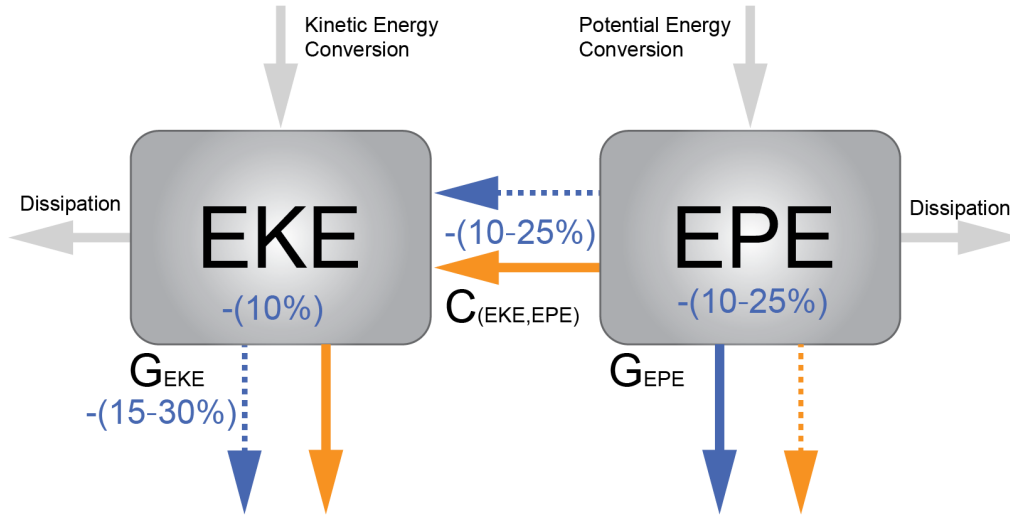
356 in SMTH experiment is likely due to a combination of the artificial suppression of TFB and the
357 increase of surface EKE due indirectly to the suppressed EPE flux.

358 A simplified Lorenz diagram summarizing the relative differences in energetics between the two
359 experiments is depicted in Fig. 9. The vertically-integrated energy fluxes, conversion rates, and
360 reservoirs of both experiments indicate that there is a loss of submesoscale eddy energy in the
361 upper ocean due to correlations between surface buoyancy anomalies and buoyancy fluxes (Fig.
362 5). This reduction of EPE then decreases the EKE indirectly through a reduction in the conversion
363 of EPE to EKE by vertical buoyancy production. Finally, the reduced EKE is associated with a
364 reduction of CFB wind work, which acts at a rate proportional to the EKE (see section 5). While
365 the magnitude of these changes in the experiments utilized here are relatively small, $O(10\%)$, they
366 are similar to changes in the energetics caused by the CFB mechanism found in prior work in this
367 region (Renault et al. 2018). We discuss the relative importance of these two mechanisms and the
368 role of temperature and salinity variability and compensation in the following section.

377 5. Discussion

378 In this work, the EPE flux mechanism is described at the submesoscale where it facilitates the
379 transfer of EPE between the ocean and the atmosphere via correlations between surface buoyancy
380 and buoyancy flux. Such mechanism is previously observed using mesoscale-resolving numerical
381 simulations as described in Ma et al. (2016); Bishop et al. (2017); Guo et al. (2022) and Renault
382 et al. (2023), which affects the energy pathways related to conversion rates and reservoirs of eddy
383 energy. The mechanism described in this work highlights the importance of submesoscale SST
384 variability in driving air-sea fluxes at the same scales and how that may affect the estimation of
385 energy conversion rates, sinks, reservoirs when using numerical simulations. In this section, the
386 limitations of reproducing the EPE flux in numerical models, and the importance of this mechanism
387 relative to other air-sea feedbacks, are discussed.

388 A hierarchy of coupling parameterizations is used in numerical models in order to reproduce the
389 air-sea fluxes, however some of the strategies may underestimate or even fail to generate surface
390 EPE fluxes. Coupled numerical simulations that use a responsive atmosphere and bulk formulae to
391 reproduce air-sea fluxes that rely on similarity theory (Monin and Obukhov 1954) can reproduce
392 the mechanism studied in this work (e.g., the FULL simulation). Uncoupled models that use a



369 FIG. 9. Sea surface temperature variability at the submesoscale alters the pathways and reservoirs of eddy
 370 energy. Simplified Lorenz diagram of the differences in eddy energy reservoirs, fluxes, and conversions. The
 371 FULL (SMTH) experiment is illustrated in blue (orange). Differences between the experiments in each component
 372 are depicted in terms of the FULL spectra decrease. Fluxes of eddy potential energy (i.e., G_{EPE}) and eddy
 373 kinetic energy (i.e., G_{EKE}) are represented by the downward arrows. The reservoirs of eddy potential energy
 374 (i.e., EPE) and eddy kinetic energy (i.e., EKE) are represented by the gray boxes. The conversion of EKE to EPE
 375 (i.e., $C_{EPE,EKE}$) is represented by the horizontal arrows. Grey arrows represent the cross-scale conversions and
 376 dissipation components of energy that are not the focus of this work.

393 fixed atmosphere, but calculate buoyancy fluxes using parameterizations that depend on SST will
 394 likewise generate EPE fluxes, however, it is possible that this flux may not be entirely accurate as
 395 the atmosphere cannot evolve in response to these fluxes. However, uncoupled models that use
 396 prescribed heat fluxes (a common approach for regional ocean or idealized numerical simulations)
 397 fail to generate the mechanism since surface buoyancy fluxes will not respond to surface buoyancy
 398 anomalies. In this case, it is anticipated that the modeled submesoscale will be overly energetic
 399 (section 4).

400 One of the approximation strategies for air-sea fluxes used in uncoupled ocean-only models relies
 401 on the linearization of parameters, such as heat flux, into climatological (background) and local
 402 anomalies (perturbation) components. The climatological components in the heat flux can then be
 403 prescribed based on available data or reanalyses, whereas the heat flux anomalies are parameterized

404 as proportional to modeled surface temperature anomalies (Barnier et al. 1995; Ma et al. 2016;
 405 Moreton et al. 2021). This linearization is particularly amenable to simple implementation in
 406 ocean-only models and may provide a simpler diagnosis of the impact of SST anomalies in the
 407 EPE flux. Here, approximations of the heat flux anomaly as a function of SST are obtained in this
 408 region at the submesoscale. This linearization of the heat flux anomaly as proportional to the SST
 409 anomaly then allows for a further approximation of EPE flux mechanism, described below.

410 Fig. 10 shows the joint probability distribution for SST and surface heat flux anomalies over the
 411 simulation period. For the California Current region, the SST anomalies are mostly correlated to
 412 latent and sensible heat flux anomalies at the submesoscale, that is, over 50% of the variance of
 413 the portion of the buoyancy flux correlated to buoyancy anomalies is explained by the heat flux
 414 components. This allows for the approximation

$$Q'_{net} \approx -(Q'_{SH} + Q'_{LH}), \quad (17)$$

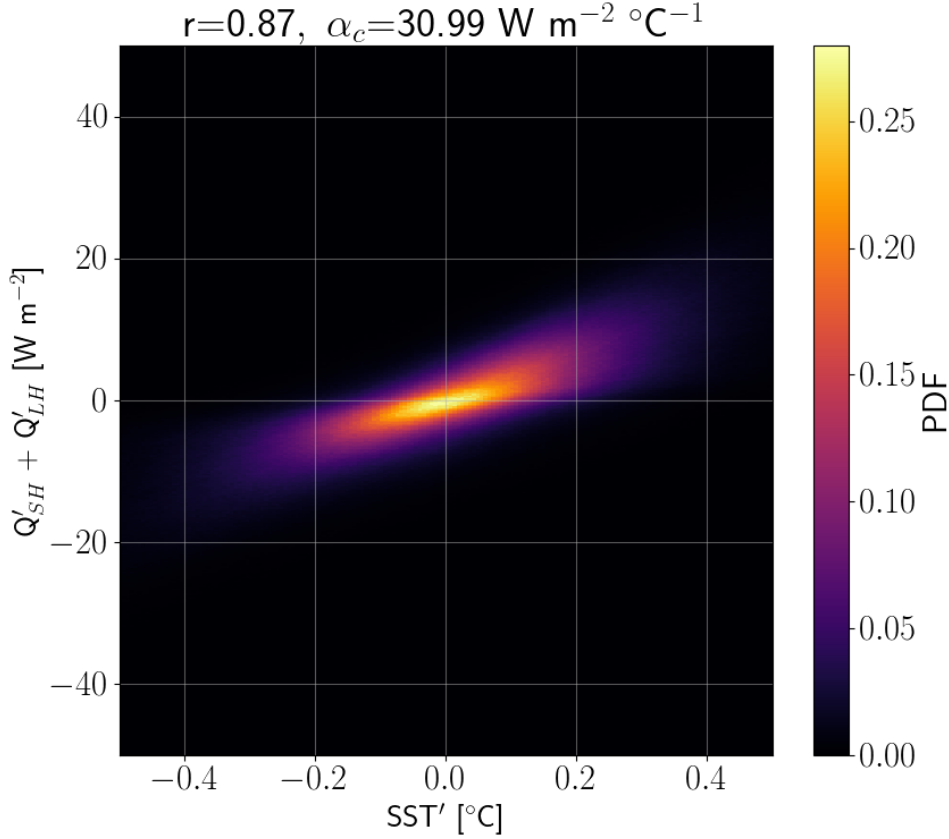
415 where Q'_{SH} and Q'_{LH} are the sensible and latent heat flux components, respectively. The spatial
 416 anomalies (') obtained in this analysis are computed from the subtraction of a spatial low-pass
 417 filter, similar to what is applied in the SST field as described in section 3c, to the variable. The
 418 coupling coefficient α_c is computed as the linear regression fit slope from the approximated heat
 419 flux (17) and SST spatial anomalies, as shown in Fig. 10. In this work, $\alpha_c=31 \text{ W m}^2 \text{ }^\circ\text{C}^{-1}$, which
 420 is similar to previous linearizations for the same region at larger scales (Barnier et al. 1995).

425 As analyzed in section 4b, the correlation between heat flux and surface buoyancy anomalies has
 426 the greatest contribution to the submesoscale EPE flux. By invoking the approximation of EPE
 427 flux in physical space (Storch et al. 2012) and the linearization of the heat flux obtained in this
 428 work (17), an approximate form of the EPE flux is given by

$$G_{EPE} = -\frac{\alpha_\theta \alpha_c g}{N_r^2 \rho_o C_p} b'_o T'_o. \quad (18)$$

429 This approximation describes EPE flux as the product of surface buoyancy, and can also be further
 430 manipulated by approximating buoyancy by the linear equation of state giving

$$G_{EPE} \approx -\frac{1}{N_r^2} \frac{\alpha_c \alpha_\theta^2 g^2}{\rho_o C_p} \left(1 - \frac{1}{R}\right) T_o'^2, \quad (19)$$



421 FIG. 10. Two-dimensional histogram of SST and sensible and latent components of the heat flux (Q_{SH+LH})
 422 anomalies in the FULL simulation setup. The impact of SST anomalies in the EPE flux at the submesoscale may
 423 be linearized using a coupling coefficient derived from anomalies of SST and non-solar heatflux – proportional
 424 to surface buoyancy flux.

431 where R is the surface density ratio defined as:

$$R = \frac{\alpha_\theta T'_o}{\beta_S S'_o}. \quad (20)$$

432 This ratio, R , indicates how temperature and salinity anomalies contribute to the decrease or
 433 increase of buoyancy simultaneously. When $R < 0$, the contribution of temperature and salinity
 434 anomalies to modulate buoyancy are positively correlated. This scenario favors loss of EPE to the
 435 atmosphere as heat flux tends to dissipate buoyancy anomalies (as in Fig. 5), and indeed (21) is
 436 strictly negative for $R < 0$. When $R > 0$, the contribution of temperature and salinity anomaly in

437 buoyancy anomalies are negatively correlated, that is, density compensation occurs (Rudnick and
 438 Ferrari 1999). Observations suggest some degree of density compensation is ubiquitous in regions
 439 with active submesoscales (Rudnick and Martin 2002; Drushka et al. 2019). In compensated fronts,
 440 the sign of G_{EPE} is dependent on the relative magnitude of the thermal and salinity components
 441 of buoyancy in R . If temperature anomalies determine buoyancy anomalies ($R > 1$) then there is
 442 a loss of EPE (as in the simulations here where the median value of the ratio $(1 - R^{-1})^{-1} \approx 0.5$).
 443 Conversely, if the salinity component of buoyancy dominates in a compensated front ($0 < R < 1$),
 444 such that surface buoyancy fluxes act to increase the density anomalies across the front (ie. the
 445 dense side of the front is associated with warm anomalies that are cooled by surface heat fluxes), the
 446 EPE will increase due to surface fluxes. This suggests that in some regimes, such as high-latitude
 447 β oceans or coastal regions with significant freshwater fluxes, the EPE flux may act as a source of
 448 submesoscale energy.

449 Finally, we note it is also possible to describe G_{EPE} as proportional to the surface EPE reservoir
 450 (detailed derivation in Appendix B)

$$G_{EPE} \approx \frac{1}{(1 - \frac{1}{R})} \frac{2s_b}{\rho_o} EPE_o, \quad (21)$$

451 where $s_b = -\frac{\alpha_c}{C_p}$ [$\text{kg m}^{-2} \text{s}^{-1}$] is the EPE flux coupling coefficient, and EPE_o is the surface
 452 EPE. This form is useful for comparison with the CFB EKE flux, which is proportional to EKE
 453 (Renault et al. 2017). The ratio between the two mechanisms can therefore be scaled as

$$\frac{G_{EPE}}{G_{EKE}} \sim \left(\frac{s_b}{s_\tau}\right) \frac{1}{(1 - \frac{1}{R})} \frac{EPE_o}{EKE_o}, \quad (22)$$

454 where $s_\tau = -3/2\rho_a C_D |U_a|$ is the wind stress coupling coefficient, C_D is the drag coefficient
 455 and $|U_a|$ is the surface wind magnitude. This ratio indicates that the relative impact between
 456 the two mechanisms is a function of: (i) the magnitude of both coupling coefficients, (ii) the
 457 surface density ratio, and (iii) the ratio of the surface eddy energy reservoirs of the system. The
 458 coupling coefficients s_b and s_τ are of similar magnitude considering previous estimates of s_τ using
 459 observations (Renault et al. 2017) and of α_c from computations in this work (Fig. 10).

460 The ratio of EKE_o and EPE_o is scale- and season-dependent due to mesoscale and submesoscale
 461 dynamics. For instance, EKE and EPE spectra of Western Boundary Currents such as the Gulf

462 Stream show that strong baroclinic currents have EKE_o and EPE_o reservoirs of similar magnitude
463 for the winter, whereas in the summer, EKE is larger (Callies et al. 2015). These differences are
464 in part related to mixed-layer instabilities amplified in the wintertime as the mixed-layer depth
465 increases (Fox-Kemper et al. 2008). Observations from the eastern subtropical North Pacific also
466 show EKE and EPE magnitudes to be similar at the mesoscale and submesoscale (Callies and
467 Ferrari 2013). Thus, (22) suggests that the results found here – where the direct EPE flux alters
468 submesoscale energetics in a manner that is quantitatively similar to the surface EKE flux – may be
469 found elsewhere when the EPE_o/EKE_o ratio is large or there is substantial density compensation.

470 **6. Summary and conclusion**

471 In this manuscript, the impact of submesoscale SST variability on the flux of EPE is assessed
472 using two configurations of a fully-coupled model with submesoscale-permitting resolution in the
473 ocean, where one of the numerical experiments (SMTH) suppresses submesoscale SST anomalies
474 in the computation of air-sea fluxes. Comparative analysis between the experiments indicates that
475 modifications to the surface buoyancy flux induced by submesoscale SST variability generate an
476 EPE flux at the air-sea interface which acts as a sink of eddy energy in the upper ocean. In these
477 simulations, this leads to a reduction of the EPE reservoir of 10 – 25% at the submesoscale and
478 the small mesoscale. Associated with this, the rate of conversion to EKE by the vertical buoyancy
479 production ($C_{(EPE,EKE)}$) also decreases by 10 – 25%. This in turn leads to an approximately 10%
480 reduction of submesoscale EKE, and consequently a change in the surface wind work (i.e., CFB;
481 Renault et al. 2018) of 15 – 30%. These changes to submesoscale energy are similar in magnitude
482 to those induced in the same region by the CFB, as well as at larger scales globally (Renault et al.
483 2018; Bishop et al. 2020).

484 Linearizations of the turbulent heat flux as a function of SST perturbations at the submesoscale
485 (coupling coefficient α_c) allow for the scaling of the EPE flux at the submesoscale in terms of
486 surface buoyancy and temperature anomalies. The EPE flux may then be described as a function of
487 the surface EPE, analogous to scaling arguments for EKE flux being proportional to EKE reservoir
488 (Renault et al. 2017, 2018), with relative magnitude also dependent on the degree of density
489 compensation (Rudnick and Ferrari 1999). A ratio between the EPE and EKE fluxes results in
490 a term proportional to the ratio between the eddy energy reservoirs, suggesting that the relative

491 importance of the EPE flux and CFB mechanisms in reducing eddy energy will be dependent on the
 492 relative sizes of the surface EPE and EKE. In this work considering the California Current region,
 493 the EPE flux is a sink of surface EPE at the same magnitude of the CFB mechanism for surface
 494 EKE, despite the counteracting effect of the partial salinity compensation found in this region (eg.,
 495 Fig. 7). In regions where salinity dominates in the density compensation (e.g., $0 < R < 1$ as found
 496 at high latitudes or regions with strong freshwater influence), EPE flux may contribute to a gain of
 497 EPE hence energizing the submesoscale.

498 Finally, we note that changes between simulations at scales larger than the SST filter scale were
 499 also observed in these experiments, which could indicate a change in the upscale flux of energy from
 500 the submesoscale to the mesoscale. However, the limited domain size and integration time period
 501 of the numerical model considered here do not allow a robust characterization of changes at larger
 502 scales. Looking forward, similar submesoscale-resolving experiments in a larger domain, with
 503 varying scales of SST filtering, would provide useful insight into both the direct and cross-scale
 504 effects of the EPE flux.

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 510 TOSCA-CNES projects.

511 *Data availability statement.* No observational data was used in this work. Analysis scripts will
 512 be made available via <http://github.com> upon publication.

513 APPENDIX A

514 Eddy Potential Energy Equation in Spectral Space

515 Consider the buoyancy equation (neglecting horizontal mixing)

$$\frac{\partial b}{\partial t} = -\mathbf{u}_h \cdot \nabla_h b - wN^2 + \frac{\partial}{\partial z} \kappa \frac{\partial b}{\partial z}, \quad (\text{A1})$$

516 where $\mathbf{u}_h=(u,v)$ is the horizontal velocity vector and ∇_h is the horizontal gradient operator. De-
 517 composing the equation in spectral space in respect to the horizontal wavenumber, (A1) becomes

$$\widehat{\frac{\partial b}{\partial t}} = -\widehat{\mathbf{u}_h \cdot \nabla_h b} - \widehat{wN^2} + \frac{\partial}{\partial z} \widehat{\kappa \frac{\partial b}{\partial z}}. \quad (\text{A2})$$

518 Multiplying (A2) by the complex conjugate of buoyancy, \widehat{b}^* , and dividing by a spatially and
 519 temporally averaged reference buoyancy frequency, N_r^2 , yields

$$\frac{\widehat{b}^* \widehat{\frac{\partial b}{\partial t}}}{N_r^2} = -\frac{\widehat{b}^* \widehat{\mathbf{u}_h \cdot \nabla_h b}}{N_r^2} - \widehat{w} \widehat{b}^* + \frac{\widehat{b}^* \partial}{N_r^2 \partial z} \widehat{\kappa \frac{\partial b}{\partial z}}. \quad (\text{A3})$$

520 where the assumption that horizontal variations of N^2 can be neglected in the calculation of the
 521 vertical buoyancy production yielding

$$\frac{\widehat{b}^* \widehat{wN^2}}{N_r^2} \approx \widehat{w} \widehat{b}^*. \quad (\text{A4})$$

522 This assumption was tested using the numerical simulation and major results were found to be
 523 robust to this approximation (Storch et al. 2012). The potential energy equation is then found by
 524 taking the real component of (A3) giving

$$\frac{\partial}{\partial t} EPE = \mathbb{R} \left[-\frac{\widehat{b}^* \widehat{\mathbf{u}_h \cdot \nabla_h b}}{N_r^2} - \widehat{w} \widehat{b}^* + \frac{\widehat{b}^* \partial}{N_r^2 \partial z} \widehat{\kappa \frac{\partial b}{\partial z}} \right]. \quad (\text{A5})$$

525 For simplicity, two different terms are obtained by further manipulating the last term of the (A5)
 526 as follows:

$$\frac{\partial}{\partial t} EPE = \mathbb{R} \left[-\frac{\widehat{b}^* \widehat{\mathbf{u}_h \cdot \nabla_h b}}{N_r^2} - \widehat{w} \widehat{b}^* - \frac{1}{N_r^2} \frac{\partial \widehat{b}^*}{\partial z} \widehat{\kappa \frac{\partial b}{\partial z}} + \frac{\partial}{\partial z} \left(\frac{\widehat{b}^* \widehat{\kappa \frac{\partial b}{\partial z}}}{N_r^2} \right) \right]. \quad (\text{A6})$$

527 The two last terms in A6 represent the dissipation rate and diffusive vertical transport of EPE,
 528 respectively. The mixed-layer balance of potential energy is obtained by integrating (A6) from the
 529 surface to the mixed layer depth z_m and using the following surface boundary condition:

$$\widehat{\kappa \frac{\partial b}{\partial z}} \Big|_{z=0} = \widehat{B}_o. \quad (\text{A7})$$

530 Finally,

$$\int_{z_m}^0 \frac{\partial}{\partial t} EPE dz = \mathbb{R} \left\{ \int_{z_m}^0 \left(-\frac{\widehat{b}^*}{N_r^2} \widehat{\mathbf{u}_h \cdot \nabla_h b} - \widehat{w} \widehat{b}^* - \frac{1}{N_r^2} \frac{\partial \widehat{b}^*}{\partial z} \kappa \frac{\partial \widehat{b}}{\partial z} \right) dz + \frac{\widehat{b}_o^* \widehat{B}_o}{N_r^2} - \frac{\widehat{b}^*}{N_r^2} \kappa \frac{\partial \widehat{b}}{\partial z} \Big|_{z_m} \right\}. \quad (\text{A8})$$

531

APPENDIX B

532

Approximated form of the EPE flux at the submesoscale

533 The EPE flux is defined as the product of the surface buoyancy and the buoyancy flux anomalies
534 as follows (Bishop et al. 2020; Storch et al. 2012):

$$G_{EPE} = \frac{b'_o B'_o}{N_r^2}, \quad (\text{B1})$$

535 where b'_o and B'_o are defined respectively as:

$$b'_o = \alpha_\theta g T'_o - \beta_s g S'_o, \quad (\text{B2})$$

$$B'_o = \frac{\alpha_\theta g}{\rho_o C_p} Q'_{net} - \beta_s g S'_o [E' - P'], \quad (\text{B3})$$

536 where the prime symbol (') denotes the anomaly of a given variable. It convenient to describe
537 surface buoyancy perturbations in terms of temperature as follows:

$$b'_o = \alpha_\theta g T'_o \left(1 - \frac{1}{R}\right), \quad (\text{B4})$$

538 where R is the density ratio (defined in (20) and see also Rudnick and Ferrari 1999). Since at
539 the submesoscale the EPE flux from ocean to atmosphere primarily generated by surface heat flux
540 anomalies (Fig. 7), we can combine (B4) and (B3) in (B1) to yield

$$G_{EPE} \approx \frac{1}{N_r^2} \frac{\alpha_\theta^2 g^2}{\rho_o C_p} \left(1 - \frac{1}{R}\right) T'_o Q'_{net}. \quad (\text{B5})$$

541 As described in section 5, it is further possible to approximate the correlated component of the
 542 heat flux anomaly in terms of a coupling coefficient. Q'_{net} may then be described as

$$Q'_{net} \approx -\alpha_C T'_o \quad (\text{B6})$$

543 Thus G_{EPE} is approximately:

$$G_{EPE} \approx -\frac{1}{N_r^2} \frac{\alpha_C \alpha_\theta^2 g^2}{\rho_o C_p} \left(1 - \frac{1}{R}\right) T_o'^2. \quad (\text{B7})$$

544 Using the definition of eddy potential energy (EPE) in terms of density ratio:

$$EPE_o = \frac{1}{2} \frac{\alpha_\theta^2 g^2}{N_r^2} \left(1 - \frac{1}{R}\right)^2 T_o'^2 \quad (\text{B8})$$

545 and multiplying the term $(1 - \frac{1}{R})$ in the numerator and denominator of (B7), the equation can be
 546 manipulated further in terms of $b_o'^2$. Thus, (B7) becomes:

$$G_{EPE} \approx \frac{1}{\left(1 - \frac{1}{R}\right)} \frac{2s_b}{\rho_o} EPE_o, \quad (\text{B9})$$

547 where $s_b = -\alpha_C / C_p$ $\text{kg m}^{-2} \text{s}^{-1}$.

548 This takes a form similar to the current feedback effect on the wind work, which can be expressed
 549 as (Renault et al. 2017):

$$G_{EKE} \approx \frac{2s_\tau}{\rho_o} EKE_o, \quad (\text{B10})$$

550 where $s_\tau \approx -3/2 \rho_a C_D |U_a|$ [$\text{kg m}^{-2} \text{s}^{-1}$]. Notably, both G_{EPE} and G_{EKE} can thus be seen to
 551 act as linear damping terms in the potential and kinetic energy equations, respectively. The ratio
 552 between EPE and EKE flux at the submesoscale is

$$\frac{G_{EPE}}{G_{EKE}} \sim \frac{2\alpha_C}{3\rho_a C_D C_p |U_a|} \frac{1}{(1 - \frac{1}{R})} \frac{EPE_o}{EKE_o}. \quad (\text{B11})$$

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