

1 **Shallow storage conditions at Krafla IDDP-1 revealed by rhyolite-MELTS**
2 **geobarometry, and implications for global shallow magmatism**

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7 **Abstract**

8 Identifying the storage depths of melt-dominated magma bodies prior to eruption is critical for
9 understanding magma transport, eruption hazards, and magma body longevity. Rhyolite-MELTS
10 has been used effectively to calculate pre-eruptive storage pressures for silicic magma bodies in
11 the upper crust (~100-350 MPa), but its precision and accuracy in very low-pressure systems
12 (<100 MPa) has not been sufficiently investigated. During the recent Krafla IDDP-1 drilling
13 project, magma was surprisingly intersected at 2.1 km depth. Here, we test the use of rhyolite-
14 MELTS geobarometry for this very low-pressure system, using natural Krafla IDDP-1
15 compositions that were stored at a known depth. We input the composition of the melt (preserved
16 as glass) and search in pressure and temperature space at a range of oxygen fugacity (f_{O_2}) to
17 model the storage conditions of the Krafla magma. For the average composition of the drilled
18 melt, rhyolite-MELTS yields reasonable storage pressures (~40-50 MPa). After converting
19 calculated pressure to depth, the calculated depths are 1.6-1.9 km. These estimates are only 0.2-
20 0.5 km different from that of the intersected magma, showing that rhyolite-MELTS provides
21 excellent estimates for very shallow magma storage, further strengthened by results from a
22 Monte Carlo analysis. The agreement between rhyolite-MELTS pressures and the drilled depth
23 of the Krafla magma supports the previously calculated very shallow storage pressures in other
24 locations, like the Taupō Volcanic Zone (TVZ), Aotearoa New Zealand. This shallowest storage
25 zone of melt-dominated magmas has significant implications for modeling volcanic unrest and
26 evaluating geothermal and economic resource potential.

27 **Keywords**

28 Shallow magma, rhyolite-MELTS, geobarometry, Krafla

29 **Introduction**

30 *Motivation*

31 Rhyolitic magmas are integral to the formation and differentiation of the crust, especially of the
32 upper crust. They often erupt explosively and are responsible for some of the largest eruptions on
33 record (Lowenstern *et al.*, 2006; Self, 2006; Miller and Wark, 2008). Identifying where melt-
34 dominated magma bodies of rhyolite form and are stored in the crust is critical for understanding
35 magma body generation and longevity (Gualda *et al.*, 2012b; Pamukçu *et al.*, 2015a, 2021; Till *et al.*,
36 2015; Bachmann and Huber, 2016; Townsend and Huber, 2020), magma decompression and
37 ascent (Putirka, 2016a; Polacci *et al.*, 2017; Cassidy *et al.*, 2018; Caricchi *et al.*, 2021), and
38 eruption hazards (Tilling, 2008; Aspinall and Blong, 2015; Cassidy and Mani, 2022).

39 Substantial effort and progress have been made to determine the depths of pre-eruptive magma
40 storage via erupted products using a variety of methods, for instance: clinopyroxene and
41 amphibole geobarometry (Blundy and Holland, 1990; Thomas and Ernst, 1990; Schmidt, 1992;
42 Nimis and Ulmer, 1998; Putirka *et al.*, 2003; Putirka, 2008, 2016b; Mutch *et al.*, 2016; Neave
43 and Putirka, 2017; Petrelli *et al.*, 2020; Jorgenson *et al.*, 2022; Wieser *et al.*, 2022, 2023), fluid-
44 inclusion geobarometry (Moore, 2008; Wallace *et al.*, 1999; Anderson *et al.*, 2000), and field
45 work in volcanic-plutonic systems (Hogan and Gilbert, 1995; du Bray and Pallister, 1999; Bachl
46 *et al.*, 2001; Ferguson *et al.*, 2013; Deering *et al.*, 2016; Chen *et al.*, 2021; Wallrich *et al.*, 2023).
47 As well, indirect observations document magma and magma bodies via geophysical methods
48 (Lees, 2007; Lerner *et al.*, 2020; Paulatto *et al.*, 2022).

49 Pre-eruptive magma storage depths of rhyolites are typically considered to be ≥ 3 km, with the
50 depth of 5-10 km commonly considered the dominant depth range of upper crustal magmas that
51 feed eruptions to the surface (Cashman and Giordano, 2014; Cashman *et al.*, 2017; Tramontano
52 *et al.*, 2017; Huber *et al.*, 2019). However, there are also multiple examples of inferred present-
53 day shallow magma storage, at depths < 3 km, for instance: Larderello geothermal field, Italy
54 (Cameli *et al.*, 1993; Manzella *et al.*, 2017; Rochira *et al.*, 2018) and Dabbahu volcano, Afar,
55 Ethiopia (Wright *et al.*, 2006; Ebinger *et al.*, 2008). Several instances where magma has been
56 directly intersected at much shallower depths via drilling document unequivocally the existence
57 of very shallow silicic magma storage. They are: 1) Puna Geothermal field, Hawai'i at 2.5 km in
58 2005 (Teplow *et al.*, 2009); 2) Menengai Caldera, Kenya at 2.1 km and 2.2 km in 2011-2014

59 from well MW-04 and MW-06 (Mbia *et al.*, 2015); and 3) Krafla geothermal system, Iceland at
60 2.6 km in 2008 from well KJ-39 in the Suðurhliðar well field (Mortensen *et al.*, 2010) and at 2.1
61 km in 2009 from the Krafla IDDP-1 well (Elders *et al.*, 2011; Zierenberg *et al.*, 2013).

62 Thermodynamic phase-equilibria modeling via rhyolite-MELTS has been utilized to determine
63 pre-eruptive storage conditions, including depth and f_{O_2} , in predominantly silicic systems
64 (Gualda *et al.*, 2012a; Bégué *et al.*, 2014b; Gualda and Ghiorso, 2014; Pamukçu *et al.*, 2015b;
65 Harmon *et al.*, 2018). Here, we investigate the efficacy of rhyolite-MELTS geobarometry to
66 infer the presence of very shallow magma bodies. Krafla IDDP-1 glass compositions provide a
67 natural laboratory to test the validity and sensitivity of this approach for evaluating shallow
68 storage depths of silicic magmas.

69 *Krafla*

70 The Krafla central volcano is located within the Northern Rift Zone of Iceland. Krafla has
71 erupted predominantly basalt throughout its ~300 ka volcanic history (Thorarinsson, 1979;
72 Sæmundsson, 1991), and there have been at least 8 rhyolitic eruptions (Sæmundsson, 1991;
73 Jonasson, 1994; Rooyackers *et al.*, 2021). The Krafla caldera collapsed at ~110 ka (Calderone *et*
74 *al.*, 1990; Sæmundsson and Pringle, 2000; Rooyackers *et al.*, 2019), and the modern-day
75 geothermal field is located within this caldera structure. The most recent eruption is the Krafla
76 Fires, which continued intermittently from 1975 to 1984, and did not contain rhyolite (Einarsson,
77 1978; Hollingsworth *et al.*, 2012). The most recent rhyolitic eruption, the 1724 C.E. Mývatn
78 Fires rifting episode that formed the Víti maar is compositionally similar to the IDDP-1 glass
79 studied here. These compositions are distinct from previous rhyolitic eruption from Krafla
80 (Rooyackers *et al.*, 2021)

81 There has been substantial geothermal exploration since 1974, which has revealed extensive
82 active geothermal systems and refined the understanding of the subsurface geology (Árnadóttir *et*
83 *al.*, 1998; Árnason *et al.*, 2008; Kennedy *et al.*, 2018, 2018b; Árnason, 2020).

84 In 2009, during the IDDP drilling project, the IDDP-1 well was designed to drill to 4-5 km depth
85 (Frioleifsson *et al.*, 2014), but surprisingly intercepted magma at 2.1 km, despite substantial
86 seismic and subsurface imaging (Elders *et al.*, 2011). Quenched glass fragments were recovered
87 in the drill cuttings and allow observation and analysis of magma intercepted at depth (Elders *et*
88 *al.*, 2011; Zierenberg *et al.*, 2013; Masotta *et al.*, 2018). The magma is rhyolitic (~76.1-77.3 wt%

89 SiO₂) with ~1.8 wt% dissolved volatiles of mixed H₂O-CO₂ (Elders *et al.*, 2011), suggesting
90 saturation pressures of ~40 MPa (Zierenberg *et al.*, 2013). Intersected rhyolite was in contact
91 with a felsite host (Elders *et al.*, 2011; Zierenberg *et al.*, 2013; Masotta *et al.*, 2018). Isotopically,
92 the magma is likely formed from melting hydrothermally altered basalt as opposed to forming
93 via fractional crystallization (Elders *et al.*, 2011; Zierenberg *et al.*, 2013; Kennedy *et al.*, 2018).
94 Previous work on the IDDP-1 drill cuttings includes major-element glass analyses, bubble
95 texture analyses, and experimental studies (Elders *et al.*, 2011; Zierenberg *et al.*, 2013; Masotta
96 *et al.*, 2018; Rooyackers *et al.*, 2021; Saubin *et al.*, 2021). We point the reader to several
97 excellent studies for further information, including: Elders *et al.* (2011), Saubin *et al.* (2021),
98 Zierenberg *et al.* (2013), and Masotta *et al.* (2018).

99 *Rhyolite-MELTS*

100 Rhyolite-MELTS is an internally consistent phase-equilibria thermodynamic model. We utilize
101 rhyolite-MELTS to determine the magmatic conditions at which a melt of a given composition is
102 in equilibrium with a known/observed mineral assemblage (Gualda *et al.*, 2012a; Bégué *et al.*,
103 2014b; Gualda and Ghiorso, 2014; Pamukçu *et al.*, 2015b; Harmon *et al.*, 2018, 2024; Smithies
104 *et al.*, 2023). We use the composition of natural volcanic glass from Krafla as a proxy for the
105 magmatic melt, and we search for the conditions at which melt of the given composition is in
106 equilibrium with an expected mineral assemblage. The pressure calculated by rhyolite-MELTS is
107 the pressure at which the melt was last in equilibrium with the mineral assemblage, which is
108 interpreted to be the pre-eruptive magma storage pressure. This is particularly appropriate for
109 Krafla IDDP-1 magmas, given that they were sampled at depth, without any effects of ascent and
110 eruption.

111 Rhyolite-MELTS geobarometry has been applied to a variety of volcanic systems (Gualda and
112 Ghiorso, 2013a; Bégué *et al.*, 2014b; Gualda *et al.*, 2018; Foley *et al.*, 2020; Pamukçu *et al.*,
113 2021; Pitcher *et al.*, 2021; Seropian *et al.*, 2021; Smithies *et al.*, 2023; Harmon *et al.*, 2024). For
114 the most part, these studies have shown that pre-eruptive storage predominately takes place
115 under pressures of ~100-300 MPa. Nonetheless, work focusing on the Taupō Volcanic Zone
116 (TVZ) has revealed a subset of rhyolite-MELTS pressures of ~50-75 MPa (Bégué *et al.*, 2014b,
117 2014a; Gualda *et al.*, 2018; Pamukçu *et al.*, 2020; Harmon *et al.*, 2024). Because such shallow
118 pressures were not as well constrained as others (see Bégué *et al.*, 2014b), these results – and

119 their potential significance – have not been the focus of any of these studies. However, the
120 existence of these very shallow pre-eruptive magma storage pressures (<100 MPa) hint that 1)
121 there is a very shallow depth of pre-eruptive magma storage that has been relatively unexplored;
122 and 2) rhyolite-MELTS has the potential to resolve very shallow magma storage.

123 The well-studied Krafla system provides an opportunity to test the efficacy and precision of
124 rhyolite-MELTS storage pressures for very shallow magma bodies, given that magma depth is
125 known. Confirmation of rhyolite-MELTS pressures in the Krafla IDDP-1 case study would lend
126 support to very shallow pressures indicated elsewhere in the world, with potentially important
127 implications for our understanding of the architecture of magmatic systems that feed eruptions to
128 the surface.

129 **Materials and Methods**

130 *Krafla Glass Compositions*

131 We use the 31 natural quenched rhyolite glass compositions from Masotta *et al.* (2018) for
132 Krafla IDDP-1 magmas, and we also calculate an average composition based on these 31
133 compositions. These compositions are the rhyolite (“RHL”) compositions from Masotta *et al.*
134 (2018), which are most similar to the Melt 1 compositions reported by Zierenberg *et al.* (2013),
135 interpreted to be generated by partial melting at depth (much deeper than the depth of
136 intersection) of a hydrothermally altered basaltic crust (Zierenberg *et al.*, 2013). We take these
137 RHL compositions from Masotta *et al.* (2018) to be best estimates of Krafla melt compositions,
138 and we use them as input for rhyolite-MELTS calculations (see supplementary data for
139 compositions). The compositions are retrieved from mm-sized glass pieces that rapidly quenched
140 during interaction with the drilling fluids. The observed mineral assemblage in the mm-sized
141 shards is plagioclase + augite + pigeonite + magnetite ± apatite (Zierenberg *et al.*, 2013; Masotta
142 *et al.*, 2018), and the samples have <3% crystals (Zierenberg *et al.*, 2013). Volatile contents
143 measured using FTIR yield average total H₂O content of ~ 1.8 wt% and CO₂ content of ~ 85 ppm
144 (Zierenberg *et al.*, 2013). Directly adjacent to the melt-rich magma body was a partially melted
145 felsite with a mineralogy of plagioclase + alkali feldspar + quartz + augite + magnetite + zircon ±
146 apatite (Elders *et al.*, 2011; Zierenberg *et al.*, 2013).

147 While there has not been a direct estimate of the f_{O_2} for the magma intersected by the IDDP-1
148 drilling, the f_{O_2} conditions of Krafla basalts have been calculated to be at the QFM buffer

149 (Nicholson, 1990) or just above it ($\Delta QFM = +0.6-0.7$) (Shorttle *et al.*, 2015; Hartley *et al.*, 2017;
150 Halldórsson *et al.*, 2018), which is ~ 1 log unit below the NNO buffer ($\Delta NNO = -1$).

151 *Projection onto the Haplogranitic Ternary Diagram*

152 We project all Krafla IDDP-1 compositions onto the quartz-albite-orthoclase haplogranitic
153 ternary using the projection scheme of Blundy and Cashman (2001), which yields a first order
154 estimate of crystallization pressures for melts in equilibrium with quartz and feldspar (Gualda
155 and Ghiorso, 2013b). For magmas in which quartz is absent, such as Krafla magmas, pressures
156 obtained with this projection represent maximum pressures of magma storage.

157 For comparison, we also project all TVZ compositions that produced pressures ≤ 100 MPa from
158 Bégué *et al.* (2014) and Gualda *et al.* (2018) onto the haplogranitic ternary. In this case, quartz is
159 generally observed, which suggests that projection onto the haplogranitic ternary yields best-
160 estimates of storage pressures.

161 *Rhyolite-MELTS Calculations*

162 Rhyolite-MELTS geobarometry yields the pressures at which the input glass composition (the
163 best proxy for the magmatic melt composition) was last in equilibrium with the mineral
164 assemblage. We use MELTS_Excel (Gualda and Ghiorso, 2015), following the methods detailed
165 in Gualda and Ghiorso (2014) for all pressure calculations. The inputs for rhyolite-MELTS are
166 the compositions of the quenched rhyolite glass retrieved from the Krafla IDDP-1 drilling
167 (Masotta *et al.*, 2018). Rhyolite-MELTS is not given any other information a priori, including the
168 mineral phases of interest. We search through pressure, temperature, and f_{O_2} space.

169 The ranges of pressure (200-10 MPa in 10 MPa steps), temperature (1100-700 °C in 1 °C steps),
170 and f_{O_2} ($\Delta NNO = -3, -2.5, -2, -1.5, -1.25, -1, -0.75, -0.5, \text{ and } 0$) were chosen to explore very
171 shallow storage pressures, from above the liquidus to near-solidus temperatures, and over the
172 possible range of f_{O_2} expected for the system (Elders *et al.*, 2011; Zierenberg *et al.*, 2013). We
173 run all compositions at fluid saturated conditions, using a pure-H₂O fluid model. At such shallow
174 conditions, the Krafla magma is likely fluid saturated (Zierenberg *et al.*, 2013). Gualda and
175 Ghiorso (2014) (2014) and Ghiorso and Gualda (2015) demonstrate that water has a small effect
176 on calculated pressures for magmas that are rich in water.

177 Rhyolite-MELTS models the saturation surfaces of solid phases in equilibrium with the given
178 melt composition. For the Krafla IDDP-1 compositions, phases that rhyolite-MELTS predicts to
179 be possibly saturated under the conditions considered include quartz, plagioclase (labeled as
180 feldspar1 in MELTS_Excel results), orthopyroxene, and magnetite (labeled as spinel in
181 MELTS_Excel results); clinopyroxene is not predicted to be potentially saturated. The saturation
182 surfaces of quartz and feldspar are not sensitive to f_{O_2} , in contrast to the saturation surfaces of
183 orthopyroxene and magnetite, due to the presence of iron in orthopyroxene and magnetite.
184 Therefore, pressure calculations that include orthopyroxene and, particularly, magnetite are
185 highly dependent on f_{O_2} , which allows us to estimate f_{O_2} in addition to pressure (Harmon *et al.*,
186 2018; Pamukçu *et al.*, 2021). The differences between the observed and modeled mineral
187 assemblages are discussed in further detail in the Discussion section.

188 Following Foley *et al.* (2020) and Gualda *et al.* (XX in revision), we extend the methods of
189 Gualda and Ghiorso (2014) and Harmon *et al.* (2018) to calculate equilibrium pressures of the
190 simultaneous saturation in quartz, plagioclase, orthopyroxene, and magnetite (which we label
191 P_4 QFOM). We also calculate pressures based on the simultaneous saturation in quartz,
192 plagioclase, and orthopyroxene (P_3 QFO); and based on the simultaneous saturation in quartz,
193 plagioclase, and magnetite (P_3 QFM). We note that all three conditions (P_4 QFOM, P_3 QFO,
194 and P_3 QFM) imply saturation in quartz and plagioclase, which are independent of f_{O_2} .

195 For P_4 QFOM pressures, we calculate a pressure when the residual temperature (the minimum
196 range in saturation temperature for all phases at a single pressure, calculated by subtracting the
197 saturation temperatures of the mineral phase with the highest saturation temperature from the
198 mineral phase with the lowest saturation temperature at a single pressure) is less than 10 °C
199 (Figure 1). If no acceptable P_4 QFOM pressure is found, we calculate P_3 QFM and P_3 QFO
200 pressures if the residual temperature is less than or equal to 5 °C, as in Gualda and Ghiorso
201 (2014). The larger residual cutoff for P_4 QFOM is used due to uncertainties inherent in the
202 exercise of finding an intersection between 4 saturation surfaces, particularly when two of them
203 are affected by f_{O_2} .

204 To convert the calculated pressures to depths, we use $h = P/(\rho * g)$ where h is the depth (m), ρ is
205 the density of the crust (estimated to be 2,500-2,700 kg/m³), P is the calculated pressure (Pa), and
206 g is the acceleration due to gravity (9.8 m/s²). We report the depths in units of km, so that we do

207 not overinterpret the precision of the rhyolite-MELTS results. This represents the lithostatic
208 pressure, which is a maximum depth assuming no hydrostatic component of the system.

209 *Monte Carlo Analysis*

210 To determine the reproducibility, variability, and precision of the rhyolite-MELTS pressure and
211 f_{O_2} estimates, we conduct a Monte Carlo analysis (Gualda and Ghiorso, 2014; Pamukçu *et al.*,
212 2021; Pitcher *et al.*, 2021; Smithies *et al.*, 2023) using 600 synthetic glass compositions based on
213 the average Krafla composition and uncertainty around this composition. The f_{O_2} was allowed to
214 vary in 0.5 ΔNNO steps from $\Delta NNO = -3$ to 0. The rhyolite-MELTS calculations in the Monte
215 Carlo simulations explore the same pressure and temperature ranges detailed above.

216 **Results**

217 *Projection onto the Haplogranitic Ternary Diagram*

218 The average Krafla rhyolite composition plots approximately on the 50 MPa cotectic curve in the
219 Qz'-Ab'-Or' ternary diagram (Figure 2). Individual compositions have values that range from
220 just above the 50 MPa cotectic (i.e, shifted towards the Qz' vertex) to just below the 100 MPa
221 cotectic (i.e., shifted towards the Ab'-Or' join). These results correspond to maximum pressures
222 that range from slightly less than 50 MPa to slightly more than 100 MPa. The haplogranitic Qz'-
223 Ab'-Or' ternary gives a first order estimate of the equilibration pressures. We emphasize that the
224 absence of quartz simply imply that pressures should be lower than estimated from the diagram,
225 suggesting that Krafla magmas equilibrated at very low pressures, most likely <50 MPa.

226 As originally observed by Bégué *et al.* (2014a), when plotted on the Qz'-Ab'-Or' ternary
227 diagram, TVZ compositions that yield the lowest pressures exhibit low pressures (~50-150
228 MPa), but which are slightly higher than Krafla IDDP-1 pressures. The TVZ samples all contain
229 quartz, indicating that these pressures can be unambiguously interpreted to be storage pressures.

230 *Rhyolite-MELTS Pressures*

231 A total of 26 of the 31 individual RHL Krafla compositions (84%) yield pressures for at least one
232 of the nine f_{O_2} values we explored. Results from individual Krafla compositions are summarized
233 in Table 1 and Figure 3. A total of 143 of the 279 rhyolite-MELTS runs (51%) yield storage
234 pressures. Individual compositions can yield more than one pressure in the cases in which
235 pressures are calculated for multiple f_{O_2} values.

236 Individual compositions return P_4 QFOM pressures for f_{O_2} values of $\Delta NNO = -1$ to -0.5 , with
 237 the most pressure calculations confined to the even narrower f_{O_2} range of $\Delta NNO = -1$ and -0.75 .
 238 The average pressure for all P_4 QFOM pressures (regardless of f_{O_2}) is 50 ± 11 MPa (1-sigma),
 239 which corresponds to a depth of $1.9\text{-}2.0 \pm 0.4$ km.

240 For more reducing conditions (f_{O_2} equal to $\Delta NNO = -3$ to -1), individual compositions return
 241 predominantly P_3 QFO pressures. A total of 117/279 (42%) of the rhyolite-MELTS runs return
 242 P_3 QFO pressures. The average pressure for all P_3 QFO pressures (regardless of f_{O_2}) is 44
 243 MPa ± 11 MPa ($1.7\text{-}1.8$ km ± 0.4 km).

244 For more oxidizing conditions (f_{O_2} equal to $\Delta NNO = -1$ to 0), individual compositions return
 245 predominantly P_3 QFM pressures. A total of 46/279 (16%) of the rhyolite-MELTS runs return
 246 P_3 QFM pressures. The average pressure for all P_3 QFM pressures (regardless of f_{O_2}) is 47
 247 MPa ± 32 MPa ($1.8\text{-}1.9$ km $\pm 1.2\text{-}1.3$ km).

248 The average Krafla IDDP-1 composition returns a P_4 QFOM pressure of 42 MPa ($1.6\text{-}1.7$ km)
 249 for $\Delta NNO = -0.75$ and 46 MPa ($1.7\text{-}1.9$ km) for $\Delta NNO = -1$. These are the only f_{O_2} values that
 250 return a P_4 QFOM pressure for the average Krafla IDDP-1 composition, indicating that – as
 251 expected – the calculations of storage pressure using four phases are highly dependent on f_{O_2} .
 252 The sensitivity of the rhyolite-MELTS calculations to f_{O_2} , temperature, and pressure are
 253 highlighted in Figure 4, where there is only a small “valley” in the average composition data that
 254 produces P_4 QFOM pressures. We highlight the average Krafla calculations, but the shape of
 255 the surface in Figure 4 is similar for all individual rhyolite-MELTS calculations.

256 Only a narrow range of f_{O_2} values produce P_4 QFOM pressures. The P_3 QFM and P_3 QFO
 257 pressures show somewhat larger ranges of pressure and f_{O_2} (Figures 3 and 4). The overall
 258 distribution of pressures is similar for all assemblages and f_{O_2} values.

259 *Monte Carlo Simulations*

260 Results from the Monte Carlo simulations are summarized in Figure 5 and Table 2. A total of
 261 373/600 (62%) of rhyolite-MELTS runs return a valid pressure. The phase assemblages
 262 considered are the same assemblages considered for the individual compositions (P_4 QFOM,
 263 P_3 QFO, and P_3 QFM). There are 27 compositions (5%) that return P_QFOM pressures. The
 264 f_{O_2} values that produced these P_4 QFOM pressures are equal to $\Delta NNO = -1$ (20 compositions)

265 and -0.5 (7 compositions). The average pressure for all P_4 QFOM pressures (regardless of f_{O_2})
266 is $48 \text{ MPa} \pm 8 \text{ MPa}$, which corresponds to a depth of $1.8\text{-}2.0 \text{ km} \pm 0.3 \text{ km}$ assuming lithostatic
267 pressure.

268 There are 317/600 compositions (53%) that return P_3 QFO pressures and 83 compositions
269 (14%) return P_3 QFM pressures. The average pressure for all P_3 QFO pressures (regardless of
270 f_{O_2}) is $46 \text{ MPa} \pm 20 \text{ MPa}$ ($1.7\text{-}1.9 \text{ km} \pm 0.8 \text{ km}$) and the average pressure for all P_3 QFM
271 pressures (regardless of f_{O_2}) is $37 \text{ MPa} \pm 13 \text{ MPa}$ ($1.4\text{-}1.5 \text{ km} \pm 0.5 \text{ km}$). A similar trend to the
272 individual compositions is observed, as P_3 QFO pressures are calculated for more reducing
273 conditions and P_3 QFM pressures are calculated for more oxidizing conditions for the Monte
274 Carlo results.

275 **Discussion**

276 *Hydrostatic vs lithostatic pressures*

277 The quenched glass fragments retrieved from 2.1 km depth at the Krafla IDDP-1 well allow us to
278 test the reliability of rhyolite-MELTS pressures for very low pressures ($\sim 50 \text{ MPa}$), with
279 important implications for our understanding of magma storage conditions. While the depth of
280 the magma intersected by Krafla IDDP-1 is known, determining the *in situ* pressure conditions
281 experienced by the magma is more nuanced. The bounds of acceptable pressures are defined by
282 the lithostatic pressure and by the hydrostatic pressure. The lithostatic pressure for this magma at
283 2.1 km is $\sim 51\text{-}56 \text{ MPa}$ using a rock density of $2,500\text{-}2,700 \text{ kg/m}^3$. In contrast, the hydrostatic
284 pressure is calculated to be $\sim 16 \text{ MPa}$ (Zierenberg *et al.*, 2013). Based on H_2O and CO_2
285 concentrations in glass determined with FTIR, Zierenberg *et al.* (2013) calculate an $\text{H}_2\text{O}\text{-CO}_2$
286 saturation pressure of $\sim 35\text{-}45 \text{ MPa}$ using VolatileCalc and assuming a temperature of $900 \text{ }^\circ\text{C}$
287 (Newman and Lowenstern, 2002). Rhyolite-MELTS calculations return pressures between
288 lithostatic and the $\text{H}_2\text{O}\text{-CO}_2$ saturation pressure, suggesting that the Krafla magma body was
289 nearly fluid saturated, but possibly at a pressure slightly lower ($<10 \text{ MPa}$) than the lithostatic
290 pressure, similar to the results of Zierenberg *et al.* (2013). To summarize, the pressures
291 calculated via VolatileCalc and rhyolite-MELTS yield pressures greater than the hydrostatic
292 pressure and lower than the lithostatic pressure. In both cases, calculated pressures are closer to
293 the lithostatic pressure ($50\text{-}57 \text{ MPa}$) than to the hydrostatic pressure (16 MPa).

294 *Constraints from glass compositions*

295 When represented on a normalized anhydrous basis, the SiO₂ contents of the Krafla glass
296 compositions have relatively low SiO₂ concentrations (i.e., 76.7 wt% SiO₂), much lower than
297 what would be expected from the correlation between SiO₂ and pressure found by Gualda and
298 Ghiorso (2013b), which would predict SiO₂ values >78 wt% for such low pressures. This is
299 easily explained by the much higher FeO values (2.9 wt% FeO for the average Krafla glass; see
300 supplementary material) observed in the tholeiitic compositions when compared to more typical
301 calc-alkaline systems studied by Gualda and Ghiorso (2013b), which only have 0.55-1.29 wt%
302 FeO. In this case, the SiO₂ concentration alone cannot be used for estimation of crystallization
303 pressure using the relationship of Gualda & Ghiorso (2013b XX).

304 The projection scheme of Blundy and Cashman (2001) circumvents the issue of high FeO values
305 of the Krafla compositions by employing normative values for Qz', Ab', Or', and An (effectively
306 this means compositions are considered on an FeO-free, MgO-free basis). While the Qz'-Ab'-
307 Or' ternary projection diagram is a somewhat crude measure of pressure, the average Krafla
308 IDDP-1 composition lies on the 50 MPa (1.9-2.0 km) cotectic (Figure 2), which indicates that the
309 cotectic is in good agreement with the natural samples intercepted at 2.1 km. Importantly,
310 pressures estimated using the Qz'-Ab'-Or' ternary are maximum pressures, given that the
311 presence of quartz is not observed in the samples.

312 *Constraints from rhyolite-MELTS geobarometry*

313 For the rhyolite-MELTS results, the average P₄ QFOM pressure for the individual
314 compositions is in excellent agreement with the drilling depth. The average P₃ QFO pressure
315 and P₃ QFM pressure for the individual compositions are also in very good agreement with the
316 observed depth. In addition to the pressure measurements, rhyolite-MELTS estimates the f_{O_2} of
317 the Krafla system to be between $\Delta NNO = -0.5$ and -1 , which agrees with the reducing f_{O_2}
318 estimates for Krafla magmas (Nicholson, 1990; Shorttle *et al.*, 2015; Hartley *et al.*, 2017). The
319 results from the individual and average compositions are supported by the Monte Carlo results,
320 which yield P₄ QFOM pressures exclusively for f_{O_2} of $\Delta NNO = -1$ and -0.5 . In summary,
321 rhyolite-MELTS calculations indicate storage pressures of <55 MPa, under reducing ($\Delta NNO = -$
322 1 to -0.5) conditions. Furthermore, comparison of rhyolite-MELTS pressures with the fluid-

323 saturation pressures of Zierenberg *et al.* (2013) suggests that Krafla IDDP-1 magmas were fluid-
324 saturated, or very nearly so.

325 There are two notable discrepancies between the reported mineral assemblage in the Krafla
326 IDDP-1 samples and rhyolite-MELTS calculations. First, the P₄ QFOM pressure results
327 suggest that quartz is in equilibrium with the Krafla magmas despite no quartz being observed;
328 as well, the pyroxene predicted by rhyolite-MELTS is orthopyroxene instead of the observed
329 augite + pigeonite assemblage (Zierenberg *et al.*, 2013; Masotta *et al.*, 2018). While quartz has
330 not been reported as part of the phenocryst assemblage within the mm-sized glass shards
331 (Zierenberg *et al.*, 2013; Masotta *et al.*, 2018), the small glass shards could make it very difficult
332 to find any phenocrystic quartz crystals. However, we do not suggest that Zierenberg *et al.* (2013
333 and Masotta *et al.* (2018) missed the presence of quartz. Instead, a plausible alternative is that
334 quartz was very near saturation and/or was saturated and kinetically suppressed. The fragments
335 of the felsite crust directly adjacent to the melt-rich, crystal-poor magma body contain quartz and
336 alkali feldspar (and plagioclase, augite, and titanomagnetite) (Zierenberg *et al.*, 2013),
337 suggesting that Krafla magmas were in contact with quartz-bearing rocks. This proximity to a
338 quartz-bearing felsite could have altered the composition of the melt during storage. The
339 coincidence of the plagioclase and quartz saturation curves for the Krafla IDDP-1 compositions
340 indicates that both minerals are saturated in rhyolite-MELTS calculations. Importantly, the only
341 pressures that are consistent with the rhyolite-MELTS calculations are pressures lower than the
342 calculated pressures. As such, similarly to the case of the haplogranitic ternary, we conclude that
343 pressures are <55 MPa.

344 Second, the pyroxenes observed in the Krafla IDDP-1 samples (augite + pigeonite) are both
345 clinopyroxene. However, the rhyolite-MELTS calculations do not produce clinopyroxene for any
346 conditions tested. Instead, the orthopyroxene saturation curve is ubiquitously present in rhyolite-
347 MELTS calculations. For P₄ QFOM, the orthopyroxene saturation curve is coincident with the
348 quartz + plagioclase + magnetite saturation curves. We conducted several rhyolite-MELTS runs
349 on the average composition while suppressing orthopyroxene. In the case of f_{O_2} equal to ΔNNO
350 = -1, the four-phase intersection of quartz, plagioclase, clinopyroxene, and magnetite (P₄
351 QFCM) is 50 MPa, which is in excellent agreement with the P₄ QFOM pressures. Therefore,
352 we suggest that the orthopyroxene predicted by rhyolite-MELTS is a good proxy for the low-Ca
353 clinopyroxene present in the natural samples in Krafla magmas. Energetically, they are almost

354 indistinguishable by rhyolite-MELTS for these Krafla compositions. It is also relevant to note
355 that recent experimental results suggest that the clinopyroxene model in rhyolite-MELTS may be
356 inaccurate for high-silica rhyolite compositions (Brugman and Till, 2019). Interestingly, our
357 results suggest that – in this case – orthopyroxene is a useful proxy for clinopyroxene stability.

358 *Very shallow magma bodies as part of transcrustal magmatic systems*

359 There is a growing body of evidence to suggest that magmatic systems span a significant range
360 of storage depths within the crust – leading to the idea of transcrustal magma systems (Annen *et al.*,
361 2006; Annen, 2009; Cashman and Giordano, 2014; Menand *et al.*, 2015; Mutch *et al.*, 2019;
362 Svoboda *et al.*, 2021; Giordano and Caricchi, 2022). The existence of very shallow magma
363 bodies indicates that we must extend our model of storage to a shallower level.

364 It is interesting to consider why these bodies have been largely underappreciated to date, with
365 much more attention being paid to deeper levels of rhyolitic magmatic systems. The lifetimes of
366 shallow melt-dominated magma bodies are short (Gualda *et al.*, 2012b, 2018; Cooper and Kent,
367 2014; Pamukçu *et al.*, 2015a, 2021; Gualda and Sutton, 2016; Pitcher *et al.*, 2021), especially
368 when a geothermal reservoir is present (Kelly *et al.*, 2021). We speculate that very shallow
369 magma bodies could commonly be a minor contribution to major eruptions. Their proximity to
370 the Earth's surface makes them economically and societally important, especially from a
371 volcanic hazards perspective.

372 The very shallow storage pressures calculated here are from Krafla in the plume-affected mid-
373 ocean rift of Iceland. Shallow magmas are also inferred to exist in the rifted arc of the TVZ in
374 Aotearoa New Zealand (Bégué *et al.*, 2014b; Gualda *et al.*, 2018; Harmon *et al.*, 2024). For both
375 Krafla and the TVZ, the magma systems are long-lived and indicate that there is a consistent
376 high heat flux to the shallow crust (Jonasson, 1994; Wilson, 1996; Mutch *et al.*, 2019; Kelly *et al.*,
377 2021), consistent with involvement of greater depths of the crust in the generation and
378 storage of eruption-forming magma bodies (Gualda *et al.*, 2018, 2019; Smithies *et al.*, 2023).

379 *Implications for global shallow magmatism*

380 In addition to drilling evidence of very shallow magma at depths between 2.1 and 2.6 km at
381 Puna, Hawaii (Teplow *et al.*, 2009), Menengai Caldera, Kenya (Mbia *et al.*, 2015), and Krafla
382 geothermal system, Iceland (Mortensen *et al.*, 2010; Elders *et al.*, 2011; Zierenberg *et al.*, 2013),

383 the potential presence of very shallow silicic magma has also been captured by 2D and 3D
384 seismic exploration at the Larderello geothermal field, Italy (Cameli *et al.*, 1993; Manzella *et al.*,
385 2017; Rochira *et al.*, 2018) and by InSAR data at the Dabbahu volcano, Afar, Ethiopia (Ebinger
386 *et al.*, 2008). Geophysical methods, including seismic data and InSAR, are likely critical for
387 determining the presence of current, very shallow magma bodies.

388 In addition to the direct evidence at drilled magma sites, the rhyolite-MELTS evidence from
389 previous eruptions in the TVZ, and the indirect observations by geophysical methods, there is
390 substantial field and petrologic evidence (mostly Al-in-hornblende and Qz'-Ab'-Or' barometry,
391 as well as common granophyric textures) for very shallow magma bodies in the plutonic record.
392 We briefly summarize some examples below.

393 In the case of the Searchlight pluton in Nevada (Bachl *et al.* 1991, Wallrich *et al.* 2023), which is
394 exposed along steeply tilted crustal sections, the upper units of the pluton intruded into roughly
395 coeval volcanic rocks, suggesting very shallow emplacement. Further, Al-in-hornblende
396 geobarometry suggests that the roof of the pluton was at a depth of ~3 km (Bachl *et al.*, 2001;
397 Wallrich *et al.*, 2023). The Geysers in California is one of the world's most active geothermal
398 fields and has a plutonic history that includes shallow intrusions ~0.2-2.8 km depth,
399 exemplifying the connection between shallow plutonism and geothermal energy production
400 (Angeles-De La Torre, *et al.*, 2023). In the case of the Turkey Creek and Silver Creek calderas in
401 Arizona, resurgent monzonite and granite intruded at very shallow levels (<~2 km) (du Bray and
402 Pallister, 1999; Ferguson *et al.*, 2013; McDowell *et al.*, 2014; Deering *et al.*, 2016). The Mount
403 Scott Granite, Oklahoma, is interpreted to have been emplaced into comagmatic Carlton Rhyolite
404 at a depth of no more than ~1.5 km (Hogan and Gilbert, 1995; Hogan *et al.*, 1998). The
405 emplacement depth of Cretaceous granitoid plutons in eastern Zhejiang, China is estimated to
406 have been at 50-100 MPa (~2-4 km) based on the Qz'-Ab'-Or' ternary (Chen *et al.*, 2021). Al-in-
407 hornblende geobarometry yields emplacement depths of ~4-5 km at the top of the Rayo Bisco-
408 Huemul plutonic complex, Chile (Nelson *et al.*, 1999; Schaen *et al.*, 2017). The granitic
409 intrusions of eastern Iceland (Austurhorn, Vesturhorn, Slaufudalur, and Reyðarártindur plutons)
410 are all interpreted to have been emplaced into slightly older basaltic strata at depths of <~2 km,
411 based on depths to reconstructed paleosurfaces and metamorphism (Walker, 1960, 1974; Blake,
412 1966; Furman *et al.*, 1992; Burchardt *et al.*, 2012; Padilla *et al.*, 2016; Twomey *et al.*, 2020;
413 Rhodes *et al.*, 2021). Granophyric lithic blocks within pumiceous pyroclastic flows are

414 interpreted to have formed in the magma body beneath Alid Volcanic Center, Eritrea, at a depth
415 of 2-4 km based on geological constraints and CO₂-H₂O concentrations in melt inclusions
416 (Lowenstern *et al.*, 1997). It is noteworthy that many of these localities are significantly affected
417 by extension, which could suggest that extension and rifting may facilitate establishment of these
418 very shallow magma bodies. Overall, the evidence above suggests that very shallow rhyolitic
419 magma bodies are relatively common, and it is critical to better understand them for hazard
420 assessment and economic exploration. As well, these very shallow magma bodies are probably
421 important components of transcrustal magma systems that deserve further study.

422 Methods to petrologically and geophysically monitor the existence and the potential hazards
423 associated with these very shallow magma bodies should be established to mitigate volcanic
424 hazards. The substantial petrologic evidence that these magma bodies not only exist but are
425 relatively widespread should be a call for the combined use of geophysics, hazard assessment,
426 and petrology to properly assess the presence, properties, and potential societal impacts of very
427 shallow magma bodies, especially as we begin to directly probe these magmas (Eichelberger *et*
428 *al.*, 2018; Lavallée *et al.*, 2023). This environment has been underappreciated as a potential
429 magma storage zone, with implications for volcanic hazards and geothermal resources.

430 *Conclusions*

431 In this study, we test rhyolite-MELTS geobarometry on the shallow magma intersected and
432 retrieved from Krafla IDDP-1 geothermal well. Using the glass compositions as the input, we
433 search for mineral-melt equilibrium conditions in temperature, pressure, and f_{O_2} space to
434 calculate the storage conditions of these magmas. Rhyolite-MELTS returns four-phase (quartz,
435 plagioclase, orthopyroxene, magnetite; P_4 QFOM) pressures of 42 MPa (1.6-1.7 km) for
436 $\Delta NNO = -0.75$ and 46 MPa (1.7-1.9 km) for $\Delta NNO = -1$, which is in excellent agreement with
437 the drilled depth of 2.1 km. While rhyolite-MELTS calculates both quartz and orthopyroxene
438 (instead of augite + pigeonite) to be in equilibrium with melt of the input composition, these
439 phases are not observed in the samples. Despite these discrepancies, we argue that rhyolite-
440 MELTS performs with high precision for this system.

441 The fact that we can obtain rhyolite-MELTS pressures consistent with the drilled depth lends
442 support to rhyolite-MELTS results in other systems that produce very shallow pressures, like the
443 Taupō Volcanic Zone, Aotearoa New Zealand. We show that rhyolite-MELTS can be used to

444 calculate not only the pressures of storage based on a variety of mineral assemblages, but that it
445 can also help constrain other intensive parameters such as f_{O_2} and H₂O-saturation during pre-
446 eruptive storage.

447 Contemporary very shallow magma bodies are confirmed in several locations, including the
448 Krafla IDDP-1 well, and they may be relatively common features of transcrustal magmatic
449 systems, including from the TVZ.

450 Coupling the magmatic information revealed by very shallow magma bodies with geophysical
451 models and volcanic hazard assessments has broad implications for both geothermal energy
452 production and volcanic hazards. Further, if we know the depths and conditions of these magma
453 bodies (e.g., by using rhyolite-MELTS), we can better understand the arrangement of melt-
454 dominated magma bodies during ongoing eruptions. This updated perspective will aid in
455 reconfiguring conceptual models and refine the focus of eruption monitoring platforms for silicic
456 magma storage in the upper crust.

457 **Acknowledgments**

458 We thank John Eichelberger for the inspiration for this project. Thanks to Mark Ghiorso for
459 comments on an early draft and for discussion of some of the results. Thank you to the
460 PUMMUS group for comments on early drafts and support.

461 **Funding**

462 There was no external funding for this project.

463 **Author contributions**

464 Conceptualization: BMW, GARG, LJH, CFM

465 Methodology: LJH, GARG, BMW

466 Investigation: LJH, GARG, BMW, CFM

467 Visualization: LJH, GARG

468 Supervision: GARG, LJH

469 Writing—original draft: LJH

470 Writing—review & editing: LJH, GARG, CFM, BMW

471 **Competing interests**

472 Authors declare that they have no competing interests.

473 **Data and materials availability**

474 All data are available in the main text or the supplementary materials.

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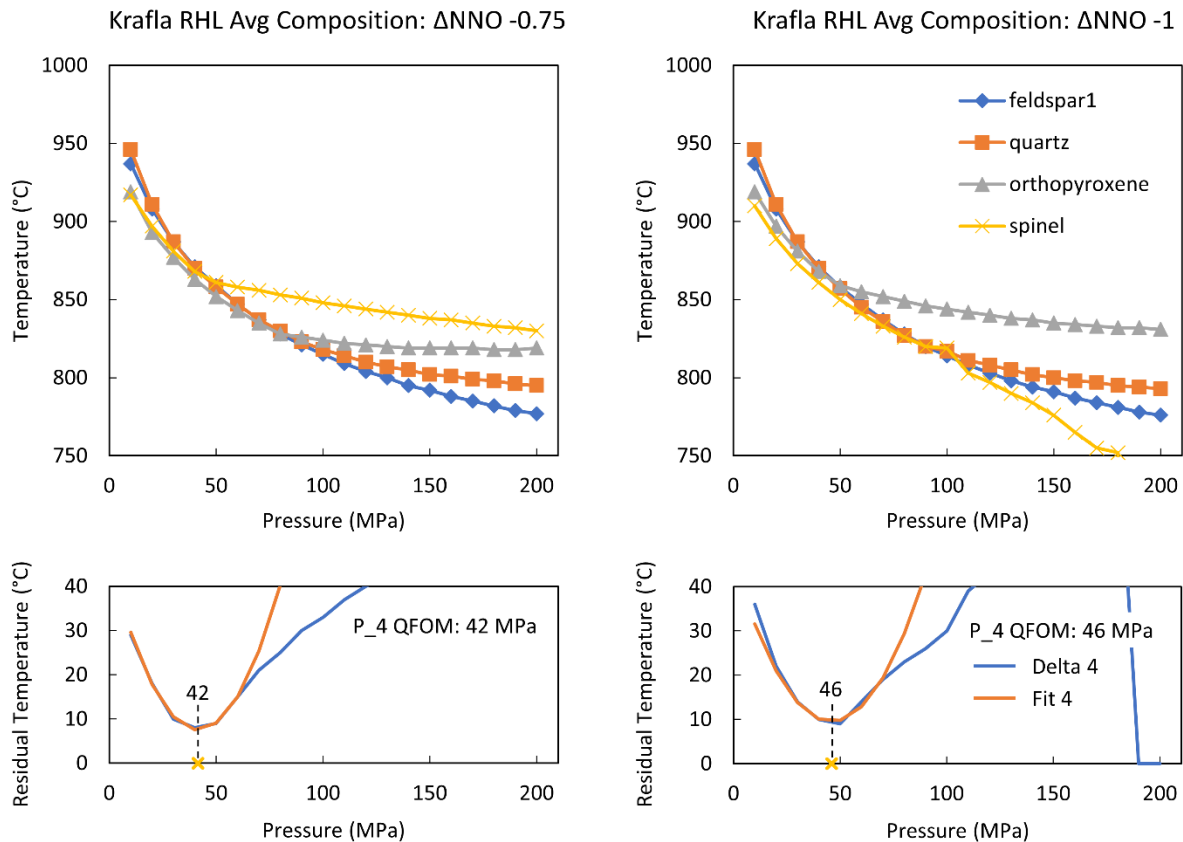
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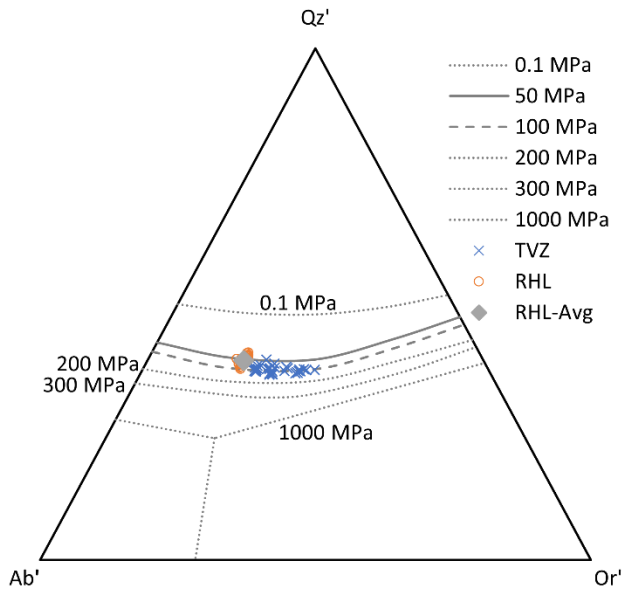
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756 **Figure Captions and Tables**

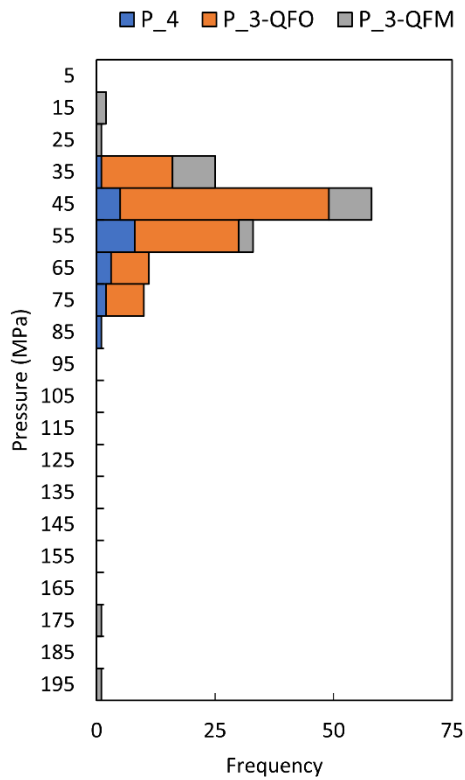
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 758 **Fig. 1. Rhyolite-MELTS storage pressures for the Krafla RHL average composition as a**
 759 **function of oxygen fugacity.** Storage pressures indicate the pressure and temperature estimates
 760 for which quartz, plagioclase, orthopyroxene, and spinel (magnetite) are simultaneously
 761 saturated (P₄ QFOM). The oxygen fugacity (f_{O_2}) estimates of $\Delta NNO = -0.75$ (left panel) and
 762 $\Delta NNO = -1$ (right panel) are the only f_{O_2} values that yield pressures for the Krafla RHL average
 763 composition, as these are the only simulations in which all four mineral phases saturate within 10
 764 °C of each other (the residual temperature cutoff we use for a 4-phase solution, shown by the
 765 blue “Delta 4” in the lower diagrams). The pressure is calculated by fitting a parabola (“Fit 4”) to
 766 the residual temperature points around the minimum residual temperature (“Delta 4”); see
 767 Gualda and Ghiorso (2014) for details. The stability of both orthopyroxene and spinel
 768 (magnetite) depends on f_{O_2} due to the presence of Fe. The estimated pressures are 42 MPa and 46
 769 MPa, respectively. Using a density of 2,500 kg/m³, these pressures are calculated to be 1.6 km
 770 and 1.7 km; using a density of 2,700 kg/m³, these pressures are calculated to be 1.7 km and 1.9
 771 km.

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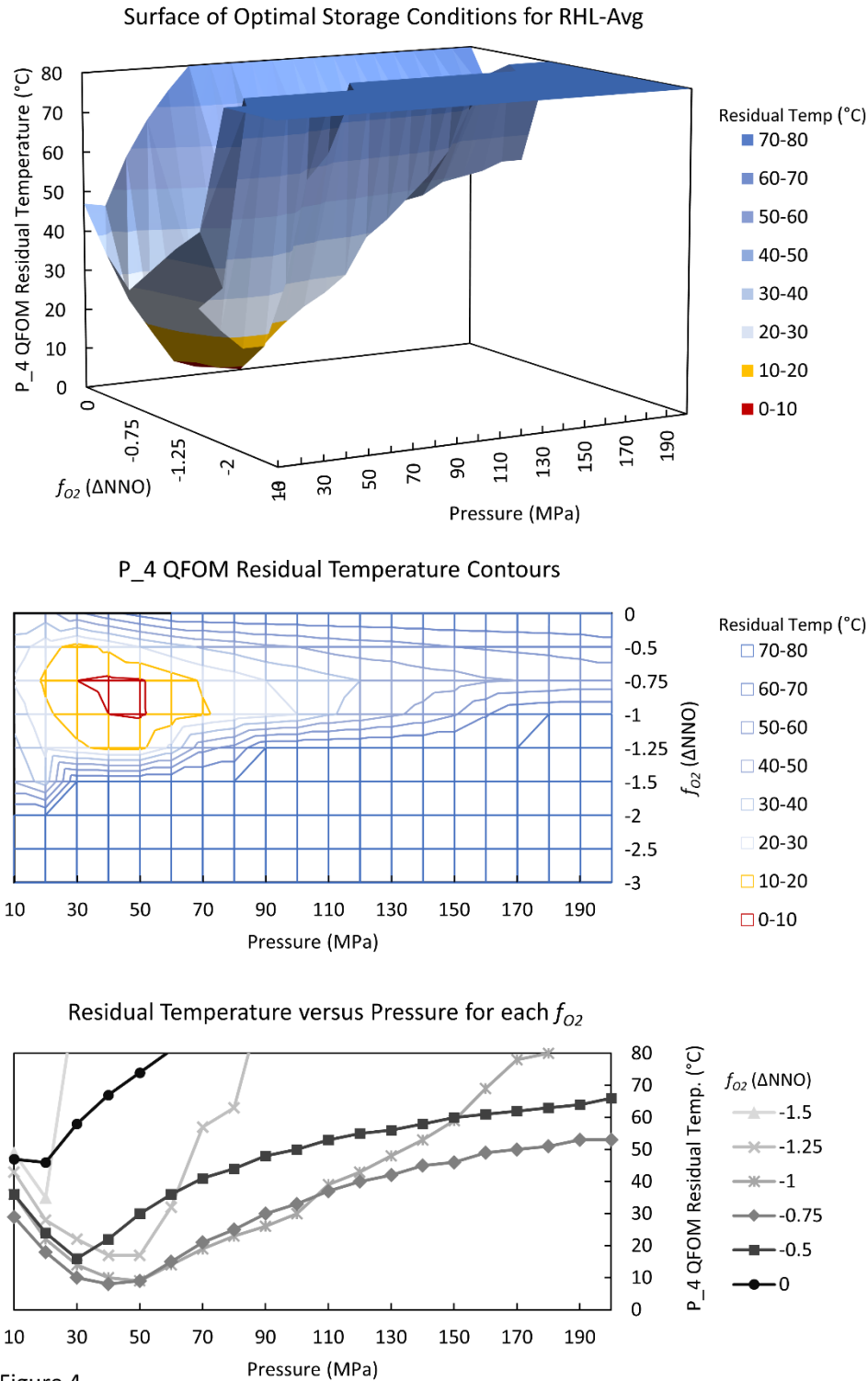
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774 **Fig. 2. Projection of glass compositions onto the quartz-albite-orthoclase (Qz'-Ab'-Or')**
 775 **ternary diagram.** Projection of individual glass compositions and the average glass composition
 776 from Krafla RHL glass shards (Masotta *et al.*, 2018) and glass compositions from TVZ
 777 ignimbrite pumice (Bégué *et al.*, 2014; Gualda *et al.*, 2018) onto the quartz-albite-orthoclase
 778 (Qz'-Ab'-Or') ternary diagram using the projection scheme of Blundy and Cashman (2001),
 779 which relies only on the glass compositions. The Krafla RHL compositions plot mostly between
 780 the 50 MPa and 100 MPa cotectic, with the average composition plotting at ~ 50 MPa. The TVZ
 781 compositions plot mostly between 50 and ~150 MPa, predominantly around the 100 MPa
 782 cotectic. Note that the Krafla RHL compositions plot slightly shallower than TVZ compositions.
 783 Importantly, inferred pressures are maximum pressures if quartz is not in equilibrium with the
 784 melt.



785

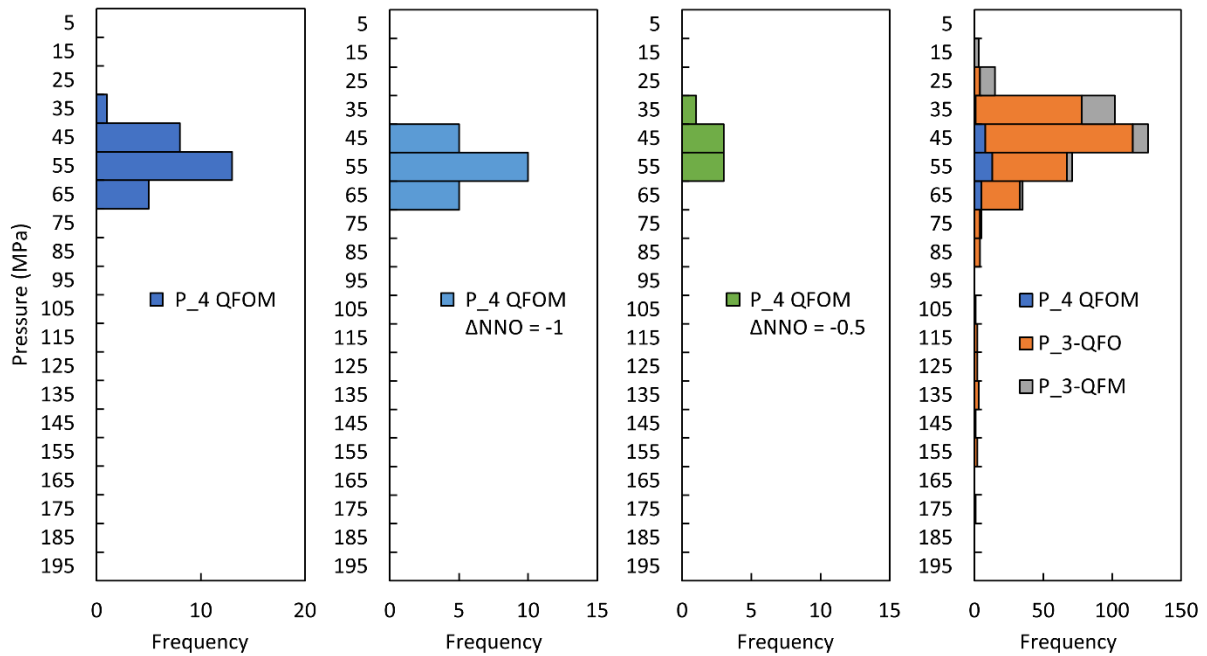
786 **Fig. 3. Results from rhyolite-MELTS pressure modeling for all individual Krafla RHL**
 787 **compositions** (Masotta *et al.*, 2018). We report the P_4 QFOM, P_3 QFO, and P_3 QFM
 788 pressures for all f_{O_2} considered. Note the narrow range of pressures that return a P_4 QFOM
 789 pressure, which is the most sensitive to f_{O_2} . There is a strong mode at ~45-55 MPa for all
 790 assemblages, with slightly higher values for P_4 QFOM pressures.



791 Figure 4

792 **Fig. 4. Rhyolite-MELTS storage pressures for the Krafla RHL average glass composition,**
 793 **showing the effect of f_{O_2} on pressure calculations.** In the top panel, the central red divot (~40
 794 MPa and $\Delta NNO = \sim -0.75$) shows the conditions that retrieves the best pressure and f_{O_2} estimate

795 based on the lowest residual temperature when all four phases are considered (P_4 QFOM). The
796 middle panel shows the same data as a 2D contour plot. For readability, residual values > 80 °C
797 are shown as 80 °C. The bottom panel shows the residual temperatures (the difference between
798 the saturation curves of different phases) for individual rhyolite-MELTS calculations, which is
799 compiled to create the two upper panels. For most f_{O_2} values, particularly between $\Delta NNO = -$
800 1.25 to -0.75, estimated pressures are between 40 and 50 MPa. Calculations outside this f_{O_2}
801 interval ($\Delta NNO < -1.5$ and $\Delta NNO > -0.75$) yield no pressures.



802

803 **Fig. 5. Rhyolite-MELTS results for the Monte Carlo analysis.** The Monte Carlo analysis
 804 included 600 synthetic compositions that varied about the mean of the Krafla RHL composition,
 805 using the calculated standard deviation, with f_{O_2} values from $\Delta NNO = -2$ to 0 distributed
 806 uniformly in $0.5 f_{O_2}$ steps. Left panel shows rhyolite-MELTS P_4 QFOM pressure results from
 807 the Monte Carlo analysis. The only f_{O_2} values that resulted in a P_4 QFOM pressure calculation
 808 are f_{O_2} equal to $\Delta NNO = -1$ and $\Delta NNO = -0.5$, shown in the middle two panels. This indicates we
 809 can determine the pressure and the f_{O_2} of the system. The right panel shows rhyolite-MELTS P_4
 810 QFOM, P_3 QOF and P_3 QFM pressure results from the Monte Carlo simulations for all f_{O_2}
 811 values considered. In all cases, modes of the distributions are in the 45-55 MPa bins.

(A) Mean pressure calculation (MPa) and depth (km)¹

	-3 ΔNNO	-2.5 ΔNNO	-2 ΔNNO	-1.5 ΔNNO	-1.25 ΔNNO	-1 ΔNNO	-0.75 ΔNNO	-0.5 ΔNNO	0 ΔNNO
P_4 QFOM	-	-	-	-	-	57 MPa (2.1; 2.3 km)	47 MPa (1.8; 1.9 km)	35 MPa (1.3; 1.4 km)	-
P_3 QFO	40 MPa (1.5; 1.6 km)	42 MPa (1.6; 1.7 km)	45 MPa (1.7; 1.8 km)	47 MPa (1.8; 1.9 km)	48 MPa (1.8; 2.0 km)	49 MPa (1.9; 2.0 km)	46 MPa (1.7; 1.9 km)	35 MPa (1.3; 1.4 km)	-
P_3 QFM	-	-	-	-	-	55 MPa (2.1; 2.2 km)	60 MPa (2.3; 2.5 km)	32 MPa (1.2; 1.3 km)	9 MPa (0.4; 0.4 km)

(B) Standard deviation pressure calculation (MPa) and depth (km)²

P_4 QFOM	-	-	-	-	-	11 MPa (0.4; 0.5 km)	5 MPa (0.2 km)	2 MPa (0.1 km)	-
P_3 QFO	9 MPa (0.3; 0.4 km)	9 MPa (0.4; 0.4 km)	11 MPa (0.4; 0.5 km)	11 MPa (0.4; 0.5 km)	12 MPa (0.5; 0.5 km)	11 MPa (0.4; 0.5 km)	5 MPa (0.2 km)	2 MPa (0.1 km)	-
P_3 QFM	-	-	-	-	-	11 MPa (0.4; 0.5 km)	47 MPa (1.8; 1.9 km)	5 MPa (0.2 km)	1 MPa (0 km)

(C) Number of compositions that return pressures for each f_{O_2} ³

P_4 QFOM	0	0	0	0	0	9	9	2	0
P_3 QFO	14	16	17	20	20	9	1	0	0
P_3 QFM	0	0	0	0	0	2	8	14	2

812 **Table 1. Pressure calculations for Krafla RHL individual compositions.**

813 ¹Mean of P_4 QFOM, P_3 QFO, and P_3 QFM storage pressures for Krafla RHL individual composition for different f_{O_2} reported in MPa and km; the
814 densities used to convert pressure to depth are 2,700 kg/m³ and 2,500 kg/m³

815 ²Standard deviation of P_4 QFOM, P_3 QFO, and P_3 QFM storage pressures for Krafla RHL individual compositions reported in MPa and km

816 ³ Number of compositions that returned storage pressures for each f_{O_2} . Note that the averages of P_3 QFO and P_3 QFM include the P_4 QFOM pressures in
817 the average and standard deviation calculations since the P_4 pressure will also return a P_3 pressure.
818

(A) Mean pressure calculation (MPa) and depth (km) ¹

	-3 ΔNNO	-2.5 ΔNNO	-2 ΔNNO	-1.5 ΔNNO	-1 ΔNNO	-0.5 ΔNNO	0 ΔNNO
P_4 QFOM	-	-	-	-	50 MPa (1.9; 2.0 km)	41 MPa (1.6; 1.7 km)	-
P_3 QFO	36 MPa (1.4; 1.5 km)	40 MPa (1.5; 1.6 km)	41 MPa (1.6; 1.7 km)	45 MPa (1.7; 1.8 km)	48 MPa (1.8; 2.0 km)	91 MPa (3.4; 3.7 km)	126 MPa (4.8; 5.1 km)
P_3 QFM	-	-	-	-	51 MPa (1.9; 2.1 km)	34 MPa (1.3; 1.4 km)	17 MPa (0.6; 0.7 km)

(B) Standard deviation pressure calculation (MPa) and depth (km) ²

P_4 QFOM	-	-	-	-	6 MPa (0.2; 0.3 km)	8 MPa (0.3; 0.4 km)	-
P_3 QFO	9 MPa (0.3; 0.4 km)	9 MPa (0.3; 0.4 km)	9 MPa (0.3; 0.4 km)	10 MPa (0.4; 0.4 km)	11 MPa (0.4; 0.5 km)	46 MPa (1.8; 1.9 km)	25 MPa (0.9; 1.0 km)
P_3 QFM	-	-	-	-	8 MPa (0.3; 0.3 km)	8 MPa (0.3; 0.3 km)	5 MPa (0.2; 0.2 km)

(C) Number of compositions that return pressures for each f_{O_2} ³

P_4 QFOM	0	0	0	0	20	7	0
P_3-QFO	51	66	60	68	32	10	3
P_3-QFM	0	0	0	0	5	42	9
Number of comps	85	83	90	102	72	74	94

819 **Table 2. Pressure calculations for Monte Carlo compositions from Krafla average composition.**

820 ¹ Mean of P_4 QFOM, P_3 QFO, and P_3 QFM storage pressures for Monte Carlo Krafla RHL compositions for different f_{O_2} reported in MPa and km; the
821 densities used to convert pressure to depth are 2,700 kg/m³ and 2,500 kg/m³

822 ²Standard deviation of P_4 QFOM, P_3 QFO, and P_3 QFM storage pressures for Monte Carlo Krafla RHL compositions reported in MPa and km

823 ³ Number of compositions that returned storage pressures for each f_{O_2} . Note that the averages of P_3 QFO and P_3 QFM include the P_4 QFOM pressures in
824 the average, standard deviation calculations since the P_4 pressure will also return a P_3 pressure.