

1 **Constraining Earth Properties using Deterministic** 2 **Backus-Gilbert Inferences**

3 Adrian M. Mag¹, Christophe Zaroli², Paula Koelemeijer¹

¹ *Department of Earth Sciences, University of Oxford, Oxford, United Kingdom*

² *Institut Terre et Environnement de Strasbourg, Université de Strasbourg, EOST, CNRS, UMR 7063,*

5 rue Descartes, Strasbourg F-67084, France

4 Received date; in original form date

5 **SUMMARY**

6 Seismic tomography is routinely used to image the Earth's interior using seismic data. How-
7 ever, in practice, data limitations lead to discretised inversions or the use of regularisations,
8 which complicates tomographic model interpretations. In contrast, Backus-Gilbert methods
9 make it possible to infer properties of the true Earth, providing useful insights into the internal
10 structure of our planet. Here, we show how in the absence of data uncertainties the Backus-
11 Gilbert-based Subtractive Optimally Localized Averages (SOLA) method can be utilised as a
12 deterministic linear inference (DLI) method to derive a new method: SOLA-DLI. SOLA-DLI
13 enables us to interpret results through the target kernels, rather than the imperfect resolving
14 kernels. This also allows us to build families of models, rather than just properties. These ad-
15 vantages are illustrated in this contribution using three case studies. In the first, we illustrate
16 how properties such as various local averages and gradients can be obtained, including asso-
17 ciated bounds on these properties and resolution information. Our second case study shows
18 how trade-offs between physical parameters are naturally included in SOLA-DLI, which is
19 particularly relevant to inferences of Earth structure using normal-mode data. Using our final

20 case study, we demonstrate that SOLA-DLI can be utilised to obtain the coefficients of basis
21 function expansions, which leads to discretised models with specific advantages compared to
22 classical least-squares solutions. This publication is accompanied by a SOLA-DLI software
23 package that allows the interested reader to reproduce our results and to utilise the method for
24 their own research.

25 **Key words:** Inverse theory, Seismology, Seismic tomography, Structure of the Earth, Surface
26 waves and free oscillations.

27 **1 INTRODUCTION**

28 Seismic tomography relies on mathematical inversions (Rawlinson et al. 2010; Nolet 2008) to
29 model Earth’s interior from collected data. Resulting tomography models highlight persistent fea-
30 tures, such as subducted plates, rising plumes, and large scale velocity anomalies, believed to
31 mirror real Earth characteristics (Ritsema & Lekić 2020). Improving these models often involves
32 the development of new models with different data or methods to enhance the resolution of certain
33 features or to reduce uncertainties. However, seismic inversions encounter a major challenge: data
34 scarcity. This leads to non-uniqueness in solutions (e.g. Backus & Gilbert 1970), which often is
35 mitigated using regularisation. Yet, such prior information might inadvertently impose unrealistic
36 constraints or introduce artefacts in the models (e.g. Nolet 2008; Zaroli et al. 2017). While incor-
37 porating such new information is not inherently wrong, it must be accurate and well-understood
38 to avoid misinterpretations of the resulting seismic tomography models.

39 Due to the inherent non-uniqueness of seismic tomography, the chosen data often provide
40 insufficient constraints for the entire model. They may however offer useful constraints for certain
41 parts or properties of the model. The approach of seeking specific properties rather than inverting
42 for the entire model falls under the domain of mathematical inference (Valentine & Sambridge
43 2023; Tsai 2023). If all properties are known, the complete model is also known. As this is typically
44 not possible, the data limitations lead to compromises. Rather than inverting for a complete model,
45 we can aim to directly constrain a set of properties. This means we utilise all information provided

46 by the data in an optimal way, which is the core idea behind the Backus-Gilbert SOLA (Subtractive
47 Optimally Localized Averages) linear inference method (Pijpers & Thompson 1994; Zaroli 2016).

48 The Backus-Gilbert method was originally used to only compute local averages of the un-
49 known model from the data (Backus & Gilbert 1967, 1968, 1970). The issue of non-uniqueness is
50 addressed by averaging, removing the need for regularisation. In the 1970s, Backus extended this
51 to compute more general “properties” of the unknown model by fitting linear combinations of the
52 sensitivity kernels to some other predefined kernel (later termed “Target kernel” by some authors)
53 (Backus 1970a,b), but this earlier work does not delve deeper into the kind of properties one could
54 analyse. The ideas of Backus and Gilbert have been popularised and practically applied in different
55 areas, including helioseismology (e.g. Pijpers & Thompson 1994, 1992, who also coined the name
56 SOLA), deconvolution (Oldenburg 1981), geomagnetism (Hammer & Finlay 2019; Hammer et al.
57 2021), and (1D, 2D and 3D) seismic tomography (Nolet 1985; Masters & Gubbins 2003; Zaroli
58 2016, 2019; Lau & Romanowicz 2021; Latallerie et al. 2022; Amiri et al. 2023; Restelli et al.
59 2024). However, these studies have mostly concentrated on extracting local average information
60 about the unknown model. A notable exception is the paper of Pijpers & Thompson (1994) who
61 used Gaussian derivatives to extract information about gradients. Past studies have also overlooked
62 the importance of placing a norm bound on the model space, as data alone cannot sufficiently
63 constrain any property (Backus 1970a,b; Backus & Gilbert 1970). Recently, Al-Attar (2021) has
64 generalised the original methods of Backus and Gilbert by developing a rigorous mathematical
65 framework for Backus-Gilbert inferences and by highlighting the importance of the model norm
66 bound for obtaining model properties.

67 More issues arise in practical applications of BG SOLA theory. First of all, results obtained
68 using the classic method must be interpreted using the resolving kernels, which are almost always
69 distinct from one location to another, and sometimes may not resemble the target kernels at all.
70 Furthermore, with classic BG-SOLA we obtain properties of the true model rather than the true
71 model itself (or an approximation of it). In situations where our data depend on multiple physical
72 parameters (e.g. when using normal modes), the sensitivity to parameters not of interest has typi-
73 cally been accounted for by additional noise (Masters & Gubbins 2003; Restelli et al. 2024), but

74 this means prior information is assumed in the inference problem. Furthermore, given that Backus-
 75 Gilbert methods provide model properties, it is thought that the method cannot be used to develop
 76 models that cover the full spatial domain (Valentine & Sambridge 2023), making it difficult to
 77 compute the fit to data and to utilise the models in forward simulations. These issues, together
 78 with the developments on general deterministic linear inferences by Al-Attar (2021) motivate this
 79 work.

80 The paper is structured as follows. In Section 2, we present the theoretical framework, includ-
 81 ing an overview of the general theory of deterministic linear inferences (Section 2.1), an overview
 82 of the classic BG SOLA method (Section 2.2), followed by a description of the modified SOLA-
 83 DLI method that incorporates model norm bounds in a more general mathematical framework
 84 (Section 2.3). Furthermore, we show how SOLA-DLI can be expanded by considering other types
 85 of target kernels (Section 2.4) and how families of discretised models can be obtained through
 86 an appropriate choice of target kernels (Section 2.5). In Section 3, we illustrate several aspects of
 87 the SOLA-DLI method through three case studies. Section 3.1 serves to illustrate how different
 88 properties are retrieved in a fully synthetic example, Section 3.2 shows how a simple resolution
 89 and trade-off analysis can be performed using synthetic data with real sensitivity kernels, and
 90 Section 3.3 demonstrates how discretised models can be obtained with SOLA-DLI. Finally, in
 91 Section 4 we discuss the advantages and limitations of SOLA-DLI and summarise our findings.

92 **2 THEORY**

93 **2.1 Deterministic Linear Inferences**

94 Relationships between data and model parameters are often linear in seismology, or can be lin-
 95 earised using perturbation theory (Tromp et al. 2005), provided that some assumptions about the
 96 size of the perturbations apply. Therefore, let us assume that we have some error-free data d that
 97 are related to some model m via a linear relation G :

$$G(m) = d. \quad (1)$$

98 We will call such model-data relationships “deterministic data constraints”, because we assume
 99 the data to be known exactly. The model is part of the model space \mathcal{M} , and the data are part of
 100 the data space \mathcal{D} . In inversions, we attempt to find the solution model from the data by inverting
 101 the forward relation (Equation 1). However, in most cases, the forward relation cannot be inverted
 102 due to insufficient or inadequate data. For continuous models, this leads to either no solutions, or
 103 infinitely many solutions (Backus & Gilbert 1967). In the absence of measurement errors, the only
 104 case when there are no solutions, is when the data are outside the range of the forward operator and
 105 thus incompatible with the physical law. Typically, in such a situation we use a different forward
 106 relation. Throughout this paper, we will therefore assume that the data are compatible with the
 107 forward relation. Thus, we have infinitely many solutions (see Fig. 1a), which we will denote
 108 by $\{m\}$. Inversions can be performed by adding constraints (regularisations) to the model space
 109 \mathcal{M} until a single model \tilde{m} is “selected”. For example, one may choose the model that has the
 110 smallest average gradient, which corresponds in some way to the smoothest model. However, if
 111 the assumptions built into the chosen regularisation are not correct, the model will not be a good
 112 representation of reality.

113 We often seek specific properties of the true model \bar{m} rather than the entire model itself. These
 114 properties, for example the average structure over some volume within the Earth or the depths of
 115 discontinuities, belong to a distinct space known as the property space \mathcal{P} , following the work of
 116 Al-Attar (2021). Therefore, we can define a new (inference) problem as:

Given that:

$$G(\bar{m}) = d \quad (2)$$

Find:

$$\mathcal{T}(\bar{m}) = \bar{p} \quad (3)$$

117 where \mathcal{T} (the property mapping) is a linear relation that extracts a property of any model, and
 118 $\bar{p} \in \mathcal{P}$ represents the value extracted by \mathcal{T} (property) when applied to the true model \bar{m} . It can
 119 be shown that in most practical situations the desired property \bar{p} can be anything given a finite
 120 number of deterministic data constraints (Al-Attar 2021; Backus 1970a). In other words, given the

121 data constraints, \bar{p} may take any value from the property space \mathcal{P} (see Fig. 1a), leaving us unable
 122 to say anything about the property of the true model \bar{m} . Backus (1970a) and subsequently Al-Attar
 123 (2021) showed that this problem can be overcome by introducing a norm bound M in the model
 124 space:

$$\|m\|_{\mathcal{M}} \leq M \quad (4)$$

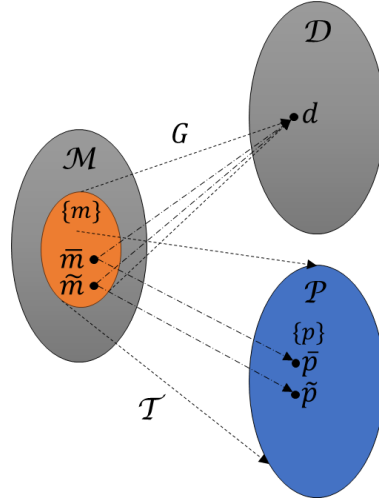
125 This constraint, applied on the model space, differs from the constraints implicitly applied during
 126 regularisations because it does not aim to single out a particular model. However, the model norm
 127 bound constraint does not lead to an invertible system, but rather limits the possible solutions to a
 128 bounded subset of \mathcal{M} . If the set of models that respect the norm bound is denoted U_M (Equation 4),
 129 then the set of solutions that respect the norm bound constraint and the data constraint is $U_M \cap \{m\}$,
 130 which is a bounded subset (Al-Attar 2021). Al-Attar (2021) further showed that such a constraint
 131 leads to the true property \bar{p} being constrained in a bounded subset $\{p\} \subset \mathcal{P}$ as long as the norm
 132 of the true model is smaller than the chosen norm bound (see Fig. 1b for a visual representation of
 133 these concepts). Without any other prior information, all properties found in $\{p\}$ are equally likely
 134 to represent the true property \bar{p} . If we choose the appropriate property mapping \mathcal{T} and prior model
 135 norm bound, we will be able to constrain the set of possible properties sufficiently for us to infer
 136 useful information about \bar{p} . Choosing the appropriate \mathcal{T} can be seen as finding the observable that
 137 can be best constrained by the data. For a more detailed and mathematical description of these
 138 ideas, the reader is referred to Section 2 of Al-Attar (2021).

139 2.2 Classic Backus-Gilbert SOLA Theory

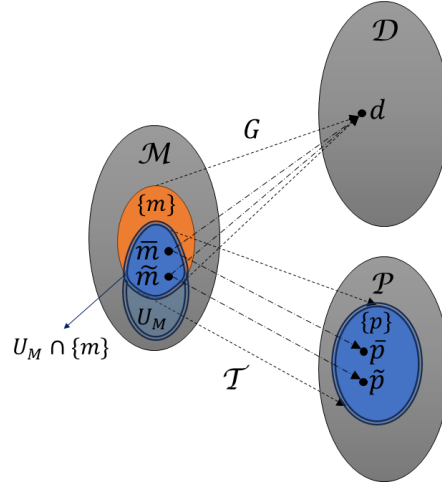
140 Equation 3 describes the general linear inference problem. In realistic applications, we will always
 141 have a finite number of data N and a finite number of physical parameters N' . In the special case
 142 of classic BG SOLA, the relationship between model and data is assumed to be:

$$d_i = G(m^j) = \sum_{j=1}^{N'} \int_{\Omega_j} K_i^j m^j d\Omega_j \quad (5)$$

143 where d_i are the (error-free) data, m^j are the various physical parameters that the data depend
 144 on (such as density, shear and compressional wave speeds, locations of discontinuities, etc.), and



(a) No norm bound applied on the model space.



(b) With a norm bound applied on the model space.

Figure 1. Schematic of general linear inference problems, illustrating the effect of bounding the model space (see Table A1 for symbols definitions). Sets with thin lines for margins represent unbounded sets, while sets with double line margins represent bounded sets. The true model is denoted by \bar{m} and the least norm model solution is denoted by \tilde{m} . When no bounds are imposed, the property of a model that respects the data constraint may take any value in the property space, which is an unbounded set (see Al-Attar 2021, Theorem 2.2). In other words, $\mathcal{T}(\{m\}) = \{p\} = \mathcal{P}$. Applying the norm bound on the model space leads to the intersection between $\{m\}$ and U_M to be bounded, which gets mapped under \mathcal{T} to a bounded subset of \mathcal{P} . In other words, $\mathcal{T}(U_M \cap \{m\}) = \{p\} \subset \mathcal{P}$

145 K_i^j are all the sensitivity kernels associated with m^j (e.g. Tromp et al. 2005; Liu & Tromp 2008;
 146 Fichtner et al. 2006). In the continuous BG SOLA formalism, we consider the physical parameters
 147 m^j and the sensitivity kernels K_i^j to be functions of space. The sensitivity kernels and the corre-
 148 sponding physical parameters are usually defined on some regions Ω_j of the Earth, which may be
 149 1D, 2D, or 3D and they live in function spaces \mathcal{M}_j . The data are always real, and the integrals are
 150 computed over the entire spatial domains over which the physical parameters are defined.

151 The Backus-Gilbert SOLA (BG SOLA) theory developed by Pijpers & Thompson (1994),
 152 Zaroli (2016, 2019), and others (Oldenburg 1981), has been used to find approximate local aver-
 153 ages of some physical parameter. Let $\tilde{p}^{l,(k)}$ be an approximate local average centered around the
 154 spatial location $\mathbf{r}^{(k)}$ of the l^{th} physical parameter m^l . This approximation is obtained only from
 155 linear combinations of the data (Equation 5):

$$\tilde{p}^{l,(k)} := \sum_{i=1}^N x_i^{(k)} d_i = \sum_{j=1}^{N'} \int_{\Omega_j} \sum_{i=1}^N x_i^{(k)} K_i^j(\mathbf{r}) m^j(\mathbf{r}) d\Omega_j \quad (6)$$

156 where $x_i^{(k)}$ are some unknown scalars that optimise the sensitivity to m^l , \mathbf{r} is the position vector,
 157 and $d\Omega_j$ is the volume/surface/line element of the domain on which m^j is defined. We further
 158 define $A^{j,(k)}$:

$$A^{j,(k)}(\mathbf{r}) := \sum_{i=1}^N x_i^{(k)} K_i^j(\mathbf{r}) \quad (7)$$

159 to be the “resolving kernel” (Zaroli 2019; Pijpers & Thompson 1994) of the physical parameter
 160 m^j at the spatial location $\mathbf{r}^{(k)}$.

161 The functions that we call resolving kernels have typically been referred to as *averaging ker-*
 162 *nels* in past studies, but we prefer to use the more general term *resolving kernel*. Our reasons for
 163 this will become more apparent in Section 2.4. If one finds the coefficients $x_i^{(k)}$ such that $A^{l,(k)}$
 164 resembles a local averaging function $T^{(k)}$ for physical parameter m^l , then $\tilde{p}^{l,(k)}$ will represent the
 165 approximate local average of the true physical parameter \bar{m}^l around the point $\mathbf{r}^{(k)}$. $T^{(k)}$ are called
 166 target kernels and represent the perfect resolving kernels we aim to obtain.

167 In practice, it is impossible to obtain $A^{l,(k)} = T^{(k)}$, but we may obtain very good approxima-
 168 tions $A^{l,(k)} \approx T^{(k)}$ when we have sufficient data coverage. If we lack sufficient data, then $A^{l,(k)}$
 169 may be significantly different from $T^{(k)}$, because the target kernels are typically linearly inde-

pendent of the sensitivity kernels. Furthermore, the resolving kernels $A^{j \neq l, (k)}$ for other physical parameters that affect the data will typically not be 0. Thus, $\tilde{p}^{l, (k)}$ will also contain contributions from physical parameters that are not of interest to us. This “contamination” is the mathematical manifestation of the trade-offs that exist between various physical parameters, such as volumetric heterogeneity and topography on internal boundaries inside the Earth.

In general, we therefore find

Approximate property of true physical parameter \bar{m}^l

$$\begin{aligned} \tilde{p}^{l, (k)} = & \int_{\Omega_l} A^{l, (k)}(\mathbf{r}) \bar{m}^l(\mathbf{r}) d\Omega_l + \\ & \sum_{j \neq l}^{N'} \int_{\Omega_j} A^{j, (k)}(\mathbf{r}) \bar{m}^j(\mathbf{r}) d\Omega_j \end{aligned} \quad (8)$$

when we really want to find

True property of true physical parameter \bar{m}^l

$$\bar{p}^{l, (k)} = \int_{\Omega_l} T^{(k)}(\mathbf{r}) \bar{m}^l(\mathbf{r}) d\Omega_l \quad (9)$$

and

$$\tilde{p}^{l, (k)} \neq \bar{p}^{l, (k)} \neq \bar{m}^l(\mathbf{r}^{(k)}). \quad (10)$$

In practical applications, for example inversions of normal mode splitting, the contaminant terms in Equation 2.2 have typically been accounted for either by scaling the contaminant physical parameters and adding them into the term containing the desired physical parameter (i.e. scaled sensitivity kernels Ritsema et al. 1999; Moulik & Ekström 2014) or by using the so called “3D noise” approach” (Restelli et al. 2024; Masters & Gubbins 2003; Lau & Romanowicz 2021). On the other hand, the complications arising from imperfect resolving kernels have been treated mostly by being cautious about the meaning of the values of $\tilde{p}^{l, (k)}$.

2.3 SOLA-DLI: Backus-Gilbert SOLA theory through linear inferences

Here, we present several modifications of the classic Backus-Gilbert (BG) SOLA method by applying the general theory of deterministic linear inferences (see Section 2.1 and Al-Attar (2021))

187 to the specific problem that SOLA is concerned with, leading to a new method that we call SOLA-
 188 DLI. By doing this, we will alleviate the issues of imperfect resolving kernels and contaminant
 189 terms at the cost of imposing a weak prior model constraint. For this, it is useful to first introduce
 190 some more general concepts.

191 An element of the model space \mathcal{M} is a tuple containing all the physical parameters considered
 192 $m = (m^0, m^1, \dots)$ (for example $m = (\rho, v_s, v_p)$). These parameters m^j are continuous variables,
 193 typically functions of space defined over some domains Ω_j . Each physical parameter lives in some
 194 space \mathcal{M}_j and for our purposes it suffices to assume that the spaces \mathcal{M}_j are Hilbert spaces. The
 195 total model space \mathcal{M} is defined as the direct sum of all constituent model spaces \mathcal{M}_j , which is
 196 also a Hilbert space. Similarly, the sensitivity kernels are also continuous variables of the same
 197 type as the physical parameters they are associated with. We can thus generalise Equation 5 to:

$$d_i = G(m) = \sum_j \langle K_i^j, m^j \rangle_{\mathcal{M}_j} \quad (11)$$

198 where $\langle \cdot, \cdot \rangle_{\mathcal{M}_j}$ are inner products defined on each space \mathcal{M}_j . We have dropped the summation
 199 limits here for brevity, and in the remainder of the manuscript we will assume that the indices of
 200 our data and physical parameters run over the following sets $i \in \{1, 2, \dots, N\}, j \in \{1, 2, \dots, N'\}$,
 201 respectively. Most often the spaces \mathcal{M}_j will be function spaces with inner products defined by:

$$\langle f, g \rangle := \int_{\Omega} f g d\Omega. \quad (12)$$

202 In some cases the physical parameter and its sensitivity kernels are represented by real numbers in
 203 which case the inner product becomes a real number multiplication (i.e. for sensitivities at internal
 204 boundaries within a 1D Earth model (Lau & Romanowicz 2021)). In this notation, a property of a
 205 model $m = (m^0, m^1, \dots)$ is defined using target kernels as:

$$p^{(k)} = \mathcal{T}(m) := \sum_j \langle T^{j,(k)}, m^j \rangle_{\mathcal{M}_j} \quad (13)$$

206 where a target kernel $T^{j,(k)}$ must be defined for each physical parameter and \mathcal{T} is the property
 207 mapping. Note that unlike in the previous subsection, we do not use the l index in $\tilde{p}^{l,(k)}$. We reserve
 208 that index for the cases when only one physical parameter (denoted m^l) has non-zero target kernels
 209 associated with it.

210 With our generalised notation, the inference problem (Equation 11) can be rewritten as:

Given

$$d_i = \sum_j \langle K_i^j, \bar{m}^j \rangle_{\mathcal{M}_j} \quad (14)$$

Find

$$\bar{p}^{(k)} = \sum_j \langle T^{j,(k)}, \bar{m}^j \rangle_{\mathcal{M}_j}. \quad (15)$$

211 Classic BG SOLA methods (assuming no data errors) would find the coefficients $x_i^{(k)}$ or simply
 212 X in matrix form (see Equation 6) such that the following cost function is minimised for all target
 213 kernels (Lau & Romanowicz 2021; Zaroli et al. 2017):

$$\Phi = \sum_j \|T^{j,(k)} - A^{j,(k)}\|_{\mathcal{M}_j}^2. \quad (16)$$

214 The true properties $\bar{p}^{(k)}$ are estimated by $\tilde{p}^{(k)}$, which are linear combinations of the data d_i weighted
 215 by the coefficients $x_i^{(k)}$ (Equation 6). The interpretation of these estimates is tied to their corre-
 216 sponding resolving kernels $A^{j,(k)}$. However, these resolving kernels often show ringing effects
 217 that make their interpretation as averaging weight functions difficult (Masters & Gubbins 2003).
 218 Furthermore, if multiple properties with identical target kernels but different spatial locations are
 219 computed, their resolving kernels will usually also be different. This leads to non-uniform inter-
 220 pretations of the properties.

221 These difficulties can be eliminated if we interpret the BG SOLA results through the target
 222 kernels, rather than the resolving kernels, as these remain unchanged. This makes us revisit the
 223 distinction between the true property $\bar{p}^{(k)}$ and the property estimate $\tilde{p}^{(k)}$. The true property has a
 224 clear meaning that is defined by the target kernels, but it cannot be found in practice. The estimate,
 225 on the other hand, can be found, but it is often difficult to interpret, since its meaning is defined
 226 by the resolving kernel. As discussed in Section 2.1, imposing a prior norm bound on the model
 227 space enables us to find an interval in which $\bar{p}^{(k)}$ is contained (Al-Attar 2021; Backus 1970a). By
 228 doing this, we effectively pay the price of imposing prior information and lose the convenience
 229 of having a single estimate for the true property, but we gain exact interpretability of our results.

230 In this manuscript, we use these concepts to modify BG SOLA giving rise to the new SOLA-DLI
231 method.

232 Following Al-Attar (2021), we assume that the norm of each true physical property is bounded:

$$\|\tilde{m}^j\|_{\mathcal{M}_j} \leq M^j \quad (17)$$

233 which leads to a norm bound for the true model \tilde{m} :

$$\|\tilde{m}\|_{\mathcal{M}} \leq \sum_j M^j, \quad (18)$$

234 where the norm on the direct sum $\mathcal{M} = \bigoplus_j \mathcal{M}_j$ is defined by $\|m\|_{\mathcal{M}} = \sum_j \|m^j\|_{\mathcal{M}_j}$ for any
235 $m \in \mathcal{M}$. Examples of such model norm bounds will be given in Section 3.

236 Imposing a norm bound leads to the following solution estimate for the true property vector \tilde{p}
237 (see Appendix A and Al-Attar (2021)):

Solution

$$\tilde{p}^{(k)} \in [\tilde{p}^{(k)} - \epsilon^{(k)}, \tilde{p}^{(k)} + \epsilon^{(k)}] \quad (19)$$

Definitions

$$\tilde{p} = \mathcal{A}(\tilde{m}) = Xd = \mathcal{T}(\tilde{m}) \quad (20)$$

$$\tilde{m} = G^*(GG^*)^{-1}d \quad (21)$$

$$\epsilon^{(k)} = \sqrt{(M^2 - \|\tilde{m}\|_{\mathcal{M}}^2) \mathcal{H}_{kk}} \quad (22)$$

238 where \mathcal{A} is the ‘‘approximate mapping’’ that maps the true model through the resolving kernels to
239 the property space (see equation A.14), \mathcal{H} effectively quantifies the difference between \mathcal{A} and \mathcal{T}
240 (see Figure A1), \tilde{m} is the least norm solution and G^* is the adjoint of G . The solution estimate
241 is a set of intervals that define the values that our property can take. For example, if the property
242 that we search is a local average at 10 spatial locations, then the solution estimate will be one
243 interval for each of the 10 local averages. Any average value within such an interval is equally
244 likely to be the real average value. Furthermore, each interval is centered on the property of the
245 least norm solution $\tilde{p}^{(k)}$. The solution intervals are dependent on the prior norm bound and the
246 difference between the target kernels and the resolving kernels as quantified by the matrix \mathcal{H} . The

247 larger their differences, the larger the intervals of our solution. For an overview of the derivations
 248 of these equations, see appendix A.

249 The introduction of the prior norm bound allows us to shift the interpretation from the resolving
 250 kernels to the target kernels. In addition, it presents an alternative to the 3D noise method for
 251 estimating the contaminant terms (Equation 2.2) that does not require an estimate of errors in the
 252 physical parameters.

253 2.4 Choice of Target Kernels

254 In this section, we will discuss how different information about the unknown model can be ex-
 255 tracted by choosing appropriate target kernels. For this, we assume the model to be a triplet of
 256 piece-wise continuous and bounded functions $m = (m^1, m^2, m^3)$ defined on the interval $[0, 1]$.
 257 This leads to a 1D inference problem, i.e. $\mathbf{r}^{(k)} = r^{(k)}$. However, the results can be easily gener-
 258 alised. The true model is assumed to be known and is plotted in Fig. 2. This model is arbitrary and
 259 has no physical significance.

260 2.4.1 Local Average Targets

261 Previous studies have primarily used the box car function as a target kernel for its simplicity and
 262 ease of interpretation - it gives a uniform local average (Restelli et al. 2024; Masters & Gubbins
 263 2003). However, many other types of target kernels could be used to obtain local averages. Here,
 264 we use three different averaging target kernels:

Uniform Local Average (for reference):

$$T_U^{(k)}(r) := \begin{cases} C & r \in V_k \\ 0 & \text{else} \end{cases} \quad (23)$$

265

Gaussian Local Average:

$$T_G^{(k)}(r) := C \exp \left[-\frac{\|r - r^k\|_2^2}{2\sigma^2} \right] \quad r \in \Omega \quad (24)$$

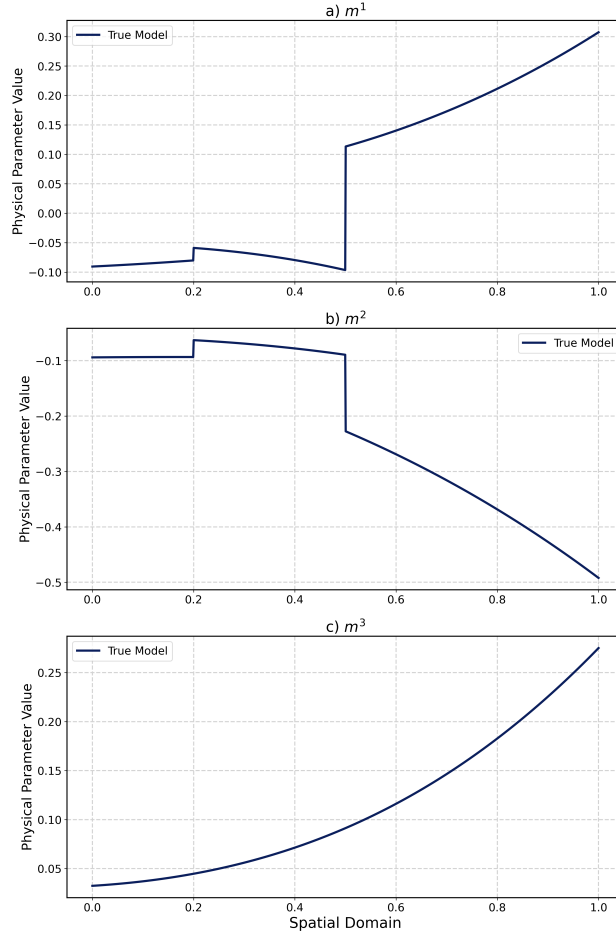


Figure 2. An arbitrary synthetic “true model”. m^j denotes the physical parameters of the model.

Bump Function Average:

$$T_B^{(k)}(r) := \begin{cases} C \exp \left[\frac{w^2}{2(r-r^{(k)})^2 - w^2} \right] & r \in V_k \\ 0 & \text{else} \end{cases} \quad (25)$$

266 where

$$V_k = \left[r^{(k)} - \frac{w}{2}, r^{(k)} + \frac{w}{2} \right] \quad (26)$$

267 is the compact support of the boxcar and bump function with width w , σ is the standard deviation of the Gaussian, and C is in each case an appropriate normalisation constant that ensures the
 268 unimodularity of each target kernel. Examples of these averaging kernels are plotted in Fig. 3. Us-
 269 ing these target kernels we can define, for example, the following property mappings for physical
 270

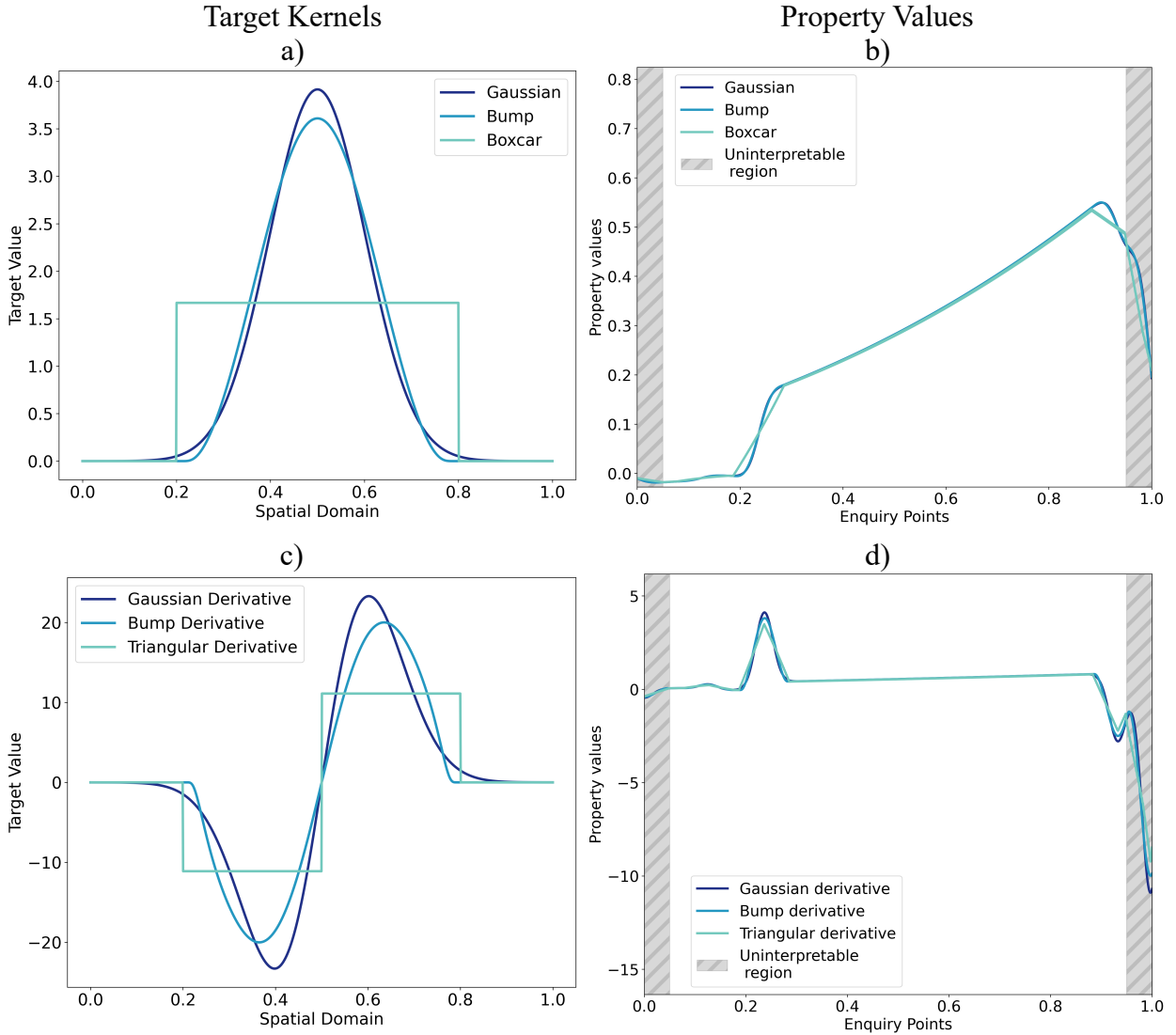


Figure 3. Examples of averaging (a) and gradient (c) target kernels and corresponding properties (b and d) obtained using these target kernels. First column: averaging and gradient target kernels with width 0.6. Second column: property values as a function of the 1000 enquiry points for different local averages and gradients of the true model using target kernels with width 0.1. The grey hatched regions represent parts of the domain where the target kernels are clipped (half-width of the target kernels).

271 parameter m^1 :

$$\bar{p}_{U/G/B}^{1,(k)} = \mathcal{T}_{U/G/B}(m) = \int_0^1 T_{U/G/B}^{1,(k)}(r) m^1(r) dr. \quad (27)$$

272 The property vector extracted by each such property mapping (uniform/Gaussian/bump) is a vector
 273 of local averages centered at a set of points $\{r^{(k)}\}$ that we call “enquiry points”. In Fig. 3 we also
 274 plot the property vectors $\bar{p}_{U/G/B}^{1,(k)}$ for 1000 evenly spaced enquiry points (right column). The grey
 275 hatched regions are parts of the domain where the target kernels $T_{U/G/B}^{1,(k)}$ are clipped and therefore

276 uninterpretable. The target kernels for physical parameters m^2, m^3 are 0 since we are not interested
 277 in these parameters.

278 Each of these target kernels has advantages and disadvantages. The boxcar function (Restelli
 279 et al. 2024) is simple and has compact support, providing a clear interpretation of the resolution of
 280 these kernels. However, most sensitivity kernels used in seismology are smooth, giving rise to poor
 281 resolving kernels that do not resemble boxcar functions. Consequently, the property error bounds
 282 are large.

283 Gaussian targets are often better reconstructed and lead to better constrained property values.
 284 They are, however, not defined on a compact domain and restricting a Gaussian to a compact
 285 domain leads to clipping. A clipped Gaussian is no longer a Gaussian, and different centering
 286 of the Gaussian leads to different clipping and therefore a “non-uniform” interpretation of the
 287 property values. If most of the Gaussian is located well within the bounds of the model domain,
 288 then the errors introduced by clipping can be negligible, but not readily quantifiable. For such target
 289 kernels we define a “width” that contains some large and arbitrary percentage of the function’s
 290 weight (such as 90%) and pretend as if the entire weight of the function is concentrated in this
 291 region.

292 Bump functions are both smooth and defined on a compact support, therefore offering the
 293 advantages of both boxcars and Gaussian targets. The one shown here is just one example of a
 294 family of functions with similar characteristics.

295 2.4.2 *Local Gradient Targets*

296 If we want to obtain some local estimate of the gradient of a 1D physical parameter such as:

$$\bar{p}^{l,(k)} = \int_0^1 T^{l,(k)} \frac{dm^l}{dr} dr \quad (28)$$

297 we can use integration by parts to obtain:

$$\bar{p}^{l,(k)} = - \int_0^1 \frac{dT^{l,(k)}}{dr} m^l dr + [T^{l,(k)} m^l]_0^1. \quad (29)$$

298 For target kernels $T^{l,(k)}$ with compact support Equation 29 will reduce to:

$$\bar{p}^{l,(k)} = - \int_0^1 \frac{dT^{l,(k)}}{dr} m^l dr \quad (30)$$

299 in the interval where the target kernels are not clipped. For any other target kernel the second term
 300 will not necessarily be 0. However, it will be very close to zero if the target kernel is centered well
 301 within the bounds of the domain. Therefore, we can use Equation 30 to define new target kernels
 302 that extract local gradients of the true model.

303 In the bottom row of Fig 3 we plot the derivative of a Gaussian, Bump, and a Triangular
 304 function. The derivative of a Boxcar gives a sum of Dirac delta distributions. These cannot be used
 305 in our framework as they do not belong to a useful Hilbert space. Instead we have opted for the
 306 derivative of a triangular function, which yields a Haar function. The true property values obtained
 307 using the different gradient target kernels are shown in row 2, column 2 of the same figure.

308 The idea of using different target kernels to extract different types of information about the
 309 unknown model has been applied previously in helioseismology by Pijpers & Thompson (1994).
 310 They used Gaussian target kernels for extracting average information, and derivatives of the Gaus-
 311 sian to extract first and higher order derivatives of the model. However, as far as we are aware, this
 312 approach has not yet been used in seismic tomography. More importantly, the work of Pijpers &
 313 Thompson (1994) regard this approach as an inversion, whereas we believe it should be considered
 314 as an inference problem instead.

315 2.5 Obtaining Discretised Models through Target Kernels

316 One perceivable downside of linear inferences, such as SOLA-DLI, is the seeming impossibility of
 317 obtaining models that cover the full spatial domain (Valentine & Sambridge 2023). We illustrate
 318 here how SOLA-DLI can in fact be used to obtain discretised models by choosing appropriate
 319 target kernels.

320 Consider a model $m \in \mathcal{M}$ related to some data $d \in \mathcal{D}$ by:

$$d_i = G(m) = \langle K_i, m \rangle_{\mathcal{M}}. \quad (31)$$

321 For the reasons already mentioned in 2.1 an inversion of this equation is ill-posed leading to an

322 infinite set of solutions $\{m\}$. A common method for fixing this issue, besides regularisation, is
 323 discretisation. Typically, a set of orthonormal basis functions $\{B_l\} \in \mathcal{M}$ is chosen and any model
 324 in \mathcal{M} is projected on the subspace formed by the span of this set, leading to a parallel m^{\parallel} and
 325 perpendicular m^{\perp} component of the model (i.e. m^{\parallel} is the component that can be expressed with
 326 $\{B_l\}$ and m^{\perp} is the residual term):

$$m = m^{\parallel} + m^{\perp} \quad (32)$$

$$m^{\parallel} = \sum_l p_l B_l \rightarrow \text{projection} \quad (33)$$

$$m^{\perp} = m - m^{\parallel} \quad (34)$$

327 where p_l are the coefficients given by the projection of m onto the basis functions:

$$p_l = \langle B_l, m \rangle_{\mathcal{M}} \quad (35)$$

328 We can then reformulate the initial inverse problem as:

$$\begin{aligned} &\text{Find } \{p_l\} \text{ s.t.} \\ &G \left(\sum_l p_l B_l \right) = d_i - G(\bar{m}^{\perp}) \end{aligned} \quad (36)$$

329 The data correction term $G(\bar{m}^{\perp})$ subtracts from the original data the component corresponding to
 330 the part of the true model that is not within the span of the basis functions. In real applications, this
 331 term can never be computed since we do not know the true model \bar{m} , nor how much of it is outside
 332 the span of $\{B_l\}$. This term is therefore typically omitted and the equation solved in practice is
 333 just given by:

$$G \left(\sum_l p_l B_l \right) = d_i \quad (37)$$

334 which, combined with Equation 31, leads to the discretised inverse problem:

$$d_i = \sum_l \langle K_i, B_l \rangle p_l \quad (38)$$

335 In the seismic tomography literature, the matrix $\langle K_i, B_l \rangle$ is often denoted by G . However, we will
 336 not use that notation here since we already have a distinct (but related) use of the letter G .

337 When the number of coefficients p_l is chosen to be smaller than the number of data, such that

Equation 38 is overdetermined, it is often solved in a least square (or regularised least square) manner to produce the coefficients $\{\hat{p}_l\}$. These are systematically different from the true coefficients $\{\bar{p}_l\}$, because the correction term $G(\bar{m}^\perp)$ is ignored. Including more data while keeping the same basis functions $\{B_l\}$ will not eliminate the systematic error caused by omitting the correction term. In order to converge to the true solution $\{\bar{p}_l\}$, one has to increase the number of data and the number of basis functions in the expansion. Increasing the number of basis functions shrinks the space in which \bar{m}^\perp resides and thus decreases the size of the correction term $G(\bar{m}^\perp)$.

SOLA-DLI can be used to tackle the same inverse problem while providing some notable advantages. The crucial step required to transform the inverse problem into a SOLA-DLI problem is to note that the coefficients $\{p_l\}$ can be viewed as values of a property p that resides in a property space \mathcal{P} . The property mapping $\mathcal{T}(m)$ is defined by choosing the target kernels to be the basis functions $\{B_l\}$:

$$\mathcal{T}(m) = \langle B_l, m \rangle_{\mathcal{M}} \quad (39)$$

Subsequently, we can introduce a prior model norm bound and apply the solution presented in Equation 19 to obtain a range of possible values for the coefficients $\{p_l\}$:

$$[\tilde{p}_l - \epsilon_l, \tilde{p}_l + \epsilon_l] \quad (40)$$

where \tilde{p}_l are the coefficients of the least norm solution to Equation 31. The set of possible coefficients from Equation 40 can be inserted into Equation 33 to form sets of possible solutions for the projection of the true model on the chosen basis span.

Assuming that the prior norm bound is larger than the true model norm, the property bounds will contain the value of the true property, i.e. coefficients $\{\bar{p}_l\}$. By adding more data, the center of the property bounds (the least norm solution \tilde{p}) approaches the true property and the widths of the bounds decrease. SOLA-DLI thus provides a way to converge to the true property while maintaining the same number of basis functions. In fact, SOLA-DLI can even be used on a single coefficient - for example a particular spherical harmonic coefficient - and it will converge to the true value as more independent data are included. This contrasts with the classical inverse solution,

362 which, in the absence of the data correction term, requires an increase in both data and basis
363 functions to converge to the true coefficients.

364 **3 APPLICATIONS**

365 We use three case studies to showcase the advantages and capabilities of the SOLA-DLI method.
366 In Case 1 (Subsection 3.1), we show the effect of the prior model norm bound and target widths
367 on the solution as well as how different types of local averages can be constrained. In Case 2
368 (Subsection 3.2), we illustrate how to perform a simple resolution and trade-off analysis with
369 SOLA-DLI. In Case 3 (Subsection 3.3), we demonstrate how discretised model solutions can be
370 obtained using SOLA-DLI, comparing the results with a least square inversion solution.

371 **3.1 Case 1: General Multiparameter Model**

372 In this completely synthetic case study, we show how SOLA-DLI is used to obtain estimates for
373 three types of local averages, finding that some averages can be better constrained than others.
374 We also illustrate how the prior information and the desired resolution change the local average
375 estimates.

376 *3.1.1 Setup*

377 We consider a 1D model space containing three physical parameters m^1, m^2, m^3 , all of which are
378 piece-wise continuous functions defined on the interval $[0, 1]$. The synthetic true model (Fig. 4) is
379 generated quasi-randomly and has no physical meaning.

380 The model-data relationship for d_i with $i \in \{1, 2, \dots, N\}$ is given by:

$$d_i = G(m^1, m^2, m^3) = \int_0^1 K_i^1(r)m^1(r)dr + \int_0^1 K_i^2(r)m^2(r)dr + \int_0^1 K_i^3(r)m^3(r)dr \quad (41)$$

381 For each physical parameter, the sensitivity kernels are also produced quasi-randomly using the
382 equation:

$$K_i^j(r) = \frac{1}{\sigma\sqrt{2\pi}} \exp\left(-\frac{(r - \mu_{i,j})^2}{2\sigma^2}\right) \sin(\omega r) \sum_q c_q (r - r_q)^2 \quad (42)$$

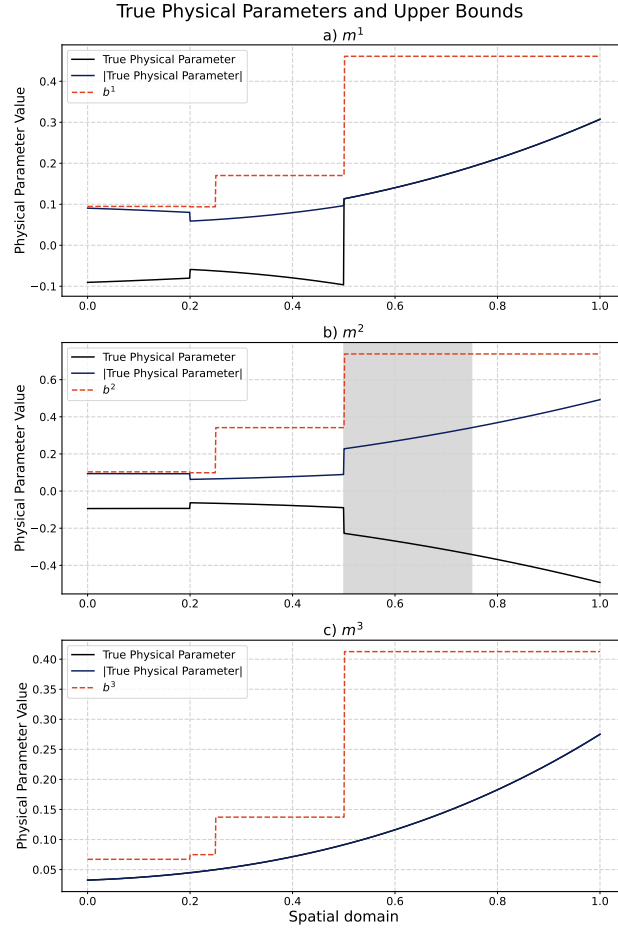


Figure 4. Case 1: True model and model norm bounds. Panels a–c show the synthetic quasi-randomly generated true model (comprised of three physical parameters) and some arbitrary piece-wise upper bound functions (b_i) used for computing the norm bound. In each panel we present both the physical parameter (black), and the absolute value of the physical parameter (blue).

383 where $\mu_{i,j}, c_q, r_q, \omega$ are randomly generated (see Fig. 5). We choose to use $N = 150$ (e.g. 150
 384 observations) with the sensitivity kernels computed for each physical parameter. To simulate the
 385 lack of data sensitivity to a particular region (e.g. no S-wave sensitivity in the Earth’s outer core),
 386 we manually set the sensitivity kernels for m^2 to zero in the interval $[0.5, 0.75]$. The synthetic
 387 (error-free) data are then produced using Equation 41 combined with the synthetic sensitivity
 388 kernels and the synthetic true model. As target kernels we use those defined in Equations 23, 24
 389 and 25 with a width of 0.2.

390 The least norm solution to this problem (Equation 41) is given by the Moore-Penrose right-

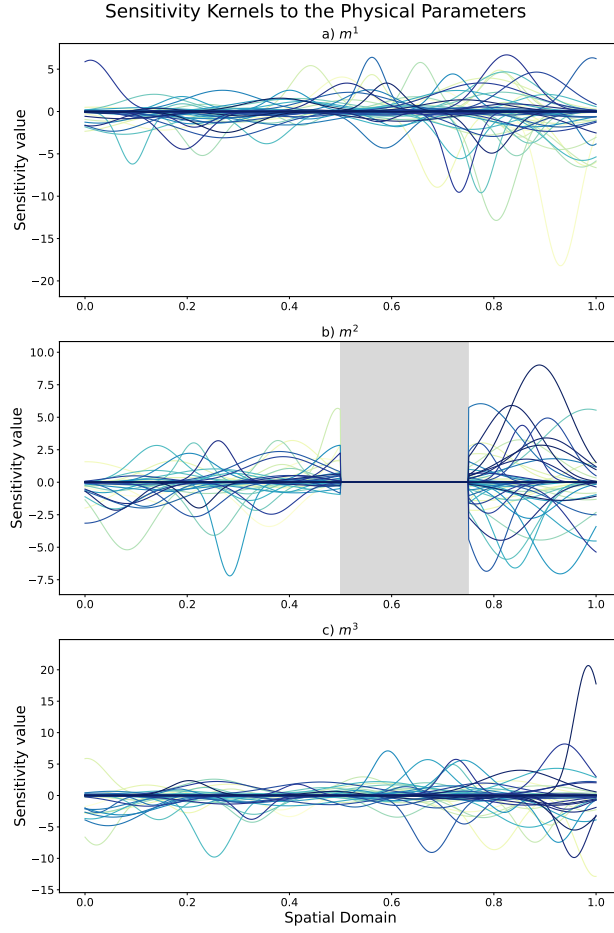


Figure 5. Case 1: Sensitivity kernels. Panels a–c show the synthetic quasi-randomly generated sensitivity kernels for physical parameters m^1, m^2, m^3 . The region with no sensitivity to m^2 is shaded in gray.

391 inverse:

$$\tilde{m} = G^*(GG^*)^{-1}d \quad (43)$$

392 and shown in Fig. 6. This is a regularized inverse solution obtained by selecting the solution with
 393 the least norm from the set of all possible solutions $\{m\}$. The lack of sensitivity in the grey region
 394 is immediately apparent in this solution, as the least norm results in a zero model value here.

395 To solve the SOLA-DLI problem, upper bound functions b^j are chosen arbitrarily (Fig. 4) such
 396 that:

$$|m^j(r)| \leq b^j(r) \quad \forall r \in [0, 1], \quad (44)$$

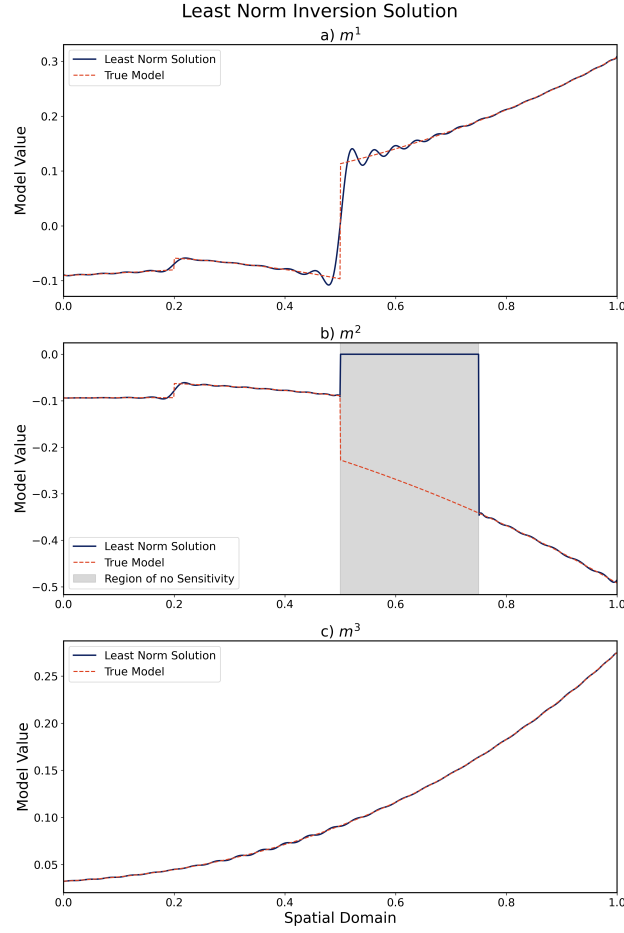


Figure 6. Case 1: Least norm solution for (m^1, m^2, m^3) obtained using Equation 41.

397 which leads to the following upper bound on the model norm:

$$\|m^j\|_{\mathcal{M}^j} = \sqrt{\int_0^1 (m^j)^2 dr} \leq \sqrt{\int_0^1 (b^j)^2 dr} = M^j \quad (45)$$

$$\|m\|_{\mathcal{M}} \leq M = M^1 + M^2 + M^3 \quad (46)$$

398 In real applications, the upper bound functions b^j should be chosen carefully based on physical
 399 arguments, for example using constraints from mineral physics.

400 3.1.2 Local averages and resolving kernels

401 In this case study, we consider three types of local averages: uniform local averages, Gaussian
 402 averages, and bump averages. Specifically, we are interested in obtaining these local averages for
 403 parameter m^2 at 100 equally spaced inquiry points in the spatial domain, with the results plotted
 404 in Fig. 7.

405 For each type of average, at each of the 100 enquiry points, the solution (Equation 19) gives an
 406 upper and a lower bound. Fig. 7 shows that the uniform local average is the least constrained prop-
 407 erty, while the Gaussian average is the best constrained property. This is not surprising considering
 408 that the sensitivity kernels are Gaussians modulated by polynomial and sinusoidal functions. If the
 409 sensitivity kernels were more similar to boxcar functions, then we may expect the uniform local
 410 averages to be better constrained. The region with no sensitivity is poorly constrained since the
 411 only constraint comes from the model norm bound.

412 In Fig. 7 we also plot, for each type of property, one set of target and resolving kernels. For
 413 Gaussian averages we observe that the resolving kernels are visually identical to the target ker-
 414 nels, which is the reason for the tight property bounds. In contrast, we retrieve visibly imperfect
 415 resolving kernels for the uniform averages, leading to large property bounds. Such large bounds
 416 ultimately make it impossible to infer any useful information about the true uniform local averages.

417 We have to be careful in comparing our SOLA-DLI results of Fig. 7 to the result of a reg-
 418 ularised inversion (e.g. the least norm solution shown in Fig. 6). The regularised solution is a
 419 member of the model space \mathcal{M} , which in this case is a tuple of piece-wise continuous and bounded
 420 functions. In contrast, the solutions obtained using SOLA-DLI are members of the property spaces
 421 \mathcal{P} (members of \mathbb{R}^{N_p} in this case, where N_p is the number of enquiry points). At each enquiry point
 422 in Fig. 7 we have a range of possible values for the desired local average of the true model solu-
 423 tion, while at each spatial point in Fig. 6 we have the tuple of values of the least norm solution.
 424 The regularised solution offers us a model spanning the full spatial domain, but it assumes that the
 425 true model has a minimal norm. The SOLA-DLI method offers bounds for several local averages
 426 of the true model solution and assumes only that the true model has a norm smaller than our norm
 427 bound.

428 *3.1.3 Effect of the prior model norm bound*

429 When we change the norm bound prior information, only the property error bounds are affected
 430 (see Equation 19). This is illustrated in Fig. 8, where we show results for three different upper
 431 bounds on the true model. Bound 3 is the most conservative, assuming a constant function three

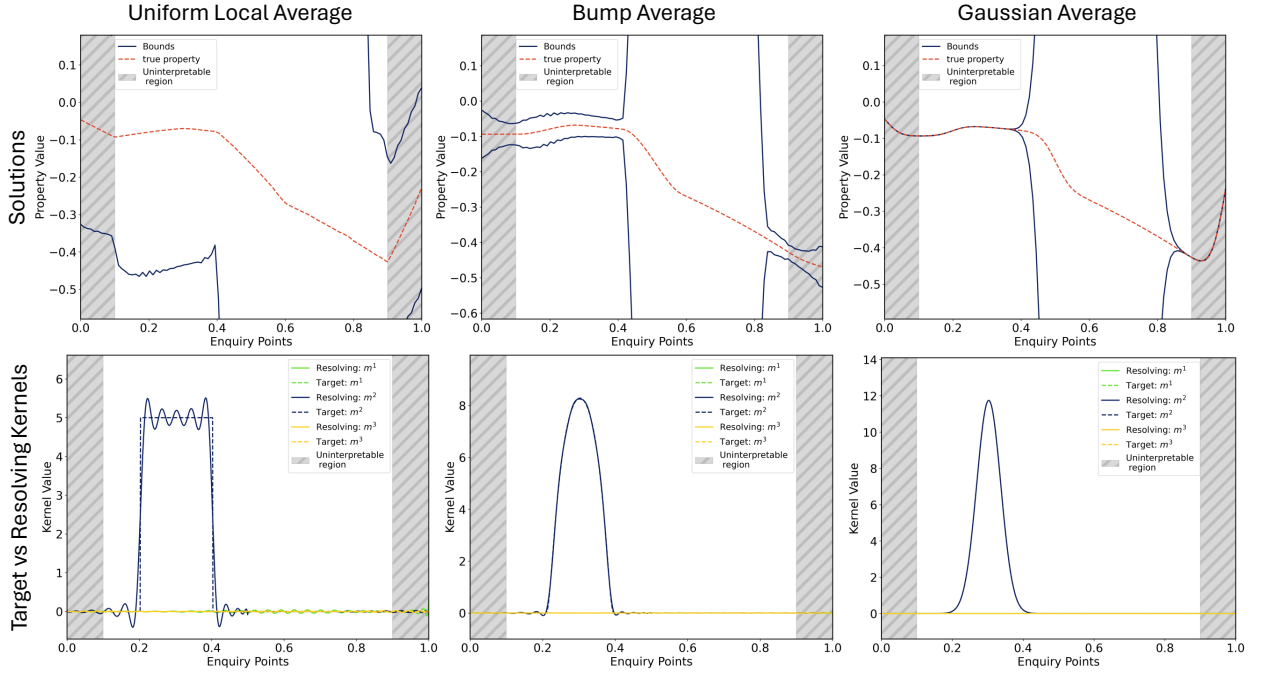


Figure 7. Case 1: SOLA-DLI solutions for three different types of local average properties. Top row: solution bounds for three types of local averages of the physical parameter m^2 at 100 evenly spaced enquiry points. Bottom row: target and resolving kernels for each type of property at the enquiry point located at $r^k = 0.3$ with width 0.2.

432 times larger than the maximum of the true physical parameter. Bounds 1 and 2 are tighter and
 433 therefore assume more prior knowledge. The bottom panel of the same figure illustrates that, as
 434 expected, tighter norm bounds lead to tighter property bounds. In all cases, the range remains
 435 centered on the true property. It is interesting to note that restricting the bounding function b^j in
 436 some local region does not lead to a tighter property bounds at an enquiry point in the same region,
 437 but rather it will lead to a uniform decrease of the property bounds at all enquiry points.

438 3.1.4 Effect of target kernel width

439 Changing the width of the target kernels can be interpreted as changing the resolution of the
 440 property evaluated at a given enquiry point. To investigate this, we have varied the target width
 441 between 1% and 100% of the domain width and computed the relative error bounds for all the
 442 enquiry points and widths. The results are plotted in Fig. 9. The relative error bound shown in the

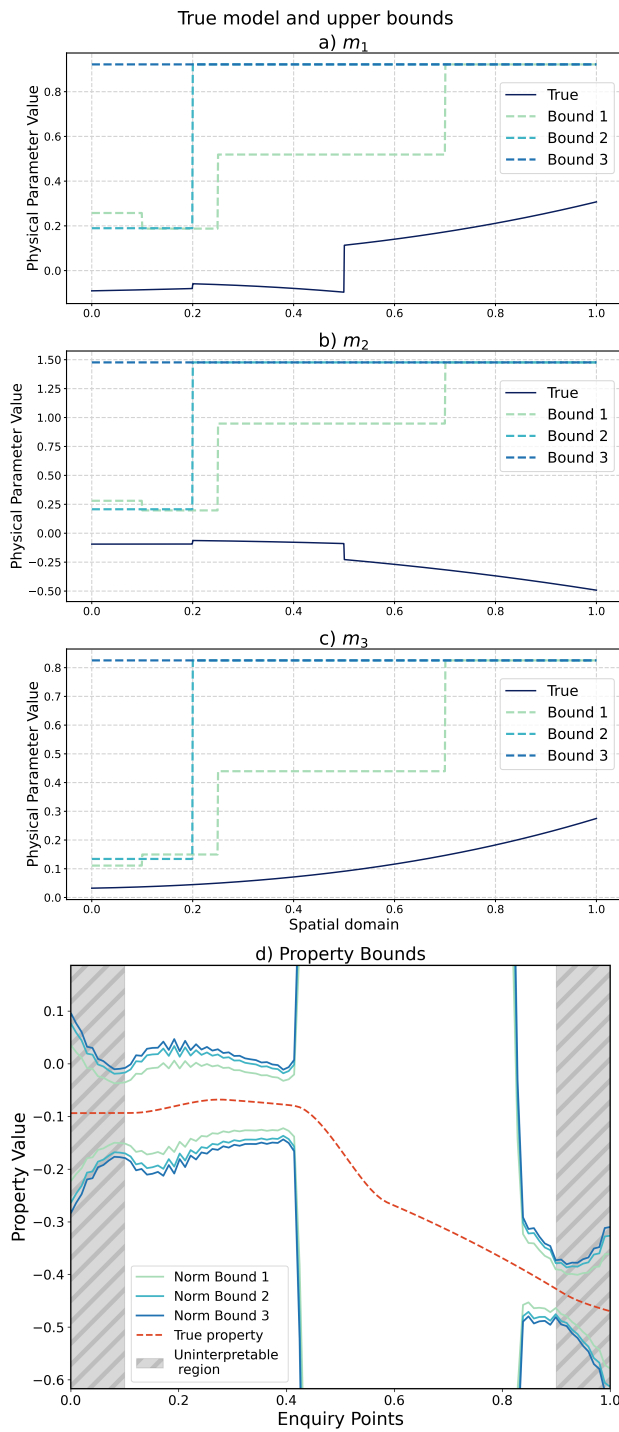


Figure 8. Case 1: Effect of prior norm bound on the property bounds. a–c) indicate the levels of three different upper bounds on the true model. Our choice of norm bound functions results in the following prior norm bounds (M^i): 2.44 (Bound 1), 3.89 (Bound 2) and 4.34 (Bound 3), which are all larger than the true model norm of 0.32. d) Solutions corresponding to the three different model upper bounds. Tighter norm bounds lead to tighter constraints on the desired properties.

443 first column is defined as:

$$e^{(k)} = \frac{\epsilon^{(k)}}{\max(\tilde{p}) - \min(\tilde{p})} \quad (47)$$

444 where \tilde{p} is the property of the least norm model solution \tilde{m} (see Equation 20). This metric has
 445 been chosen as the absolute error ϵ is not a good metric for determining whether a property is well
 446 constrained, while the classic relative error defined as ϵ/\bar{p} cannot be computed without knowing
 447 the true property \bar{p} . While there is no quantitative rule for what constitutes an unacceptable high
 448 relative property error bound, we believe any relative error higher than 100% is “certainly too
 449 high”, and relative errors less than 10% are “generally good”.

450 In general, we find that for all target kernel types the relative error bounds increase when
 451 we decrease the width of the target kernel (i.e. increase resolution). In addition, regions with no
 452 sensitivity always lead to large relative errors. As expected, the width of the uninterpretable regions
 453 at the edge of the domain increases with larger target kernel width (decreasing resolution) as the
 454 half-width of the kernel increases. Finally, in this setup, we find that particular properties, e.g.
 455 Gaussian averages, are constrained better (i.e. lower relative error bounds) than uniform or bump
 456 local averages, for all enquiry points and for target kernel widths (i.e. all resolutions), likely due
 457 to the Gaussian-like sensitivity kernels used.

458 This case study illustrates the general notion that we typically use inference methods to answer
 459 specific questions about a true model rather than finding the entire model itself. In SOLA-DLI,
 460 these questions are encoded in our chosen target kernels. The differing extent to which we are able
 461 to retrieve the target kernels effectively shows that our data can answer some questions better than
 462 others.

463 3.2 Case 2: Quasi Synthetic Normal Mode Application

464 In this quasi-synthetic case study, we show how to conduct a simple resolution analysis without
 465 real data or model values nor any prior information, based solely on the sensitivity kernels of the
 466 data set. We also illustrate how the results of such a resolution analysis can be linked to trade-offs
 467 between physical parameters.

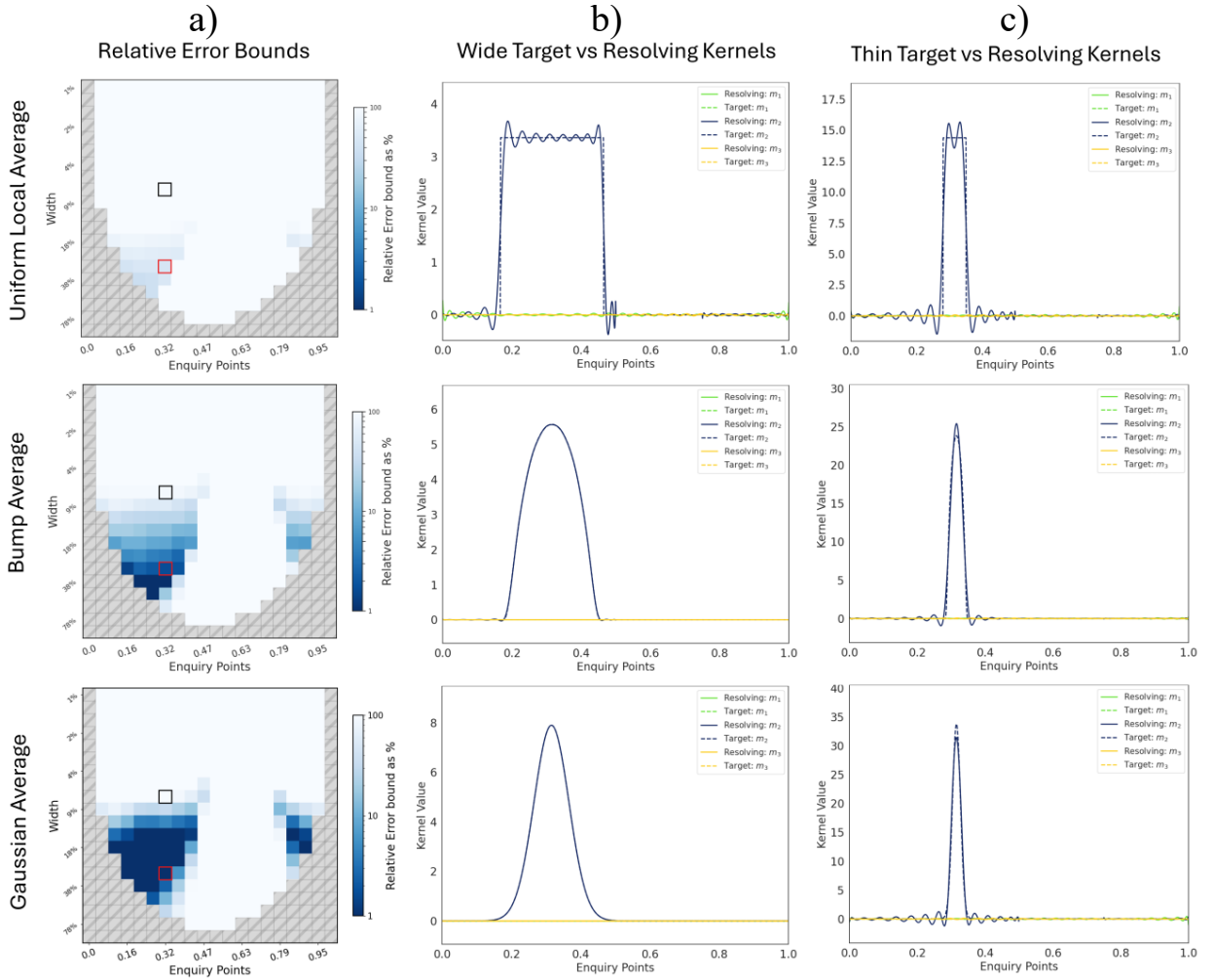


Figure 9. Case 1: Relative error bounds with examples of resolving kernels compared to their target kernels. The different rows correspond to different type of target kernels, e.g. uniform local average (top), bump average (middle) and Gaussian average (bottom). In the three columns we show the relative error bounds e (a); examples for wide target and resolving kernels, corresponding to the red squares (b) and examples for narrow target and resolving kernels (c), corresponding to the black squares.

468 3.2.1 Setup

469 Here, we consider a model formed by the triplet $m = (\delta \ln(v_s), \delta \ln(v_p), \delta \ln(\rho))$, where v_s is shear
 470 wave speed, v_p is compressional wave speed, and ρ is density (Fig. 10). Each physical parameter
 471 is assumed to be a piece-wise continuous function defined over the interval $[0, R_E]$ where R_E is
 472 Earth's radius (approximately 6371 km). We aim to constrain Gaussian averages and gradients of
 473 this synthetic true model using realistic normal mode sensitivity kernels (Woodhouse & Dahlen
 474 1978). Specifically, we select the same modes as in the SP12RTS dataset (Koelemeijer et al. 2016;

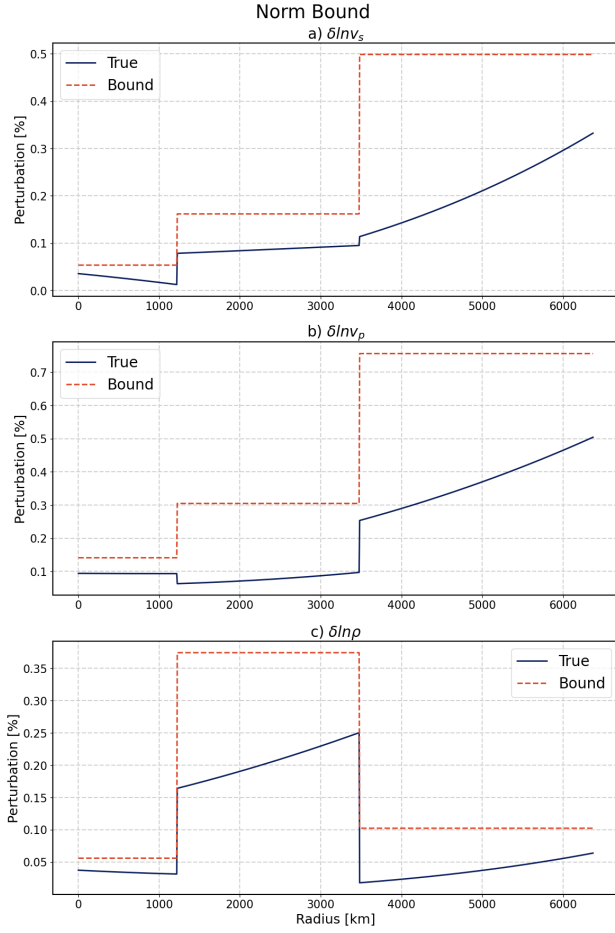


Figure 10. Case 2: Arbitrary quasi-random synthetic true model and the upper bound functions used to compute the prior upper bound norm.

475 Restelli et al. 2024), i.e. 143 modes with their sensitivity to $\delta \ln(v_s)$, $\delta \ln(v_p)$ and $\delta \ln(\rho)$ concn-
 476 trated mostly in the mantle (see Fig. 11).

477 3.2.2 Resolution Analysis

478 Before introducing any data or model values, we are able to perform a simple resolution analysis to
 479 investigate where and at what resolution our data contain information regarding the Earth model.
 480 While the SOLA-DLI solution depends on the model norm bound via M (see Eq. 19), indirectly on
 481 the data via $\|\tilde{m}\|_{\mathcal{M}}$ (see Eq. 21), and on the relationships between the target kernels and sensitivity
 482 kernels via \mathcal{H} (see Eq. A.14–A.23), the resolving kernels only depend on the data geometry, i.e.
 483 data sensitivity kernels.

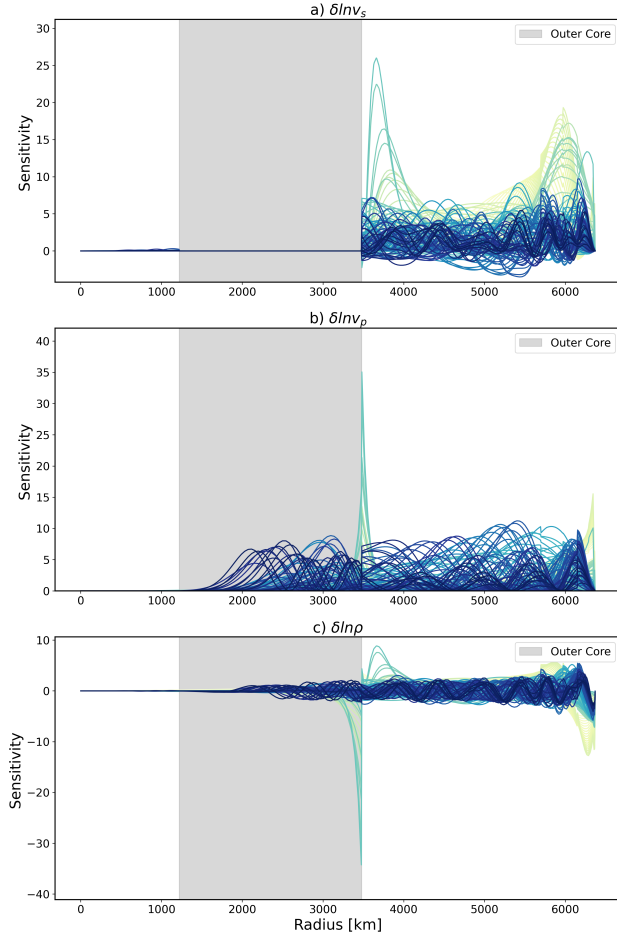


Figure 11. Case 2: Normal mode sensitivity kernels for a) $\delta \ln(v_s)$, b) $\delta \ln(v_p)$ and c) $\delta \ln(\rho)$, obtained using a modified version of OBANI. The shaded region indicates the depth range of the outer core, where the sensitivity to v_s is zero.

484 The diagonal elements of the matrix \mathcal{H} can be shown to equal:

$$\mathcal{H}_{kk} = \sum_j \|T^{j,(k)} - A^{j,(k)}\|_{\mathcal{M}_j}^2 \quad (48)$$

485 which essentially quantifies the cumulative difference between our target and resolving kernels.

486 Using \mathcal{H} we can also define the resolving misfit as a more useful metric:

$$R_k = \frac{\sqrt{\mathcal{H}_{kk}}}{\sum_j \|T^{j,(k)}\|_{\mathcal{M}_j}} = \frac{\sqrt{\sum_j \|T^{j,(k)} - A^{j,(k)}\|_{\mathcal{M}_j}^2}}{\sum_j \|T^{j,(k)}\|_{\mathcal{M}_j}} \quad (49)$$

487 which is a generalisation of the “resolution misfit” as defined in Restelli et al. (2024). The resolving

488 misfit is 0 when all the resolving kernels associated with some property evaluated at $r^{(k)}$ are equal

489 to the corresponding target kernels. This would mean that our data contain exact information about

490 the desired property and the property error bounds are 0. On the other hand, the resolving misfit

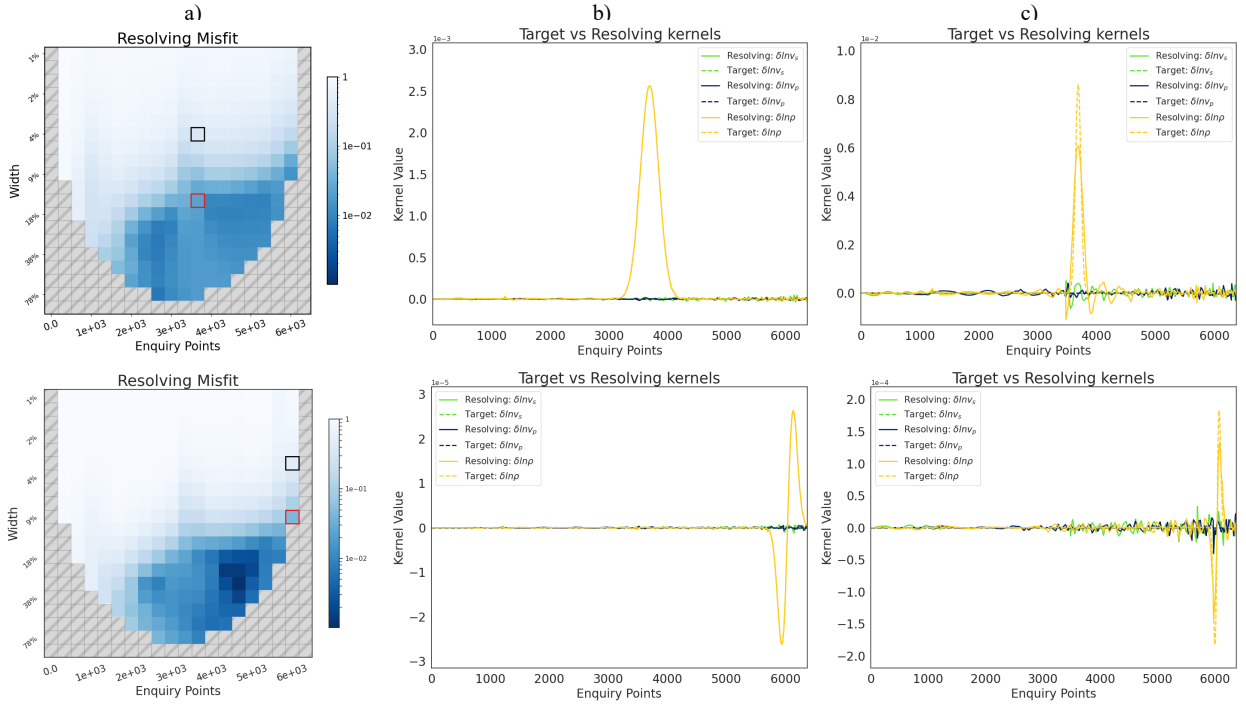


Figure 12. Case 2: Resolution analysis for a Gaussian average (top) and gradient (bottom) target for $\delta \ln(\rho)$ using realistic mode sensitivity kernels. The resolving misfit (left) can be computed without the need for data or any prior norm bound information. The middle and right panels illustrate the target and resolving kernels for a wide and thin target, including the contaminant kernels that indicate trade-offs between parameters.

491 is equal to 1 when our resolving kernels are zero, which would correspond to a complete lack of
 492 sensitivity of our data to the desired property. It is important to note that the computation of the
 493 resolving misfit does not use the data vector d nor any prior model information, it only uses the
 494 “geometry of the data set”, i.e. the sensitivity kernels (Latallerie et al. 2024). Fig. 12 illustrates the
 495 information that is provided by the resolving misfit (left column). As indicated by a low resolving
 496 misfit (darker shades of blue), our data mostly contain information in the mantle, as expected for
 497 this set of sensitivity kernels. The resolving misfit is also typically low for wide target kernels
 498 (over 18% domain width, or over 1000 km). Wide gradient kernels can be better recovered in the
 499 mantle, while wide averaging kernels can be better recovered in the lower outer core, indicating
 500 that our choice of target (i.e. property) is important.

3.2.3 Trade-offs between physical parameters

When our data are sensitive to two or more physical parameters, it may become difficult or impossible to obtain properties of a single parameter in isolation from the others. These trade-offs between physical parameters pose problems for interpretations, particularly in regions such as the lower mantle where the sensitivity of normal modes to seismic velocities and density is similar.

Our setup with SOLA-DLI, where we explicitly set the target kernels for parameters not of interest to zero, enables us to easily visualise and consider model parameter trade-offs. Suppose we are interested in some local property of $\delta \ln(\rho)$, for example the Gaussian local average in the deep mantle or Gaussian gradient in the mantle transition zone (Fig. 12). If we choose low resolution (wide) target kernels (middle column), we find that the resolving kernels for $\delta \ln(\rho)$ match the target kernels well. Furthermore, the resolving kernels for $\delta \ln(v_s)$ and $\delta \ln(v_p)$ also match their respective target kernels, which are just zero. Such zero or near zero resolving kernels indicate that the trade-off between the physical parameter of interest and the other physical parameters is small. However, if we choose higher resolution (thin) target kernels (right column), we notice that the resolving kernels are struggling to match their respective target kernels. The resolving kernels for $\delta \ln(v_p)$ and particularly $\delta \ln(v_s)$ are far from zero, indicating significant trade-offs with the desired property of $\delta \ln(\rho)$, which are regarded as *contaminants*. Such trade-offs between physical parameters are naturally taken into account by SOLA-DLI and typically result in higher error bounds on the property. If instead we would account for the sensitivity to $\delta \ln(v_p)$ and $\delta \ln(v_s)$ by scaling the sensitivity kernels, we would obtain tighter bounds, at the expense of assuming more prior information.

3.3 Case 3: Discretised inversions using continuous SOLA-DLI

This case study serves to illustrate how we can obtain a family of discretised model solutions using SOLA-DLI, which we compare with a typical least-squares inversion model solution.

525 3.3.1 Setup

526 Here, we consider a model m with only one physical parameter, denoted also m (see the true
527 model in Fig. 13 a)). Our model space \mathcal{M} is $PCb[0, 1]$ and the data are given by:

$$d_i = \langle K_i, m \rangle_{\mathcal{M}} \quad (50)$$

528 where K_i are some quasi-randomly functions, generated again using Eq. 42 (see Fig. 13 b)).

529 In this setup, we choose to discretise the model using a Fourier expansion. The resulting basis
530 functions are (see Fig. 13 c)):

$$B_l(r) = \begin{cases} 1, & l = 0 \\ \sqrt{2} \sin(2\pi \frac{l+1}{2} r), & l \text{ odd} \\ \sqrt{2} \cos(2\pi \frac{l}{2} r), & l \text{ even} \end{cases} \quad (51)$$

531 and a possible model expansion with Fourier coefficients p_l is given by:

$$m(r) \approx \sum_l p_l B_l(r). \quad (52)$$

532 The discretised model–data relation used for the least-squares inversion is:

$$d_i = \sum_l \langle K_i, B_l \rangle_{\mathcal{M}} p_l = \sum_l \Gamma_{il}^* p_l, \quad (53)$$

533 where (see also Appendix A2):

$$\Gamma_{ij}^* = \langle K_i, B_j \rangle_{\mathcal{M}}. \quad (54)$$

534 This leads to the following least-squares solution for p_l :

$$\hat{p} = (\Gamma \Gamma^*)^{-1} \Gamma d \quad (55)$$

535 Using the least-squares solution $\{\hat{p}_l\}$, we can thus find the corresponding model solution by using
536 the Fourier expansion:

$$\hat{m} = \sum_l \hat{p}_l B_l. \quad (56)$$

537 To obtain the SOLA-DLI solution, we consider the Fourier coefficients p_l to be elements of a
538 property vector obtained from the property mapping:

$$p_l = \mathcal{T}(m) = \langle B_l, m \rangle_{\mathcal{M}}. \quad (57)$$

539 We also introduce a prior model norm bound (see Fig. 13 a)). This leads to the following SOLA-
540 DLI problem:

Given

$$d_i = \langle K_i, m \rangle_{\mathcal{M}} \quad (58)$$

Find

$$p_l = \langle B_l, m \rangle_{\mathcal{M}}. \quad (59)$$

541 This problem is readily solved using Eq. 19 to obtain upper and lower bounds for the possible
542 values of the Fourier coefficients:

$$\bar{p}_l \in [\tilde{p}_l - \epsilon_l, \tilde{p}_l + \epsilon_l]. \quad (60)$$

543 In this case \tilde{p}_l are the Fourier coefficients of the least norm solution to Eq. 50, which are not to
544 be confused with the least-squares solution \hat{p}_l . The property bounds obtained from SOLA-DLI
545 (Equation 60 offer a family of solutions that can be sampled.

546 3.3.2 *Discretised least-squares vs SOLA-DLI solution*

547 We compute both the least-squares and discretised SOLA-DLI solution using different number of
548 data points (50, 70 or 100), solving for 29 Fourier coefficients. SOLA-DLI provides bounds on the
549 Fourier coefficients, and we therefore have to draw samples from the distribution for each Fourier
550 coefficient distribution to obtain a model solution. Fig. 14 shows the results for both the Fourier
551 coefficients and the resulting model solutions, including two random samples for the SOLA-DLI
552 solution. When using few data (Fig. ??, the least-squares inversion generally struggles to retrieve
553 Fourier coefficients close to the true ones, while the bounds of the SOLA-DLI solution always
554 encompass the true property. However, for certain Fourier coefficients, and in certain parts of the
555 model, the least-squares solution appears to approach the true property (Fourier coefficients) and
556 the true model better than the SOLA-DLI solution. Increasing the number of data to 70 leads to a
557 better least-squares solution, especially for the first 10 Fourier coefficients, and tighter bounds of
558 the SOLA-DLI solution. However, it now becomes clear that the SOLA-DLI bounds offer more
559 accurate information, always encompassing the true Fourier coefficients and better resembling the

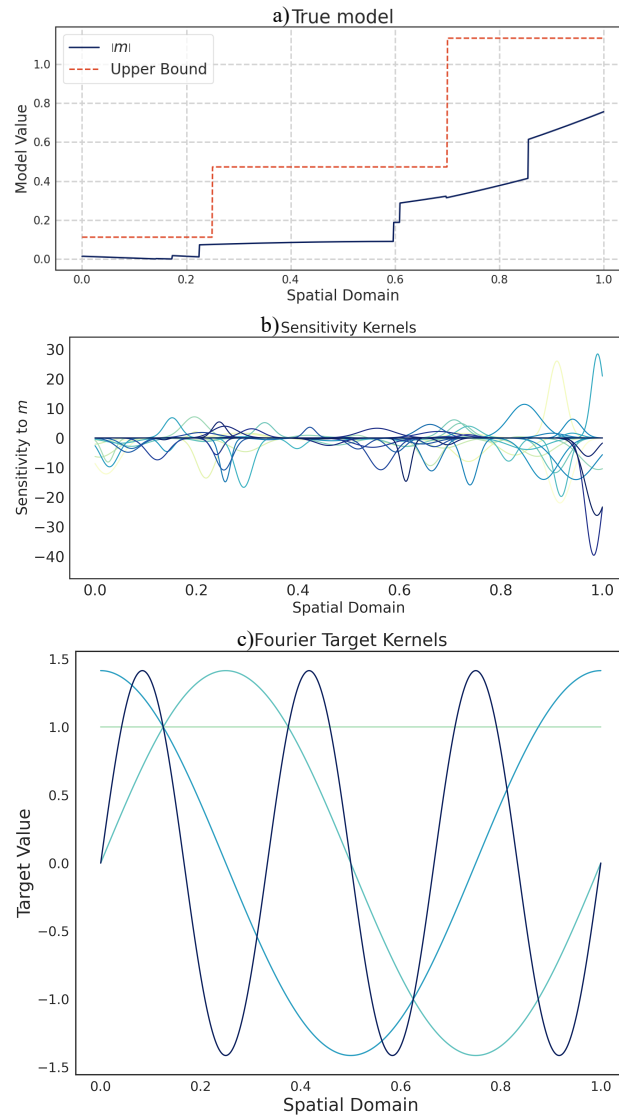


Figure 13. Case 3: True model m with the upper bound function used to compute the prior model norm bound. Case 3: Synthetic quasi-random sensitivity kernels. Case 3: Examples of four Fourier basis functions.

560 true model. When we further increase the number of data points to 100 (see Figs. ?? and ??), we
 561 note that the SOLA-DLI solution converges to the true Fourier coefficients and model, while the
 562 least-squares inversion systematically deviates.

563 In our synthetic setup, it is possible to compute the data correction term, that captures the
 564 components of the true model that are not within the span of the basis functions (see Section 2.5,
 565 Eq. 2.5). When we correct the data, using our knowledge of the true model, we find that the least-
 566 squares inversion solution converges to the true property (Fourier coefficients), even for few data
 567 (Fig. 15). This demonstrates the equivalence of the discretised least-squares and SOLA-DLI solu-

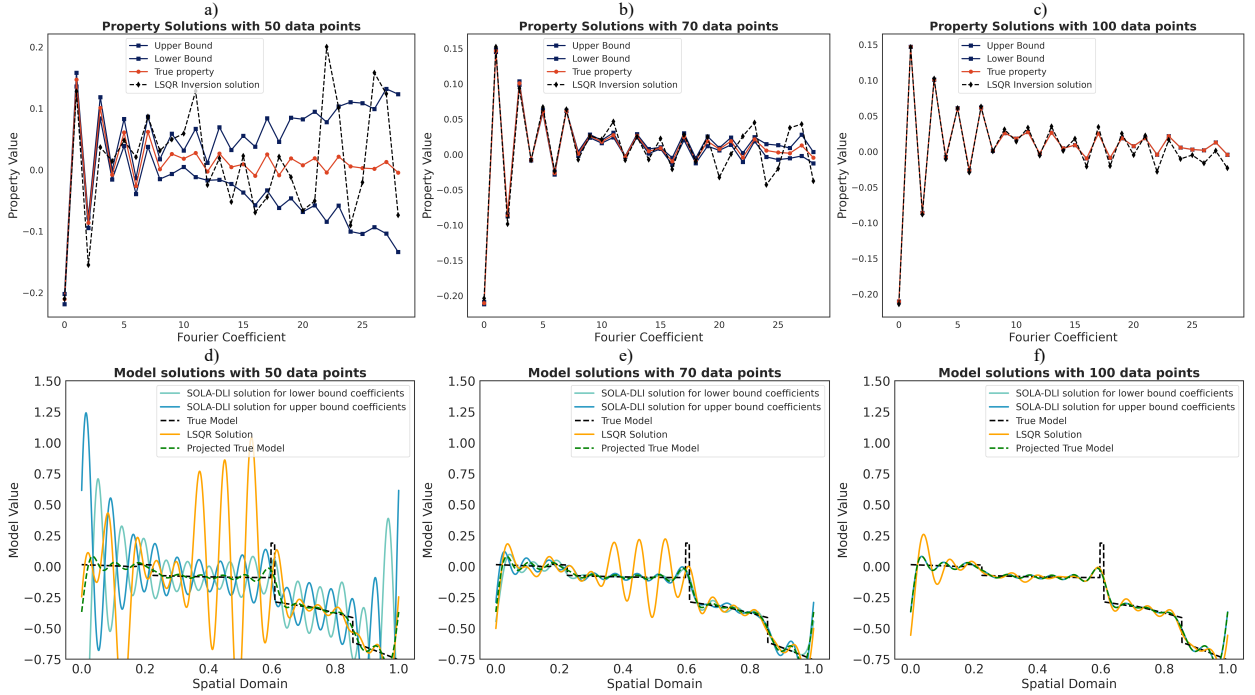


Figure 14. Case 3: Comparison between least-squares and SOLA-DLI solution for a model discretised using Fourier basis functions. a–c) Fourier coefficients from discretised least-squares inversion and SOLA-DLI using a) 50 data points, b) 70 data points and c) 100 data points. d–f) Discretised model solution from discretised least-squares and SOLA-DLI inversion using d) 50 data points, e) 70 data points and f) 100 data points

568 tions. However, in real world applications, when the true model is unknown, this data correction
 569 term cannot be computed. Consequently, the SOLA-DLI solution is preferred over the discretised
 570 least-squares inversion method.

571 4 DISCUSSION

572 Inference methods such as Backus-Gilbert and SOLA-DLI are centered around specific questions
 573 that we wish to answer, in contrast to inversion methods that typically aim to obtain the entire
 574 model. The questions we wish to answer guide our choices of specific target kernels. We have
 575 illustrated several possibilities in this contribution (e.g. uniform averages, Gaussian averages or
 576 gradients), but there are many other useful choices of target kernels. An obvious next step is the
 577 implementation of 2D and 3D target kernels that extend beyond the classic 2D and 3D uniform
 578 averaging kernels (e.g. Zaroli 2016; Latallerie et al. 2022; Freissler et al. 2024). Utilising ideas

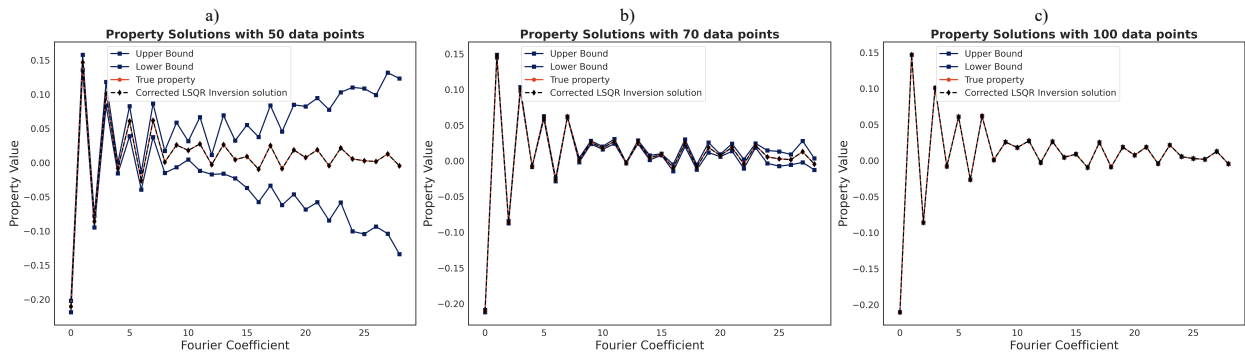


Figure 15. Case 3: Comparison between Fourier coefficients obtained using SOLA-DLI and the discretised least-squares method with a data correction term, using a) 50 data points, b) 70 data points and c) 100 data points. A comparison with the least-squares solutions in Fig. 14 indicates that the data correction leads to the systematic error in the Fourier coefficients.

579 from signal processing, we can think of targets intended to detect 3D spatial gradients, spherical
 580 harmonic coefficients, or unique target kernels to detect shape specific features inside the Earth.
 581 We also foresee applications of the theory in other geophysical disciplines, e.g. geomagnetic and
 582 geodetic imaging studies.

583 While the choice of target kernels depends in first instance on the question we wish to answer,
 584 it is also important to consider the geometry of our data (i.e. sensitivity kernels). Certain types of
 585 local averages can be better constrained depending on the details of our sensitivity kernels, e.g.
 586 uniform local averages are difficult to constrain with smooth normal mode kernels (Section 3.2).
 587 It then depends on the specifics of the question we want to answer, whether we may be able to
 588 find slight variations of the question whose associated target kernels are more easily reconstructed
 589 from linear combinations of the sensitivity kernels. While this might lead to more mathemati-
 590 cally complex target kernels, the interpretation of the results will always remain well defined. One
 591 could even use families of target kernels that contain hyperparameters found by optimising the
 592 ability of our sensitivity kernels to reconstruct them. Such approaches can be extended and lead to
 593 applications in observable optimisation - finding the observable best constrained by our data.

594 The introduction of prior model information via the model norm bound is of great importance
 595 in the SOLA-DLI method. The model norm bound (L_2 norm) chosen here is the most common due
 596 to its mathematical simplicity, but as pointed out by Al-Attar (2021), there might be better prior

607 constraints. Other norms for our model space can be thought off that allow to place bounds on the
608 maximum point-value of the true model, or its gradients. Such modifications necessitates the use
609 of more general spaces than Hilbert spaces, which add significant theoretical complications. We
600 refer the interested reader to Al-Attar (2021) for the theoretical modifications needed.

601 An important aspect that has not been used or addressed in this contribution is the presence
602 of data errors, i.e. an assumption of perfect (true) data with zero uncertainty. In such a situation,
603 the data provide deterministic constraints, hence the name “DLI”. In contrast, if there exists a
604 probability measure defined on the data space and the true data are uncertain, then we would be
605 in a probabilistic constraint case. Past applications have included data measurement errors using
606 various variations of classic Backus-Gilbert methods (Backus & Gilbert 1970; Zaroli 2016, 2019;
607 Freissler et al. 2020; Latallerie et al. 2022; Restelli et al. 2024). However, tackling them requires
608 added mathematical complexity which, we believe, would have distracted from the main points of
609 this paper. We believe that the main ideas presented here will not change by the introduction of
610 data errors, but the output of SOLA-DLI problems will be affected. Notably, the property bounds
611 will increase or might be replaced by probability density functions, as they will contain the effects
612 of imperfect resolving kernels and propagated data errors. Future work will focus on the modi-
613 fications to the theory that are required for dealing with real data, whose uncertainties cannot be
614 neglected.

615 **5 CONCLUSION**

616 In this contribution, we have presented the theory and possible applications of SOLA-DLI, which
617 combines the Backus-Gilbert SOLA method with deterministic linear inferences. Through three
618 synthetic case studies with error-free data, we have illustrated the abilities and advantages of
619 SOLA-DLI over other inverse and inference methods. Specifically, we have shown how the method
620 can be used to constrain different properties of an unknown model, defined by different target ker-
621 nels. By using a norm bound on the model space, the values of these properties can be constrained
622 and interpreted through the lens of the target kernels, rather than the resolving kernels. Using a set
623 of realistic normal-mode data sensitivity kernels, we have also shown how a resolution analysis can

624 be performed using SOLA-DLI and how trade-offs between physical parameters can be visualised
625 and quantified. Finally, by selecting a set of basis function coefficients as our target properties, we
626 have demonstrated that we can infer robust bounds on the coefficients of the unknown model's
627 projection onto these basis functions. This enables us to generate families of discretised model
628 solutions to an equivalent inversion problem, showing that the development of Earth models span-
629 ning the full spatial domain can be obtained using Backus-Gilbert methods. While the presented
630 work assumes error-free data, data errors can be included into the framework in several ways and
631 are expected to primarily lead to an increase in the property error bounds.

632 **6 ACKNOWLEDGEMENTS**

633 AMM is funded by a NERC DTP studentship NE/S007474/1 and gratefully acknowledges their
634 support. AMM also received funding from Royal Society grant RF\ERE\210182 awarded to PK.
635 CZ acknowledges financial support from ITES (Institut Terre et Environnement de Strasbourg,
636 UMR 7063) for a research visit to Oxford. PK acknowledges financial support from a Royal So-
637 ciety University Research Fellowship (URF\1\180377). Normal mode sensitivity kernels were
638 computed using a modified version of OBANI, which is based on the work of Woodhouse &
639 Dahlen (1978). The authors thank Sam Scivier and Franck Latallerie for fruitful theoretical dis-
640 cussions. The following Python packages were used extensively for producing and plotting the
641 synthetic data: Scipy (Virtanen et al. 2020), Numpy (Harris et al. 2020) and Matplotlib (Hunter
642 2007). For the purpose of Open Access, the author has applied a CC BY public copyright licence
643 to any Author Accepted Manuscript (AAM) version arising from this submission.

644 **7 DATA AVAILABILITY**

645 The sensitivity kernels and all the codes used to produce the figures in this paper can be found at
646 https://github.com/Adrian-Mag/SOLA_DLI (Mag et al. 2024).

647 **APPENDIX A: DERIVATION OF THE SOLA-DLI SOLUTION**

648 In this appendix we derive the equations and equalities needed for the solution of the SOLA-DLI
649 method presented in Section 2.3.

650 **A1 Data-model relationships**

651 Given a model $m = (m^1, m^2, \dots)$, the relationship between the model and data is defined by:

$$d_i = G(m) = \sum_j \langle K_i^j, m^j \rangle_{\mathcal{M}_j}. \quad (\text{A.1})$$

652 It is useful to define:

$$G^j(m) = \langle K_i^j, m^j \rangle_{\mathcal{M}_j} \quad (\text{A.2})$$

$$G = \sum_j G^j. \quad (\text{A.3})$$

653 The adjoint of G is defined by:

$$\langle G(m), d' \rangle_{\mathcal{D}} = \langle m, G^*(d') \rangle_{\mathcal{M}} \quad (\text{A.4})$$

654 for all $m \in \mathcal{M}$ and $d' \in \mathcal{D}$. We expand the LHS:

$$\sum_i \sum_j \langle K_i^j, m^j \rangle_{\mathcal{M}_j} d'_i = \langle m, G^*(d') \rangle_{\mathcal{M}} \quad (\text{A.5})$$

655 For the RHS, we use the formula for the inner product in the direct sum space \mathcal{M} :

$$\langle a, b \rangle_{\mathcal{M}} = \sum_j \langle a^j, b^j \rangle_{\mathcal{M}_j} \quad (\text{A.6})$$

656 where a, b are some members of \mathcal{M} . Therefore, we write

$$\sum_i \sum_j \langle K_i^j, m^j \rangle_{\mathcal{M}_j} d'_i = \sum_j \langle m^j, G^{j*}(d') \rangle_{\mathcal{M}_j} \quad (\text{A.7})$$

657 Taking the sum over i inside, we can also write:

$$\sum_j \left\langle \sum_i d'_i K_i^j, m^j \right\rangle_{\mathcal{M}_j} = \sum_j \langle m^j, G^{j*}(d') \rangle_{\mathcal{M}_j} \quad (\text{A.8})$$

658 and we identify:

$$G^{j*}(d') = \sum_i d'_i K_i^j \quad (\text{A.9})$$

$$G^* = (G^{1*}, G^{2*}, \dots) \quad (\text{A.10})$$

659 G^* maps elements from the data space to elements (tuples) in the model space. A similar approach
660 shows that the adjoint of the property mapping \mathcal{T} is given by:

$$\mathcal{T}^{j*}(p) = \sum_k p^{(k)} T^{j,(k)} \quad (\text{A.11})$$

$$\mathcal{T}^* = (\mathcal{T}^{1*}, \mathcal{T}^{2*}, \dots) \quad (\text{A.12})$$

661 A2 H matrix

662 The \mathcal{H} matrix introduced in Section 2.3 quantifies the difference between the target and resolving
663 kernels. It is defined by Al-Attar (2021, see Equation 2.84) as:

$$\mathcal{H} = HH^* \quad (\text{A.13})$$

$$H = \mathcal{T} - \mathcal{A} \quad (\text{A.14})$$

664 where \mathcal{A} is the “approximate mapping”, given by:

$$\mathcal{A} = \mathcal{T}G^*(GG^*)^{-1}G \quad (\text{A.15})$$

665 This mapping takes any model $m \in \mathcal{M}$ into the data space $d \in \mathcal{D}$, then finds the least norm
666 solution to $G(m) = d$ and maps this least norm solution into the property space. When applied to
667 one of the possible model solutions $U_M \cap \{m\}$, it gives the property of the model solution that has
668 the smallest norm. Combining (A.13), (A.14), and (A.15) we obtain:

$$\mathcal{H} = (\mathcal{T} - \mathcal{A})(\mathcal{T} - \mathcal{A})^* \quad (\text{A.16})$$

$$\mathcal{H} = \mathcal{T}\mathcal{T}^* - \mathcal{T}G^*(GG^*)^{-1}G\mathcal{T}^* \quad (\text{A.17})$$

669 Let us denote

$$\Lambda := GG^*. \quad (\text{A.18})$$

670 Using a simple application of Equation (A.1) and (A.10), we can then easily obtain:

$$\Lambda_{iq} = \sum_j \langle K_i^j, K_q^j \rangle_{\mathcal{M}_j} \quad (\text{A.19})$$

671 Similarly, we denote:

$$\chi := \mathcal{T}\mathcal{T}^* \quad (\text{A.20})$$

$$\chi_{k,l} = \sum_j \langle T^{j,(k)}, T^{j,(l)} \rangle_{\mathcal{M}_j} \quad (\text{A.21})$$

672 and

$$\Gamma := \mathcal{T}\mathcal{G}^* \quad (\text{A.22})$$

$$\Gamma_{ki} = \sum_j \langle T^{j,(k)}, K_i^j \rangle_{\mathcal{M}_j} \quad (\text{A.23})$$

673 Using the definitions of Λ, χ, Γ we can write (A.17) as:

$$\mathcal{H} = \chi - \Gamma\Lambda^{-1}\Gamma^T. \quad (\text{A.24})$$

674 Figure A1 provides a visualisation of the ellipse in the property space as determined by \mathcal{H} when
675 only two properties are considered.

676 **A3 Error Bounds**

677 The error bounds defined in Equations (19) and (22) are derived from the property bounds defined
678 by Al-Attar (2021, see Equation 2.84) as:

$$\langle \mathcal{H}^{-1}(p - \tilde{p}), p - \tilde{p} \rangle_{\mathcal{P}} \leq M^2 - \|\tilde{m}\|_{\mathcal{M}}^2 \quad (\text{A.25})$$

679 Equation (A.25) describes a hyperellipsoid centered on \tilde{p} with major axes given by the eigenvalues
680 of \mathcal{H}^{-1} scaled by $\sqrt{M^2 - \|\tilde{m}\|_{\mathcal{M}}^2}$. If the matrix \mathcal{H} is diagonal, then the inverse is trivial to find
681 and the hyperellipsoid has its major axes aligned with the coordinate axes of the property space.
682 In all other cases, the hyperellipsoid will have some arbitrary orientation and \mathcal{H} will be difficult to
683 invert numerically.

684 To avoid numerical complications, we use here a different, more relaxed approximation for the

685 error bounds given in the form:

$$\|p - \tilde{p}\|_{\mathcal{P}}^2 \leq (M^2 - \|\tilde{m}\|_{\mathcal{M}}^2) \text{diag}(\mathcal{H}) \quad (\text{A.26})$$

686 where $\text{diag}(\mathcal{H})$ is the diagonal of \mathcal{H} . We can also write this in component form:

$$\|p^{(k)} - \tilde{p}^{(k)}\|_{\mathcal{P}}^2 \leq (M^2 - \|\tilde{m}\|_{\mathcal{M}}^2) \mathcal{H}_{kk} \quad (\text{A.27})$$

687 Inequality (A.27) describes a hyperparallelepiped that contains the error bounds of (A.25) with
 688 sides parallel to the coordinate axes of \mathcal{P} (see Fig. A1). As this approximation does not require
 689 the inversion of the \mathcal{H} matrix, it is computationally advantageous. Visually, the hyperellipsoid
 690 fits “perfectly” inside the hyperparallelepiped (Fig. A1), but the error bounds of the hyperparal-
 691 lelepiped are easier to visualise in a static plot (see for example first row of Fig. 7). The hyperellip-
 692 soid encodes the correlations between the error bounds of the various components of the property
 693 vector (such as the correlation between the error bounds of two different local averages). Plotting
 694 the bounds for each component of the property vector simultaneously would therefore be very
 695 difficult, since the error bounds of each property component would depend on the values of the
 696 bounds on all other property components. The hyperparallelepiped ignores these correlations, sim-
 697 plifying thus the plotting. However, it overestimates the property bounds, which will likely make
 698 it more difficult to interpret the property values.

699 To show how (A.27) arises from (A.25), we need to prove the following:

Given that

$$x^T A^{-1} x \leq b \quad (\text{A.28})$$

Show that

$$x_k^2 \leq b A_{kk} \quad (\text{A.29})$$

700 where $x = \tilde{p} - \epsilon$, $A = \mathcal{H}$, and $b = M^2 - \|\tilde{m}\|_{\mathcal{M}}^2$. To prove this, we start by finding the maxi-
 701 mum extent of the hyperellipsoid (A.28) along the k^{th} coordinate axis, which can be described

702 mathematically as:

$$\text{Find } \max(c^T x) \quad (\text{A.30})$$

$$\text{Given that } x^T A^{-1} x \leq b. \quad (\text{A.31})$$

703 where c^T will be chosen later to be a vector with all entries 0 except the k^{th} one. We shall use
704 the Lagrangian approach to solve this problem. We introduce the slack constant s and use it to
705 transform the inequality A.31 into an equality (slack constraint):

$$x^T A^{-1} x - b + s^2 = 0 \quad (\text{A.32})$$

706 Let λ be a Lagrange multiplier. The problem then becomes finding the extremum points of the
707 Lagrangian:

$$f(x, \lambda) = c^T x + \lambda(b - x^T A^{-1} x - s^2). \quad (\text{A.33})$$

708 Differentiating f with respect to x and setting the result to zero leads to:

$$c - 2\lambda A^{-1} x = 0. \quad (\text{A.34})$$

709 Notice that $\lambda = 0$ leads to $c = 0$, which is a contradiction. Therefore we must have $\lambda \neq 0$ and
710 $s^2 = 0$, which means that our constraint is active. Assuming A^{-1} to be invertible, we obtain:

$$x = \frac{Ac}{2\lambda}. \quad (\text{A.35})$$

711 We next differentiate f with respect to λ (using $s^2 = 0$ since we have shown the constraint to be
712 active) to obtain the second Lagrange equation. Setting the result equal to zero leads to:

$$b - x^T A^{-1} x = 0. \quad (\text{A.36})$$

713 Substituting (A.35) into (A.36) and rearranging for b , we obtain:

$$b = \left(\frac{Ac}{2\lambda} \right)^T A^{-1} \frac{Ac}{2\lambda}. \quad (\text{A.37})$$

714 Since A is symmetric this leads to:

$$\lambda^2 = \frac{c^T A c}{4b}. \quad (\text{A.38})$$

715 Assuming that A is positive definite (its eigenvalues give the lengths of the hyperellipsoids' major

716 axes), we must have

$$\lambda = \frac{\sqrt{c^T A c}}{2\sqrt{b}}. \quad (\text{A.39})$$

717 Finally, using (A.38) and (A.35) the optimal vector solution x can be expressed for any vector
718 c as:

$$x = \frac{A c}{2\lambda} = \frac{A c}{\frac{\sqrt{c^T A c}}{\sqrt{b}}} = \frac{\sqrt{b} A c}{\sqrt{c^T A c}}, \quad (\text{A.40})$$

719 and the maximal value of $c^T x$ is thus:

$$\sqrt{b c^T A c}. \quad (\text{A.41})$$

720 Now, we consider a fixed index k between 1 and N , and we define c to be the following vector:

$$c := (\delta_{ik})_{1 \leq i \leq N}, \quad (\text{A.42})$$

721 where δ_{ik} is 1 if $i = k$, and 0 if $i \neq k$. Substituting this for c in (A.41), we obtain:

$$\max(x_k) = \sqrt{b A_{kk}} \quad (\text{A.43})$$

722 or equivalently:

$$x_k \leq \sqrt{b A_{kk}} \quad (\text{A.44})$$

723 If instead we choose $c = -\delta_{ik}$, then we have:

$$x_k \geq -\sqrt{b A_{kk}} \quad (\text{A.45})$$

724 These two inequalities can be summarised in the final answer:

$$(x_k)^2 \leq b A_{kk} \quad (\text{A.46})$$

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Table A1. Table summarising the main mathematical symbols used in the manuscript. Elements are grouped on columns and rows depending on the relationships between them. For example, \bar{m} is part of the model space \mathcal{M} and is related to \bar{d} through G , which is determined by K . Similarly, \bar{p} is related to \bar{m} through \mathcal{A} and to \tilde{m} through \mathcal{T} .

\mathcal{M}	\mathcal{D}	\mathcal{P}		
Model space	Data space	Property space	Mapping	Kernels
\bar{m} : TRUE model	\bar{d} : TRUE data		G : Forward mapping	K : Sensitivity kernels
\bar{m} : TRUE model		\bar{p} : TRUE property	\mathcal{T} : Property mapping	T : Target kernels
\tilde{m} : TRUE model		\tilde{p} : Aproximate property	\mathcal{A} : Approximate mapping	A : Resolving kernels
\tilde{m} : Least norm model		\tilde{p} : Aproximate property	\mathcal{T} : Property mapping	T : Target kernels
		ϵ : Property error	\mathcal{H} : Hyperellipsoid matrix	
j : Physical parameter index	i : Data index	k : Property index		
N' : Number of physical parameters		N : number of data		

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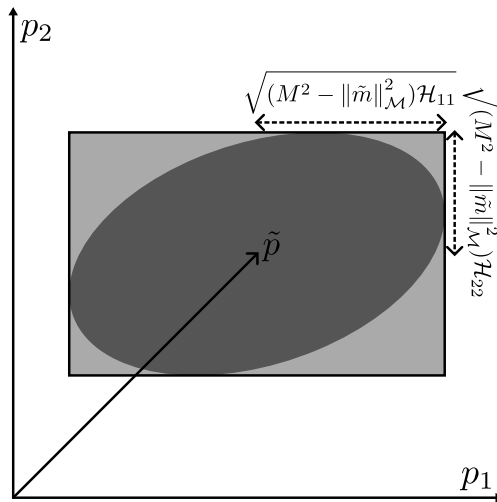


Figure A1. Illustration of the relationship between the hyperellipsoid defined in Equation A.25 and the hyperparallelepiped defined in Equation A.27 for the case when the property space is two-dimensional. p_1 and p_2 could represent, for example, two local averages at two different spatial locations. The dark shaded ellipse contains all the possible combinations of these two properties given by the tighter inequality of Equation A.25. The lighter gray shaded rectangle contains all the possible combinations of these two properties under the simplified inequality in Equation A.27.

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