

# Insights into upper lithosphere detachment and exhumation through numerical modelling

\*Dimitrios Papadomarkakis, Constantin D. Athanassas

*Department of Geological Sciences, School of Mining and Metallurgical Engineering, National  
Technical University of Athens, 9 Iroon Polytechniou Str., GR15780, Zografou Campus, Athens,  
Greece*

Dear Editor,

I am here submitting a research paper entitled “Insights into upper lithosphere detachment  
and exhumation through numerical modeling” to be considered for publication in *EarthArxiv*.  
This article has not been submitted for peer review in any journal.  
We hope it will be of interest to the readership of *EarthArxiv*.

On behalf of my co-authors,

Sincerely yours,

Dimitrios Papadomarkakis

Constantin D. Athanassas email address: [athanassas@central.ntua.gr](mailto:athanassas@central.ntua.gr)

Dimitrios Papadomarkakis email address: [papadomarkakisdimitrios@gmail.com](mailto:papadomarkakisdimitrios@gmail.com)

# Insights into upper lithosphere detachment and exhumation through numerical modelling

\*Dimitrios Papadomarkakis, Constantin D. Athanassas

*Department of Geological Sciences, School of Mining and Metallurgical Engineering, National Technical University of Athens, 9 Iroon Polytechniou Str., GR15780, Zografou Campus, Athens, Greece*

## Abstract

The mechanisms underlying exhumation have been a topic of debate among researchers for many decades, prompting the development of numerous computational models aimed at elucidating the processes that initiate exhumation. However, a key gap in the literature lies in understanding how segments of the subducting lithospheric plate detach and subsequently exhumate after extended periods. Specifically, there has been limited investigation into whether material is stripped from the upper portion of the downgoing plate at relatively shallow depths, approximately 30 km, and, if so, whether these segments can eventually reach the Earth's surface. Furthermore, no existing model demonstrates that material can detach from the subducting plate as early as 15,000 years after the subduction is initiated. This study seeks to examine whether such a phenomenon occurs, using the subduction system of the Mediterranean Ridge as a case study. The process was simulated by extensively modifying an established thermomechanical visco-elasto-plastic code, named I2ELVIS, initially introduced by Gerya (2010). Lastly, macroscopic observations from the broader Hellenides region were employed to ascertain whether any such metamorphic rocks had indeed surfaced, thus confirming their exhumation. In conclusion, this research serves as a foundational investigation into a subject that warrants further exploration and detailed analysis in the future.

*Keywords:* 2-D geodynamic modelling, subduction zones, exhumation mechanisms, mantle rheology, detachment mechanisms.

\*Corresponding author, E-mail address: papadomarkakisdimitrios@gmail.com

## 85 1. Introduction

86 Subduction zones are regions where ageing lithospheric material is recycled back  
87 into the mantle (Gerya et al., 2002). These zones, comprising a subducting lithospheric plate  
88 beneath an overlying plate, have been modelled extensively by numerous researchers using  
89 numerical methods (Neil and Houseman, 1999; Gerya et al., 2002; Gerya and Yuen, 2003a;  
90 Raimbourg et al., 2007; Warren et al., 2008a, 2008b; Beaumont et al., 2009; Li et al., 2012).  
91 One of the key geodynamic processes that has gathered significant attention in subduction  
92 settings is the exhumation of lithospheric rocks. The most commonly examined phenomenon  
93 of the latter is the kinetics and mechanisms of either ultra high-pressure metamorphic rocks  
94 or either high-pressure metamorphic rocks, both of which have been extensively researched  
95 (Platt, 1993; Chemenda et al., 1996; Gerya et al., 2002; Warren et al., 2008a, 2008b;  
96 Beaumont et al., 2009).

97 In general, the exhumation mechanisms and processes of (ultra)-high and high  
98 pressure metamorphic rocks have been widely debated over the past decades, with many  
99 researchers supporting that the formation of the gravitational instability, also known as the  
100 Rayleigh-Taylor instability, is the main driving force for the exhumation of deep buried  
101 material (Neil et al., 1999; Gerya and Yuen, 2003a). This instability is characterized by its  
102 diapiric formations, also known as “plumes”, that require millions of years to fully develop.  
103 When fully developed, it is believed that they completely detach from the subducting plate  
104 and enter the flowing mantle, where they are subjected to high pressure and temperature  
105 conditions that induce their metamorphism. These metamorphic rocks are than most likely  
106 exhumed.

107 On the other hand, the same process has been extensively modelled by researchers  
108 using lubrication theory (England and Holland, 1979; Cloos, 1982; Cloos and Shreve, 1988a,  
109 1988b; Mancktelow, 1995; Raimbourg et al., 2007; Warren et al., 2008a, 2008b; Beaumont  
110 et al., 2009; Li, 2014), if it is considered that the subducting plate is governed by two  
111 competing flows, a down channel Couette flow and an up channel Poiseuille flow. In these  
112 models, the lithospheric slab is understood to exhibit two competing flows: one promoting  
113 the burial of material and the other facilitating its exhumation.

114 Another exhumation mechanism that has been suggested in the past 25 years, is a  
115 buoyancy driven exhumation, that originates from the density difference between the  
116 detached rock and the flowing mantle. Particularly, the latter mechanism is proposed to be  
117 the dominant one when the flowing mantle has a higher density value than that of the  
118 detached rock, thus creating an upwards buoyancy force that possibly exhumes the rock.  
119 This exhumation mechanism has been suggested and utilised by many researches over the  
120 years (Liou et al., 2004; Warren et al., 2008a, 2008b; Beaumont et al., 2009; Warren, 2013),  
121 in order to explain (ultra) high-pressure rock exhumations in the Alpine and Himalayan  
122 regions. Additionally, it has also been suggested that rock exhumation, which is governed by  
123 buoyancy, allows the detached rocks to reside in the (ultra) high pressure conditions for a  
124 very short time, less than 5 Myr (Rubatto and Hermann, 2001).

125 These studies have consistently shown that the exhumation process can take over  
126 25 million years (Gerya and Yuen, 2003a; Li et al., 2012), with the detachment of exhumed  
127 material typically occurring in the lower parts of the asthenosphere at depths of 80-120 km

128 (Gerya and Yuen, 2003a), or even deeper, at 180-200 km (Warren et al., 2008a, 2008b).  
129 However, there have been no previous modelling attempts that have explored whether  
130 lithospheric material could detach from the subducting plate at much shallower depths,  
131 particularly in the upper parts of the lithosphere at depths as shallow as 30 km, and whether  
132 this detached material could eventually reach the surface. Additionally, it has not been  
133 investigated whether this detachment from the subducting plate could occur over a relatively  
134 short geological timescale, specifically in less than 20,000 years.

135 Despite this large existing gap in the available literature, previous macroscopic  
136 observations conducted by Thomson et al. (1998) in the island of Crete provided clear  
137 evidence that HP/LT exhumed metamorphic rocks had been detached from the downgoing  
138 plate from the shallow depths of around 30-35 km. These detached rocks from the  
139 subducting plate were subjected to high pressures and relatively low temperatures, thus  
140 causing the metamorphism of the granite rock into gneiss, which is mostly composed of  
141 quartz. Overall, such HP/LT exhumed metamorphic rock have been widely studied and  
142 observed in the Greek island of Crete (Fassoulas et al., 1994; Thomson et al., 1998;  
143 Chatzaras et al., 2006; Ring et al., 2010). Thus, although evidence from previous  
144 macroscopic geological observations suggests that lithospheric material can detach from the  
145 subducting plate at shallow depths of approximately 30 km, no numerical model has been  
146 developed to definitively confirm this phenomenon or to determine the time required for its  
147 occurrence.

148 To sum up the aim of the study was firstly to investigate whether at the shallow  
149 depths of the lithosphere, particularly close to 30 km, if the upper deformed part of the  
150 downgoing plate endures such increased strain values that the detachment of rock  
151 segments occurs and to observe if the detachment can take place within a relatively short  
152 timescale, as early as 15,000 years from the initiation of the subduction. In order to properly  
153 address the above-mentioned questions a computational model was developed for the  
154 subduction zone of the Mediterranean Ridge, located south of Crete. This was achieved  
155 using a two-dimensional thermomechanical visco-elasto-plastic code originally introduced by  
156 Gerya (2010), named I2ELVIS. Finally, macroscopic geological observations from the  
157 Hellenides were utilized in order to assess whether these shallow depth detached  
158 lithospheric rocks were exhumed.

159

## 160 **2. 2-D Model Setup**

### 161 **2.1. Structure of the Model**

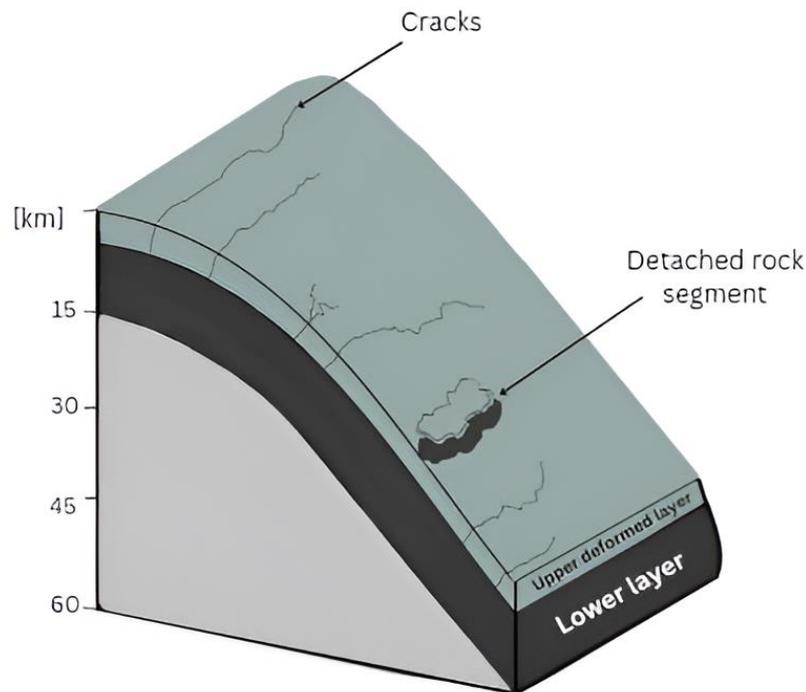
162 In the original computational model created by Gerya (2010), three different layers  
163 were present. First, the "weak medium" layer represented the flowing mantle, which was  
164 adapted in later versions of the code. Second, the "block" layer was modified to resemble  
165 the thin upper layer of the subducting lithospheric slab, which had undergone significant  
166 tectonic deformation. The existence of this thin upper rock layer has been proposed by  
167 various researchers (Capitanio et al., 2007; Royden and Papanikolaou, 2011; Goes et al.,  
168 2011). Finally, the "weak inclusion" layer was removed entirely, as it served no purpose in  
169 the current geodynamic context. A new layer, the "lower weak medium," was added by the

170 authors. This layer represented the lower part of the subducting lithospheric plate,  
171 characterised by much stiffer properties compared to the thin upper layer. The latter acted as  
172 the interface between the flowing mantle and the rest of the slab.

173 The initial model size of the computational setting was 1,000 km in both the  
174 horizontal and vertical axes. Additionally, 51 nodes were placed along both axes, and 40,000  
175 markers were randomly distributed throughout the model (Gerya, 2010). To improve the  
176 visualisation of the results, the horizontal length was reduced by 800 km, resulting in a new  
177 length of 200 km, while the vertical length was shortened to 80 km. Despite these changes,  
178 the number of nodes and markers remained the same, maintaining 51 nodes in both  
179 directions and 40,000 randomly distributed markers.

180 Overall, the computational model's structure consisted of a flowing mantle and a  
181 subducting lithospheric plate divided into two uneven layers. The upper thin layer was more  
182 heavily deformed due to its constant interaction with the mantle, while the lower, thicker  
183 layer, which made up most of the plate, exhibited less deformation. Figure 1 provides a  
184 visual representation of the computational model. The subducting system was modelled after  
185 the Hellenic subduction zone, commonly referred to as the Mediterranean Ridge, located  
186 south of Crete. Geophysical investigations have confirmed that the Moho zone's depth in this  
187 region begins at approximately 25–30 km below the surface (Makris, 2010; Makris et al.,  
188 2013). Consequently, it was assumed that the subduction process initiates at this same  
189 depth, aligning with the Moho zone's starting depth. This approximation was consistent with  
190 the resemblance between the map of the Moho zone surface, as derived from geophysical  
191 data (Makris et al., 2013), and the typical topography of a subducting and overriding plate.

192



193

194 **Figure 1:** In this figure, the subducting plate is depicted, highlighting its two distinct, uneven  
 195 layers. Additionally, select outcomes from the computational simulation are illustrated,  
 196 particularly the onset of cracking in the upper deformed layer of the slab and the eventual  
 197 detachment of a rock segment. These results will be comprehensively presented and  
 198 analysed in the relevant sections of this study.

199 In the computational model, only a section of the subducting plate is visualised.  
 200 Additionally, a justified approximation was made by the authors to simplify the visualisation  
 201 of the results at each timestep. Specifically, no subduction rate was assigned to the slab,  
 202 effectively treating it as a stagnant plate. The reasoning behind this approximation lies in the  
 203 relatively short simulation time of around 15,000 years, which is considered brief in  
 204 comparison to the much longer timescales over which geodynamic processes typically  
 205 occur. As such, it was assumed that the slab did not move significantly enough during this  
 206 period to substantially influence the results. Therefore, the subduction rate was entirely  
 207 neglected in the current simulation.

208

209

## 210 2.2. Material Properties and Boundary Conditions

211 The material properties of each of the aforementioned layers were significantly  
 212 modified to meet the objectives and requirements of this research. Specifically, Table 1  
 213 presents all the values utilised for the material properties. Most of these values have been  
 214 extensively studied and recommended by previous researchers (Capitanio et al., 2007;  
 215 Makris, 2010; Royden and Papanikolaou, 2011; Makris et al., 2013).

216 **Table 1** In this table the values of all the material properties are presented.

Parameters	Flowing Mantle	Stagnated Slab	Thin Rock Layer
Viscosity (Pa s)	$3 \times 10^{20}$	$10^{24}$	$10^{22}$
Density (kg/m <sup>3</sup> )	3300	2800	2800
Thermal Conductivity (W/mK)	2.55	2.4	2.55
Thermal Expansion (1/K)	$5.8 \times 10^{-8}$	$10^{-7}$	$5.8 \times 10^{-8}$
Radioactive Heating (W/m <sup>3</sup> )	$10^{-6}$	$10^{-7}$	$10^{-6}$
Heat Capacity (J/kg)	750	650	750
Shear Modulus (Pa)	$10^{10}$	$15 \times 10^9$	$22.8 \times 10^9$
Cohesion (Pa)	$28.4 \times 10^6$	$11.5 \times 10^6$	$10^{11}$
Friction Coefficient	0.3	0.6	0.7
Temperature (K)	520	400	400

217 Firstly, the viscosity was initially set at  $10^{23}$  Pa·s for the subducting slab and at  
 218  $10^{17}$  Pa·s for the flowing mantle. The value for the lower part of the subducting slab was  
 219 increased, as this section was expected to be stiff rather than viscous. Additionally, the  
 220 original density values were  $4000 \text{ kg/m}^3$  for the stagnated slab and  $1 \text{ kg/m}^3$  for the flowing  
 221 mantle. These values were completely revised since they represented materials not typically  
 222 present in geodynamic processes. The rationale for the chosen density values can be  
 223 explained by the chemical composition of the flowing mantle and the subducting plate.  
 224 Particularly, the flowing mantle was considered to be primarily composed of peridotite, thus  
 225 justifying the density value of  $3,300 \text{ kg/m}^3$ . This density value was also used by Royden and  
 226 Papanikolaou (2011) in their modelling of the segmentation of the Hellenic arc and later  
 227 verified by geophysical investigations (Makris et al., 2013). As for the subducting slab, both  
 228 the lower, thicker layer and the upper, thinner layer were assumed to be predominantly  
 229 composed of granite, giving them a density value of approximately  $2,800 \text{ kg/m}^3$  (Makris,  
 230 2010; Royden and Papanikolaou, 2011; Makris et al., 2013).

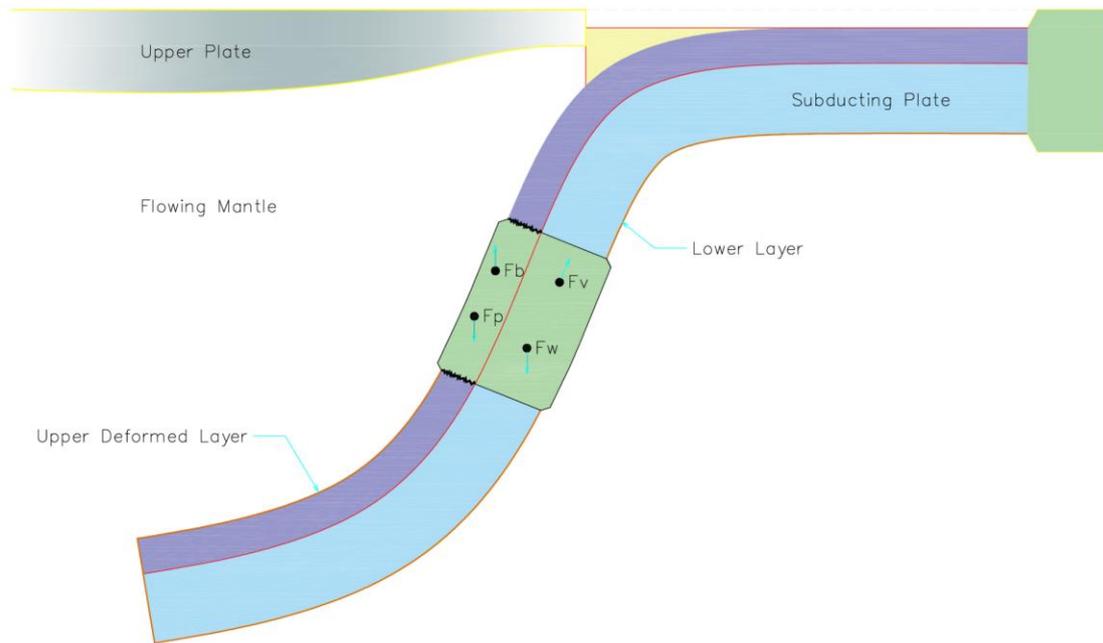
231           Moreover, several thermomechanical parameters were altered, including thermal  
232 conductivity, thermal expansion, radioactive heating, and heat capacity. The shear modulus  
233 values were also adjusted, as the original code assigned a uniform value of  $10^{10}$  Pa for both  
234 the mantle and the slab. Consequently, the cohesion and friction coefficient values were  
235 extensively modified, with the cohesion of the lower part of the plate being significantly  
236 increased to ensure its impermeability against the flowing mantle. Finally, the original mantle  
237 and slab temperatures were set at 1000 K and 1300 K, respectively. Both were reduced due  
238 to the relatively shallow depth of approximately 30 km at which the process occurs.

239           In general, the temperatures at the boundaries of the computational model, including  
240 the upper, lower, right, and left boundaries, were held constant. Specifically, the upper and  
241 right boundaries were set to a constant temperature of 520 K, consistent with that of the  
242 mantle, while the lower and left boundaries were set to 400 K, mirroring the temperature of  
243 the slab. These boundaries were characterised as insulating, thus implying that no heat flux  
244 occurred through the boundary, meaning no temperature gradient existed across it.

245           In addition to the no heat flux condition, a lateral symmetry condition is applied, which  
246 is widely used in both 2D and 3D Cartesian geodynamic models (Gerya, 2010).  
247 Furthermore, to simplify the fluid flow analysis, a free-slip boundary condition was employed,  
248 ensuring that no frictional forces act parallel to the boundary surface.

### 249   **2.3. Driving Subduction and Detachment Forces**

250           The primary forces exerted on the thin upper layer of the plate include the buoyancy  
251 force ( $F_b$ ), which acts in the direction of the surface (Capitanio et al., 2007; Goes et al.,  
252 2011). This buoyancy force arises because the slab is essentially submerged in the flowing  
253 mantle, a viscous fluid, and as a natural response, the fluid exerts an upward force to make  
254 the thin layer float. Additionally, the weight force ( $F_w$ ) acts in opposition to the buoyancy  
255 force. Furthermore, a lithostatic pressure force ( $F_P$ ) is also applied, with the same downward  
256 direction as the weight force (Turcotte and Schubert, 1982). Lastly, there is a viscous force  
257 ( $F_V$ ) due to the flowing mantle, which acts parallel to the slab's surface and upwards,  
258 resisting the plate's subduction (Royden and Papanikolaou, 2011). All these forces are  
259 depicted in Figure 2.



260

261 **Figure 2:** This figure depicts the forces exerted on a rock segment situated within the upper  
 262 deformed layer of the subducting plate, highlighting those that promote the burial of the  
 263 segment (weight and pressure forces) and those that drive it upwards (buoyancy and  
 264 viscous forces).

265

266 In summary, the driving forces behind subduction are the weight force and the  
 267 pressure force, which together promote the downward motion of the plate. Conversely, the  
 268 buoyancy force and the viscous force resist the slab's subduction. The balance between  
 269 these driving and resisting forces governs the subduction process (Forsyth and Uyeda,  
 270 1975; Vlaar and Wortel, 1976; Chapple and Tullis, 1977; Davies, 1988; Capitanio et al.,  
 271 2007).

## 272 2.4. Potential Development of the Rayleigh–Taylor Gravitational 273 Instability

274 An important phenomenon that warrants discussion is the Rayleigh-Taylor gravitational  
 275 instability. This instability occurs when a heavier fluid (in this case, the mantle) overlies a  
 276 lighter fluid (the lithospheric slab), leading to a gravitational imbalance. The interface  
 277 between the two fluids becomes distorted, causing fluid motions, and ultimately, the heavier  
 278 mantle is penetrated by the lighter, underlying slab. The resulting effect is the formation of  
 279 characteristic diapiric structure, often described as a "mushroom-like" shape, as the lighter  
 280 fluid rises through the heavier fluid (Turcotte and Schubert, 1982). The Rayleigh-Taylor  
 281 instability has been extensively modelled by various researchers (Turcotte and Schubert,  
 282 1982; Neil et al., 1999; Gerya and Yuen, 2003a; Kaus and Becker, 2006; Ghosh et al., 2020)  
 283 because it is believed that the resulting "cold" plumes produced by this instability have a  
 284 strong likelihood of being exhumed. Despite the fact that the appropriate conditions for the  
 285 initiation of the Rayleigh-Taylor gravitational instability were present at this computational

286 setting, the phenomenon was completely dismissed from the simulation. The reason for its  
 287 exclusion is grounded in prior studies (Neil et al., 1999; Gerya and Yuen, 2003a), which  
 288 demonstrate that the full development of the instability, including the diapiric formation that  
 289 may eventually be exhumed, typically requires around 25 million years (Gerya and Yuen,  
 290 2003a) to form. Given that the simulations conducted in the present study spanned only  
 291 15,000 years, a brief period in geological time, there was insufficient time for the Rayleigh-  
 292 Taylor instability to develop to a stage where it would affect the outcomes of the model. As  
 293 such, this phenomenon was neglected to simplify the overall computational framework.

### 294 **3. Analytical Rheological Modelling**

#### 295 **3.1. Down-Channel Couette Flow**

296 Firstly, within the “flowing” subducting channel, exists a down-channel Couette flow,  
 297 commonly referred to as the subducting flow, which promotes the burial of the upper  
 298 deformed material (Raimbourg et al., 2007; Warren et al., 2008a, 2008b; Beaumont et al.,  
 299 2009; Li, 2014). This downward Couette flow is given by the following equation (Raimbourg  
 300 et al., 2007; Warren et al., 2008a, 2008b; Beaumont et al., 2009; Li, 2014):

$$301 \quad u_{couette} = U(1 - \frac{y}{h}) \quad (9),$$

302 where,  $u_{couette}$ : is the downgoing Couette flow velocity, U: denotes the subduction velocity of  
 303 the underlying lithosphere, x: is measured in the down-dip direction, y: is the position in the  
 304 channel measured normal to the base and h: is the channel (lithosphere) thickness.

#### 305 **3.2. Up-Channel Poiseuille Flow**

306 Secondly, the competing flow within the subducting channel is an up-channel  
 307 Poiseuille flow, commonly referred to as the exhumation flow, which promotes the upward  
 308 movement of the deformed upper section of the subducting plate (Raimbourg et al., 2007;  
 309 Warren et al., 2008a, 2008b; Beaumont et al., 2009; Li, 2014). This up-channel flow is  
 310 primarily driven by the buoyancy of the low-density subducted material, in contrast to the  
 311 denser mantle, and by the viscous resistance of the mantle, which counteracts the  
 312 downward motion of the plate (Raimbourg et al., 2007; Warren et al., 2008a, 2008b;  
 313 Beaumont et al., 2009; Li, 2014). The Poiseuille flow can be described by the following  
 314 equation (Raimbourg et al., 2007; Warren et al., 2008a, 2008b; Beaumont et al., 2009; Li,  
 315 2014):

$$316 \quad u_{Poiseuille} = (\frac{1}{2n})(\frac{\partial P}{\partial x})(yh - y^2) \quad (10),$$

317 where,  $u_{Poiseuille}$ : is the up-channel Poiseuille flow velocity, n: is the uniform viscosity and  
 318  $\frac{\partial P}{\partial x}$ : denotes the effective down-channel pressure gradient.

#### 319 **3.3. Lubrication Theory**

320 These two competing flows, which are present within the viscous channel of the  
 321 subducting plate, have been extensively studied by numerous researchers in the past

322 through the application of lubrication theory (England and Holland, 1979; Cloos, 1982; Cloos  
 323 and Shreve, 1988a, 1988b; Mancktelow, 1995; Raimbourg et al., 2007; Warren et al., 2008a,  
 324 2008b; Beaumont et al., 2009; Li, 2014). Under the assumptions of lubrication theory  
 325 (Pozrikidis, 2001), which postulates that the overlying lithospheric plate remains stationary,  
 326 the velocity of flow within the channel is expressed by the following equation:

$$327 \quad u(x, y) = -\left(\frac{1}{2n}\right)\left(\frac{\partial P}{\partial x}\right)(yh - y^2) + U\left(1 - \frac{y}{h}\right) \quad (11).$$

328 Additionally, when the non-dimensional variables:  $u' = u/U, h' = h/H, y' = y/h$  and  $x' =$   
 329  $x/h$  are used, Eq. (11) is reduced to:

$$330 \quad u' = -Eh'^2 \frac{(y' - y'^2)}{2 + (1 - y')} \quad (12),$$

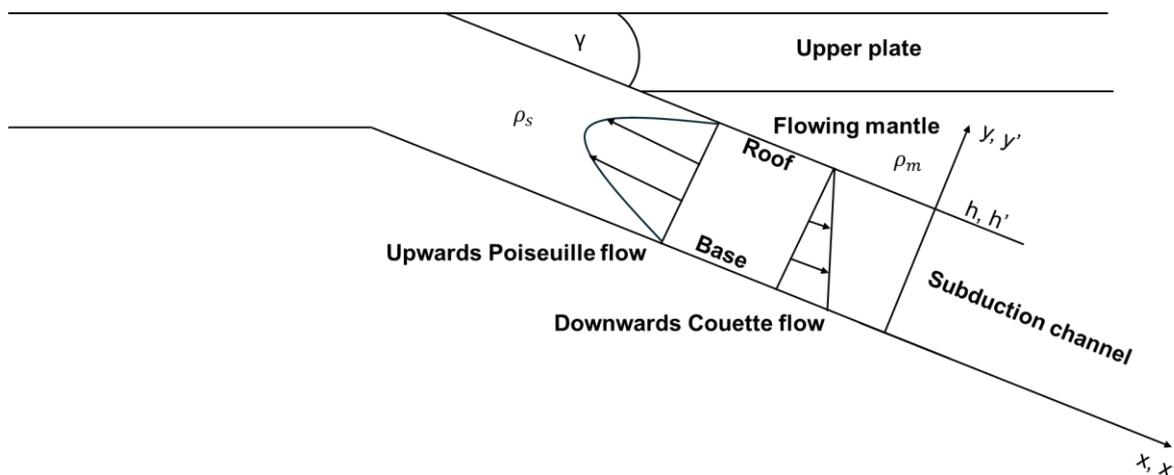
331 where,

$$332 \quad E = (H^2 \times \frac{\partial P}{\partial x}) / (n_{eff} U) \quad (13),$$

$$333 \quad H = \left(\frac{2n_{eff} U}{\partial P / \partial x}\right)^{1/2} \quad (14),$$

334 where,  $n_{eff}$ : represents the effective viscosity and the parameter H: is the characteristic  
 335 channel (slab) thickness for  $E=1$ , the balance point between the downward and returning  
 336 flows. All of the parameters along with the flow directions can be seen in Figure 3.

337



338

339 **Figure 3:** This figure illustrates the two competing flows, particularly, the downgoing Couette  
 340 flow and the up-channel Poiseuille, along with other key parameters.

341 However, the key parameter in Eq. (12) is the exhumation number (E) (Warren et al.,  
 342 2008a, 2008b; Beaumont et al., 2009; Li, 2014), which corresponds to the  $\alpha$  parameter as  
 343 described by Raimbourg et al. (2007). This parameter fundamentally quantifies the  
 344 competition between the up-channel Poiseuille flow, and the down-channel Couette flow.

345 In cases where the channel walls are deformable, as in the present model, and there  
 346 is no tectonic overpressure or underpressure, the pressure gradient depends on the density  
 347 difference between the subducted material (granite) within the slab (channel) and the  
 348 surrounding mantle (peridotite) and it is thus given by the following equation (Warren et al.,  
 349 2008a, 2008b; Beaumont et al., 2009; Li, 2014):

$$350 \frac{\partial P}{\partial x} = (\rho_m - \rho_s)g \sin(\gamma) \quad (15),$$

351 where,  $\rho_m$  and  $\rho_s$ : are the density of the mantle and the lithospheric slab respectively and  $\gamma$ :  
 352 represents the dip of the subduction channel.

353 Overall, there are three distinct scenarios that can arise depending on the  
 354 exhumation number (E). The first scenario, occurs when the density of the subducting  
 355 material ( $\rho_s$ ) exceeds that of the surrounding mantle ( $\rho_m$ ), thus  $\rho_s > \rho_m$ . In this case, the  
 356 pressure gradient, as defined by Eq. (15), decreases along the channel, resulting in both  $\frac{\partial P}{\partial x}$   
 357 and E being negative. As a consequence, the down-channel Couette flow dominates,  
 358 leading to a low likelihood of exhumation.

359 The second scenario, occurs when the density of the subducting material equals the  
 360 density of the surrounding mantle, thus  $\rho_s = \rho_m$ . In this instance, the pressure gradient and  
 361 exhumation number equal zero, or more conveniently,  $E \sim 1$ . This represents the balance  
 362 point between the downward subducting flow and the upward return flow, leaving the slab  
 363 stagnant.

364 Finally, the third scenario takes place when the density of the subducting material is  
 365 less than that of the surrounding mantle, thus  $\rho_s \leq \rho_m$ , which is also the occurring scenario.  
 366 Here, the pressure gradient increases along the channel, resulting in both  $\frac{\partial P}{\partial x}$  and E being  
 367 positive. Buoyancy drives the upward flow in the channel, making the Poiseuille flow  
 368 dominant (Raimbourg et al., 2007; Warren et al., 2008a, 2008b; Beaumont et al., 2009; Li,  
 369 2014) and significantly increasing the likelihood of exhumation. In this case, the greater the  
 370 positive value of E, the higher the exhumation velocity.

371 However, it is important to note that buoyancy alone is a necessary but not sufficient  
 372 condition for exhumation to occur (Raimbourg et al., 2007; Warren et al., 2008a, 2008b;  
 373 Beaumont et al., 2009; Li, 2014). Other controlling factors, such as a decrease in effective  
 374 viscosity ( $n_{eff}$ ) play a critical role in driving E beyond the exhumation threshold. Generally,  
 375 the exhumation number (E) should be viewed as a measure of the local exhumation  
 376 potential. Even when the local threshold value is exceeded ( $E > 1$ ), efficient exhumation may  
 377 still be hindered by constrictions (low h) or high effective viscosities ( $n_{eff}$ ) further up the  
 378 channel (Raimbourg et al., 2007; Warren et al., 2008a, 2008b; Beaumont et al., 2009; Li,  
 379 2014).

380 To sum up, although it is not certain that exhumation will occur, it can be clearly  
 381 observed that buoyancy is the driving force of the whole process thus further increasing the  
 382 odds of detachment.

## 383 4. Numerical Modelling

### 384 4.1. Governing equations

385 Firstly, the x,y-Stokes equations for slow, viscous, incompressible flow in a uniform  
386 gravity field are expressed by the following formulas:

$$387 \quad n\left(\frac{\partial^2 u_x}{\partial x^2} + \frac{\partial^2 u_x}{\partial y^2}\right) - \frac{\partial P}{\partial x} = \rho g_x \quad (1)$$

$$388 \quad n\left(\frac{\partial^2 u_y}{\partial x^2} + \frac{\partial^2 u_y}{\partial y^2}\right) - \frac{\partial P}{\partial y} = \rho g_y \quad (2)$$

389 where, Eq. (1) represents the x-Stokes equation and Eq. (2) the y-Stokes equation  
390 respectively. Additionally, x and y denotes, respectively, the horizontal and vertical  
391 coordinates, the coefficient n: represents the effective viscosity (in Pa s), that depends on  
392 pressure, temperature and the velocity, P: stands for the pressure (in Pa),  $\rho$ : is the density  
393 (in kg/m<sup>3</sup>),  $g_x$  and  $g_y$ : denote components of the vector of acceleration within the gravity field  
394 (in m/s<sup>2</sup>) for the x–y 2D coordinate system and finally,  $u_x$  and  $u_y$ : represent the velocity  
395 components (m/s) in the x and y direction respectively. For simplification purposes in the  
396 computational calculations, the values of  $\rho g_x$  and  $\rho g_y$  were neglected as they remained  
397 constant throughout the simulation and did not significantly influence the overall results. This  
398 allowed for a reduction in computational complexity without affecting the accuracy of the final  
399 outcomes.

400 Furthermore, another important equation is that of mass conservation, which is  
401 primarily described by the continuity equation. It is also worth noting that the density ( $\rho$ ) was  
402 assumed to be constant in all terms, an approximation commonly used in numerical  
403 geodynamic modelling, known as the Boussinesq approximation (Moresi and Solomatov,  
404 1995; Trompert and Hansen, 1996; Albers, 2000; Gerya and Yuen, 2003b; Gerya, 2010).  
405 The continuity equation can be expressed in the following form, assuming incompressible  
406 flow:

$$407 \quad \frac{\partial u_x}{\partial x} + \frac{\partial u_y}{\partial y} = 0 \quad (3)$$

408 Finally, another fundamental equation used extensively during the computational  
409 steps was the two-dimensional temperature equation. This equation is expressed using the  
410 Lagrangian form of the heat conservation equation, which accounts for radioactive, adiabatic  
411 and shear heating production:

$$412 \quad \rho \times C_p \times \frac{DT}{Dt} = -\frac{\partial q_x}{\partial x} - \frac{\partial q_y}{\partial y} + H_r + H_\alpha + H_s \quad (4),$$

$$413 \quad q_x = -k \times \frac{\partial T}{\partial x} \quad (5),$$

$$414 \quad q_y = -k \times \frac{\partial T}{\partial y} \quad (6),$$

415  $H_r = const,$

416  $H_a = T\alpha[u_x(\frac{\partial P}{\partial x}) + u_y(\frac{\partial P}{\partial y})] = T\alpha\rho(u_x g_x + u_y g_y) (7),$

417  $H_s = n(2\frac{\partial^2 u_x}{\partial x^2} + \frac{\partial^2 u_y}{\partial y^2} + [\frac{\partial u_x}{\partial y} + \frac{\partial u_y}{\partial x}]^2) (8),$

418 where,  $C_P$ : is the isobaric heat capacity (in J/kgK),  $q_x$  and  $q_y$ : are horizontal and vertical heat  
 419 fluxes (in W/m<sup>2</sup>),  $k$ : is the variable thermal conductivity coefficient (in W/mK) and  $H_r$ ,  $H_a$ ,  $H_s$ :  
 420 denote, respectively, radioactive, adiabatic and shear heating production (in W/m<sup>3</sup>). It is also  
 421 important to mention that for the sake of simplicity in the calculation of  $H_a$ , slight deviations in  
 422 the dynamic pressure gradients  $\frac{\partial P}{\partial x}$  and  $\frac{\partial P}{\partial y}$  from the values  $\rho g_x$  and  $\rho g_y$  were neglected.

423

## 424 **4.2. Characteristics Based Marker-In-Cell Method with Conservative** 425 **Finite Differences Schemes**

426 Overall, Eqs. (1)-(4) represent the primary equations employed to establish the  
 427 computational model of the I2ELVIS code that was utilised. Additionally, these equations  
 428 were solved at every timestep to ensure the accurate representation of the system's  
 429 behaviour. The structure of the computational code can be broken down into nine steps,  
 430 which involve the solution of the momentum, continuity and temperature equation, Eq. (1)-  
 431 (4), by using finite-differences and the characteristic based marker-in-cell technique. These  
 432 nine steps are not presented in this paper, since they are out of its primary scope, however  
 433 they are introduced and consequently analysed and explained extensively in previous  
 434 studies by Gerya and Yuen (2003b) and Gerya (2010).

## 435 **4.3. Simulation Parameters**

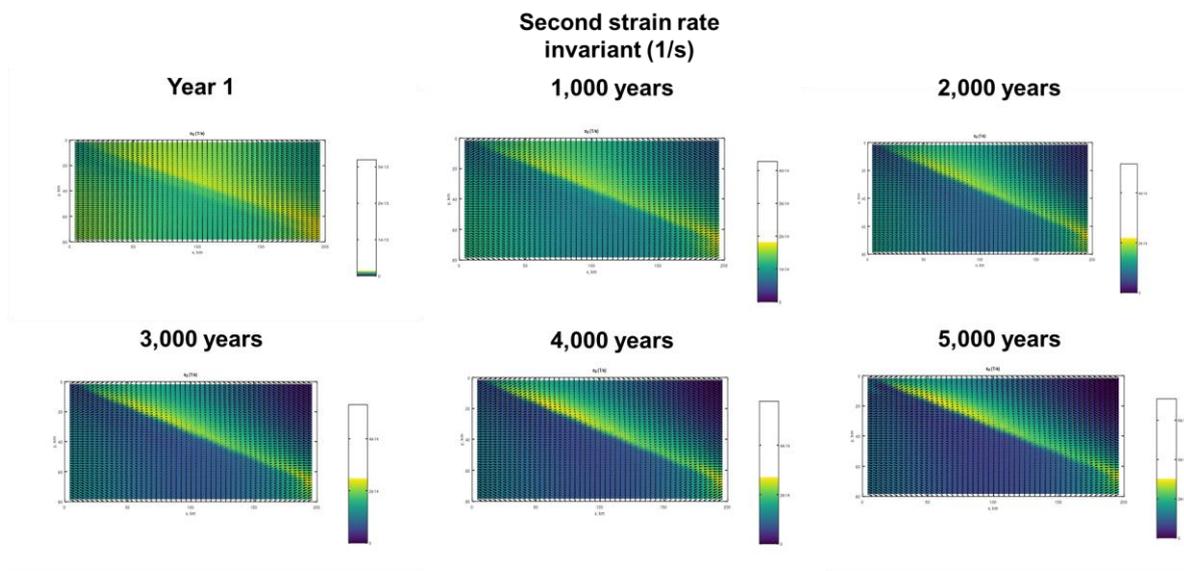
436 The simulation was executed twice to account for two different thicknesses of the  
 437 upper deformed layer of the subducting slab. In the first simulation (experiment 1), the upper  
 438 layer thickness was set at 5 km. Each simulation spanned 150 timesteps, representing a  
 439 period of 15,000 years. During each timestep, three critical parameters were calculated and  
 440 visualised on the computational map: viscosity ( $\eta$ ), the second stress invariant ( $\sigma_{II}$ ) and the  
 441 second strain rate invariant ( $\underline{\varepsilon}_{II}$ ). In the second simulation (experiment 2), the thickness of  
 442 the upper deformed layer was increased to 10 km. Similar to the first simulation, during each  
 443 timestep, the three key parameters, viscosity ( $\eta$ ), second stress invariant ( $\sigma_{II}$ ) and second  
 444 strain rate invariant ( $\underline{\varepsilon}_{II}$ ), were calculated and visualised on the computational map.

## 445 **5. Model Results**

### 446 **5.1. Second Strain Rate Invariant**

447 First and foremost in the following Figures, the visualisation of the second strain rate  
 448 invariant is showcased for the whole model, as well as the distribution of the second strain

449 invariant values, for various time steps of the simulation, across the 49 vertical available  
 450 nodes that were located inside the upper thin part of the lithospheric slab, respectively. In the  
 451 three-part Figure 4 the second strain rate invariant values for the whole computational model  
 452 are presented, the thickness of the upper deformed part of the plate was set at 5 km:

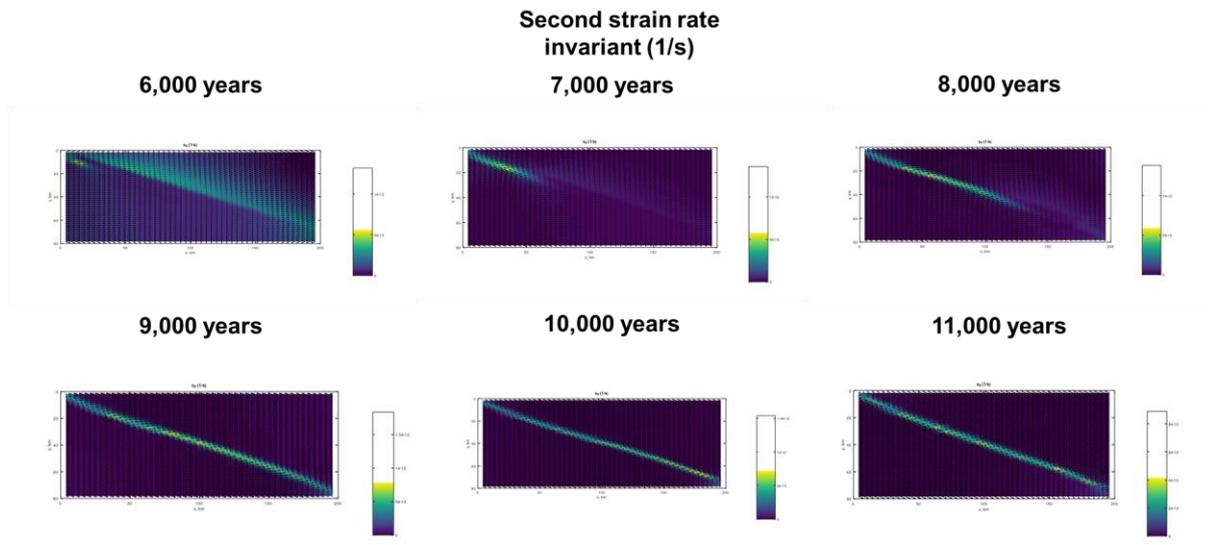


453

454 **Figure 4:** In part I of this figure the second strain rate invariant values are showcased for the  
 455 computational model, when the thickness of the upper deformed layer was set at 5 km,  
 456 particularly for the first year of the simulation, after 1,000 years, 2,000 years, 3,000, 4,000  
 457 and 5,000 years respectively.

458 In general, it can be seen that for the first 5,000 years of the simulation, a relatively  
 459 high value of strain rate was present throughout the whole surface of the lithospheric slab and  
 460 particularly in its upper left side, which is likely caused by the constant viscous flowing mantle.

461



### Part II

462

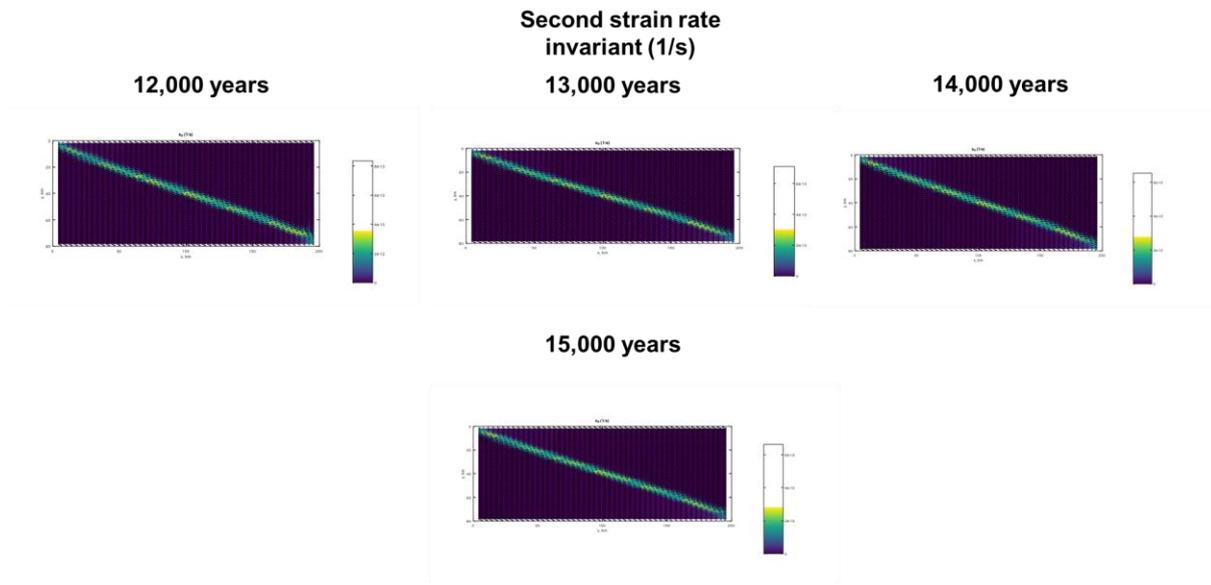
463 **Figure 4:** In part II of this figure the second strain rate invariant values are showcased for  
 464 the computational model, when the thickness of the upper deformed layer was set at 5 km,  
 465 particularly for the first year of the simulation, after 6,000 years from the initiation of the  
 466 simulation, after 7,000 years, 8,000 years, 9,000, 10,000 and 11,000 years respectively.

467 After just 6,000 years from the start of the simulation, significant deformation is  
 468 observed in the upper left section of the slab, where high strain rates persisted for nearly  
 469 5,000 years. This prolonged strain on the upper layer induced substantial deformation,  
 470 potentially initiating cracks and possibly detaching sections of the material. Such damage  
 471 and particularly the emergence of cracks may have allowed the infiltration of the viscous,  
 472 flowing mantle into the upper deformed layer. Consequently, over the next 3,000 years,  
 473 these cracks and the potential mantle intrusion propagated throughout the entire upper layer  
 474 of the plate, establishing sustained high strain rates across its length.

475 In the final 4,000 years of the simulation (11,000-15,000 years), distinct patterns in  
 476 strain rate values emerge within the heavily deformed upper layer, fluctuating between high  
 477 and low values.

478

479



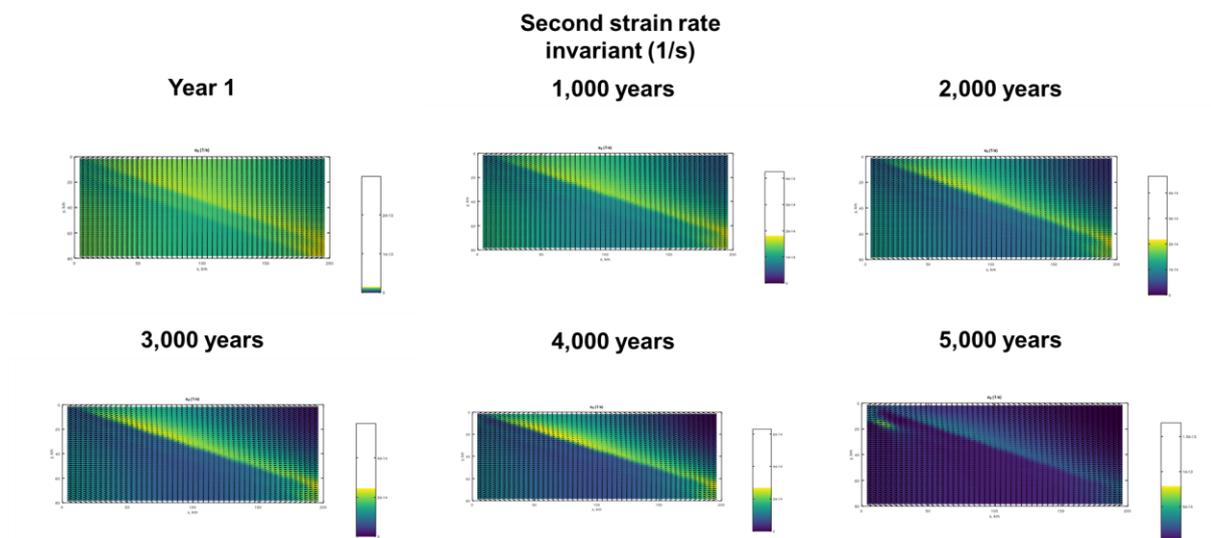
**Part III**

480

481 **Figure 4:** In part III of this figure the second strain rate invariant values are showcased for  
 482 the computational model, when the thickness of the upper deformed layer was set at 5 km,  
 483 particularly for the first year of the simulation, after 12,000 years from the initiation of the  
 484 simulation, after 13,000 years, 14,000 and 15,000 years respectively.

485 Moreover, in the three-part Figure 5 the second strain rate invariant values for the  
 486 whole computational model are presented, when the thickness of the upper deformed part of  
 487 the plate was set at 10 km:

488

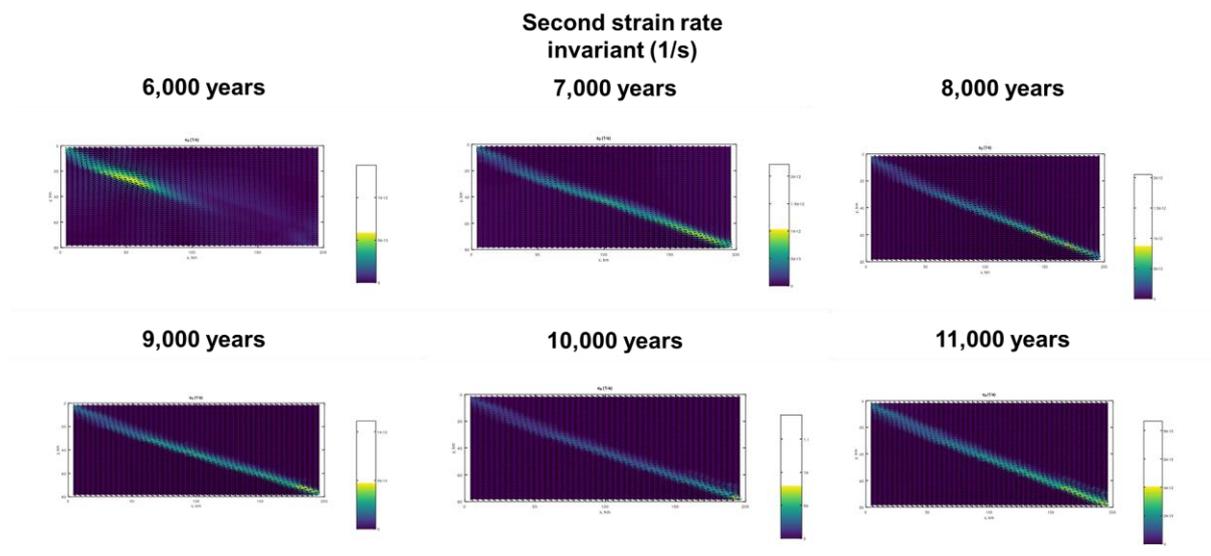


**Part I**

489

490 **Figure 5:** In part I of this figure the second strain rate invariant values are showcased for the  
 491 computational model, when the thickness of the upper deformed layer was set at 10 km,  
 492 particularly for the first year of the simulation, after 1,000 years, 2,000 years, 3,000, 4,000  
 493 and 5,000 years respectively.

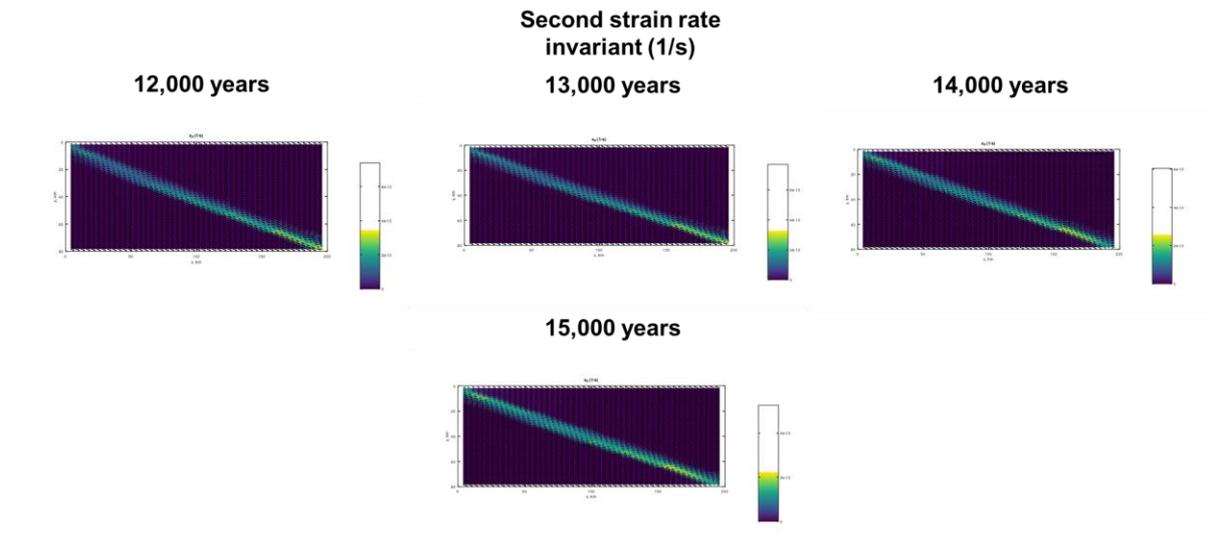
494 As with the case of the 5 km thickness, the same high values of strain rate can be  
 495 observed throughout the whole surface of the upper layer of the plate, again with the left side  
 496 showcasing the larger values. The main difference with the 5 km model is that in the 10 km  
 497 one the upper deformed layer displayed major deformation from the high strain rate values,  
 498 and thus crack initiation, and consequently possible material detachment and entrance of the  
 499 viscous flowing mantle, after only 5,000 years, almost 1,000 years earlier than the 5 km  
 500 model.



501

502 **Figure 5:** In part II of this figure the second strain rate invariant values are showcased for  
 503 the computational model, when the thickness of the upper deformed layer was set at 10 km,  
 504 particularly after 6,000 years from the initiation of the simulation, after 7,000 years, 8,000  
 505 years, 9,000, 10,000 and 11,000 years respectively.

506 Consequently, similar to the 5 km model, the cracks and potential infiltration of the  
 507 flowing mantle extended throughout the entire upper layer of the plate. However, in the 10  
 508 km model, this process occurred over a much shorter time span, specifically within 2,000  
 509 years. From this point onward, significant strain rate values emerged along the length of this  
 510 layer. Finally, consistent with the findings of the 5 km model, distinct patterns in strain rate  
 511 values appear in the last 4,000 years of the simulation (11,000-15,000 years), fluctuating  
 512 between high and low values within the heavily deformed upper layer.

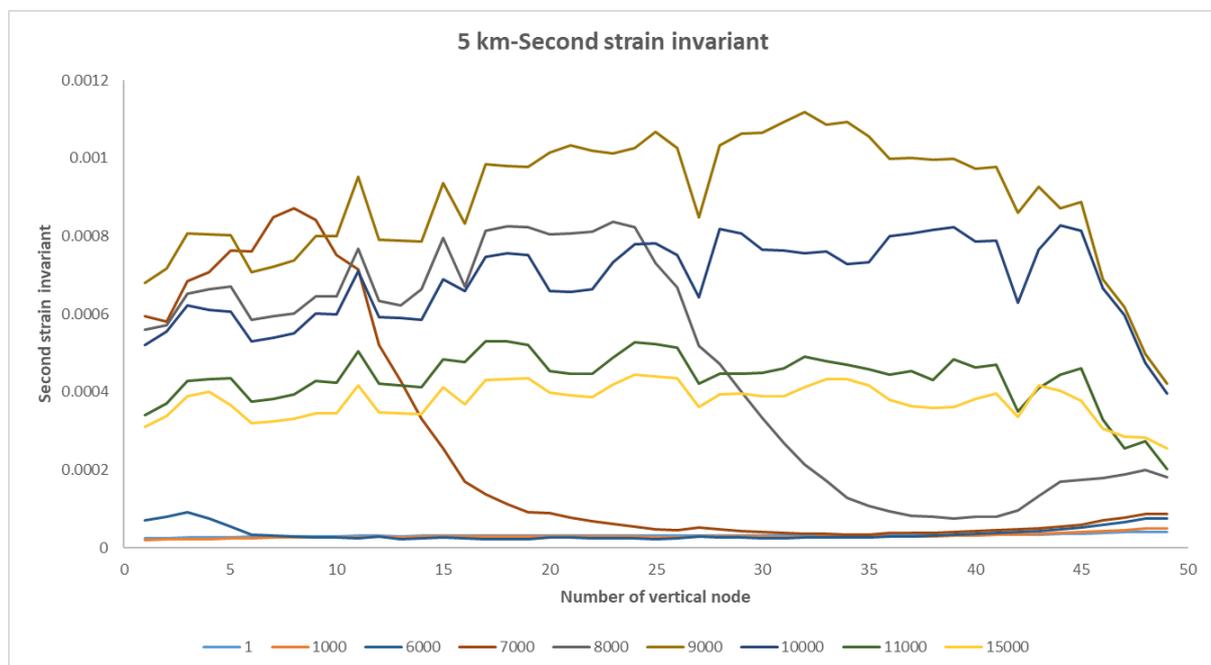


**Part III**

513

514 **Figure 5:** In part III of this figure the second strain rate invariant values are showcased for  
 515 the computational model, when the thickness of the upper deformed layer was set at 10 km,  
 516 particularly after 12,000 years from the initiation of the simulation, after 13,000 years, 14,000  
 517 and 15,000 years respectively.

518 Figure 6 illustrates the values of the second strain invariant for the upper layer of the  
 519 plate with the thickness set at 5 km, evaluated at various significant time steps of the  
 520 simulation. These values are analyzed across the 49 vertical nodes distributed throughout the  
 521 entire length of the upper deformed layer of the slab. Specifically, the selected time intervals  
 522 correspond to key moments of interest, namely the first year, the 1,000th year, and the 6,000th  
 523 through 11,000th years in 1,000-year increments, as well as the 15,000th year.



524

525 **Figure 6:** In this figure the second strain invariant values are presented, when the thickness  
526 of the upper layer of the plate was set at 5 km, for various important time moments of the  
527 simulation, in relation to the 49 vertical nodes that are located inside the upper deformed  
528 part of the lithospheric slab.

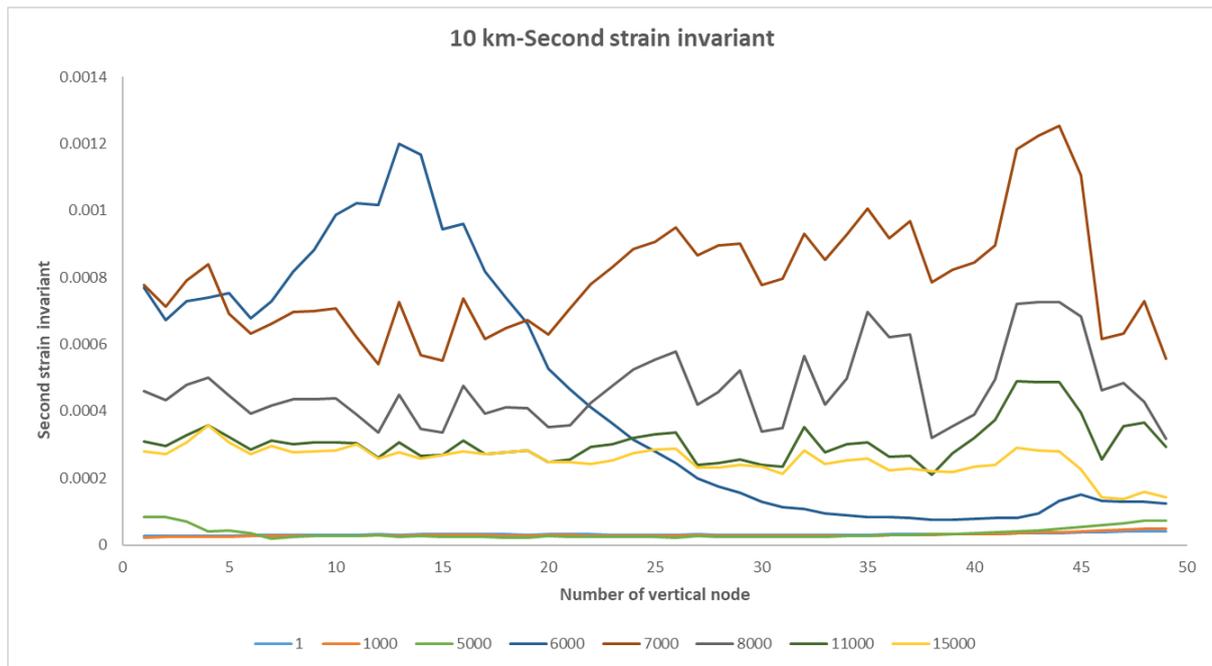
529 As previously observed in Fig. 4 (Part I), the strain values within the upper layer of the  
530 slab remain relatively low during the first 5,000 years, as significant strain is exerted primarily  
531 on the surface of the plate by the underlying mantle flow. However, by the 6,000th year, a  
532 slight increase in these low values becomes evident, attributed to the initiation of cracks in the  
533 left section of the slab, resulting in a minor spike in strain values. Subsequently, during the  
534 7,000th and 8,000th years, the strain values increase significantly, affecting larger portions of  
535 the slab as the cracks propagate across most of the upper layer's length.

536 In the 9,000th year, the strain values reach their peak for the entire simulation, likely  
537 due to the cracks permeating the entire slab, suggesting material removal throughout the  
538 upper layer and mantle infiltration through the cracks. During the final 5,000 years of the  
539 simulation (10,000–15,000 years), the strain values exhibit a consistent pattern of local  
540 maxima and minima with minor variations, broadly aligning with the trends observed in Fig. 4  
541 (Part III).

542 Accordingly, Figure 7 presents the second strain invariant values for an upper layer  
543 thickness of 10 km, evaluated at various significant time steps of the simulation. The  
544 selected time steps correspond to moments of particular interest, specifically the first year,  
545 the 1,000th year, and the 5,000th through 8,000th years in 1,000-year increments, followed  
546 by the 11,000th and 15,000th years.

547 Consistent with the observations derived from Fig. 5 (Part I), during the initial 4,000  
548 years of the simulation, the strain values within the upper layer of the plate remain relatively  
549 low, as significant strain is exerted predominantly on the slab's surface by the mantle. By the  
550 5,000th year, minor increases in strain are evident in the left portion of the slab, likely  
551 resulting from the initiation of cracks and the onset of lithospheric material detachment. By  
552 the 6,000th year, strain values have significantly increased across nearly half of the upper  
553 layer, driven by the propagation of cracks and the advancing infiltration of mantle material.

554 In the 7,000th year, the entire upper layer exhibits severe deformation, with strain  
555 values reaching their highest levels observed in the simulation. This deformation may  
556 indicate the detachment of material from the slab. During the remaining simulation period  
557 (8,000–15,000 years), the strain values display a generally consistent pattern of local  
558 maxima and minima, with only minor variations.



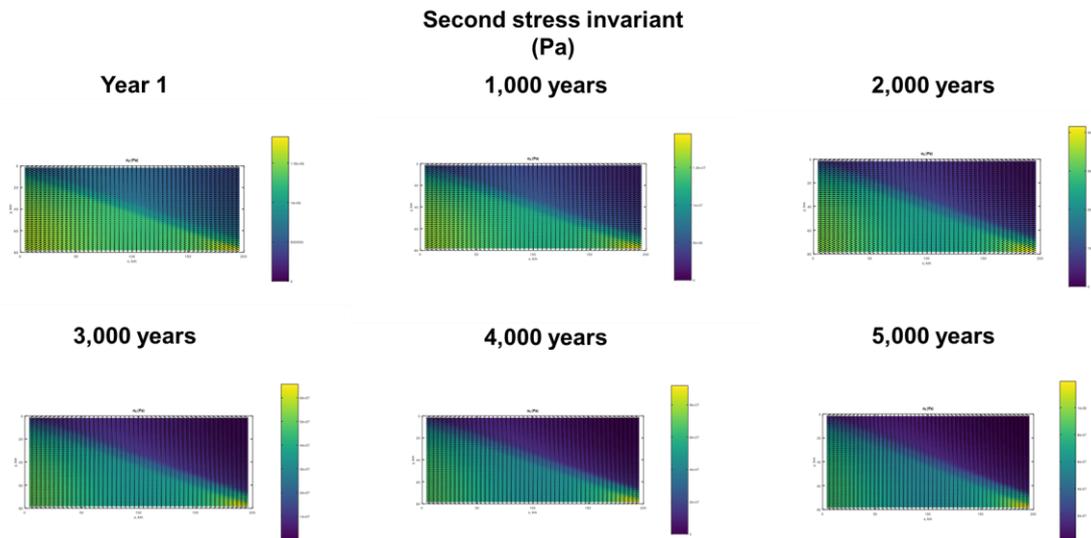
559

560 **Figure 7:** In this figure the second strain invariant values are presented, when the thickness  
 561 of the upper layer of the plate was set at 10 km, for various important time moments of the  
 562 simulation, in relation to the 49 vertical nodes that are located inside the upper deformed  
 563 part of the lithospheric slab.

564

## 565 5.2. Second Stress Invariant

566 In the following three-part Figure 6 the second stress invariant values for the whole  
 567 computational model are presented, when the thickness of the upper deformed part of the  
 568 plate was set at 5 km:

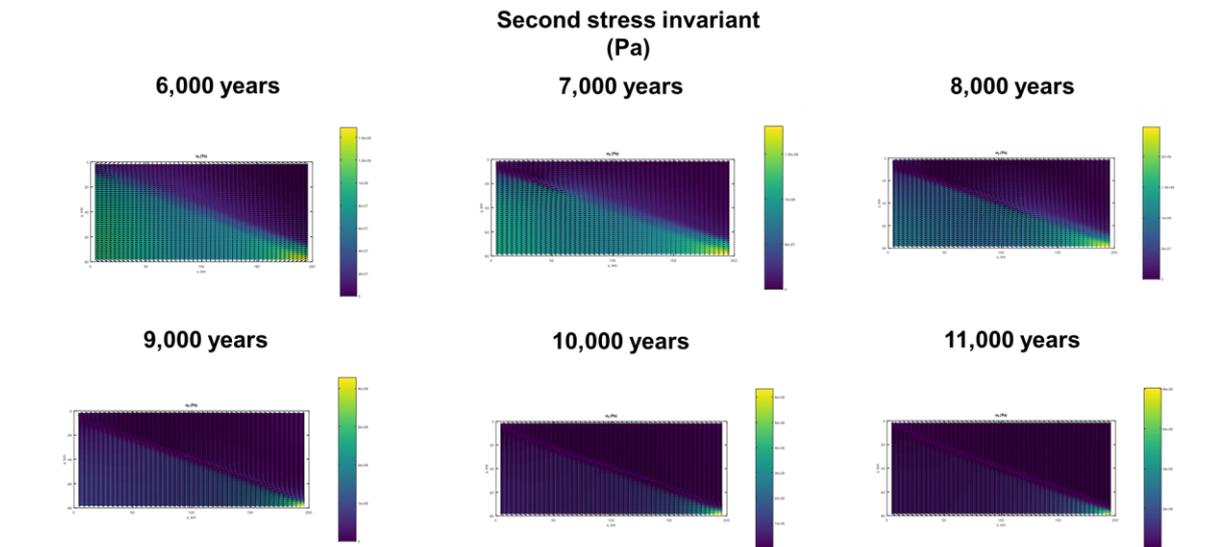


569

570 **Figure 8:** In part I of this figure the second stress invariant values are showcased for the  
 571 computational model, when the thickness of the upper deformed layer was set at 5 km,  
 572 particularly for the first year of the simulation, after 1,000 years, 2,000 years, 3,000, 4,000  
 573 and 5,000 years respectively.

574 In contrast, with the strain rate values, the stress values in the surface of the upper  
 575 part of the plate are generally low, with the highest stress values mostly located in the lower,  
 576 more stiffer parts of the plate. Additionally, it can be seen that the stress values have a slight  
 577 increasing tendency with the passage of time, throughout the early part of the computational  
 578 model, and especially, in the lower part of the slab. This increasing path can be seen from  
 579 the first year of the simulation, till the 8,000th year.

580 This observation demonstrates that the relationship between stress and strain is non-  
 581 linear. In fact, it appears inverse, as periods and locations where the upper layer exhibits  
 582 high strain values correspond to relatively low stress levels.



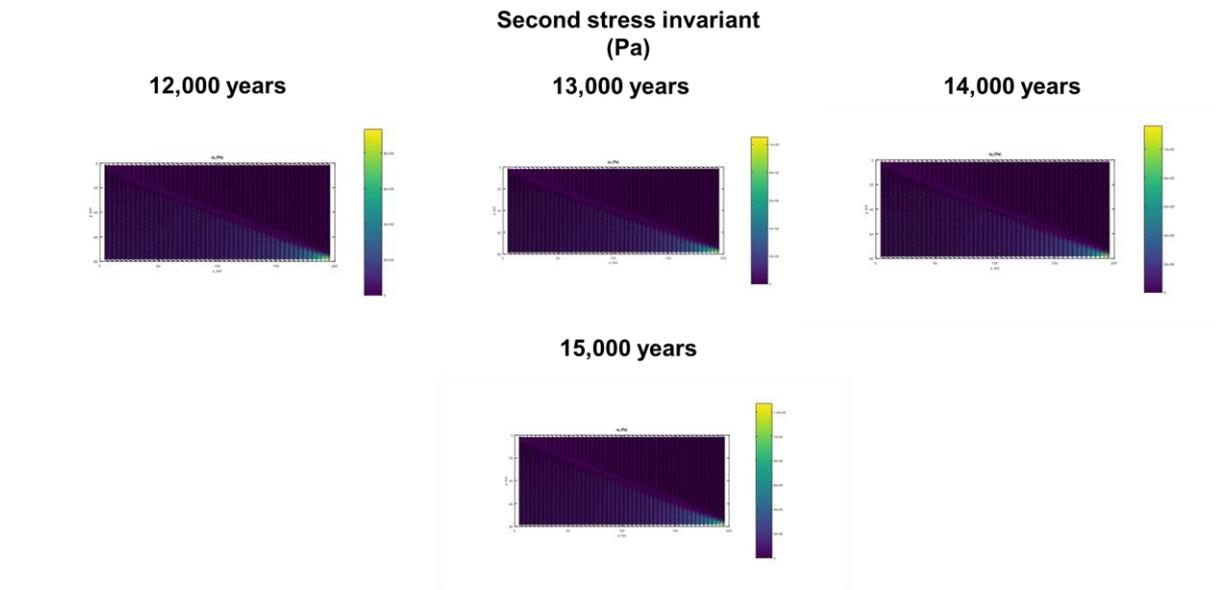
### Part II

583

584 **Figure 8:** In part II of this figure the second stress invariant values are showcased for the  
 585 computational model, when the thickness of the upper deformed layer was set at 5 km,  
 586 particularly after 6,000 years from the initiation of the simulation, after 7,000 years, 8,000  
 587 years, 9,000, 10,000 and 11,000 years respectively.

588 Additionally, for the final 6,000 years, 9,000-15,000, of the simulation the stress  
 589 values throughout almost the whole computational model are low and mostly constant with  
 590 only slight deviations. The latter note makes it clear that the high strain values that can be  
 591 seen inside the upper deformed layer, after the spreading of the cracks, in the 9,000th year,  
 592 is most likely caused by deformation creep.

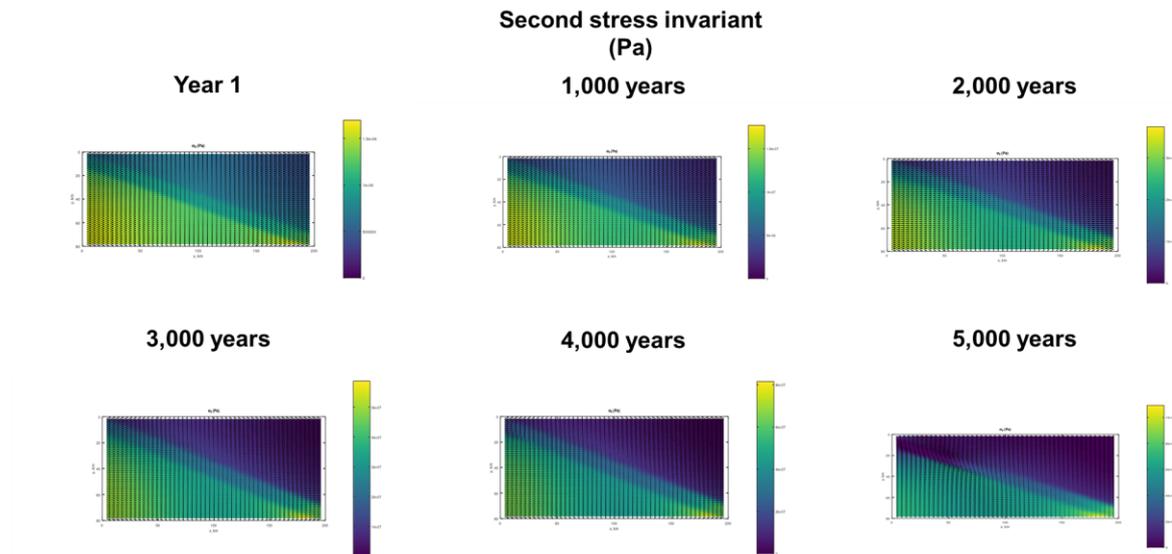
593



594

595 **Figure 8:** In part III of this figure the second stress invariant values are showcased for the  
 596 computational model, when the thickness of the upper deformed layer was set at 5 km,  
 597 particularly after 12,000 years from the initiation of the simulation, after 13,000 years, 14,000  
 598 and 15,000 years respectively.

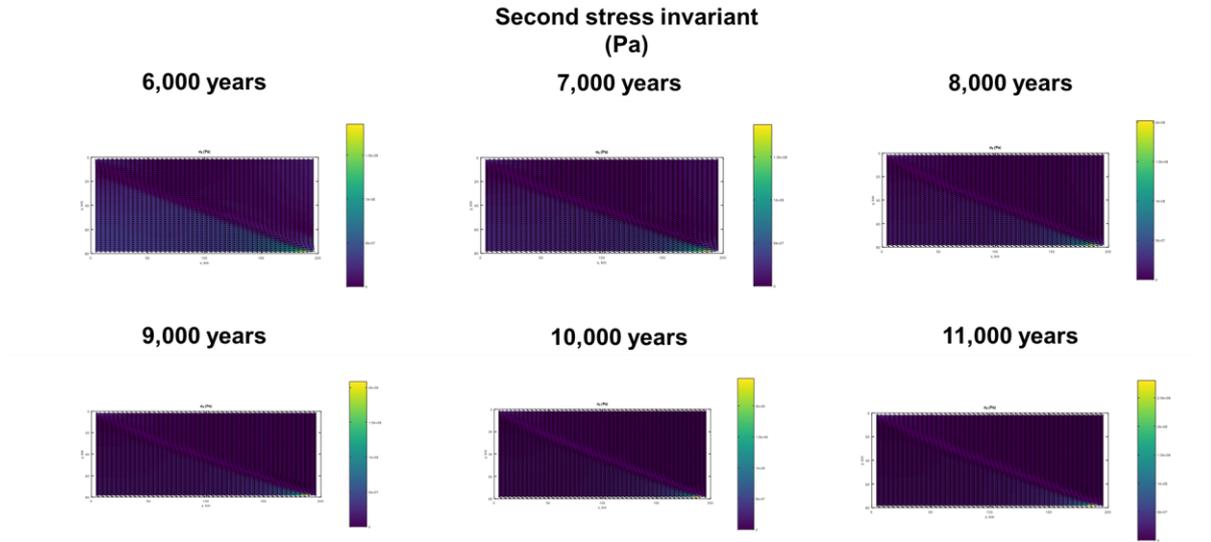
599 Furthermore, in the three-part Figure 7, the second stress invariant values for the  
 600 whole computational model are presented, when the thickness of the upper deformed part of  
 601 the plate was set at 10 km. As with the 5 km model the stress values in the surface of the  
 602 upper part of the plate are generally low, with the highest stress values mostly located in the  
 603 lower, more stiffer parts of the plate. Moreover, it can be seen that the stress values have  
 604 again a slight increasing tendency with the passage of time, throughout most of the  
 605 computational model, and especially, in the lower layer of the slab. However, in contrast with  
 606 the 5 km model, in this case the increasing value motive stops much earlier particularly in  
 607 the 5,000th year.



608

609 **Figure 9:** In part I of this figure the second stress invariant values are showcased for the  
 610 computational model, when the thickness of the upper deformed layer was set at 10 km,  
 611 particularly for the first year of the simulation, after 1,000 years, 2,000 years, 3,000, 4,000  
 612 and 5,000 years respectively.

613 From the 6,000th year till the end of the simulation, 6,000-15,000, the stress values  
 614 throughout almost the whole computational model are again low and mostly constant. Overall,  
 615 the primary difference between the 5 km and 10 km models is that, in the 10 km model,  
 616 comparable results are observed in a significantly shorter timeframe for both strain and stress  
 617 measurements. Consistent with the 5 km model, it is evident that the high strain values within  
 618 the upper deformed layer, following crack propagation in the 7,000th year, are likely  
 619 attributable to deformation creep.

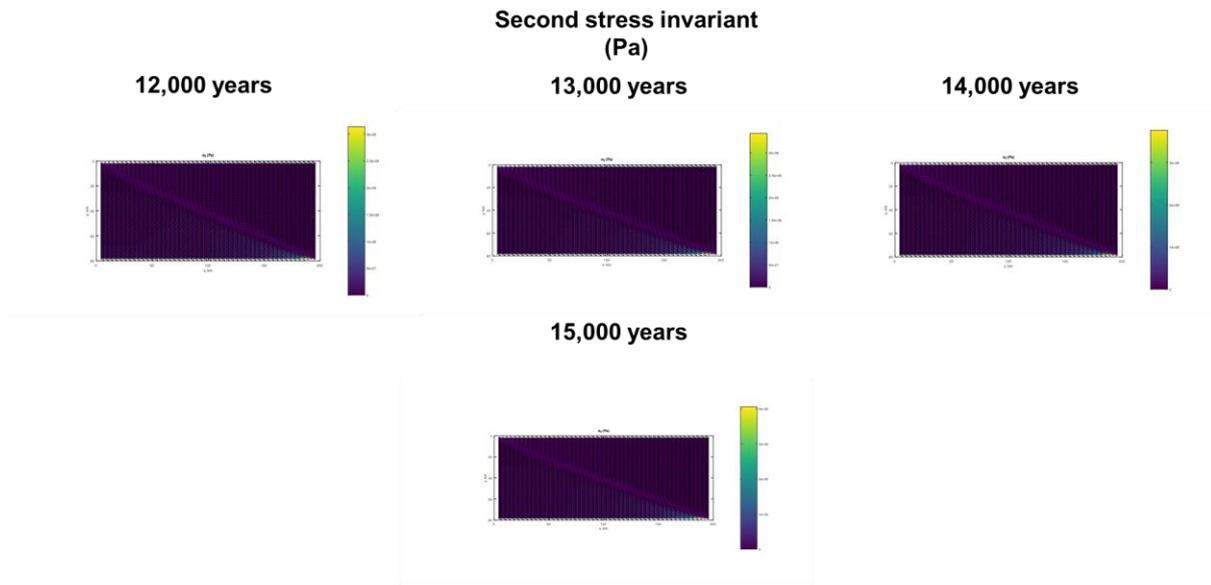


620

**Part II**

621 **Figure 9:** In part II of this figure the second stress invariant values are showcased for the  
 622 computational model, when the thickness of the upper deformed layer was set at 10 km,  
 623 particularly after 6,000 years from the initiation of the simulation, after 7,000 years, 8,000  
 624 years, 9,000, 10,000 and 11,000 years respectively.

625



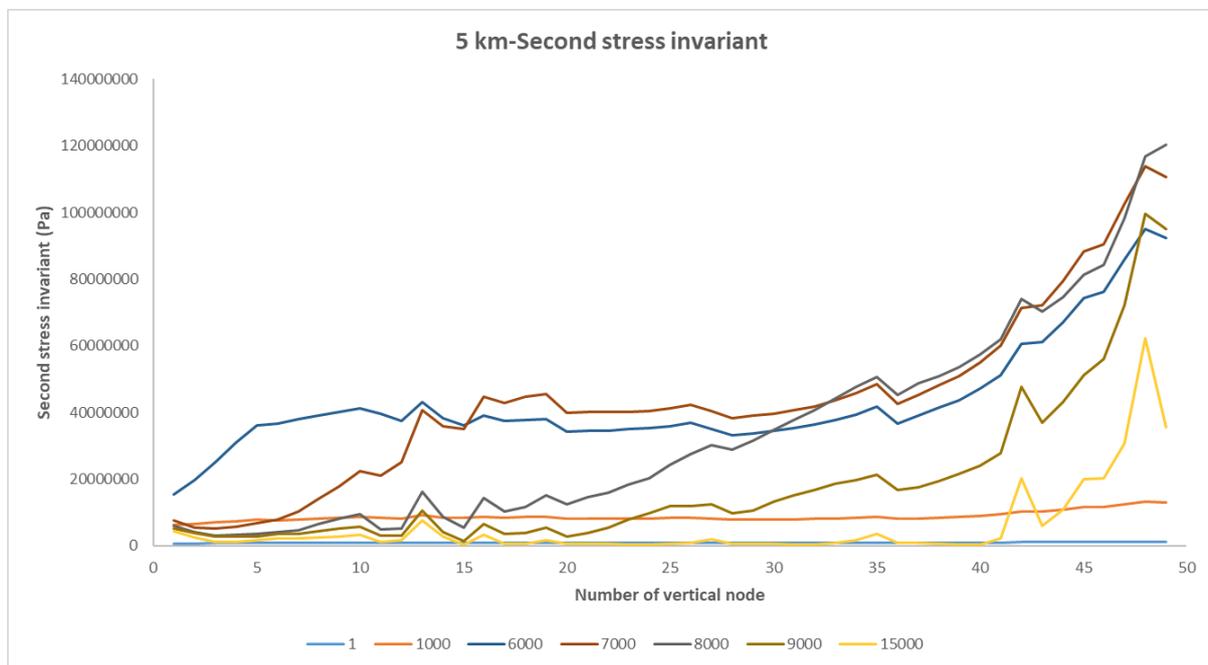
626

**Part III**

627 **Figure 9:** In part III of this figure the second strain rate invariant values are showcased for  
 628 the computational model, when the thickness of the upper deformed layer was set at 10 km,

629 particularly after 12,000 years from the initiation of the simulation, after 13,000 years, 14,000  
 630 and 15,000 years respectively.

631 Additionally, Figure 10 depicts the second stress invariant values at specific time  
 632 steps of the simulation, corresponding to the 49 vertical nodes situated within the upper thin  
 633 layer of the plate, with a thickness set at 5 km. The analyzed time steps include the first  
 634 year, the 1,000th year, and the 6,000th through 9,000th years in 1,000-year intervals, as well  
 635 as the 15,000th year.

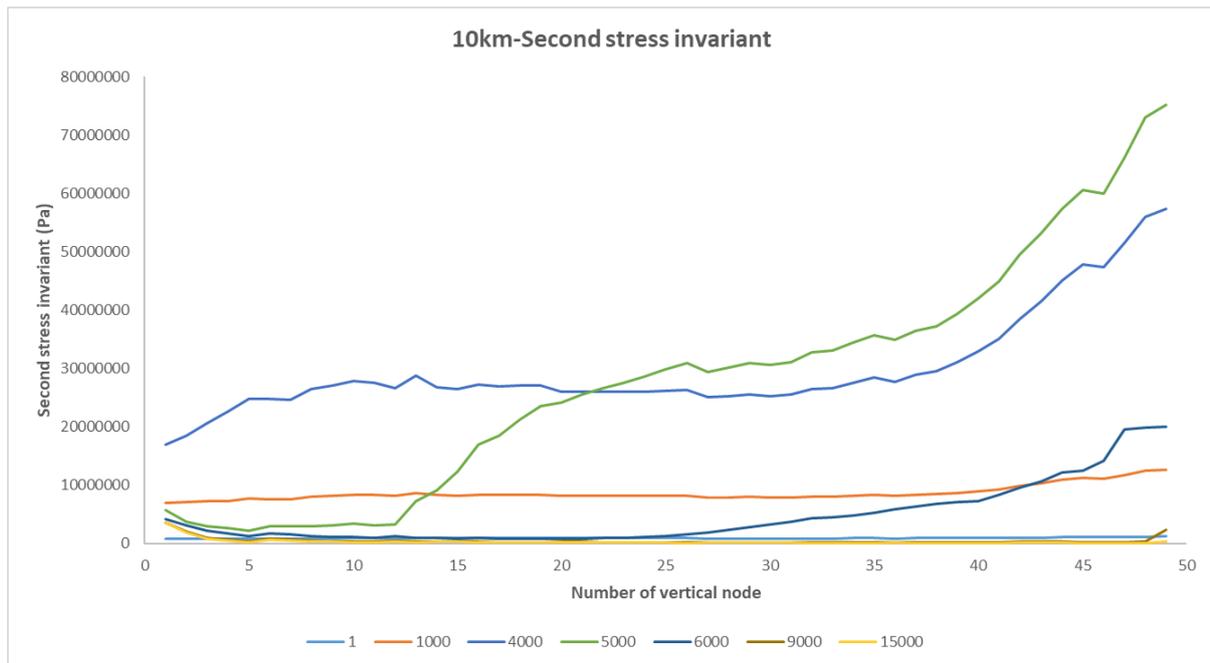


636

637 **Figure 10:** In this figure the second stress invariant values are presented, when the  
 638 thickness of the upper layer of the plate was set at 5 km, for various important time moments  
 639 of the simulation, in relation to the 49 vertical nodes that are located inside the upper  
 640 deformed part of the lithospheric slab.

641 Generally, from the first year of the simulation until the 8,000th year, the stress values  
 642 exhibit a consistent, gradual increase, particularly in the lower right section of the slab, with  
 643 the values reaching their peak at the 8,000th year. In alignment with Fig. 8 (Part II), from the  
 644 9,000th year onward, the stress values are mostly stable and constant.

645 Similarly, Figure 11 presents the second stress invariant values, analyzed in relation  
 646 to the 49 vertical nodes situated within the upper deformed layer of the plate. In this case, the  
 647 slab thickness is set at 10 km. The time steps depicted in Fig. 11 correspond to the first year,  
 648 the 1,000th year, the 4,000th year, and subsequent key intervals at the 5,000th, 6,000th,  
 649 9,000th, and 15,000th years after the start of the simulation.



650

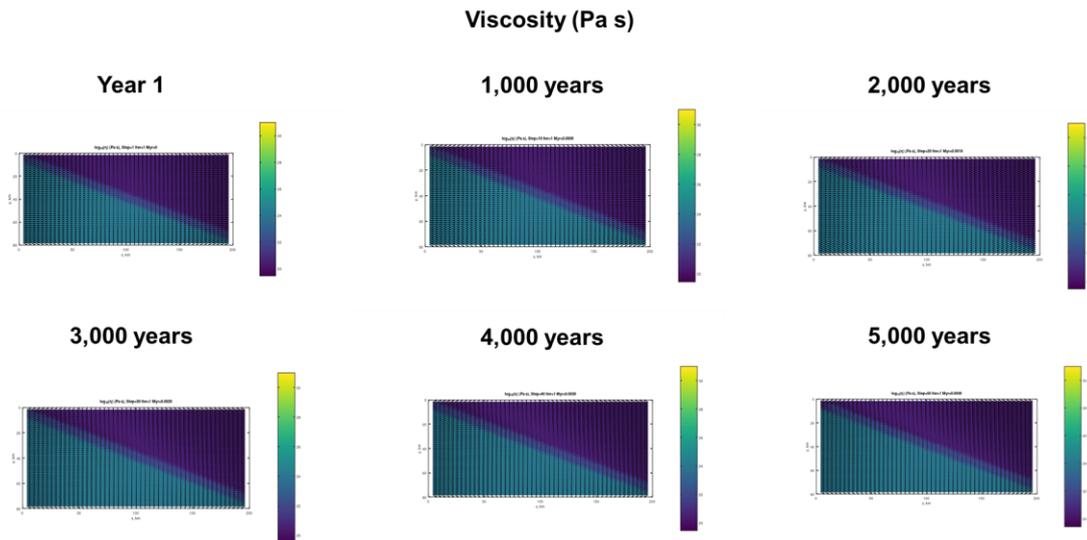
651 **Figure 11:** In this figure the second stress invariant values are presented, when the  
 652 thickness of the upper layer of the plate was set at 10 km, for various important time  
 653 moments of the simulation, in relation to the 49 vertical nodes that are located inside the  
 654 upper deformed part of the lithospheric slab.

655 Similarly to the 5 km model, an initial slight increase in stress values is observed  
 656 throughout the entire length of the upper layer during the first 5,000 years, coinciding with the  
 657 onset of crack propagation and the potential detachment of material from the slab. From the  
 658 6,000th year onward, the stress values are mostly low and constant in nearly the entire  
 659 computational model, particularly within the upper portion of the plate.

660

### 661 5.3. Viscosity

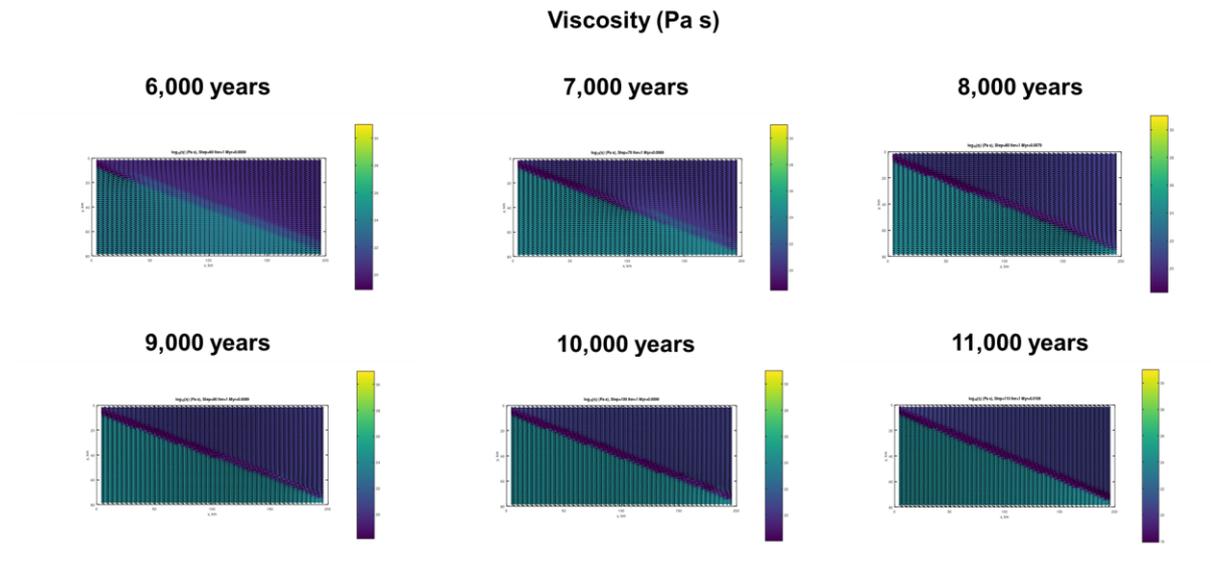
662 Overall, viscosity values exhibited only minor fluctuations over time. Figure 12,  
 663 presented in three parts, shows the viscosity map of the computational model when the  
 664 thickness of the plate's upper deformed layer was set to 5 km.



665

666 **Figure 12:** In part I of this figure the viscosity values are showcased for the computational  
 667 model, when the thickness of the upper deformed layer was set at 5 km, particularly for the  
 668 first year of the simulation, after 1,000 years, 2,000 years, 3,000, 4,000 and 5,000 years  
 669 respectively.

670 For the first 5,000 years of the simulation the viscosity values of the mantle, the upper  
 671 deformed layer and the lower stiffer layer of the plate mostly maintain their original given  
 672 values constant. However, in the 6,000th year, the same year that the major cracks start to  
 673 emerge in the upper left part of the plate, the viscosity values in that same area, become much  
 674 lower, and closely match that of the flowing mantle, indicating that the mantle has infiltrated  
 675 the upper deformed layer, through the newly formed cracks.

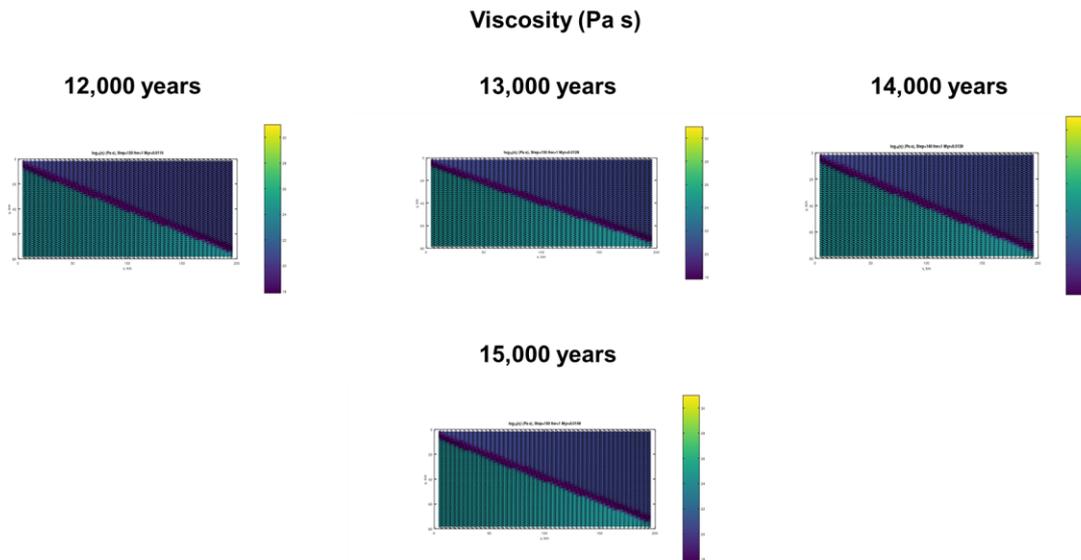


676

**Part II**

677 **Figure 12:** In part II of this figure the viscosity values are showcased for the computational  
 678 model, when the thickness of the upper deformed layer was set at 5 km, particularly after  
 679 6,000 years from the initiation of the simulation, after 7,000 years, 8,000 years, 9,000,  
 680 10,000 and 11,000 years respectively.

681 This decrease in viscosity values spreads throughout the whole length of the upper  
 682 layer, for the next 2,000 years, matching the propagation of the cracks, and perhaps the  
 683 detachment of lithospheric material, and thus making the mantle enter the rest of the layer.  
 684 Finally, the values of the viscosity again stabilise at the 9,000th year and remain mostly the  
 685 same for the rest of the simulation.



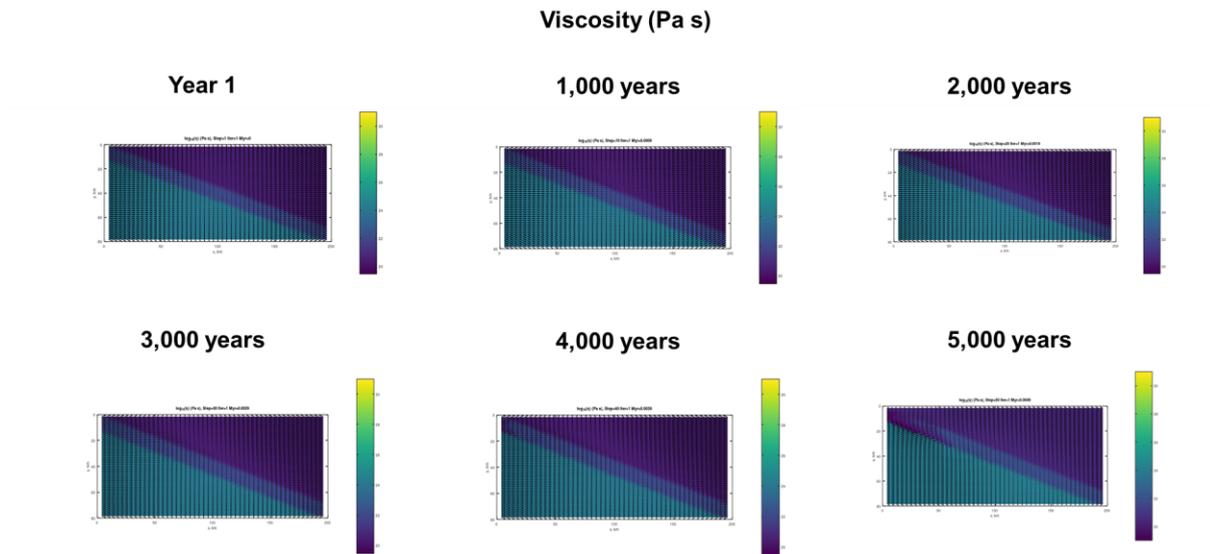
### Part III

686

687 **Figure 12:** In part III of this figure the viscosity values are showcased for the computational  
 688 model, when the thickness of the upper deformed layer was set at 5 km, particularly after  
 689 12,000 years from the initiation of the simulation, after 13,000 years, 14,000 and 15,000  
 690 years respectively.

691 Overall, in the three-part Figure 13, which corresponds to the viscosity values of the  
 692 computational model when the thickness of the upper deformed layer was set at 10 km, the  
 693 same observations can be made with the only difference being the different time moments  
 694 that the same events appear. For the first 4,000 years of the simulation the original given  
 695 values for the viscosity remain relatively the same. Following the 5,000th year, coinciding  
 696 with the onset of crack formation and the subsequent entry of the viscous mantle into the  
 697 plate's upper layer, the viscosity values in this layer begin to decrease. These reduced  
 698 viscosity values, closely resembling those of the flowing mantle, spread to encompass the  
 699 entire length of the heavily deformed upper layer within only 1,000 years. Consequently,  
 700 from the 6,000th year until the end of the simulation, viscosity values remain consistently  
 701 lower than their initial levels, as the mantle has fully infiltrated through the cracks

702

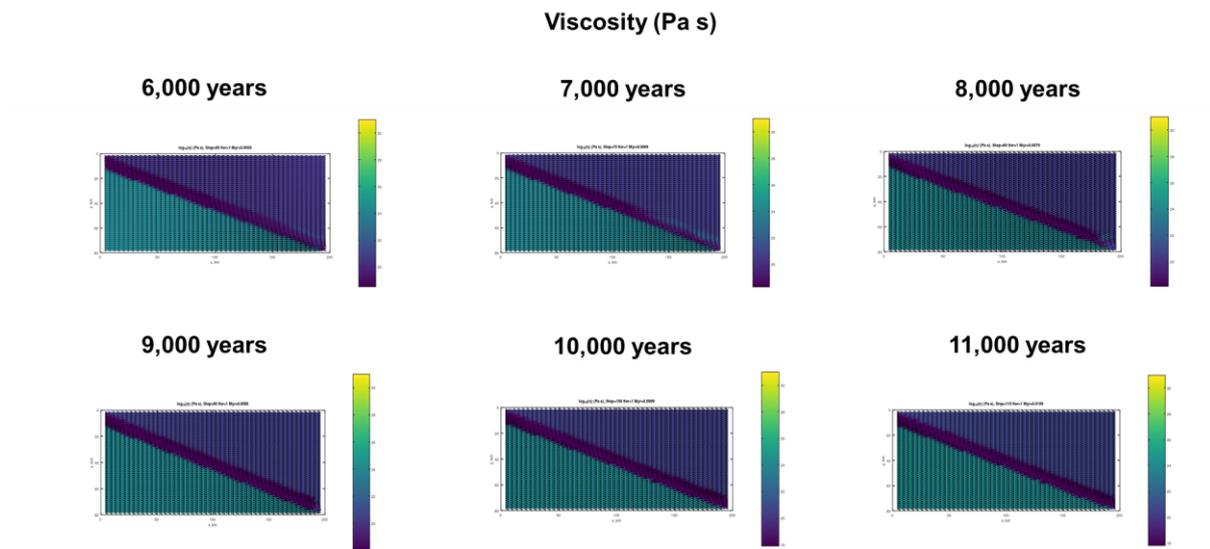


703

**Part I**

704 **Figure 13:** In part I of this figure the viscosity values are showcased for the computational  
 705 model, when the thickness of the upper deformed layer was set at 10 km, particularly for the  
 706 first year of the simulation, after 1,000 years, 2,000 years, 3,000, 4,000 and 5,000 years  
 707 respectively.

708



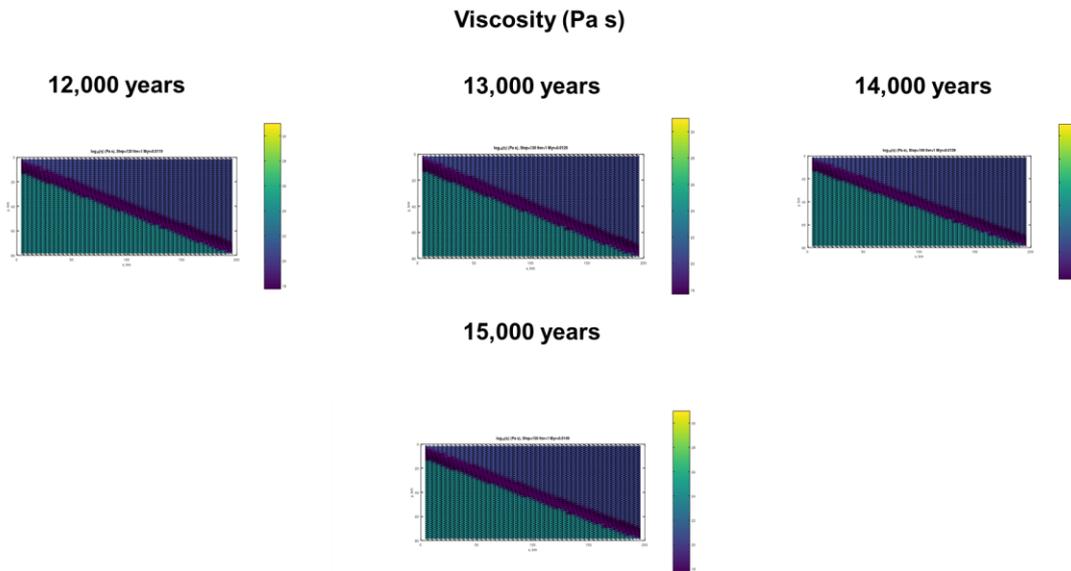
709

**Part II**

710 **Figure 13:** In part II of this figure the viscosity values are showcased for the computational  
 711 model, when the thickness of the upper deformed layer was set at 10 km, particularly after

712 6,000 years from the initiation of the simulation, after 7,000 years, 8,000 years, 9,000,  
 713 10,000 and 11,000 years respectively.

714



715

### Part III

716 **Figure 13:** In part III of this figure the viscosity values are showcased for the computational  
 717 model, when the thickness of the upper deformed layer was set at 10 km, particularly after  
 718 12,000 years from the initiation of the simulation, after 13,000 years, 14,000 and 15,000  
 719 years respectively.

720 A key observation that further supports the hypothesis of lithospheric material  
 721 detachment due to extensive cracking and mantle infiltration within the plate's upper layer  
 722 appears in the 7,000th year. Close examination reveals two rock segments on the upper  
 723 right side of the slab's heavily deformed layer, each approximately 30 km in length, that  
 724 retain their original viscosity values. This suggests that the cracks may not have yet  
 725 propagated to this section of the layer, while the rest of the layer shows full crack formation  
 726 and consequent mantle infiltration. By the 8,000th year, however, only a small portion,  
 727 around 10 km, of the previously unaffected segments remained unchanged. This remaining  
 728 segment appears semi-detached from the plate and shows an upward orientation.

729 In sum, this detail reinforces that the cracks initiated in both the 5 km and 10 km  
 730 models create rock segments, ranging from 10 to 30 km in length, that detach from the  
 731 upper plate layer, thereby allowing the mantle to fully infiltrate the surrounding highly  
 732 fractured and deformed material.

733

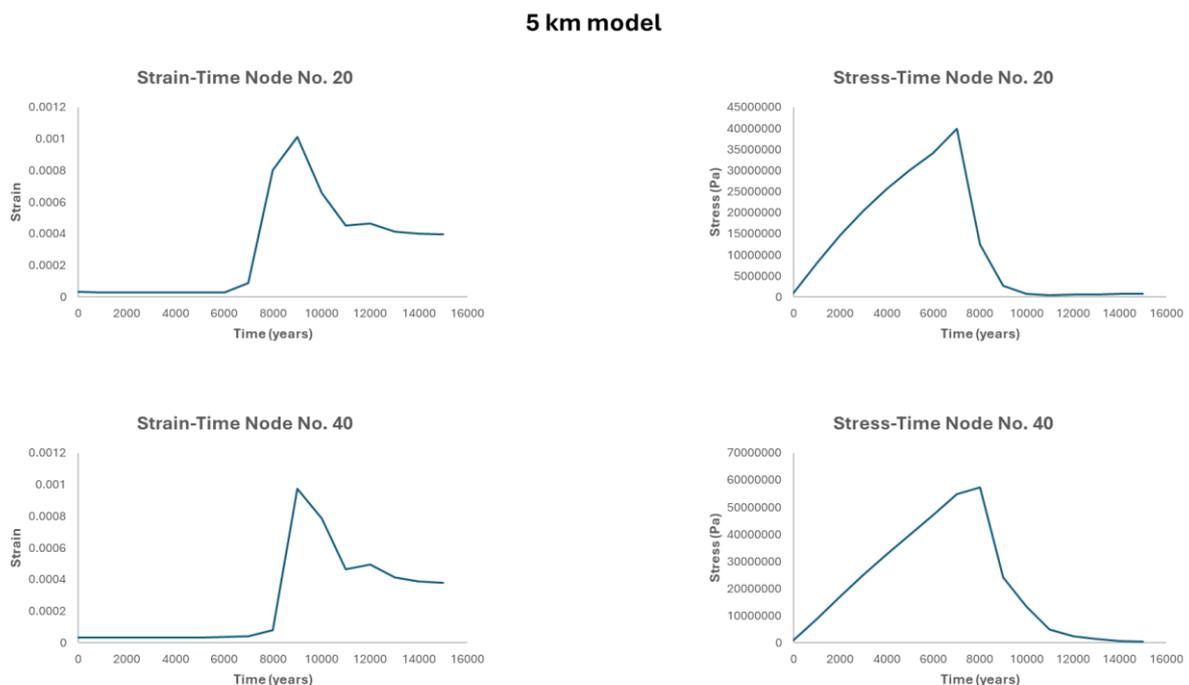
734

## 735 6. Remarks

### 736 6.1. Stress-Strain Nonlinear Relation

737 In this section the nonlinear relation between arguably the two most important  
 738 parameters of the computational model, the second strain invariant and the second stress  
 739 invariant is further analysed. As it has been already stated there appears to be a paradoxical  
 740 relation between the stress and the strain, particularly, when the stress values are high, the  
 741 strain values do not follow the same pattern, but quite the opposite and the same applies for  
 742 the contrary.

743 Overall, this important observation can be further seen in the following Figures 14  
 744 and 15, when the thickness of the upper layer of the plate was set at 5 and 10 km  
 745 respectively. In the latter figures the stress and strain values are presented in relation to the  
 746 simulation time, for one out of the 49 vertical nodes that are located inside the upper  
 747 deformed layer of the slab.



748

749 **Figure 14:** In this figure the stress and strain values are depicted in relation to the simulation  
 750 time for the 20th and 40th node, while the thickness of the upper layer of the slab was set at  
 751 5 km.

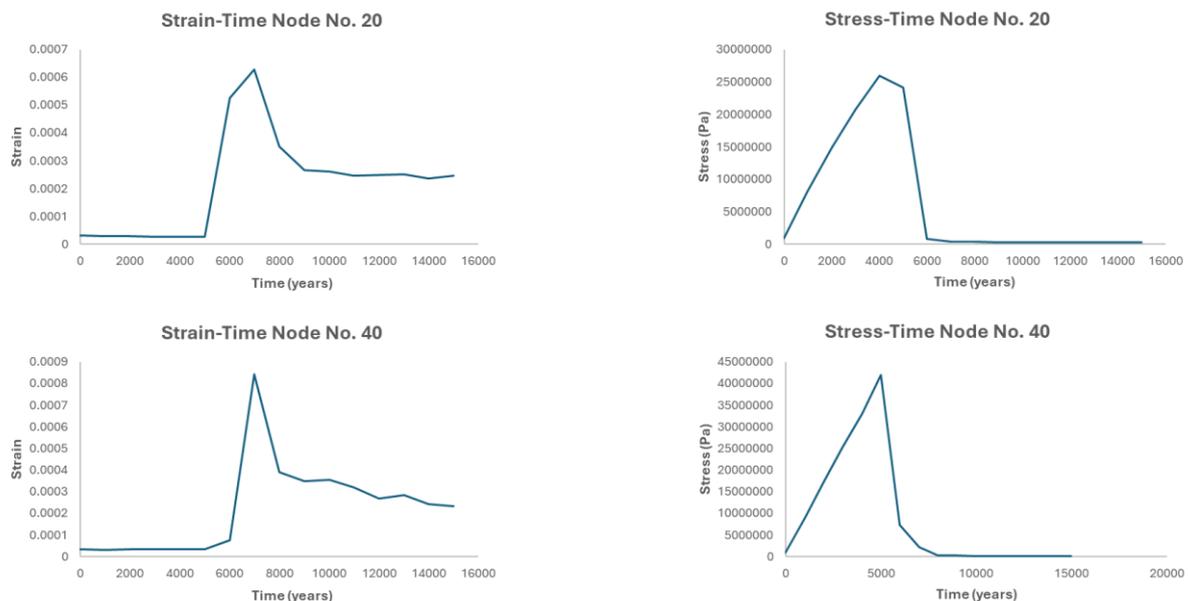
752 Firstly, all of the nodes display somewhat the same results, making most of the graphs  
 753 similar, and hence the 20th and 40th nodes were chosen completely randomly. As for Fig. 14,  
 754 which represents the 5 km model, it can be clearly seen that for the first 6,000 years of the  
 755 simulation the stress values are constantly increasing almost linearly, while at the same time  
 756 the strain values inside the upper layer of the plate are almost zero. As it has been previously  
 757 stated this can be easily explained because the stress that is exerted inside the layer may not  
 758 be sufficient to cause any major, or noteworthy strain, since the stress values have a range of

759 5-40 MPa, which are relatively really low values when taking into account the geodynamic  
760 process that is taking place.

761 However, after the 7,000th year for the 20th node, it appears that the upper layer  
762 developed enough stress, to cause cracks, and to generally generate deformation, hence the  
763 increasing strain values. The cracks needed around 1,000 years to propagate to the 40th  
764 node, thus explaining that the values of strain increased there almost 1,000 years later.  
765 Generally, with the initiation of the cracks and commencement of the increase in strain values,  
766 the stress values suddenly decrease fast, further explaining and justifying that major cracks  
767 occurred in the length of the upper layer of the plate, because with the development of cracks  
768 comes stress release.

769 Despite the sudden decrease in the stress values, and especially, in the node numbers  
770 10-40, the strain values maintain a constant high value. This can be only attributed to the  
771 development of diffusion creep (Gerya et al., 2002; Gerya, 2010). In general, creep  
772 deformation is time-dependent and occurs under constant stresses (Gariboldi and Spigarelli,  
773 2019), just like in the present case. Creep deformation is a very common phenomenon in  
774 subducting plates that has been observed and studied by many previous researchers  
775 (Stöckhert, 2002; Wassmann and Stöckhert, 2013; Agard et al., 2020).

#### 10 km model



776

777 **Figure 15:** In this figure the stress and strain values are depicted in relation to the simulation  
778 time for the 20th and 40th node, while the thickness of the upper layer of the slab was set at  
779 10 km.

780 Similarly, in Figure 15, which corresponds to the 10 km model, observations align  
781 with those of the 5 km model, with the primary distinction being that these events occur  
782 earlier. Specifically, stress decreases while strain increases, beginning at the 20th node  
783 around the 4,000th year. By approximately the 7,000th year, stress reaches a low, stable  
784 value, and creep deformation initiates.

## 785 **6.2. Effect of the Thickness of the Upper Deformed Layer of the** 786 **Subducting Plate**

787 Overall, another key insight drawn from the simulation results suggest that the  
788 thickness of the upper deformed layer of the subducting plate only influences the timing of  
789 which the same events will occur. Specifically, by increasing the thickness of the layer, its  
790 structural strength is reduced. In the 10 km model, cracks, potential material detachment,  
791 and viscous mantle infiltration occur within 5,000 years and subsequently propagate along  
792 the entire layer. In contrast, in the 5 km model, this sequence begins 1,000 years later, in the  
793 6,000th year.

794

## 795 **7. Comparison with Observations from the Hellenides**

796 The final step in supporting the aforementioned assumptions involves examining  
797 whether metamorphic rocks of this nature are present in the broader Hellenides region and  
798 attempting to identify the depths from which they originated. Furthermore, it is crucial to note  
799 that after detachment from the upper portion of the subducting plate, it is highly probable that  
800 such rocks would require millions of years to re-emerge at the Earth's surface. Initially, these  
801 rocks are expected to remain within the viscous flowing mantle, where the high-pressure,  
802 low-temperature conditions facilitate their metamorphism from granite to quartzites and  
803 phyllites occurs. At the typical detachment depth of approximately 30 km, mantle  
804 temperatures are around 520 K (approximately 250°C). The lithostatic pressure at this depth  
805 can be estimated using the following equation (Turcotte and Schubert, 1982):

$$806 \quad P_L = \rho g y \quad (20),$$

807 where,  $P_L$ : is the lithostatic pressure,  $\rho$ : is the density of the detached segment and  $y$ : is the  
808 depth from the surface. Consequently, the pressure exerted on the segment is  
809 approximately 824 MPa. These pressure-temperature (P-T) conditions are classified as high-  
810 pressure and low-temperature (HP/LT). Thus, it is essential to investigate whether any  
811 HP/LT exhumed metamorphic rocks exist within the Hellenides and to estimate the time  
812 required for their exhumation. However, it is critical to establish that, if such exhumed rocks  
813 are identified, they originated from the same detachment depth as those observed in the  
814 simulation. This would confirm that shallow-depth rocks, detached from the upper part of the  
815 subducting lithospheric plate in under 10,000 years, can indeed undergo exhumation.

816 A comprehensive study by Thomson et al. (1998) highlights the presence of HP/LT  
817 exhumed metamorphic rocks within the Hellenides, particularly in Crete. The study further  
818 indicates that these HP/LT rocks predominantly consist of quartz and phyllite, two common  
819 minerals found in gneiss. Moreover, it is reported that these rocks were originally detached  
820 from the lower subducting lithospheric plate at depths of approximately 30–35 km, where  
821 they also underwent metamorphism shortly after detachment (Thomson et al., 1998). Finally,  
822 the P-T conditions at which the metamorphism occurred match that of the computational  
823 setting, particularly, the maximum temperature was found to be around  $300 \pm 50^\circ\text{C}$  and the  
824 pressure  $800 \pm 300 \text{ MPa}$  (Thomson et al., 1998). It should be noted that various others  
825 researchers have investigated the HP/LT exhumed metamorphic rocks that appear in Crete

826 (Fassoulas et al., 1994; Chatzaras et al., 2006; Ring et al., 2010), with the most complete  
827 and systematic study being that of Thomson et al. (1998).

## 828 **8. Discussion and Conclusions**

829 In this study a two-dimensional thermomechanical visco-elasto-plastic code,  
830 originally created by Gerya (2010), was heavily modified in order to simulate the subducting  
831 plate of the Mediterranean ridge. The aim of the study was firstly to investigate whether at  
832 the shallow depths of the lithosphere, particularly close to 30 km, if the upper deformed part  
833 of the downgoing plate endures such increased strain values that the detachment of  
834 segments occurs. The simulation results clearly showcased that major cracks begin to  
835 propagate (see Fig. 4 part II for the 5 km model and Fig. 5 part I for the 10 km model) for  
836 both model sizes, and in general after a very short geological time. For the 5 km model (see  
837 the three-part Fig. 4) the cracks initiated after only 6,000 years, as for the 10 km model (see  
838 the three-part Fig. 5) the cracks initiated 1,000 years earlier, after only 5,000 years. After the  
839 initiation of the cracks, the latter slowly propagated towards the lower parts of the upper  
840 section of the plate (see Fig. 4 part II, III and Fig. 5 part II), until they had covered the whole  
841 length of the layer.

842 Additionally, it was observed that the viscosity values of the “broken” upper layer  
843 showcased a significant drop (see Fig. 12 part II for the 5 km model and Fig. 13 part I for the  
844 10 km model), and the values closely aligned with those of the flowing mantle, thus  
845 indicating that the latter infiltrated the newly formed cracks and started flowing through the  
846 cracks of the layer. Furthermore, another important assumption that was derived from the  
847 simulation results was the non-linear relation between the stress and strain values on the  
848 surface of the plate and inside the upper deformed layer of the plate (see Fig. 14 for the 5  
849 km model and Fig. 15 for the 10 km model). At the initial stages of the simulation, this  
850 paradoxical relation was attributed to the very low stress values, around 5-40 MPa, that can  
851 be deemed insufficient to cause any major strain deformation on the subducting slab. As for  
852 the later stages of the computational model, this non-linear relation can be easily attributed  
853 due to the development of diffusion creep inside the heavily broken and deformed upper part  
854 of the plate, a phenomenon that has been studied and observed by various researchers over  
855 the past decades (Gerya et al., 2002; Stöckhert, 2002; Gerya, 2010; Wassmann and  
856 Stöckhert, 2013; Agard et al., 2020).

857  
858 Moreover, an additional significant observation from the computational model, which  
859 supports the authors hypothesis, that during the initiation and the propagation of the cracks,  
860 respectively, medium to small size segments from the upper part of the layer were  
861 completely detached, was seen in the 10 km model after 7,000 years that the simulation had  
862 started (see Fig. 13 part II). Particularly, it was clearly seen that small rock segments, having  
863 a length of around 10-30 km, had kept their original given viscosity value and they were  
864 semi-detached from the plate, while at the same time having an upwards motion. After all the  
865 aforementioned assumptions, the natural question that arose was if these detached  
866 segments of rock ever made it back to the surface, thus getting exhumed.

867 Another very common and more simpler exhumation mechanism that has been  
868 suggested is that the detached rock is driven upwards, and consequently exhumed, due to  
869 the density difference between the flowing mantle and the density of the segment (Liou et

870 al., 2004; Warren et al., 2008a, 2008b; Beaumont et al., 2009; Warren, 2013), thus  
871 producing a upwards buoyancy driven force, capable of exhuming the detached rock. The  
872 latter model can be easily considered suitable for the current setting since all the necessary  
873 prerequisites were present in the computational model. Finally, having proven that segments  
874 are detached from the downgoing plate, at a relatively shallow depth of around 30 km, and  
875 knowing that the Mediterranean ridge setting favours one of the many proposed exhumation  
876 mechanisms, the only important question remaining is whether rocks of that types have ever  
877 been observed in the surface of the Hellenides.

878 The detached rocks from the subducting plate were subjected to high pressure and  
879 relatively low temperature conditions for an extended period of time, thus making them  
880 HP/LT metamorphic rocks. Overall, such HP/LT exhumed metamorphic rock have been  
881 widely studied and observed in the Greek island of Crete (Fassoulas et al., 1994; Thomson  
882 et al., 1998; Chatzaras et al., 2006; Ring et al., 2010). Particularly, the in-depth study by  
883 Thomson et al. (1998) clearly indicated that these HP/LT exhumed metamorphic rocks had  
884 been detached from the downgoing plate from the shallow depths of around 30-35 km. In  
885 addition, by using thermochronology the researchers identified that the detachment of the  
886 segments occurred around 24 to 19 Ma ago.

887 To sum up, in this study a combination of computational, theoretical techniques and  
888 macroscopic observations were utilised to fully investigate whether shallow depth rocks are  
889 detached from the subducting plate, in the Mediterranean ridge, and consequently, if they  
890 ever reach the surface of the Earth again, in the broader area of the Hellenides. The general  
891 results and assumptions of the study indicated and proved that the latter process does occur  
892 in the subduction setting of the Mediterranean ridge. Overall, due to the limited available  
893 literature on the subject of modelling and mechanisms of shallow depth detachment of  
894 material from a subducting plate, the current simulation results and general efforts serve as  
895 a stepping stone to a process that requires more analysis and investigations from  
896 researchers in the future. An important future step, and consequent idea for research,  
897 towards the better understanding of the mechanisms behind the latter process is to fully  
898 incorporate the subducting rate parameter of the plate in future computational models, in  
899 order to make the model even more realistic, a parameter that was neglected in the present  
900 study due to the short geological time of the simulation.

## 901 **Appendix**

902 Additional material that further explains the structure and mechanisms of the utilised code,  
903 as well as access to the modified version of the code used by the authors during the  
904 simulation will be provided upon request to the corresponding author.

## 905 **Acknowledgments**

906 The authors would like to thank Maria-Sotiria Frousiou and Stefanos Petrakopoulos for  
907 designing Figure 1 and 2 respectively.

908

909

910 **References**

- 911 Agard, P., Prigent, C., Soret, M., Dubacq, B., Guillot, S., Deldicque, D., 2020. Slabification:  
912 Mechanisms controlling subduction development and viscous coupling. *Earth-Science*  
913 *Reviews* (2020), Vol. 208, 103259. doi: <https://doi.org/10.1016/j.earscirev.2020.103259>.
- 914 Albers, M., 2000. A local mesh refinement multigrid method for 3D convection problems with  
915 strongly variable viscosity. *J. Comput. Phys.* (2000), Vol. 160, pp. 126–150.
- 916 Beaumont, C., Jamieson, R., A., Butler, J., P., Warren, C., J., 2009. Crustal structure: A key  
917 constraint on the mechanism of ultra-high-pressure rock exhumation. *Earth and Planetary*  
918 *Science Letters* (2009), Vol. 287, pp. 116–129. doi: 10.1016/j.epsl.2009.08.001.
- 919 Capitanio, F., A., Morra, G., Goes, S., 2007. Dynamic models of downgoing plate-buoyancy  
920 driven subduction: Subduction motions and energy dissipation. *Earth and Planetary Science*  
921 *Letters* (2007), Vol. 262, pp. 284–297. doi: 10.1016/j.epsl.2007.07.039.
- 922 Chapple, W., M., Tullis, T., E., 1977. Evaluation of forces that drive plates. *J. Geophys. Res.*,  
923 Vol. 82, pp. 1967–1984.
- 924 Chatzaras, V., Xypolias, P., Doutsos, T., 2006. Exhumation of high-pressure rocks under  
925 continuous compression: a working hypothesis for the southern Hellenides (central Crete,  
926 Greece). *Geological Magazine*, Vol. 143, Issue 6, pp. 859-876, November 2006. doi:  
927 <https://doi.org/10.1017/S0016756806002585>.
- 928 Chemenda, A., I., Mattauer, M., Bokun, A., N., 1996. Continental subduction and a mechanism  
929 for exhumation of high-pressure metamorphic rocks: new modelling and field data from Oman.  
930 *Earth and Planetary Science Letters* (1996), Vol. 143, pp. 173-182.
- 931 Christensen, U., 1982. Phase boundaries in finite amplitude mantle convection. *Geophysical*  
932 *Journal of the Royal Astronomical Society*, Vol. 68, pp. 487–97.
- 933 Cloos, M., 1982. Flow melanges; numerical modeling and geologic constraints on their origin  
934 in the Franciscan subduction complex, California. *Geol. Soc. Am. Bull.*, Vol. 93, pp. 330–344.
- 935 Cloos, M., Shreve, R., L., 1988a. Subduction-channel model of prism accretion, melange  
936 formation, sediment subduction, and subduction erosion at convergent plate margins; Part I,  
937 Background and description. *Pure Appl. Geophys.*, Vol. 128, pp. 455–500.
- 938 Cloos, M., Shreve, R., L., 1988b. Subduction-channel model of prism accretion, melange  
939 formation, sediment subduction, and subduction erosion at convergent plate margins; Part II,  
940 Implications and discussion. *Pure Appl. Geophys.*, Vol. 128, pp. 501–545.
- 941 Davies, G., F., 1988. Role of the lithosphere in mantle convection. *J. Geophys. Res.*, Vol. 93,  
942 pp. 10451–10466.
- 943 England, P., C., Holland, T., J., B., 1979. Archimedes and the Tauern eclogites: the role of  
944 buoyancy in the preservation of exotic tectonic blocks. *Earth Planet. Sci. Lett.*, Vol. 44, pp.  
945 287–294.

- 946 Fassoulas, C., Kiliyas, A., Mountrakis, D., 1994. Postnappe stacking extension and exhumation  
947 of high-pressure/low-temperature rocks in the island of Crete, Greece. *Tectonics*, Vol. 13, No.  
948 1, pp. 127-138, February 1994.
- 949 Forsyth, D., Uyeda, S., 1975. On the relative importance of driving forces of plate motion.  
950 *Geophys. J. R. Astron. Soc.*, Vol. 43, pp. 163–200.
- 951 Gariboldi, E., Spigarelli, S., 2019. Creep and High-Temperature Deformation of Metals and  
952 Alloys. *Metals* 2019, 9(10), 1087; doi: <https://doi.org/10.3390/met9101087>.
- 953 Gerya, T., V., 2010. *Introduction to Numerical Geodynamic Modelling*. Cambridge University  
954 Press, First Edition. doi: <https://doi.org/10.1017/9781316534243>.
- 955 Gerya, T., V., Stöckhert, B., Perchuk, A., L., 2002. Exhumation of high-pressure  
956 metamorphic rocks in a subduction channel: A numerical solution. *Tectonics* (2002), Vol. 21,  
957 No. 6, 1056, doi: 10.1029/2002TC001406.
- 958 Gerya, T., V., Yuen, D., A., 2003a. Rayleigh-Taylor instabilities from hydration and melting  
959 propel 'cold plumes' at subduction zones. *Earth and Planetary Science Letters* (2003), Vol.  
960 212, pp. 47-62. doi: 10.1016/S0012-821X(03)00265-6.
- 961 Gerya, T., V., Yuen, D., A., 2003b. Characteristics-based marker-in-cell method with  
962 conservative finite-differences schemes for modeling geological flows with strongly variable  
963 transport properties. *Physics of the Earth and Planetary Interiors* (2003), Vol. 140, pp. 293–  
964 318. doi: 10.1016/j.pepi.2003.09.006.
- 965 Ghosh, D., Maiti, G., Mandal, N., Baruah, A., 2020. Cold Plumes Initiated by Rayleigh-Taylor  
966 Instabilities in Subduction Zones, and Their Characteristic Volcanic Distributions: The Role of  
967 Slab Dip. *Journal of Geophysical Research: Solid Earth*, Vol. 125, e2020JB019811. doi:  
968 <https://doi.org/10.1029/2020JB019814>.
- 969 Goes, S., Capitanio, F., A., Morra, G., Seton, M., Giardini, D., 2011. Signatures of downgoing  
970 plate-buoyancy driven subduction in Cenozoic plate motions. *Physics of the Earth and*  
971 *Planetary Interiors* (2011), Vol. 184, pp. 1–13. doi: 10.1016/j.pepi.2010.10.007.
- 972 Gustafsson, B., 2008. *High Order Finite-Difference Methods for Time-dependent PDE*.  
973 Springer-Verlag (2008).
- 974 Kaus, B., J., P., Becker, T., W., 2007. Effects of elasticity on the Rayleigh–Taylor instability:  
975 implications for large-scale geodynamics. *Geophys. J. Int.* (2007), Vol. 168, pp. 843–862. doi:  
976 10.1111/j.1365-246X.2006.03201.x.
- 977 Li, Z., H., 2014. A review on the numerical geodynamic modeling of continental subduction,  
978 collision and exhumation. *Science China, Earth Sciences*, Vol. 57, pp. 47-69. doi:  
979 10.1007/s11430-013-4696-0.
- 980 Li, Z., H., Xu, Z., Q., Gerya, T., V., Ahmand, D., I., 2012. Numerical geodynamic modeling of  
981 continental convergent margins. *Earth Sciences*, Chapter 13.

- 982 Liou, J., G., Tsujimori, T., Zhang, R., Y., Katayama, I., and Maruyama, S., 2004. Global UHP  
983 metamorphism and continental subduction/collision, the Himalayan model. *Int. Geol. Rev.*,  
984 Vol. 46, pp. 1-27, 2004.
- 985 Lynch, D., R., 2005. *Numerical Partial Differential Equations for Environmental Scientists and*  
986 *Engineers: A Practical First Course*. Springer-Verlag (2005).
- 987 Makris, J., 2010. Geophysical studies and tectonism of the Hellenides, *Bull. Geol. Soc.*  
988 *Greece*, XLIII (1), pp. 32–45.
- 989 Makris, J., Papoulia, J., Yegorova, T., 2013. A 3-D density model of Greece constrained by  
990 gravity and seismic data. *Geophys. J. Int.* (2013), Vol. 194, pp. 1–17. doi: 10.1093/gji/ggt059.
- 991 Mancktelow, N., S., 1995. Nonlithostatic pressure during sediment subduction and the  
992 development and exhumation of high-pressure metamorphic rocks. *J. Geophys. Res.*, Vol.  
993 100, pp. 571–583.
- 994 Moresi, L., N., Solomatov, V., S., 1995. Numerical investigation of 2D convection with  
995 extremely large viscosity variations. *Phys. Fluids* (1995), Vol. 7, pp. 2154–2162.
- 996 Neil, E., A., Houseman, G., A., 1999. Rayleigh-Taylor instability of the upper mantle and its  
997 role in intraplate orogeny. *Geophys. J. Int.* (1999), Vol. 138, pp. 89-107.
- 998 Platt, J., P., 1993. Exhumation of high-pressure rocks: a review of concepts and processes.  
999 *Terra Nova*, Vol. 5, Issue 2, pp. 119-133. doi: [https://doi.org/10.1111/j.1365-  
1000 3121.1993.tb00237.x](https://doi.org/10.1111/j.1365-3121.1993.tb00237.x).
- 1001 Pozrikidis, C., 2001. *Fluid Dynamics: Theory, Computation and Numerical Simulation*. Pp.  
1002 675, Kluwer Acad., Boston, Mass.
- 1003 Raimbourg, H., Jolivet, L., Leroy, Y., 2007. Consequences of progressive eclogitization on  
1004 crustal exhumation, a mechanical study. *Geophys. J. Int.* (2007), Vol. 168, pp. 379–401. doi:  
1005 10.1111/j.1365-246X.2006.03130.x.
- 1006 Ring, U., Glondy, J., Will, T., Thomson, S., 2010. The Hellenic Subduction System: High-  
1007 Pressure Metamorphism, Exhumation, Normal Faulting, and Large-Scale Extension. *Annual*  
1008 *Review of Earth and Planetary Sciences* (2010), Vol. 38, pp. 45-76. doi:  
1009 <https://doi.org/10.1146/annurev.earth.050708.170910>.
- 1010 Royden, L., H., Papanikolaou, D., J., 2011. Slab segmentation and late Cenozoic disruption  
1011 of the Hellenic arc. *Geochem. Geophys. Geosys.* (G3), Vol. 12, No. 3, Q03010. doi:  
1012 10.1029/2010GC003280.
- 1013 Rubatto, D., Hermann, J., 2001. Exhumation as fast as subduction?. *Geology*, Vol. 29, pp. 3–  
1014 6. doi: 10.1130/0091 7613(2001)029<0003:EAFAS>2.0.CO;2, 2001.
- 1015 Schmelting, H., 1987. On the relation between initial conditions and late stages of Rayleigh-  
1016 Taylor instabilities. *Tectonophysics*, Vol. 133, pp. 65–80.

- 1017 Stöckhert, B., 2002. Stress and deformation in subduction zones: insight from the record of  
1018 exhumed metamorphic rocks. Geological Society, London, Special Publications, Vol. 200, pp.  
1019 255-274. doi: <https://doi.org/10.1144/GSL.SP.2001.200.01.15>.  
1020
- 1021 Thomson, S., N., Stöckhert, B., Brix, M., R., 1998. Miocene high-pressure metamorphic  
1022 rocks of Crete, Greece: rapid exhumation by buoyant escape. Geological Society, London,  
1023 Special Publications, Vol. 154 (1), pp. 87-107. doi: 10.1144/gsl.sp.1999.154.01.04.
- 1024 Trompert, R., A., Hansen, U., 1996. The application of a finite-volume multigrid method to  
1025 three-dimensional flow problems in a highly viscous fluid with variable viscosity. Geophys.  
1026 Astrophys. Fluid Dyn. (1996), Vol. 83, pp. 261–291.
- 1027 Turcotte, D., L., Schubert, G., 1982. Geodynamics. Cambridge University Press, Second  
1028 Edition. doi: <https://doi.org/10.1017/CBO9780511807442>.
- 1029 Vlaar, N., J., Wortel, M., J., R., 1976. Lithospheric aging, instability and subduction.  
1030 Tectonophysics, Vol. 32, pp. 331–351.
- 1031 Warren, C., J., Beaumont, C., Jamieson, R., A., 2008a. Formation and exhumation of ultra-  
1032 high pressure rocks during continental collision: role of detachment in the subduction  
1033 channel. *Geochem. Geophys. Geosys.* (G3), Vol. 9, No. 4. doi: 10.1029/2007GC001839.
- 1034 Warren, C., J., Beaumont, C., Jamieson, R., A., 2008b. Deep subduction and rapid  
1035 exhumation: Role of crustal strength and strain weakening in continental subduction and  
1036 ultrahigh-pressure rock exhumation. *Tectonics* (2008), Vol. 27, TC6002. doi:  
1037 10.1029/2008TC002292.
- 1038 Warren, C., J., 2013. Exhumation of (ultra-)high-pressure terranes: concepts and  
1039 mechanisms. *Solid Earth*, Vol. 4, pp. 75-92, 2013. doi: 10.5194/se-4-75-2013.
- 1040 Wassmann, S., Stöckhert, B., 2013. Rheology of the plate interface - Dissolution precipitation  
1041 creep in high pressure metamorphic rocks. *Tectonophysics* (2013), Vol. 608, pp. 1-29. doi:  
1042 <http://dx.doi.org/10.1016/j.tecto.2013.09.030>.
- 1043 Weinberg, R., B., Schmeling, H., 1992. Polydiapirs: multiwavelength gravity structures.  
1044 *Journal of Structural Geology*, Vol. 14, pp. 425–36.
- 1045 Woidt, W., D., 1978. Finite-element calculations applied to salt dome analysis.  
1046 *Tectonophysics*, Vol. 50 (2–3), pp. 369–86.
- 1047 Zhong, S., J., Yuen, D., A., Moresi, L., N., 2007. Numerical methods in mantle convection. In  
1048 Bercovici, D. (ed.) *Treatise in Geophysics*, Vol. 7, (editor-in-chief: Gerard Schubert), Elsevier,  
1049 pp. 227–52.
- 1050
- 1051
- 1052