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9 **On the assessment of sinking particle fluxes from in situ particle size**
10 **distributions**

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31

32 **Abstract**

33 The biological carbon pump plays a crucial role in the global carbon cycle, particularly through
34 sinking particles carrying carbon to deep waters. The Underwater Vision Profilers (UVP) is widely
35 used for studying particle properties. UVP-based particulate organic carbon (POC) flux is typically
36 derived from particle size distributions (PSDs) assuming size dependent sinking rates and carbon
37 content. This approach, the “classic UVP method”, calibrates PSD-based flux against sediment
38 trap flux data that are not necessarily co-located in space or time. We put forth a “modified UVP
39 method” that combines a large data set of UVP measurements calibrated against POC flux from
40 co-located and simultaneously collected sediment traps and thorium-234 measurements. Data were
41 collected in the North Pacific (50°N, 145°W, August 2018) and the North Atlantic (49°N, 16.5°W,
42 May 2021) as part of EXPORTS (EXport Processes in the Ocean from RemoTe Sensing), covering
43 a wide range of environmental conditions. We find that our modified UVP methods explain 80%
44 of the variance in POC flux when applied across sites, where flux values vary over orders of
45 magnitude. However, the method fails to account for smaller flux variations within a single site or
46 across depths. Reasons include undersampling rare large particles, mismatch in time and spatial
47 scales of UVPs calibrated against fluxes in traps and ²³⁴Th, and difficulties in interpreting particle
48 stock and flux changes within non-steady state conditions. To use UVP as a high-resolution POC
49 flux tool, it is recommended not to rely on a few profiles for calibration.

50 **Keywords:** biological carbon pump, sinking carbon export fluxes, particle imaging, sediment
51 traps, thorium-234, ocean carbon cycle.

52

53 1 Introduction

54 The oceanic biological carbon pump (BCP) (Eppley & Peterson, 1979) drives carbon storage in
55 the ocean's interior through the downward flux of biogenic carbon, produced by the net primary
56 production of phytoplankton. There are multiple paths for the BCP (Boyd et al., 2019) but when
57 considering transport to the deep ocean (>1000m) this process occurs primarily through the
58 gravitational sinking of Particulate Organic Carbon (POC) - phytoplankton, zooplankton, detritus,
59 and fecal pellets - from the euphotic zone to deep waters (Boyd et al., 2019; Turner, 2002; Turner,
60 2015). While the BCP significantly impacts atmospheric CO₂ over geological timescales (De La
61 Rocha & Passow, 2007), its current role in the global carbon cycle, particularly in response to
62 human-induced changes, remains uncertain. Annual carbon export estimates by the BCP vary from
63 5 to >12 Gt C yr⁻¹ (Boyd & Trull, 2007; Henson et al., 2011).

64 Traditional oceanographic techniques like sediment traps (e.g., Buesseler et al., 2007) or
65 radioactive disequilibrium methods such as thorium-234 (²³⁴Th half-life, t_{1/2} = 24.1 d; e.g.,
66 Buesseler et al., 1992), are resource-intensive and provide limited spatial, vertical, and temporal
67 coverage (Buesseler et al., 2007). The emergence of imaging techniques in recent decades has
68 transformed oceanography (Lombard & Kjørboe, 2010), offering a range of tools to study
69 properties and dynamics of particles across a range of sizes, from micrometers to centimeters, and
70 time scales of seconds depending upon imaging protocols. These techniques, deployable on CTD
71 rosettes from research vessels, autonomous floats, gliders, or moorings, enable observations on
72 oceanographically relevant temporal and spatial scales, while at the same time reducing
73 deployment costs (Giering et al., 2020). Among these, the Underwater Vision Profiler or UVP
74 (Picheral et al., 2010) has become one of the most widely used imaging techniques because of its
75 versatility (Kiko et al., 2022; Picheral et al., 2022).

76 In situ observations of particle size distributions (PSD) from imaging platforms historically served
77 three main purposes: 1) gaining knowledge of plankton, as PSD of the living fraction offers
78 insights into the structural properties of marine food webs and various ecological processes
79 (Sheldon et al., 1972); 2) inferring sinking velocities of individual particle size classes when
80 combined with particle collection traps equipped with polyacrylamide gels (McDonnell &
81 Buesseler, 2010); and 3) estimating sinking particles fluxes when PSD is combined with
82 conventional sediment trap measurements to derive particle properties that correlate PSD with

83 independent flux data. This approach was pioneered by Guidi et al. (2008) and applied by others
 84 (e.g., Iversen et al. 2010; Kiko et al. 2017). In this study, we revisit and expand this third
 85 application of PSD but, in this case, using a local calibration of flux against both traps and ²³⁴Th
 86 approaches. Each of these methods has their strengths and weaknesses including the spatial and
 87 temporal scales each represents. Here, we test the assumption that an instantaneous PSD imaged
 88 by UVP, essentially a stock measurement, can be related to POC flux, a rate of carbon removal on
 89 gravitationally sinking particles.

90 We refer to the method first proposed by Guidi et al. (2008) as the “classic UVP method” where
 91 the UVP retrieved PSD is used to estimate mass fluxes constrained by sediment trap observations
 92 of the sinking POC flux. The classic method assumes that the total mass flux (e.g., POC) can be
 93 theoretically estimated from PSDs by integrating the flux spectrum over all particle diameters (D),
 94 or:

$$95 \quad POC_{flux} = \int_{D_{min}}^{D_{max}} N(D) \rho_{POC}(D) w_s(D) \quad \text{Eq. (1),}$$

96 where $N(D)$ is the number concentration of particles as a function of size in differential form (#
 97 $L^{-1} \text{ mm}^{-1}$), ρ_{POC} is the POC content of an individual particle ($\text{mg C particle}^{-1}$), w_s is its sinking
 98 rate (m s^{-1}), and D_{min} and D_{max} are the smallest and largest particle sizes resolved, respectively
 99 (mm). However, we do not a priori know particle sinking rates and POC content, and even if we
 100 did, they are not likely to be uniform within a given size class (Laurenceau-Cornec et al., 2015).

101 Hence, Eq. (1) is often reformulated by combining the POC content and sinking rate terms into a
 102 power function of particle diameter (Alldredge, 1998; Alldredge & Gotschalk, 1988; Lombard &
 103 Kiørboe, 2010), resulting in:

$$104 \quad POC_{flux} = \int_{D_{min}}^{D_{max}} N(D) A D^B dD \quad \text{Eq. (2),}$$

105 Guidi et al. (2008) calculated POC flux using PSDs retrieved from UVPs, with A and B
 106 coefficients optimized using available sinking particle flux observations. The wide application of
 107 this approach assumes that the power law constants are valid globally.

108 Coefficient A represents the product of sinking speed and POC content, while coefficient B, is
 109 linked to the fractal dimension, which describes particle shape complexity. Under fractal scaling,
 110 where that POC content and sinking speed spectra are mathematically described by power-law

111 relationships with particle size, the value of B represents the slope of the product of POC and
112 sinking speed and is related to the fractal dimension (Guidi et al., 2008; Logan & Wilkinson, 1990).
113 Changes in B reflect changes in F and the physical characteristics of particles, such as porosity or
114 how they aggregate together as they sink through the water column. A lower fractal dimension
115 (smaller B coefficient) suggests a more rugged particle, with more intricate structures and higher
116 porosity, whereas a larger B indicates a smoother, more compact, nearly spherical shape. However,
117 this relationship is idealized, and in practice, the fit value of B is influenced by factors such as the
118 size range, instrument resolution, and specific image processing methodologies used in Bisson et
119 al. (2022).

120 In developing the classic UVP method, Guidi et al. (2008) compared UVP observations to a set of
121 moored and drifting sediment trap fluxes collected across different studies. However, the trap and
122 UVP observations were not consistently co-located, nor were they deployed in the same season or
123 at the same depth, as only a few UVP profiles were taken in conjunction with sediment trap
124 deployments. Trap data were collected from a few sites in the North Atlantic Ocean, Mediterranean
125 Sea and South Pacific Ocean (see Guidi et al. (2008), Table 3), while the UVP data were from a
126 much broader set of sites and studies (see Guidi et al. (2008), Figure 1).

127 Since the introduction of this classic UVP method, the A and B coefficients from Guidi et al.
128 (2008) have been used for the prediction of POC flux in many studies without much questioning
129 (e.g., (Forest et al., 2013; Guidi et al., 2009, 2015, 2016; Ramondenc et al., 2016). Only Iversen et
130 al. (2010) and Fender et al. (2019) estimated A and B coefficients following the classic UVP
131 method using co-located deep sediment trap fluxes off the Mauritanian coast and the California
132 Current Ecosystem, respectively. More recently, Clements et al. (2023) estimated A and B using
133 a machine learning reconstruction of global ocean PSDs from UVPs measurements tuned against
134 a global compilation of in situ sediment trap and thorium-derived particle flux observations.

135 In this study, we use a comprehensive and unique dataset of PSD observations from UVPs, along
136 with POC flux estimates from sediment traps and thorium-234 (^{234}Th) that were co-located in
137 space and time, to evaluate the performance and validity of the classic UVP method. We refer to
138 this as the “modified UVP method” henceforth in this paper, a distinguishing feature being the
139 application of a local calibration of A and B for a given set of UVP data. These UVP and co-
140 located flux data were collected together as part of the NASA-funded EXport Processes in the

141 Ocean from RemoTe Sensing (EXPORTS) project at two biogeochemically contrasting sites in
142 terms of POC flux and biogeochemical conditions (Siegel et al., 2024, 2021). This modified
143 method allows us to better quantify POC flux differences between widely varying flux settings in
144 the NP and NA, but fails to account for regional flux variations within a single site or as it changes
145 with depth.

146 **2 Materials and Methods**

147 2.1 Settings

148 The goal of the EXPORTS field campaign is to develop a predictive understanding of the export,
149 fate, and C cycle impacts of global ocean net primary production and to assess their impacts in
150 contemporary and future climates (Siegel et al., 2016). Two field campaigns were carried out in
151 two vastly different ocean ecosystems encompassing a wide range of environmental conditions: 1)
152 the North Pacific (NP) at Ocean Station Papa (OSP, 50°N, 145°W) in August-September 2018,
153 and 2) the North Atlantic (NA) in the vicinity of the Porcupine Abyssal Plains Sustained
154 Observatory (PAP, 49°N, 16.5°W) in May 2021. In both field deployments, operations were
155 conducted in three consecutive sampling cycles or “epochs” (E1, E2, and E3 from hereafter)
156 designed to constrain the pathways for organic carbon transformation and export. The length of
157 each epoch was approximately one week with the goal of completing a sequence of observations
158 that could be repeated three times during a given cruise. Conceptually, the aim is to follow how
159 surface properties might be observed to propagate to depth as part of the BCP, i.e., a particle
160 formed at the surface on day 1, would take approximately one week to reach 500 m if settling at
161 roughly 70 m/d (Siegel et al., 2021).

162 2.1.1 North Pacific (OSP, Aug-Sep 2018)

163 OSP (50°N, 145°W) can be characterized as an iron-limited, high-nutrient, low-chlorophyll
164 (HNLC) regions of the world ocean, which leads to limited phytoplankton production and surface
165 chlorophyll (Chl-a) concentrations. An overview of the NP sampling plan, including context
166 information on physical and bio-optical properties, nutrients, and phytoplankton biomass, is
167 presented in Siegel et al. (2021). Briefly, during our study period the oceanographic setting was
168 typical of late-summer conditions at OSP with low biomass, a highly recycled food web, and low
169 sinking POC export fluxes driven by largely zooplankton processes (Stamieszkin et al., 2021;
170 Steinberg et al., 2023). Weak horizontal currents and spatial gradients in biogeochemical fields

171 and low level of temporal variability characterized the three sampling epochs (E): E1 (August 14-
172 23), E2 (August 24-31), and E3 (September 1-9). In terms of POC export, the site was
173 characterized by a modest sinking carbon fluxes with an export efficiency at the base of the
174 euphotic zone of 13% + 5% (ratio of POC flux to net primary production) and 39% flux attenuation
175 in the subsequent 100 m (Buesseler et al., 2020).

176 2.1.2 North Atlantic (PAP, May 2021)

177 The PAP site (49°N, 16.5°W) is a highly advective environment dominated by eddies which can
178 cause upper ocean biogeochemical properties to evolve on time and space scales comparable to
179 those driven by biological processes. The NA operations were conducted in a coherent,
180 anticyclonic, physically retentive eddy that minimized horizontal exchanges, so that changes in
181 biological or chemical properties were dominated by local rather than advective processes. An
182 overview of the NA is available by Johnson et al. (2024). In short, the expedition sampled a dual-
183 phase spring bloom with a transition to high sinking particle export. Conditions differed across the
184 three sampled epochs. The first epoch (E1; May 5-7) was marked by a very large diatom biomass
185 in the surface ocean but no presence of aggregates in the upper mesopelagic portion of the water
186 column (Siegel et al., 2024). The second epoch (E2; May 11-20) followed a major storm (>40 kts)
187 and was marked by a large change in surface properties and the appearance of aggregates. Export
188 flux metrics remained low by most measures until the third epoch (E3; May 21-29) which large
189 increases in both sinking particle fluxes and aggregate abundances (see e.g., Brzezinski et al.,
190 2024; Clevenger et al., 2024; Meyer et al., 2024; Romanelli et al., 2024).

191 2.2 Platforms and instrumentation

192 The deployments during the NP expedition included coordinated sampling with two research
193 vessels: a Process Ship (*R/V Roger Revelle*, cruise *RR1813*) focused on sampling biogeochemical
194 stocks and fluxes, ecological abundances and rates, and optical properties following a Lagrangian
195 float, and a Survey Ship (*R/V Sally Ride*, cruise *SR1812*) that characterized the horizontal and
196 vertical distribution of properties including phytoplankton, particulate and dissolved organic
197 carbon, ²³⁴Th deficit, and other constituents surrounding the Process Ship and cross-calibrated
198 sensors onboard the ships and the autonomous platforms (Siegel et al., 2021).

199 In the NA, a third ship was added to the Survey Ship (*RSS Discovery*, cruise *DY131*) and the
200 Process Ship (*R/V James Cook*, cruise *JC214*), the *R/V Sarmiento de Gamboa* (cruise *SdG2105*),

201 that focused on plankton and metazoan imaging. Similar to the NP sensor calibration exercise, a
202 detailed intercalibration was performed on all the NA sensor observations (Siegel et al., 2023). A
203 summary of the instruments and measurements most relevant to this study is presented below.

204 2.2.1 Sediment trap fluxes

205 Two types of sediment traps with identical collection tubes (collection area = 0.0226 m²) were
206 used - neutrally buoyant sediment traps (NBSTs) and a surface-tethered sediment trap array (STT).
207 Sinking particles were collected over approximately 2 to 5-day deployments in the upper 500 m
208 of the ocean during the three epochs in each EXPORTS field campaign (Estapa et al., 2021;
209 Johnson et al., 2024). Results from these traps are discussed here without distinction of trap type.
210 Formalin-poisoned brine traps were gravity filtered through 335-micron screens, swimmers were
211 manually removed from the screens under magnification, and the remaining material was
212 recombined with < 335 μm material for bulk elemental analysis for POC. During the NP, sample
213 composition was used to perform an additional correction for POC from small swimmers that
214 could not be removed following screening (Estapa et al., 2021). POC fluxes were modest during
215 E1 and E2, and increased moderately during E3. During NA, particle export flux seen in the traps
216 increased rapidly two weeks after the experiment started, from similarly low fluxes in E1 and E2
217 to high fluxes in E3, suggesting strong temporal variability in flux (Siegel et al., 2024).

218 2.2.2 ²³⁴Th fluxes

219 ²³⁴Th disequilibrium relative to Uranium-238 (²³⁸U) in depths ranging from 0 to 500 m was used
220 during EXPORTS to estimate POC fluxes following the methodological approach proposed by
221 Buesseler et al. (1992). A description of the sampling method can be found in Clevenger et al.
222 (2021). A full description of ²³⁴Th measurements and derived fluxes from the NP experiment can
223 be found in Buesseler et al. (2020), while the NA experiment is described in Clevenger et al., 2024.
224 During the NP, the ²³⁴Th data show relatively homogenous and consistent ²³⁴Th disequilibria, with
225 higher ²³⁴Th fluxes observed at depths of 50 to 100 m, remaining relatively constant or decreasing
226 at greater depths in the water column (see Figure 1 in Buesseler et al. (2020)). ²³⁴Th-derived POC
227 fluxes showed a similar trend, but with fluxes decreasing more sharply with increasing depth
228 during all epochs. During NA, the magnitude of ²³⁴Th disequilibria varied both temporally and
229 spatially through the cruise. Existing deficits indicated that particle fluxes had already been high
230 prior to the start of the cruise (Clevenger et al., 2024). Since ²³⁴Th measurements are integrated

231 over time rather than an instantaneous representation of a system, a non-steady state model was
232 needed to derive POC fluxes that increased during the experiment.

233 2.2.3 UVPs

234 The 5th generation of underwater cameras (UVP5; Picheral et al. 2010) was used to collect all PSD
235 data used here. A total of 387 high-resolution full depth particle size spectra (0.10 – 26 mm
236 equivalent spherical diameter, ESD) were sampled with five different UVPs mounted on the
237 bottom of each ship CTD-Rosette. The smallest consistently sampled bin across all UVPs is
238 centered at 0.14 mm while the largest size sampled is a function of particle concentrations. Details
239 of these units and data provided by each are summarized in the [Supporting Information \(SI\)](#) and
240 in the EXPORTS UVP intercalibration report by Siegel et al. (2023). A full description of the
241 UVP5 can be found in Picheral et al. (2010) and a thorough review of the instrument, data
242 collection, processing and quality control can be found in Kiko et al. (2022).

243 Briefly, the UVP5 acquires and processes images in real time using two lighting units, each
244 consisting of a row of red-light emitting diodes (LEDs), that illuminate and images a well-defined
245 volume of water of ~1L sample volume with 100 μm flashes of collimated light at a maximum
246 rate of approximately 6 Hz (Picheral et al., 2010). Particles (living and non-living) are identified
247 from each image and then sized as the number of contingent pixels. The projected area for each
248 particle is then converted to equivalent spherical diameter (ESD). Particles are binned into size
249 bins and the particle size spectra, $N(D)$, is determined as the number of particles in each size bin
250 divided by the width of the bin. Data are acquired on the downcast of each CTD profile (drop rate
251 $\sim 1 \text{ m s}^{-1}$) and images are binned into 5-m depth intervals (with ~ 100 UVP scans per 5-m bin).
252 Consequently, each depth bin represents a sampling volume of about 100 L. Further, PSD data
253 from each platform were inter-calibrated to ensure comparability of data from different UVP units,
254 as detailed in the EXPORTS UVP intercalibration report by Siegel et al. (2023) and summarized
255 in the SI.

256 2.3 UVP-based POC flux estimates

257 Both the classic and our modified UVP-based POC flux estimates apply Eq. (2) to the PSD
258 observations by optimizing the A and B coefficients using independent POC flux estimates from
259 traps alone (classic UVP method) or both traps and ^{234}Th (our modified UVP method). Several
260 considerations arise concerning how data are used to fit coefficients A and B, including the UVP

261 size range selected for fitting, the spatial area considered, vertical resolution of each technique,
262 and temporal variability between epochs.

263 *Particle size range considerations* - UVPs have evolved over the years, with each model iteration
264 expanding the size spectra measured (Picheral et al., 2022). In Guidi et al. (2008), data from older
265 UVP models (UVPs 2, 3, and 4), resulted in an overlapping size range from 0.25 to 1.5 mm across
266 8 size bins (see Tables 2 and 3 in Guidi et al. (2008) for details). However, the UVP5 model covers
267 a wider size range from 128 μm to 26 mm across 23 size bins (Picheral et al., 2022). To reproduce
268 the classic UVP method used by Guidi et al. (2008) with a more recent instrument model two
269 options were considered: the classic 8 size bins (250 μm to 1.5 mm) from Guidi et al. (2008) and
270 an extended version using 23 size bins (128 μm to 26 mm). In both cases, we include all living
271 and non-living particles as was done in Guidi et al. (2008).

272 *Creating global and regional models* - Guidi et al. (2008) established a global relationship between
273 UVP-observed PSD and particle flux using UVP data from multiple locations and sediment trap
274 fluxes that were not co-located in space or time (see Table 1 and Figure S1). To more closely
275 replicate fitting at these larger global scales, we first reproduce a similar fit by calibrating our UVP
276 data against POC fluxes observations from the combined EXPORTS data set. Note that unlike
277 Guidi et al. (2008), the UVP data and POC observations are from the same time and study area.
278 We refer to this as a “global” approach. We also used a “regional” comparison by fitting A and B
279 considering UVP and flux data separately for the NP and the NA regions. Both approaches can be
280 evaluated using either the 8 or newer 23 size bin criteria.

281 *Sediment trap considerations* - For the consideration of the spatial scales for the NP traps, we only
282 matched UVP profiles and trap fluxes from the same source funnel, as described in Estapa et al.
283 (2021). In essence, choosing only UVP profiles falling within the particle source region defined
284 by local currents and assuming a particle sinking rate of 100 m d^{-1} (e.g., Siegel et al. 2008). In the
285 NA, the spatial criterion was related to the dimensions of the core of a retentive eddy that was
286 sampled during the experiment in a Lagrangian fashion. In this case, only UVP profiles and trap
287 flux within the eddy center (EC) waters (Johnson et al., 2024), defined as $< 15 \text{ km}$ from the
288 autonomous asset-defined eddy center, were used.

289 *Thorium-234 flux considerations* - As both UVP images and ^{234}Th samples were collected on the
290 same CTD cast, we could combine all of the paired UVP and ^{234}Th results on a cast-by-cast basis

291 for a global estimate of A and B using the NP and the NA results. Alternatively, these could be
292 considered separately, to derive regional coefficients for the NP and the NA. Note there was little
293 spatial variability in the NP, and all casts were included. However, for the NA, we restricted the
294 ^{234}Th and UVP matchups to the eddy center, consistent with the approach used for the trap data.

295 *Vertical Resolution Considerations* - The three data sets used all have different vertical resolutions
296 in their sampling. For comparisons to a specific trap depth, we used the average of the 5-m
297 resolution UVP data at the measurement depth and the three 5-m depth intervals above it. This
298 resulted in a 20-m depth average, corresponding to imaging roughly 400 L of seawater. See **Table**
299 **1** for the resulting number of UVP profiles and sediment traps used in these global and regional
300 comparisons.

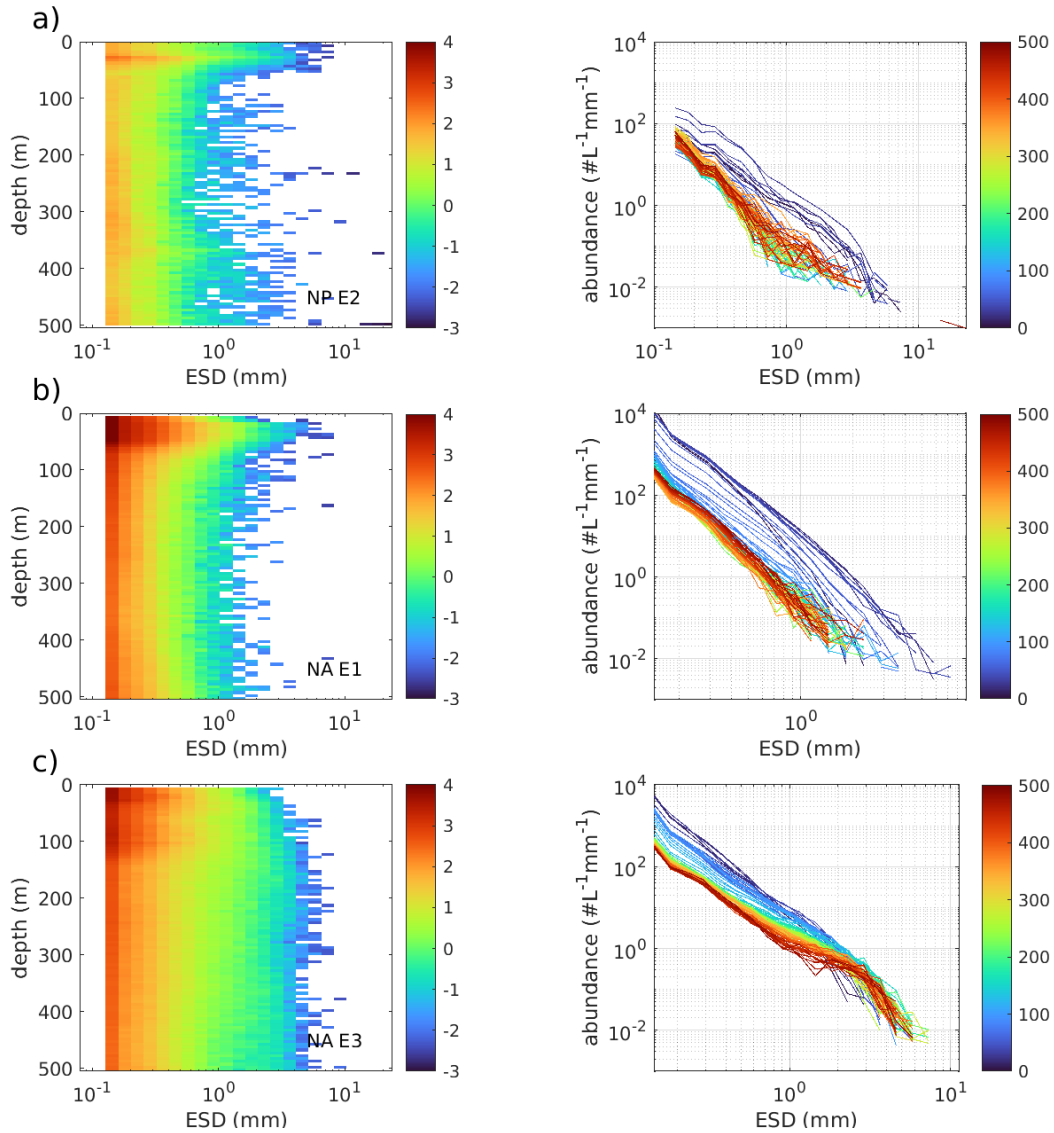
301 In the case of ^{234}Th , we again matched the UVP data by averaging the 5-m intervals containing the
302 ^{234}Th measurement and the three intervals above, similar to the traps. No ^{234}Th data in the mixed
303 layer were used. See **Table 1** for the resulting number of UVP and ^{234}Th profiles used in these
304 global and regional comparisons.

305 *Temporal variability between epochs* - In the NP experiment, variability between epochs was
306 considered minimal (Siegel et al., 2021). In the NA experiment, we clearly sampled the
307 development of the evolving system (Johnson et al., 2024; Romanelli et al. 2024). At both sites,
308 results from UVP, traps and ^{234}Th were grouped and averaged by epoch for comparison to the
309 cruise average. The number of casts broken down by these epochs is found in Table S1.

310 3 Results

311 3.1 PSD observations

312 UVP-PSDs observations of in situ particles ($\# \text{L}^{-1} \text{mm}^{-1}$) as a function of ESD (mm) binned in 5-
313 m depth intervals are shown in **Figure 1**. In the NP, UVP-PSD observations exhibit little
314 discernible temporal variations (**Figure 1a** and **S2a**). Conversely, in the NA, the PSDs changed
315 throughout the cruise (**Figure 1b, 1c, and S2b**).



316

317 **Figure 1.** Observed particle size distribution (PSD) between epochs of in situ particles plotted against aggregate
 318 equivalent spherical diameter (ESD, mm) for the a) NP and b-c) NA. For each panel, the figure on the left shows ESD
 319 versus depth, the color bar indicates particle abundance for each size (in # L⁻¹ mm⁻¹, logarithmic scale). Red indicates
 320 a higher number of particles than blue. The figure on the right shows ESD versus abundance, the color bar indicates
 321 depth (m). Red indicates deeper waters than blue. For the NP, no changes in PSD were observed over time and a
 322 random profile in E2 was chosen. For the NA, PSD evolved between epochs and representative profiles early (E1)
 323 and late (E3) in the cruise are shown.

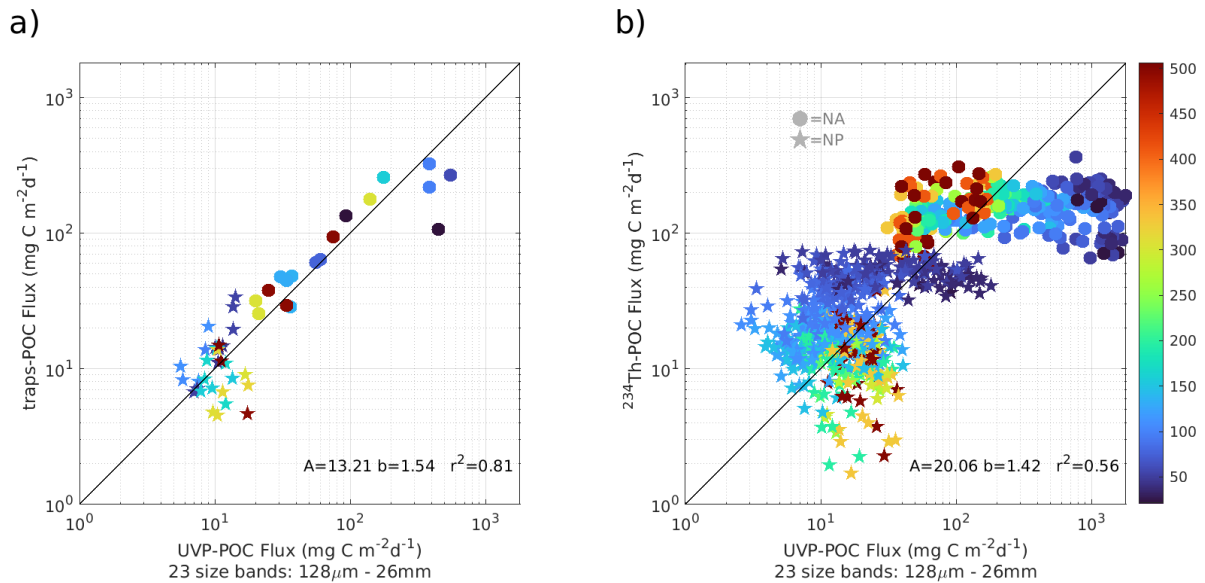
324 As expected, PSDs decline in particle abundance with increasing size across all casts in both
 325 experiments, with smaller particles being more abundant than large ones by several orders of
 326 magnitude at all depths (see left panel in the figures). However, distinct vertical patterns emerge
 327 for each experiment based on the particle size.

328 In the NP, abundances of small particles remain relatively consistent across depths and time,
329 whereas larger particles decreased in abundance with depth and were constant at lower abundances
330 below 100 m (see right panel in Figure 1a). In the NA, PSDs changed both over time and depth
331 (see Figure 1b and 1c). During the initial days of the experiment (E1), small particles were highly
332 concentrated in the surface layer (0-50 m) and decreased with depth, while deeper layers showed
333 a shift toward fewer, larger particles. In the later stages of the experiment (E3), there was a
334 substantial increase in large particles, especially at greater depth.

335 3.2 A and B coefficients using EXPORTS results

336 Table 1 displays present and prior estimates of A and B coefficients. Uncertainty bounds were
337 calculated as the half-width of the interval of the obtained values and the confidence intervals (CIs)
338 were provided as the range between their minimum and maximum values. The coefficient of
339 determination for the linear fit (r^2) is also included, calculated as $1-(SSE/SS)$, where SSE is the
340 sum of squared errors (differences between predicted and actual values), and SS is the total sum
341 of square values (difference between actual values and their mean).

342 When applying the global approach using traps, i.e. calibrating all data from both sites for a single
343 A and B value, no significant differences were observed within the uncertainty bounds in the
344 retrieved A and B coefficients using either the 8-size bins or 23-size bins (Table 1). However, the
345 ratio of uncertainty to estimate is > 1 for the 8 bins (SI, Figure S3). Hence, we only consider the
346 full 23 size bins in our comparisons below ($A = 13.1 \pm 5.5 \text{ mg C d}^{-1} \text{ mm}^{1-b}$, $B = 1.54 \pm 0.44$, $r^2 =$
347 0.81 , Figure 2a). The same global approach applied to ^{234}Th POC flux determinations is a poorer
348 fit with a larger value of A though a similar B ($A = 20.1 \pm 2.9 \text{ mg C d}^{-1} \text{ mm}^{1-b}$, $B = 1.42 \pm 0.15$,
349 $r^2 = 0.56$; Figure 2b). When applying the regional approach using traps, no significant differences
350 were observed within the uncertainty bounds of the A and B coefficients compared to the global
351 approach in the NA. However, there was no predictive power in the NP regional models. In
352 contrast, when using ^{234}Th , there was no regional predictive capability in either of the deployments
353 (Table 1).

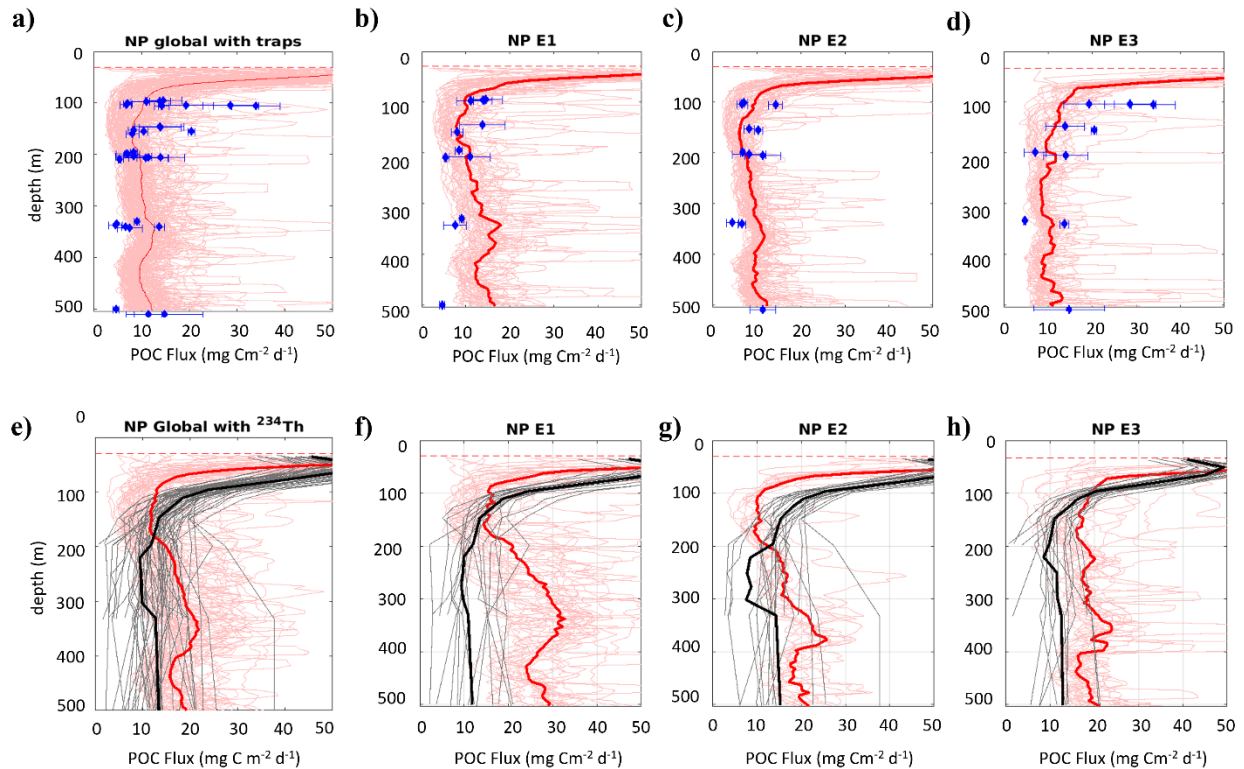


354
 355 **Figure 2.** Results of the matchups between UVP-based POC fluxes (in $\text{mg C m}^{-2} \text{d}^{-1}$, x-axis) using a) traps and b)
 356 ^{234}Th (y-axis) using the global approach. Indicated in each panel along with the r^2 of the linear fit. 95% CIs for the A
 357 and B coefficients are provided in [Table 1](#). Stars represent the NP results, and circles the NA results, with color
 358 representing depth (in m). The black line indicates the 1:1 line.

359 3.3 Vertical profiles of POC flux

360 Using the A and B values from EXPORTS, we can derive vertical profiles of POC flux from any
 361 given UVP profile in that study ([Table 1](#)). [Figures 3](#) and [4](#) show the variability in UVP derived
 362 flux profiles for both the global (panels a and e) and regional (panels b-d and f-h) fit parameters
 363 for the NP and NA, respectively. Concurrent trap and ^{234}Th flux profile observations are overlain
 364 in the upper and lower rows of [Figures 3](#) and [4](#), respectively. In each case, the flux profiles using
 365 UVP data show both extreme vertical variability and large cast to cast variability (gray lines).

366 For the NP, UVP fluxes range from 5 to 50 $\text{mg m}^2 \text{d}^{-1}$ (global calibration against all traps; [Figure](#)
 367 [3a](#)), with an average around 50 $\text{mg m}^2 \text{d}^{-1}$ below the mixed layer depth (MLD) decreasing to
 368 vertically uniform values of $\sim 10 \text{ mg m}^2 \text{d}^{-1}$ below 100 m. Note however individual UVP casts
 369 even in the upper 50-100 m can have lower or higher fluxes predicted than these averages. The fit
 370 of the UVP average with the observational trap results in the NP is good within the spread of the
 371 trap data ([Figure 3a](#)). Similar patterns are found for the UVP fluxes calibrated with the ^{234}Th
 372 observations, although the fluxes are > 1.5 times higher, reflecting the higher values of A used
 373 ([Figure 3d](#)).



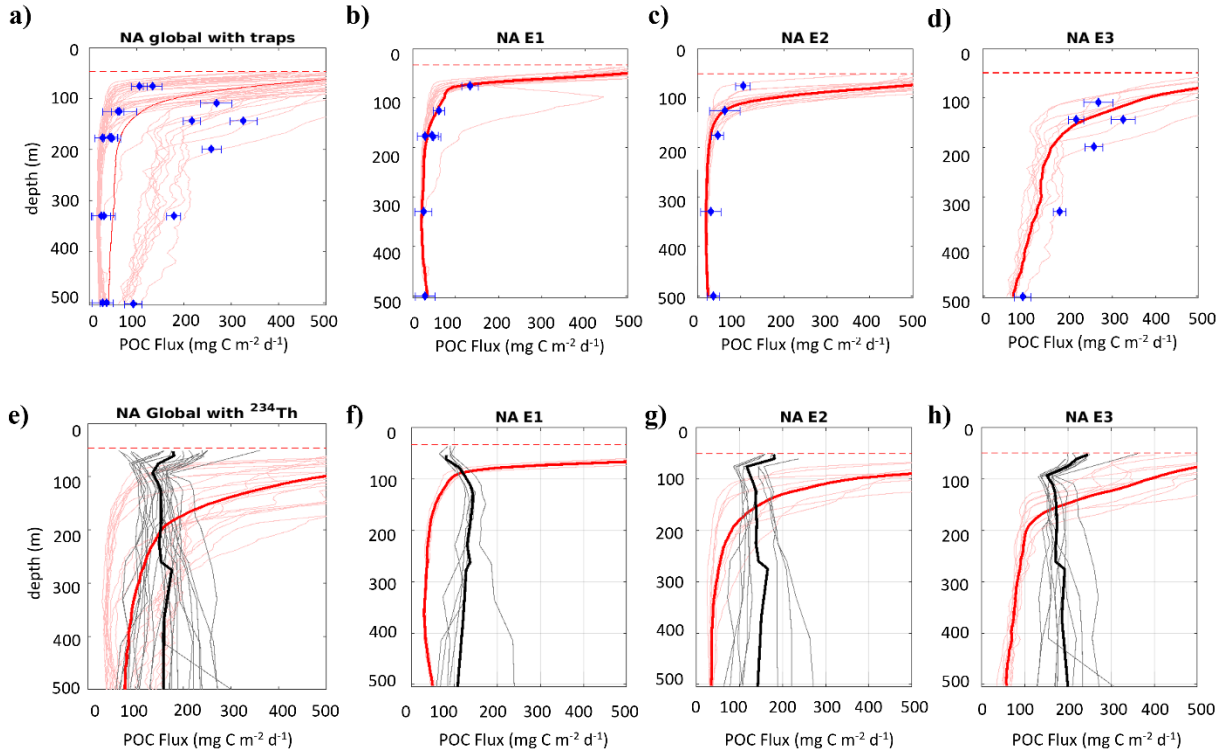
374
 375 **Figure 3.** Vertical profiles of POC fluxes for (a) the global and (b-d) regional approaches in the NP using traps grouped
 376 by epochs: E1, E2, and E3 (from left to right) where the blue diamonds denote sediment trap POC fluxes with
 377 uncertainties from Estapa et al., (2021). Panels (e-h) show the corresponding UVP fluxes optimized using ^{234}Th data
 378 while the gray lines show the ^{234}Th -derived POC flux and the fluxes for each cast, and the black lines show the mean
 379 flux. The horizontal dashed line indicates the mean mixed layer depth over the course of the experiment. Pink lines
 380 show the UVP-based POC flux for each UVP cast, calculated using the A and B coefficients derived from each
 381 approach. Only UVP casts paired with trap or ^{234}Th profiles that fall within an epoch are shown. In panels (a) and (e)
 382 the red line represents the cruise mean UVP-based POC flux of all casts used for the fitting, which includes casts
 383 within a 100 m^{-1} source funnel region (see main text for details). In the rest of the panels, the red line indicates the
 384 epoch mean.

385 In principle, a regional calibration of A and B and concurrent UVP and flux observations might be
 386 expected to improve the matchups; however, that is not the case. Using regionally derived A and
 387 B for the NP with UVP derived flux profiles separated by epoch (Figures 3b-d), the UVP POC
 388 flux exceeds the traps around 350 m in E1 and E2 (Figures 3b and 3c), and the increase in trap flux
 389 in E3 at 100 m is not captured in the E3 UVP profiles (Figure 3d). Small scale changes in flux
 390 versus depth or time are thus not improved by using a regional fit to the NP trap data. Considering
 391 the NP regional model derived from ^{234}Th , individual ^{234}Th flux profiles vary cast to cast, but not

392 over as wide a range as the UVP derived fluxes (Figures 3f-h). A decrease in flux versus depth is
393 seen in both UVP and ^{234}Th results, but with the UVP showing higher values on average at deeper
394 depths (Figure 3e). A subsurface peak at a depth of ~350 m is observed in the ^{234}Th optimized
395 UVP fluxes for all three epochs and is especially strong in the NP E1 (Figure 3f). Also, while ^{234}Th
396 fluxes generally increase to some maximum value in the subsurface (here 50 m) and then decrease,
397 UVP derived fluxes always are highest in the shallowest depths.

398 When we examine these same depth trends in the NA, the trap fluxes reach a maximum of > 500
399 $\text{mg m}^2 \text{d}^{-1}$ for UVP derived fluxes using our global fit for the traps (Figure 4a). The magnitude
400 and range of the UVP fluxes are considerably higher than in the NP. As noted previously, the
401 progression of the bloom resulted in large local changes in particle properties, and this is illustrated
402 clearly when the NA results are displayed by epoch. Using a NA regional calibration of the UVP
403 data with traps, there is a better fit between UVP fluxes and traps if broken down by epoch, for
404 which the flux values increase dramatically in E3, especially at depths > 100 m (Figures 4e versus
405 Figures 4c and 4d).

406 The global fit using ^{234}Th data in the NA results in higher UVP fluxes on average (Figure 4e) and
407 a predicted increase in flux using matched UVP profiles and regional fit in E3 (Figure 4, panels f-
408 h). However, the shape of the two flux curves differs, with a higher flux at the surface and steeper
409 flux attenuation evident in the UVP data but not in ^{234}Th results.



410

411 **Figure 4.** Corresponding figure to Figure 3 for the NA deployment. Casts used for the fitting include only those within
 412 the eddy center.

413 **4 Discussion**

414 We set out to assess the quality of POC flux estimates derived from the UVP imagery using
 415 concurrent sediment traps and ^{234}Th flux observations. The challenges are substantial given the
 416 multitude of issues presented. First, UVP images and the PSD derived from them, are at best
 417 measures of sinking and non-sinking particles and living stocks, within certain size ranges and
 418 over scales set by the number of images (Hz), imaging volume (liters) and duration of a CTD cast
 419 (hours; meters) (Picheral et al., 2022). Traps measure the gravitational sinking flux of particles,
 420 largely non-living, and originating from a large particle source area (Siegel et al., 2008) determined
 421 by sinking rates, currents and deployment durations (several days; 10^3 km^2). Biases in trap fluxes
 422 occur due to hydrodynamics, swimmers, preservation and other issues (Buesseler et al., 2007). The
 423 ^{234}Th flux method tracks small scale variations in flux (km) but averaged over days to weeks. The
 424 ^{234}Th -derived POC flux also depends on the measured ^{234}Th disequilibrium and generally ignore
 425 physical processes, but importantly here, the flux can be sensitive to whether a system is at steady
 426 state or not, over the course of its half-life (24.1 d) (Ceballos-Romero et al., 2018; Clevenger et

427 al., 2024; Savoye et al., 2006). A conversion from ^{234}Th flux to POC flux requires consideration
428 of observed variations in its ratio to POC on particles that vary with depth and location (Buesseler
429 et al., 2006). So, mismatches between UVP, trap and ^{234}Th estimates of POC flux are expected to
430 be due to a combination of methodological consideration and their respective spatial and temporal
431 averaging.

432 The EXPORTS project provides a unique opportunity to evaluate the utility of UVP-derived flux
433 estimates using an extensive set of co-located and simultaneous sediment traps and ^{234}Th
434 observations. A key finding is that UVPs can be effectively trained to translate observed PSD
435 changes into POC fluxes when the range in POC fluxes spans several orders of magnitude, and
436 when stocks and fluxes are roughly near steady state. At its most basic, and when in steady state,
437 if there are more particles observed with the UVP, there is higher flux, and vice versa. In terms of
438 flux attenuation versus depth, note that both POC stocks and POC fluxes decrease with depth (Lam
439 & Bishop, 2007; Martin et al., 1987). However, when POC flux variations are smaller than an
440 order of magnitude, i.e., at local scales, or are in non-steady state conditions, UVP approaches to
441 determine flux have large uncertainties, and particle stocks are not necessarily a reflection on in-
442 situ export. Here, we delve into the findings that led us to these insights.

443 4.1 Global performance of the modified UVP method

444 The globally optimized A and B coefficients found using traps and ^{234}Th fluxes range from 13.2
445 to 20.1 for A, and from 0.73 to 1.54 for B, respectively (8 and 23 size bins, Table 1). The other
446 global estimates from Clements et al. (2023); Guidi et al. (2008); and Kriest (2002) range from
447 12.5 to 18.0 for A, and 2.2 to 3.8 for B (Table 1). All of these have considerable uncertainty, in
448 particular our estimate using only 8 size bins.

449 Importantly, even small discrepancies in the A and B coefficients lead to significant differences in
450 POC fluxes predicted from UVP data (Figure 5). These differences are primarily driven by the B
451 coefficient, which is more sensitive to the particle size range than the A coefficient. Using the
452 classic UVP-based POC flux coefficients for A and B, it would consistently underestimate trap-
453 measured fluxes by about an order of magnitude, particularly at the lower flux ranges (Figure 5).
454 Using regionally derived A's and B's would make these POC flux differences even larger, such as
455 the regional estimates of A and B from the NP and the NA in EXPORTS, or using those from
456 Iversen et al. (2010), Fender et al. (2019), or Forest et al. (2013), which focused on the Mauritanian

457 coast, the California Current Ecosystem, and the Southeast Beaufort Sea (Arctic Ocean),
458 respectively.

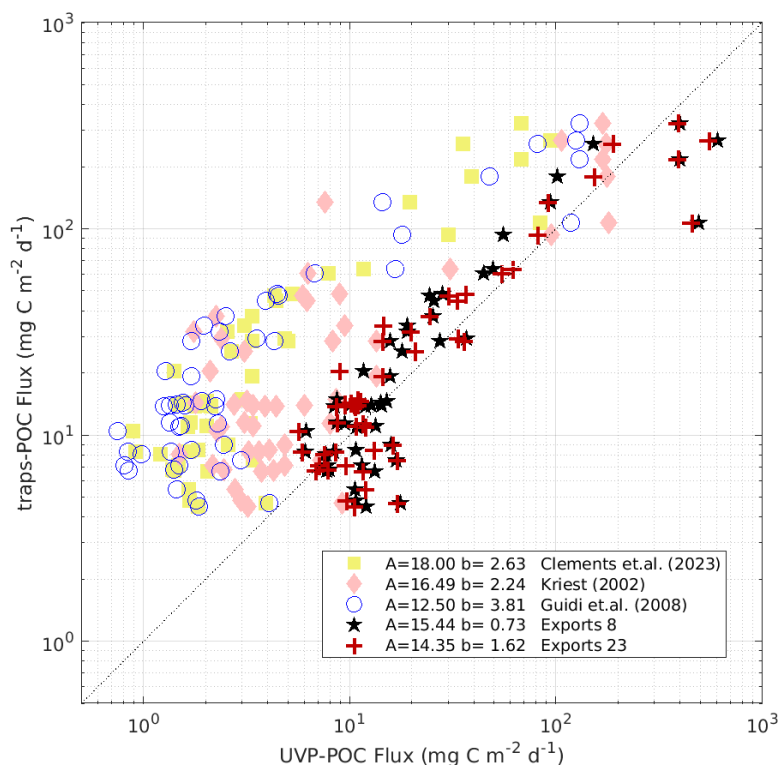
459 We attribute differences in the A and B coefficients primarily to differences in particle
460 characteristics found in the NP and during the spring bloom in the NA, compared to the prior
461 studies (SI, Figure S1). These differences are likely due to the different depth ranges considered
462 in each study, which is especially important since a single A and B values are applied at all depths.
463 Guidi et al. (2008) and Iversen et al. (2010) focused on deeper sinking aggregates (100–1000 m,
464 and 1200-1900 m depth respectively) than our study (0 -500 m depth). According to Guidi et al.
465 (2008), an increase in fractal dimension with depth implies a decrease in particle porosity, which
466 could account for their higher B values (3.8 and 4.3, respectively), as the physical compression of
467 aggregates due to fluid dynamics can compact them decrease porosity (Logan & Kilps, 1995). Our
468 EXPORTS results, which show lower B values (all values <1.6), suggest the presence of more
469 porous particles, likely aggregates, in the upper 500 m of the NP and NA.

470 We also consider whether changes in A and B reflect genuine variations in particle characteristics
471 rather than an artifact of the size range used by the imaging systems in different studies. Our 8 bin
472 size range was chosen to match Guidi et al. (2008), and yet the B in Guidi et al. (2008) is 5 times
473 higher (3.8 versus 0.7; Table 1). Therefore, we do not think size range is the primary reason here
474 for differences in A or B values, rather it is more likely to be attributed to differences in particle
475 properties.

476 Another aspect of UVP size ranges is that, at the smaller end, there can be an undercount of the
477 more abundant, but smallest sizes as one reaches the resolution limit of the UVP for identifiable
478 particles as these particles may not be detected or can be difficult to distinguish from background
479 noise (Stemmann & Boss, 2012). In this data set, there is no obvious tailing downward of PDS
480 slopes for the smaller sizes classes that would indicate this type of bias (Jackson et al., 1997). At
481 the larger size ranges, > 0.8 mm to 28 mm, the total particle counts drop off as there are very few
482 large particles in the ocean, and these typically decrease with depth. One reason we combine UVP
483 images from 5 m depth intervals is to increase imaging volume, and hence our chances of counting
484 more than one particle in these larger size bins. At an abundance of $\log 10^{-2}$ (Figure 1, y-axis), the
485 total number of particles counted would be less than 10 in our 100 L imaging volume for the size
486 bins from 0.8 mm and larger (Figure 1a). This difficulty in quantifying rare large particles, leads

487 to a high uncertainty that translates directly into large POC flux variability with depth, and between
488 casts.

489 Another consideration is that if more data are from shallow waters, we would image a higher
490 fraction of living versus non-living particles. Shallow water particles would be larger and more
491 POC rich and include more non-sinking materials. Including living organisms detected by the
492 UVP introduces uncertainty into flux calculations based on unsorted PSD data (Bisson et al.,
493 2022). Supporting this, Kiko et al. (2020) found that excluding living organisms and artifacts larger
494 than 1 mm ESD from the UVP5 dataset decreased the variability of PSD-derived POC flux
495 estimates. Clearly, while one may desire to use UVPs to estimate fluxes for shallow depths, one
496 needs to be careful of our calibration of UVPs. In EXPORTS, we only consider UVP flux estimates
497 below the MLD and match the calibration data as closely as possible to the same depths as the flux
498 observations.



499
500 **Figure 5.** Results of the matchups between UVP-based POC fluxes (in $\text{mg m}^{-2} \text{d}^{-1}$, x-axis) and sediment trap fluxes
501 (y-axis) using five different sets of A and B coefficients from previous studies and our own, applied to the entire
502 EXPORTS data set. The coefficients used are from: 1) Clements et al. (2023) (yellow squares), 2) Kriest (2002)

503 (orange diamond), 3) Guidi et al. (2008) (blue circles), and the coefficients obtained in our study for the global
504 approach using traps with 4) the same 8 size bins as used by Guidi et al. (2008) (black stars), and 5) 23 size bins (red
505 cruxes; See Table 1).

506 4.2 Using traps versus ^{234}Th for calibration if UVP fluxes

507 The ^{234}Th approach has the advantage of being relatively easier to measure and at higher
508 resolutions than traps; however, the cast by cast comparisons with UVP had low predictive power
509 (Figure 2a and 2b; Table 1). Thorium-234 also has the advantage of providing flux data at
510 shallower depths than sediment traps, the shallowest of which were deployed at 100 m in the NP
511 experiment and 75 m in the NA experiment. The differences in the depth range covered by each
512 technique likely contributes to the variations in A and B coefficients, resulting in a larger
513 coefficient A for ^{234}Th than traps (20.1 versus 13.2 in global approach), which could reflect the
514 higher number of shallow water particles considered when using UVP and ^{234}Th matchups. Since
515 the gradients in POC flux are largest below the MLD and level off around 75 – 100 m in
516 EXPORTS, the calibrations of A and B against UVP data necessitates tracers like ^{234}Th if we want
517 to estimate flux at these shallower depths. Most importantly, a significant difference is that the
518 temporal and spatial scales of integration differ for POC flux derived from traps versus ^{234}Th .
519 Specifically, traps in EXPORTS have a shorter average time scale of flux over 3-4 days, but they
520 involve larger spatial averaging over 10's of kilometers for the particle source regions (Siegel et
521 al., 2008).

522 4.3 Regional performance of the modified UVP method

523 Our assumption in creating the regional UVP flux models was that there would be regional
524 differences that influence how POC content and sinking speed relate to particle size, requiring
525 distinct A and B coefficients for PSD-flux conversions. However, optimizing the A and B
526 parameters to regional data sets led to higher uncertainty in UVP-based POC fluxes (Table 1). This
527 means that while the “modified UVP method” can still be used for regional applications, high
528 uncertainties are to be expected when flux variations used to calibrate the UVP are small, and
529 results should be interpreted carefully. For instance, we attribute the lack of predictive power in
530 the regional calibration in the NP - regardless of whether traps or ^{234}Th are used - to the minimal
531 spatial and temporal variations in POC flux during the experiment. In the NA we see a higher
532 range and variation in POC flux, which improved the flux calibrations using UVP and traps (Table
533 1). In sum, we consider UVPs well-suited for understanding large scale POC flux differences

534 across multiple orders of magnitude, as exemplified by studies like e.g., Clements et al. (2023).
535 However, UVPs are less effective at refining small-scale or short-term export rates, especially if
536 strong flux variations do not occur locally or temporally.

537 4.4 POC fluxes versus depth in the NP and NA

538 4.4.1 Insights from the NP profiles

539 The UVP fluxes, using either the global or regional models that were optimized using either the
540 trap or ^{234}Th fluxes or, performed poorly for the NP (Table 1). Temporal changes in export during
541 the NP deployment were minimal, and the system exhibited largely steady state conditions in water
542 mass properties, production rates, chlorophyll, particle stocks (McNair et al., 2023; Siegel et al.,
543 2021) and UVP PSD profiles (Figures 1 and S2). Both trap and ^{234}Th results show no noteworthy
544 changes in POC flux between E1 and E2, with only small increases detected by traps in E3
545 (Buesseler et al., 2020; Estapa et al., 2021). Figure 3b, 3c, 3d). Furthermore, the trap and ^{234}Th
546 results were similar (Estapa et al., 2021); Figure 3), suggesting that over the study area, conditions
547 were relatively constant with respect to POC flux, at least in the upper 500 m.

548 We do know, however, that sinking particle types vary with depth – from larger organic aggregates
549 and fecal pellets at shallower depths to smaller aggregates and less labile particles as depth
550 increases - and across the three deployment periods (Durkin et al., 2021). These depth-related
551 changes in particle types suggest that depth-constant coefficients A and B may not fully account
552 for differences in shallow, mid, and deeper mesopelagic particles observed in the UVP-based POC
553 flux profiles (see Figures 3b, 3c, 3d). Different size-to-carbon and size-to-sinking-speed
554 relationships (i.e., distinct A and B coefficients) at varying depths, or depth-dependent A and B
555 relationships may be required to improve upon such UVP based assessments. Numerous particle
556 features, such as gray level, circularity, skewness, or kurtosis among others, measured by UVPs,
557 are currently underutilized but could be valuable for capturing depth-related particle changes
558 (Karthäuser et al., *in review*). We recommend further investigation into these parameters to
559 improve the accuracy of UVP-based POC flux profiles in future studies.

560 Other studies from EXPORTS in the NP indicate that POC available for export out of the MLD
561 primarily exists in the form of small particles driven by mesozooplankton repackaging of biomass
562 or detritus (Durkin et al., 2021; McNair et al., 2023), as larger marine snow aggregates (particles
563 > 0.5 mm) were not detected in Marine Snow Catcher deployments between 20-500 m (Romanelli

564 et al., 2024). Durkin et al. (2021) used sediment traps equipped with polyacrylamide gel layers to
565 assess the contribution of small particles ($< 100 \mu\text{m}$) to total POC flux. They found that during the
566 NP, small particles contributed 7% to 36% (mean contribution of $17\% \pm 9\%$) of total POC flux at
567 depths between 95 and 510 m. However, Estapa et al. (2021) concluded that traps undercollected
568 small particles due to hydrodynamic effects (Buesseler et al., 2007). Thus, the contribution of
569 particles smaller than the UVP lower detection limit may have been higher than 17%. Ultimately,
570 flux calculated from UVPs is highly sensitive to small particles (Bisson et al., 2022). This could
571 explain why discrepancies between UVP-based POC fluxes and observations are even larger for
572 ^{234}Th , which should reflect the export of particles across all sizes (Buesseler et al., 2007).

573 Limitations in imaging at the larger limit of the UVP size range may also affect the interpretation
574 of fluxes. Larger particles disproportionately influence the overall UVP-based POC flux due to
575 their substantial volume and rapid sinking speeds. As particle volume scales by its diameter cubed,
576 even a small number of large particles can significantly contribute to the total flux. For instance,
577 in the NP, UVP-based POC flux estimates from particles larger than $\sim 0.1 \text{ mm}$ derive from fewer
578 than ten particles imaged per 100 L (5-m bins) (Figure 1a, right panel). If we consider the average
579 POC flux value at depths of 150 or 300 m, typically around $10 - 20 \text{ mg m}^2 \text{ d}^{-1}$, just a few additional
580 large particles can have a substantial impact on POC flux. For example, for particles in the 5 - 8
581 mm size bins, just ten particles would contribute between 10 to $> 40\%$ of the total POC flux when
582 using the UVP method calibrated with traps. Because of differences in A and B, if the calibration
583 is based on ^{234}Th , the contribution to flux of rare particles increases, accounting for 17% to 68%
584 of the total POC flux. This issue is less relevant in the NA, where POC fluxes at 150 or 300 m
585 range from 50 to $200 \text{ mg m}^2 \text{ d}^{-1}$. In the NA, particles $> 1 \text{ mm}$ would have contributed between 1%
586 to 13% of the total flux, while only being counted in fewer than 10 images.

587 These results emphasize the importance of rare large particles in the NP are supported by
588 observations related to salp fecal pellets, whose presence in small numbers in the NP has been
589 shown to account for a significant portion of POC flux (Durkin et al., 2021; Steinberg et al., 2023).
590 Missing such large particles in UVP imaging could lead to significant underestimation of POC
591 flux, leading to misleading conclusions about carbon export dynamics. The sporadic presence of
592 these large particles would also explain the fluctuating nature of UVP-derived POC flux profiles
593 versus depth and the high cast-to-cast variability. Since large particles are infrequent but have a

594 disproportionately large impact on flux, their presence would lead to spikes in the flux data. Estapa
595 et al. (2021) also extrapolated the PSD from gel traps to larger sizes (> 1 mm) and concluded that
596 traps undercollected rare larger particles, including salp fecal pellets, which would make the trap
597 POC flux results too low. The ^{234}Th distribution should reflect these rare salp pellets. However,
598 the fact that ^{234}Th and trap fluxes largely agree in the NP suggests this is not a big issue. The largest
599 difference in the POC flux profiles is from the overall larger A and B's from ^{234}Th , which we can
600 attribute to sampling shallower in the water column, leading to higher POC fluxes at all depths.

601 In the NP during E1 (Figure 3f) there is a large peak between 300 - 400m in the UVP derived
602 fluxes, which may indicate some active transport layer. Salps were observed migrating from
603 between 300 to 750 m during the day to the upper 100 m at night, (Steinberg et al., 2023). Notably,
604 during E1, salp fecal pellet export was highest, comprising 48.1%, 88%, 85%, and 57% of the
605 modeled total POC flux at 100, 208, 336, and 500 m, respectively (Steinberg et al., 2023),
606 coinciding with the largest discrepancies between UVPs and ^{234}Th and, to a lower extent, traps.
607 Migrating zooplankton actively transporting carbon could bypass sediment traps and leave behind
608 particles that are non-sinking (no increase in ^{234}Th flux at the same depths is seen) but could be
609 reflected in the UVP data. We suggest future studies analyze UVP casts separated by day and night
610 to further explore daily migrations in order to test this hypothesis.

611 4.4.2 Insights from the NA profiles

612 Discrepancies of UVP-based POC flux results relative to traps and ^{234}Th measurements in the NA
613 experiment cannot be solely explained by depth variations in A and B. Instead, we attribute the
614 variability in the different flux indicators to temporal factors (SI, Figure S4).

615 The NA cruise took place during a dual-phase spring bloom in a highly dynamic system (Johnson
616 et al., 2024; Romanelli et al., 2024; Siegel et al., 2024). Briefly, during E1 (May 5-7, storm 1 on
617 May 7-11), export fluxes were low across most measures, including sediment traps and ^{234}Th
618 (Figure 4, upper and bottom panels respectively), and as estimated by the Marine Snow Catcher,
619 which collect no aggregates (Romanelli et al., 2024). After storm 2 (May 15) in E2 (May 11-20),
620 aggregates were collected in the Marine Snow Catchers, but export fluxes remained low until E3
621 (May 21-29), when large particles substantially increased, particularly at depth (see Figure 2a in
622 (Siegel et al. 2024). Shifts in the dominant plankton community, from large diatoms to a more
623 diverse phytoplankton community were observed later in E3 (San Soucie et al. *in review*).

624 Throughout the NA experiment, particles were consistently fluffy and porous, becoming
625 progressively fluffier and more porous over time, with large particles exhibiting extreme porosity
626 that increased as the cruise progressed (Siegel et al. 2024; Soucie et al. *in review*). Analysis of the
627 morphology of individual large (≥ 1 mm) particles sampled in the images collected by the UVP
628 showed that fluffy aggregates dominated the dataset, accounting for 88% of the particles, while
629 dense aggregates (10%), fecal pellets (1%), and zooplankton pellets (1%) made up the rest (Drago,
630 2023).

631 Under these conditions, the average of the UVP profiles effectively captured the bloom stages,
632 mirroring the trap results, both of which increase in E3 (Figures 4a-d). POC fluxes from traps
633 varied by an order of magnitude between E1 and E3, which allowed us to effectively train UVPs
634 to translate observed PSD changes into POC fluxes. We attribute the good agreement between
635 techniques to the main particles contributing to flux falling within the UVP's detectable size range
636 (i.e., no significant contribution of particles < 100 μm) and a low presence of rare large particles.
637 This, together with consistently fluffy and porous particle morphologies with depth and time, led
638 to accurate calibration of A and B coefficients.

639 However, the performance of the ^{234}Th -derived POC fluxes was poor for both the global and
640 regional tuning (Figures 4e-h). Differences in between traps and ^{234}Th in the NA have previously
641 been reported at the EXPORTS sampling site and were linked to the persistence of ^{234}Th
642 disequilibrium in the water column prior to sampling (Ceballos-Romero et al., 2016, 2018).
643 Clevenger et al. (2024) suggest that the ^{234}Th profiles during the EXPORTS cruise reflect both an
644 earlier export in addition to the evolving conditions. We conclude that the prior export likely
645 contribute to the differences observed in the UVP flux calibration between traps and ^{234}Th in the
646 NA. We see this both in shallow and in deeper depths. The ^{234}Th deficit at depth indicates that
647 export had reached deeper waters from the previous time-period. This explains the much higher
648 ^{234}Th -derived POC fluxes compared to traps and UVPs at depth. The export from the first bloom
649 also would influence the calibration of A and B coefficients using ^{234}Th , at the same time leading
650 to the low predictive power for the regional NA ^{234}Th method. Fluxes based upon ^{234}Th do increase
651 in E3 but are not reflected in the UVP particle fields that are responding more quickly. Therefore,
652 in highly dynamic and non-steady state environments like the NA, changing ^{234}Th distributions to
653 the evolving particle fields measured by UVPs.

654 The evolving particle characteristics also would have impacted the accuracy of UVP fluxes given
655 that A and B coefficients were held constant across epochs. However, determining A and B
656 coefficients for each epoch was not possible given the higher uncertainties that would result with
657 even smaller data sets.

658 **5 Conclusions**

6596 Our modified UVP method successfully reflects large-scale POC flux changes globally (i.e., NP
660 versus NA) and temporally (i.e., early versus late stages of the bloom), but struggles with vertical
661 and temporal changes at regional scales. We conclude that the limited predictive power of the
662 regional approach in the NP with both traps and ^{234}Th is due to small changes in the magnitude of
663 the POC flux observations used to calibrate UVP data, coupled with the disproportionate
664 contribution of rare large particles to the flux. In the NA, the calibration using sediment traps had
665 a higher predictive power than using ^{234}Th , better capturing the larger scales and changing
666 conditions represented by both traps and UVP averages during each epoch. UVP flux calibrations
667 using ^{234}Th were ineffective in the NA due to the influence of an earlier export event on ^{234}Th
668 deficits, in addition to the cast-by-cast mismatch between changing particle stocks measured by
669 UVPs, and flux determined with non-steady state ^{234}Th models. Further studies calibrating UVP
670 against ^{234}Th are crucial to draw more definitive conclusions about the performance of the
671 approach with ^{234}Th in non-steady-state conditions, as ^{234}Th can be applied at shallower depths
672 and with higher vertical resolution. Exploring using UVPs to quantify changes in particle stocks
673 over time versus the assumptions here that higher stocks reflect higher fluxes, would be
674 informative.

675 The classic and modified UVP methods for estimating POC fluxes from UVP-based PSDs are
676 valuable but face several challenges, including limitations in imaging size and resolution, and other
677 methodological concerns. Technological advancements are needed to sample larger volumes and
678 detect a broader range of particle sizes. Coefficients A and B, derived through minimization
679 procedures, can introduce errors influenced by the particle size range, the depth range over which
680 calibrations are possible, detection limits for imaging rare large particles, and the assumption of
681 constant A and B versus location, depth, and time. Measurement uncertainties can exceed 50% for
682 PSD-based POC flux estimates (Bisson et al., 2022). Variations in A and B used in different studies
683 can result in order of magnitude differences in predicted POC flux (Figure 5).

684 Since the variability with depth and between casts in UVP data is large, UVP users should not rely
685 on a few UVP profiles to calibrate their own A and B coefficients, regardless of whether co-located
686 trap flux measurements are available. Ideally, users should fit A and B coefficients to a large
687 number of UVP profiles and co-located flux estimates for more accurate results. However, when
688 this is not feasible, for global applications in the upper 500 m, we recommend using the coefficients
689 from the “global approach using traps from this study” from Table 1, as they were co-located in
690 space and time. For regional applications in the upper 500 m in the NA, we recommend using the
691 coefficients from the “regional approach using traps NA” (Table 1). In general, for regional
692 applications, more co-located data from other regions and times are needed to assess the variability
693 of the A and B coefficients and further refine them for improved regional and global applications.

694 Moving forward, we propose exploring alternative approaches that consider particle morphologies
695 for more accurate UVP-based POC flux estimates. Relying on particle size alone neglects the
696 importance of morphology, which is likely crucial for accurately assessing export variations over
697 time and depth. Furthermore, the optical properties of particles, particularly their index of
698 refraction, are linked to their carbon content, as demonstrated in phytoplankton (Stramski, 1999),
699 and likely hold true for other types of particles such as zooplankton, detritus, etc. Incorporating
700 these optical characteristics alongside morphology could enhance our understanding of particle
701 composition and its role in carbon export. Unsorted PSD data only provides fluxes derived from
702 consideration of standing stocks, which may or may not accurately reflect real POC fluxes. At a
703 minimum, particles should be categorized as living and non-living, with ideally, a more detailed
704 classification considering different particle types across the water column. We recommend further
705 studies to compare UVP-based POC flux with estimates from gel traps and include morphological
706 sorting of sinking particles, as introduced in Trudnowska et al., 2021. This combined approach
707 could enhance the accuracy of UVP-based POC flux estimates, helping to distinguish between
708 living and non-living particles, identify dominant morphological groups specific to each sampling
709 site and time, and advance our comprehension of carbon export dynamics in marine environments.

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724 **Availability Statement**

725 NASA-funded primary data products are archived at SeaWiFS Bio-optical Archive and Storage
726 System (SeaBASS). All EXPORTS data are being archived under one digital object identifier
727 (DOI: <http://dx.doi.org/10.5067/SeaBASS/EXPORTS/DATA001>) that further expands into the
728 individual data subsets. NSF-funded. To find out information about all the data collected during
729 the EXPORTS field campaigns, their data repositories and availability, please visit:
730 <https://sites.google.com/view/oceanexports/home>.

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963 intercomparison of seven methods. *Elementa: Science of the Anthropocene*, 11(1).
964 <https://doi.org/10.1525/elementa.2022.00094>

Approach	# of size bins	size range	# camera casts	# trap deployments	trap depths (m)	#234Th profiles	type of data	A	CI	b	CI	r^2
Previous studies												
Kriest (2002) used by used by Kiko et al. (2017, 2020) for the size range 140 μm to 16.8 mm	nan	1000 μm – 10 mm	nan	nan	nan	nan	nan; nan; nan	16.49 \pm nan	nan	2.24 \pm nan	nan	nan
Guidi et al. (2008): classic UVP method	8	250 μm – 1.5 mm	118	11	100, 200, 300, 400, 1000	nan	S & C	12.50 \pm 3.40	nan	3.81 \pm 0.70	nan	0.73
Iversen et al. (2010) (using own, custom built camera system)	nan	150 μm – 6.5 mm	5	5	~ 1200, 1866	nan	S	273.8 \pm nan	nan	4.27 \pm nan	nan	0.74
Forest et al. (2013)	17	80 μm - 4.2 mm	> 154	21	~ 100, ~ 200	nan	C	28.1 \pm 6.1	nan	2.0 \pm 0.39	nan	0.68
Fender et al. (2019)	19	100 μm – 1.5 mm	~ 300	69	100-150	nan	S	15.40 \pm nan	nan	1.05 \pm nan	nan	nan
Clements et al. (2023)	23*	35 μm – 5 mm	8805	250	20-160	8805	C	18.0 \pm 2.8	nan	2.63 \pm 0.06	nan	0.51
This study												
Global approach using traps- 8 bins	8	250 μm – 1.5 mm	387	55	75-500	nan	S	15.44 \pm 21.53	-6.08, 36.97	0.73 \pm 1.55	-0.82, 2.29	0.79
Global approach using traps	23	128 μm – 26 mm	387	55	75-500	nan	S	13.21 \pm 5.52	7.69, 18.74	1.54 \pm 0.44	1.10, 1.98	0.81

Regional approach using traps NP	23	128 μ m – 26 mm	169	23	100-500	nan	S	13.89 \pm 7.75	6.14, 21.64	1.61 \pm 0.69	0.92, 2.30	0.02
Regional approach using traps NA	23	128 μ m – 26 mm	218	32	75-500	nan	S	12.46 \pm 8.35	4.11, 20.81	1.48 \pm 0.61	0.87, 2.09	0.82
Global approach using ²³⁴ Th	23	128 μ m – 26 mm	387	nan	nan	83	S	20.06 \pm 2.88	17.18, 22.94	1.42 \pm 0.15	1.26, 1.57	0.56
Regional approach using ²³⁴ Th NP	23	128 μ m – 26 mm	169	nan	nan	61	S	27.27 \pm 3.42	23.85, 30.70	1.52 \pm 0.15	1.37, 1.67	0.09
Regional approach using ²³⁴ Th NA	23	128 μ m – 26 mm	218	nan	nan	22	S	6.76 \pm 4.24	2.52, 11.01	0.82 \pm 0.48	0.34, 1.30	0.02

965

966 **Table 1.** Summary of A (in mg C m d⁻¹ mm^{-1-b}) and B coefficients with uncertainties and the r^2 -value for the linear fit from previous studies and this study. We
967 provide the values and uncertainties of coefficients as reported by other authors and calculate the uncertainty for our A and b coefficients as the half-width of the
968 interval of the values obtained. For our coefficients we provide the Confidence Interval (CI) as the ranges between the minimum and maximum values obtained
969 for the coefficients. The numbers of size bins and the size range used for each approach are also reported along with the number of camera casts, number and depths
970 of trap deployments, and number of ²³⁴Th profiles when appropriate. The type of data used is specified, distinguishing between climatological (i.e., long-term data,
971 denoted as C in the table) and simultaneous POC flux observations (denoted as S). Regarding previous studies, the A coefficient by Kriest (2002) was obtained
972 from a relationship between size and nitrogen content of *273ESD*¹⁻⁶² reported by Alldredge (1998) for miscellaneous aggregates (detritus, flagellates, etc.), while
973 the b coefficient was calculated applying Stokes law assuming a sinking speed of 2.8 md⁻¹ at 0.002 cm (see Tables 1 and 2 in Kriest (2002)). Studies by Kiko et al.
974 (2017, 2020) and Bisson et al. (2022) applied the same coefficients as by Kriest (2002). To obtain the coefficients by Guidi et al. (2008), from the 1254 UVP
975 profiles available in their database only 118 data sets had both particle size distributions from the UVP and sediment trap flux measurements, see Table 3 in Guidi
976 et al. (2008) for location, position and duration of the deployments of the sediment traps used in this study. Coefficients by Iversen et al. (2010) were obtained
977 using a Particle Camera (ParCa) that illuminated a volume of 12.4 L (12 cm width) and was deployed only at night (see Table 2 in Iversen et al. (2010) for trap
978 depth and collection periods).

979 *Forest et al. (2013) did not explicitly specify the number of UVP profiles and sediment trap deployments. However, the study covered 154 locations across the
980 Mackenzie Shelf. UVP5 was mounted on the CTD rosette and deployed multiple times at each sampling station, with the number of deployments varying by
981 station. The paper also indicates that there were 21 overlaps between sediment trap sampling and UVP5 deployments during the field campaign. Therefore, the
982 minimum number of UVP and sediment trap deployments referenced in the study is 154 and 21, respectively.

983 †Clements et al. (2023) used the 23 size bins available for UVP5s, extending them to set the minimum size class to be 35 μm - below the detection limit of the
984 UVP5 - because the power law slope has been shown to extend to this size range, and the maximum size to be 5 mm, which roughly corresponds to the size at
985 which zooplankton become significant contributors to the particle biovolume detected by UVP5 in various regions. The shadowed area indicates the methods with
986 low predictive power.

987 **Supplementary Information**

988 **Contents of this file**

989 Text S1

990 Table S1

991 Figures S1 to S6

992 **Text S1.**

993 UVP: data, sources, and calibration

994 1. UVP data

995 UVPs were deployed during EXPORTS to estimate the true particle population by means of
996 measuring the PSD in the water column. Two UVP5s in EXPORTS-NP (serial numbers SN 201
997 and 207, see calibration report for more details, Siegel et al. (2023)), and three in EXPORTS-NA
998 (SN 201, 203 and 205) provided high quality measurements of the abundance, composition, and
999 size distribution of suspended particles (including live organisms) in the water column. During
1000 EXPORTS-NP, UVP5 SN207 captured data during CTD downcast during 138 of 144 casts of
1001 cruise *SR1812*; the unsuccessful casts did not have a clear cause as to why the instrument did not
1002 work as programmed. During cruise *RR1813*, UVP5 SN201 provided a total of 84 profiles of
1003 particle size distribution as part of the activities of the survey ship. During EXPORTS-NA, the
1004 UVP5s captured data during CTD downcast during 112, 69 and 13 casts of cruises DY131
1005 (SN201), JC214 (SN203), and SdG2105 (SN205) respectively.

1006 2. Data sources

1007 NASA-funded primary data products are archived at *SeaWiFS Bio-optical Archive and Storage*
1008 *System (SeaBASS)*. All EXPORTS data are being archived under one digital object identifier (DOI:
1009 <http://dx.doi.org/10.5067/SeaBASS/EXPORTS/DATA001>) that further expands into the
1010 individual data subsets.

1011 For easier access and preliminary sharing, the 5m binned particle size data are also available in
1012 *EcoPart* (<https://ecopart.obs-vlfr.fr/>) and *EcoTaxa* (<https://ecotaxa.obs-vlfr.fr/>) with the following
1013 project names and numbers (3 digits for *EcoPart* and 4 digits for *EcoTaxa*) for the NP and NA
1014 experiments respectively:

- 1015 • NP:
- 1016 ○ cruise *RR1813*: uvp5_sn201_exports_01_filtered(228) (*EcoTaxa* project 1286),
- 1017 ○ cruise *SR1812*: uvp5_sn207_2018_exports_np_sr1812_filtered(230) (*EcoTaxa*
- 1018 project 1591),
- 1019 • NA:
- 1020 ○ cruise *JC224*: uvp5_sn201_exports02_filtered(473) (*EcoTaxa* project 5652),
- 1021 ○ cruise *DY131*: uvp5_sn203_nasa_exports_north_atlantic_2021(371) (*EcoTaxa*
- 1022 project 4661), and
- 1023 ○ cruise *SdG2105*: uvp5_sn205_exports_2021_filtered(363) (*EcoTaxa* project
- 1024 4521).

1025

1026 3. UVPs intercalibration

1027 Calibration of the UVP5 pixel to millimeter conversion is done by the Laboratoire
 1028 d'Océanographie de Villefranche sur Mer (LOV). This involves quantifying the illuminated
 1029 volume and determining the appropriate conversion between imaged particle area in pixels and
 1030 mm² (Picheral et al., 2010). The latter is initially done by measuring the size of particles with a
 1031 microscope and their corresponding pixel area in the UVP image (dropping particles one at a time
 1032 in the field of view of the camera) in order to determine the raw pixel size (pixels mm⁻¹). The
 1033 instruments used during each EXPORTS cruise were sent together to the manufacturer for pre-
 1034 cruise calibration, where each instrument underwent an in situ inter-calibration against the same
 1035 reference UVP in the Bay of Villefranche, France. From this intercalibration, a power law
 1036 relationship of the form $S_m = Aa \times S_p^{Exp}$ is used to derive calibration coefficients *Aa* and *Exp* for
 1037 each instrument that minimize the log-transformed differences between the particle area in pixels
 1038 (S_p) and particle area in mm² (S_m). These calibration coefficients are reported in the calibration
 1039 documents and used to calculate the size of each particle imaged by the UVP during the EXPORTS
 1040 program (see Siegel et al. (2023) for UVP calibration report).

1041 As several UVP units were used, an intercalibration procedure was developed to allow
 1042 comparability of data from these units. The full EXPORTS UVP intercalibration report can be
 1043 found in (Siegel et al., 2023). Very briefly, the intercalibration procedure is based on a comparison
 1044 between one or several reference units and the units to be calibrated. UVP-based PSD observations

1045 are intercompared using nearby in time and space casts. During EXPORTS-NP, UVP serial
1046 number 207 on the Survey Ship consistently overestimated the particle concentrations measured
1047 by UVP serial number 201 on the Process Ship for unknown reasons, despite being intercalibrated
1048 to a reference UVP prior to the cruise. These differences were observed across most size and depth
1049 bins. However, due to the limited number of intercomparison profiles and the fact that the
1050 instruments were not sampling the exact same parcels of water (deployed from separate ships), no
1051 specific corrections or adjustments were applied for EXPORTS-NP. An intermittent lighting issue,
1052 where at least one lighting unit failed to illuminate, was also noted for both UVP units. Bad images
1053 resulting from this issue were filtered out of the original database using a procedure developed by
1054 LOV. During EXPORTS-NA, four match-up casts were identified between *DY131* and *JC214* and
1055 two between the *DY131* and *SdG2105*. While *DY131* and *SdG2105* UVPs showed very similar
1056 PSD data; *JC214* PSD data were considerably lower, particularly for the smaller size bins, and had
1057 known issues related to the discovery of a zip tie in the field of view. *DY131* PSD data were used
1058 as the “standard” due to their consistent availability throughout the entire cruise and broad
1059 consistency with *SdG2105* PSD results. Linear regression models were employed to correct the
1060 *JC214* PSD data to best match the *DY131* data for the size bins with a linear correlation (r^2) value
1061 exceeding 0.8. The data were subsequently vertically binned into 25-m bins to enhance statistics
1062 for the largest size bins without significantly altering the values of the smaller bin corrections. The
1063 resulting remapped data show consistency among the three UVP-PSD data sources, offering a data
1064 set to explore the relationships among particle distributions and export fluxes.

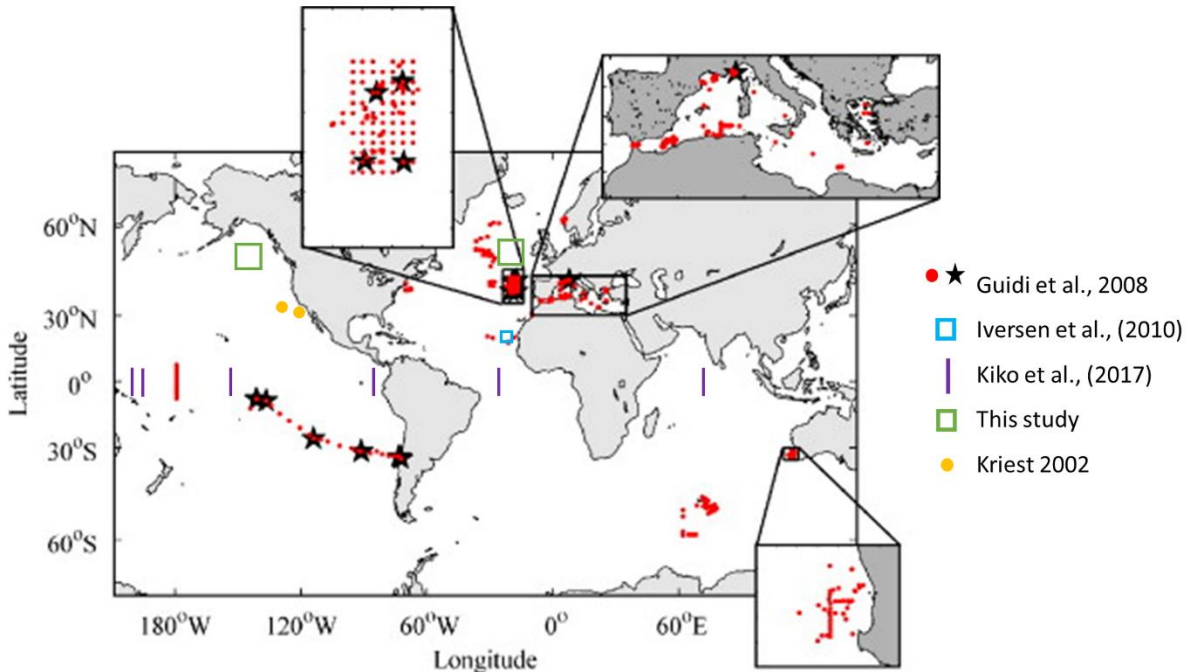
1065

epoch	NP			NA		
	traps	UVP	234Th & UVP	traps	UVP	234Th & UVP
E1	11	37	17	7	17	5
E2	11	34	20	5	18	8
E3	10	73	18	6	9	9

1066

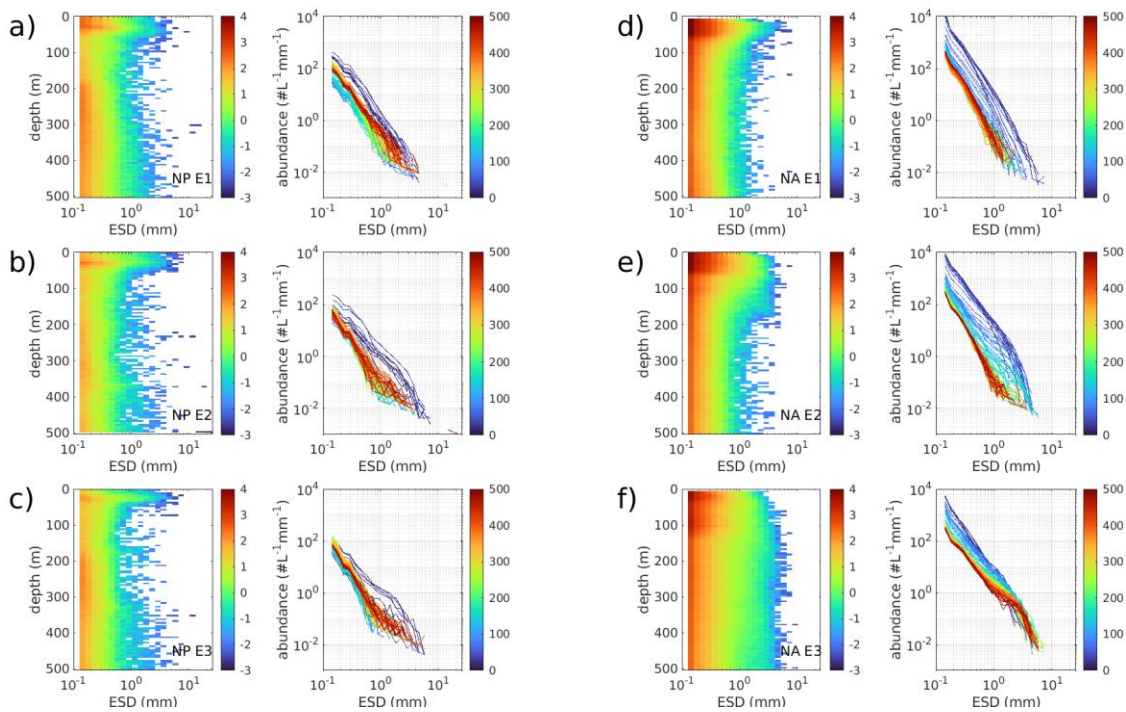
1067 **Table S1.** Summary of the number of UVP profiles, sediment traps, and ²³⁴Th profiles used in the regional matchups
 1068 by epoch.

1069



1070

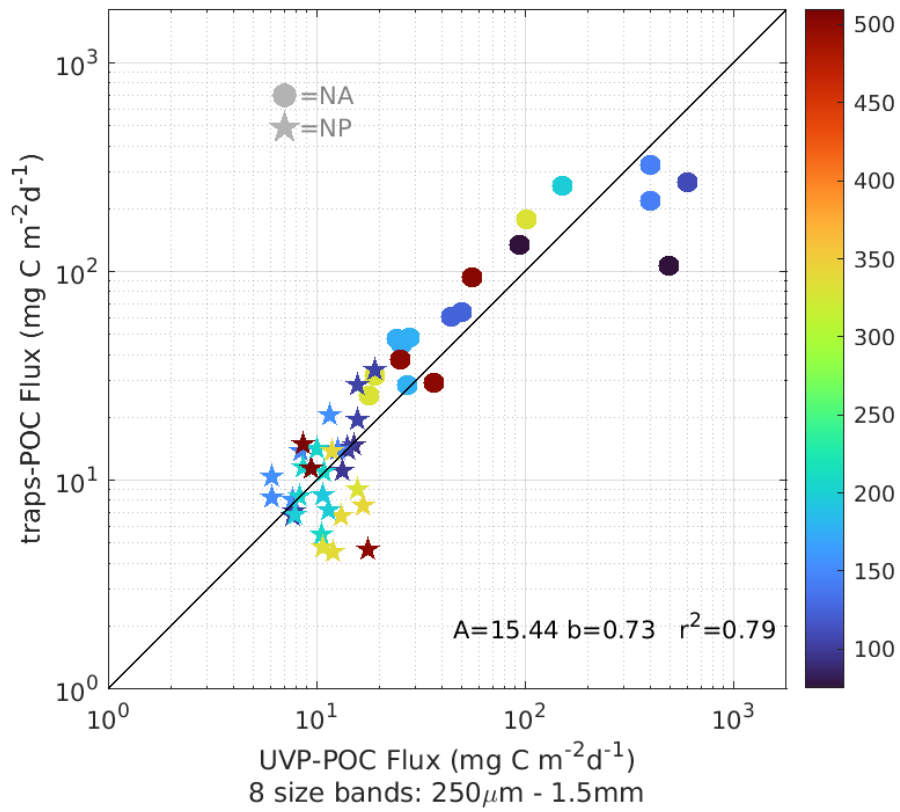
1071 **Figure S1.** Modified map based on Figure 1 from Guidi et al., (2008) to include the sampling sites of the studies
 1072 summarized in [Table 1](#) that use observations to estimate A and B coefficients. The map features: red dots and black
 1073 stars to indicate UVP and trap data respectively from Guidi et al., (2008), blue square to show study site from Iversen
 1074 et al., (2010), sampled with deep sediment traps and an underwater camera (ParCa type), purple lines to indicate
 1075 transits from Kiko et al., (2017) with UVP data, and green squares to denote sampling locations from this study, which
 1076 included UVP data, and traps and ^{234}Th measurements. Kriest, (2002) estimated A and B coefficients based on a
 1077 variety of observations about the mass and sinking speed of marine snow in relationship to its diameter from Alldredge
 1078 (1998) and Smith et al., (1998). Sampling locations of these observations are indicated with yellow and orange circles
 1079 respectively. Note that observations from Alldredge (1998) are based on undisturbed aggregates of marine snow hand-
 1080 collected at depths of 10 to 20 m in the Santa Barbara Channel (California, USA), while those from Smith et al., (1998)
 1081 are based on large detrital aggregates collected from the seafloor at an abyssal site in the NE Pacific. The observations
 1082 used by Clements et al., (2023) are very extensive, and we therefore refer to the original publications for detailed
 1083 sampling sites. Sediment traps and ^{234}Th locations used can be found in Figure 2 in Bisson et al., (2018), while the
 1084 UVP data are available in Figure 2 in Kiko et al., (2022). The profiles of PSD observations used in Bisson et al., (2022)
 1085 come from the first version of the UVP compilation by Kiko et al., (2022) accessible in (Kiko et al. (2021).



1086

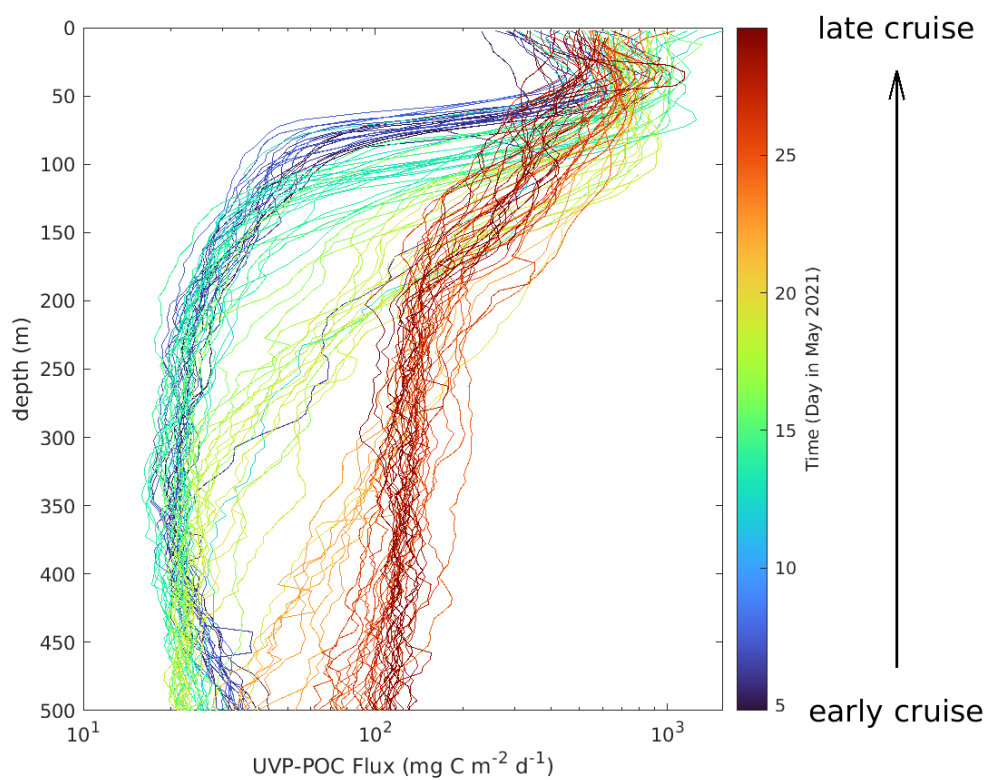
1087 **Figure S2.** Observed particle size distribution (PSD) of in situ particles plotted against aggregate equivalent spherical
 1088 diameter (ESD, mm) for the (a-c) NP and (d-f) NA experiments. For each panel, the figure on the left shows ESD
 1089 versus depth, the color bar indicates particle abundance for each size (in $\# L^{-1} mm^{-1}$, logarithmic scale). Red indicates
 1090 a higher number of particles than blue; the figure on the right shows ESD versus abundance, the color bar indicates
 1091 depth (in m). A random profile is shown for each of the epochs (E1, E2, and E3).

1092



1093

1094 **Figure S3.** Detailed figure for our reproduction of the “classic UVP method” with the EXPORTS data set (i.e., global
 1095 approach using traps with 8 bins). Color coding indicates depth.



1096
1097 **Figure S4.** Vertical profiles of POC fluxes for the global approach in the NA using trap data (Figure 4a in the main
1098 text) featuring color as time, with blue corresponding to early in the cruise and red with late in the cruise.