

1 **Disentangling impact ejecta dynamics using micro-X-ray fluorescence (μ -XRF): a case**
2 **study from the terrestrial Cretaceous-Paleogene (K-Pg) boundary**

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20 *Geosystems.*

21
22 **Key Points:**

- 23 • Micro-X-ray fluorescence offers a well-suited high-resolution mapping tool for event
24 deposits such as the Cretaceous-Paleogene boundary.
- 25 • Geochemical profiles from the Raton Basin (USA) shed new light on ejecta dynamics in
26 the first minutes to years post-Chicxulub impact.
- 27 • Prominent enrichments in both zirconium and chromium are linked to the slow
28 atmospheric settling of the Chicxulub impact dust plume.
29

30 **Abstract**

31 This study presents a non-destructive geochemical and petrographic workflow to generate
32 high-resolution chemostratigraphic records across key stratigraphic intervals, here exemplified
33 by a terrestrial Cretaceous-Paleogene (K-Pg) boundary sequence. The geochemical records
34 fingerprint specific Chicxulub related impact ejecta products and thereby further constrain the
35 timeline of ejecta deposition. High-resolution (25 μm) micro-X-ray fluorescence ($\mu\text{-XRF}$)
36 mapping and quantitative integrated-area linescans in combination with ESEM-EDS analyses of
37 the Starkville South sequence (Raton Basin, Colorado) reveal a complex microstratigraphy, in
38 which additional sublayers can be identified than the classic 'dual-layer' succession, described
39 in literature for US Western Interior K-Pg sites. First, a basal claystone is identified with
40 abundant glassy impact spherules that have been altered over time to kaolinite and jarosite due
41 to acidic and reducing conditions in a local swamp environment. This first lithology is followed
42 by a carbonaceous shale interval rich in ejected quartz grains. These two ejecta intervals are
43 interpreted to have formed by ballistic transport from the Chicxulub crater region and were
44 likely emplaced within ~ 1 hour after impact at Starkville. In the overlying lignite layer,
45 pronounced enrichments in both zirconium and chromium are detected, hinting to a triple
46 ejecta layer with a large part of the siderophile element anomaly being likely preserved in this
47 coaly interval, including the famous iridium anomaly. These enrichments are attributed to fine-
48 grained impact dust comprising of pulverized granitoid basement (Zr) and an admixture of
49 meteoritic material (Cr, Ni and likely Ir), probably deposited < 20 years after impact following
50 slow atmospheric settling.

51

52 **Plain Language Summary**

53 This study uses advanced geochemical mapping to analyze layers of debris derived from the
54 Chicxulub asteroid impact, linked to the dinosaur mass extinction 66 million years ago. In
55 sediments from the Raton Basin (Colorado, USA), we identified three sublayers containing
56 Chicxulub impact materials. These are altered melt spherules and ejected mineral grains that
57 were rapidly formed after impact, followed by a layer of ultrafine dust consisting of pulverized
58 granite and meteorite material that settled slowly out of the atmosphere. Our high-resolution
59 geochemical workflow offers new insights into the timing and mechanisms of ejecta
60 production, transport and deposition after a large meteorite impact event.

61

62 **1 Introduction**

63 The Chicxulub hypervelocity impact event on Earth ~ 66 Myr ago represents an ideal
64 case-study to better understand cratering processes on Earth and in the Solar System (Alvarez
65 et al., 1980; Smit and Hertogen, 1980; Kring et al., 2017; Morgan et al., 2022). A key element is
66 the reconstruction of the processes related to impact ejecta production, transport, and
67 deposition (Alvarez et al., 1995; Smit, 1999; Claeys et al., 2002; Artemieva and Morgan, 2009;
68 Schulte et al., 2010). The ~ 200 -km-wide Chicxulub impact structure on the Yucatán Peninsula
69 (Mexico) is unique (Hildebrand et al., 1991) as it is the only terrestrial impact structure with

70 ejecta preserved worldwide, which allows for a direct comparison between material from the
71 source area (Kring and Boynton, 1992; Swisher et al., 1992; Feignon et al., 2020; de Graaff et al.,
72 2022; Kaskes et al., 2024) and its global deposits (Bohor et al., 1987; Kamo and Krogh, 1995;
73 Claeys et al., 2002). High-resolution geochemical and petrographic analysis across these globally
74 distributed Cretaceous-Paleogene (K-Pg) boundary event deposits is crucial to disentangle the
75 sequence of geological and biological events that occurred prior to and in the direct aftermath
76 of the Chicxulub impact event, including the dynamics related to the K-Pg boundary mass
77 extinction (Kring, 2007; Schulte et al., 2010; Morgan et al., 2022; Senel et al., 2023). Recent
78 sedimentological analysis of highly expanded K-Pg boundary successions from within the
79 Chicxulub impact structure (Gulick et al., 2019; Kaskes et al., 2022) and from North-Dakota
80 (DePalma et al., 2019; Senel et al., 2023) make it possible to reconstruct impact cratering
81 processes at an unparalleled resolution of hours to days after impact.

82 The terrestrial K-Pg boundary sites in the US Western Interior (Fig. 1A) are ideal
83 candidates to further unravel the mechanisms and timing of Chicxulub impact ejecta
84 distribution and deposition in the first moments after impact. In contrast to the thick (>1 dm)
85 but high-energy, marine K-Pg deposits around the Gulf of Mexico (at < 1500 km paleodistance
86 to the Chicxulub structure; Smit et al., 1992b; Arz et al., 2022), the paludal Western Interior K-
87 Pg localities have preserved a non-disturbed chronological record of the ejecta sedimentation
88 because the material did not settle through a thick water column (Izett, 1990). In addition, the
89 intermediate paleodistance to the Chicxulub impact structure (2000-3000 km) accounts for a
90 boundary sequence of 1-3 cm thick, thereby providing a much higher temporal resolution
91 compared to the often condensed and thin (< 1 cm) distal K-Pg boundaries at > 5000 km
92 paleodistance to Chicxulub (Smit, 1999; Schulte et al., 2010).

93 Over the last decade, the rapid development of *in-situ* chemical and petrographic
94 techniques at the micro- to even nanoscale makes it possible to analyze and visualize
95 compositional variations at an ultra-high stratigraphic resolution and accuracy. This progress
96 opens a number of opportunities to reconstruct key moments in the early history of abrupt
97 geological processes such as hypervelocity impact events like Chicxulub. High-resolution, (near-
98)continuous datasets from well-preserved K-Pg impact ejecta sequences shed light on the
99 nature and chronology of Chicxulub ejecta products, such as impact spherules (microtektites
100 and microkrystites; (Smit et al., 1992a; Belza et al., 2017)), shocked minerals (Bohor et al., 1984;
101 Bohor et al., 1993), Ni-rich spinels (Robin et al., 1992) and/or iridium rich dust (Alvarez et al.,
102 1980; Smit and Hertogen, 1980; Goderis et al., 2013; Goderis et al., 2021), and may therefore
103 assist in groundtruthing ejecta emplacement models (Artemieva and Morgan, 2009; Artemieva
104 and Morgan, 2020).

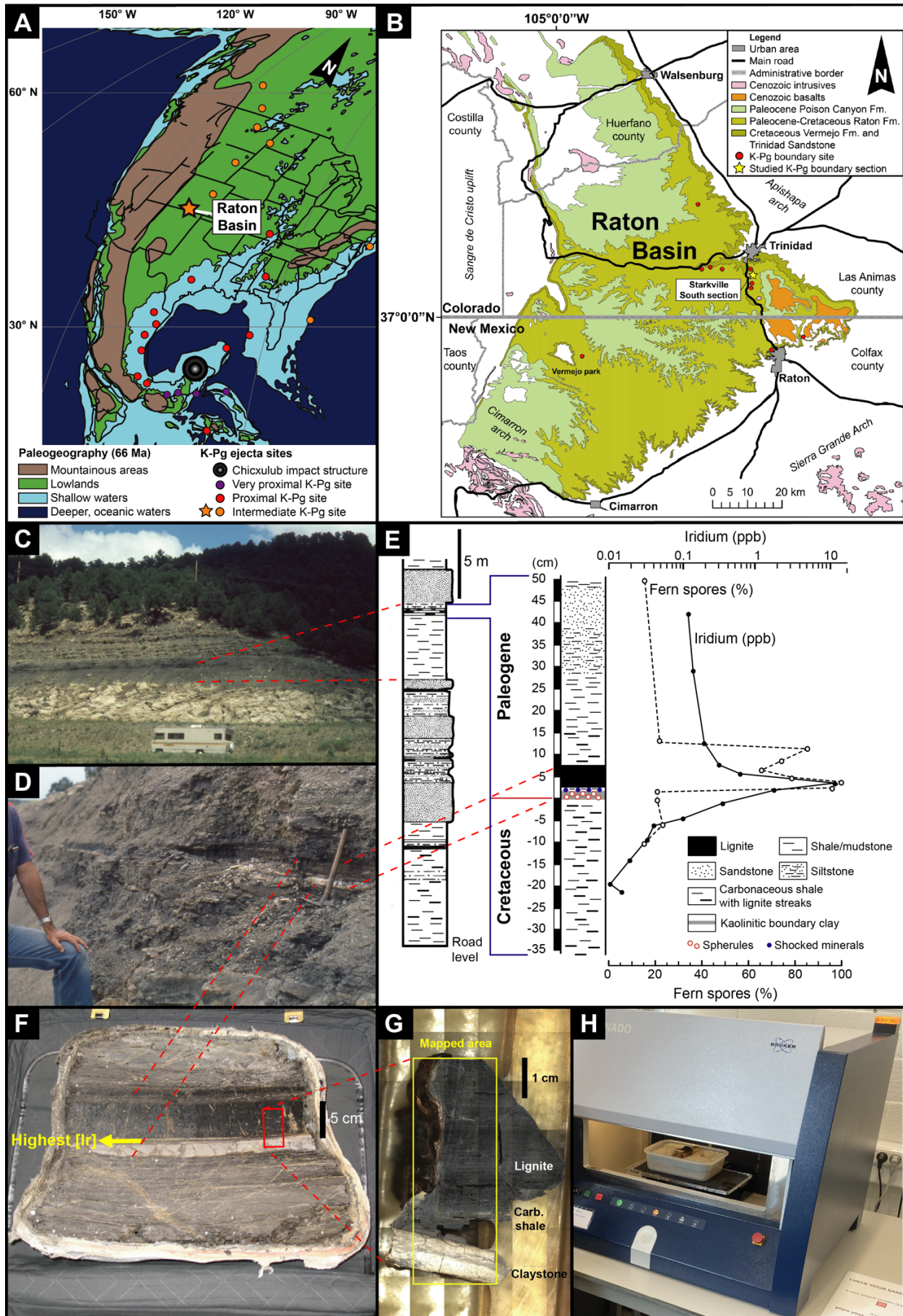
105 This study presents the results of high-resolution micro-X-ray fluorescence (μ -XRF)
106 analysis integrated with scanning electron microscopy on a K-Pg boundary microstratigraphy
107 from the Raton Basin (USA), typical for the US Western Interior intermediate K-Pg sites, to
108 elucidate the complex processes of ejecta transport and deposition linked to the Chicxulub
109 impact event. The novel, non-destructive μ -XRF workflow proposed here, also opens windows
110 to characterize impact ejecta sequences related to other impact structures and to investigate at
111 (sub-)centimeter resolution key stratigraphic intervals in general.

112 **2 Geological setting**

113 The K-Pg boundary deposits from the Raton Basin in the border region between
114 Colorado and New Mexico (USA; Fig. 1B), located in a 6400 km² wide asymmetric structural
115 depression are the most proximal terrestrial K-Pg localities that preserved large quantities of
116 relatively well-preserved Chicxulub impact ejecta material (Orth et al., 1981; Pillmore et al.,
117 1984). The paleodistance from the Raton Basin to the center of the Chicxulub impact structure
118 is estimated to be in between 2100 and 2250 km (Schulte et al., 2010). The Upper Cretaceous
119 to Paleocene stratigraphy in the Raton Basin reflects a regressive sequence consisting mainly of
120 thick packages of siliciclastic deposition at the edge of the retreating Cretaceous Western
121 Interior Seaway (Pillmore and Flores, 1987).

122 The Campanian marine Pierre Shale is followed stratigraphically by the characteristic
123 cliff-forming Trinidad Sandstone, a 25-30 m thick tabular sandstone unit reflecting coastal
124 barrier-bar environments during the eastward progradation of the Upper Cretaceous coastline
125 (Fig. 1B) (Pillmore et al., 1984). This unit is conformably overlain by the fine-grained siliciclastic
126 Vermejo Formation (Maastrichtian) followed by an erosional unconformity at the base of the
127 Raton Formation (Maastrichtian-Danian). This formation comprises conglomerate, sandstone,
128 siltstone, mudstone and coal intervals, and its thickness ranges from 335 m in the eastern part
129 of the Raton Basin to more than 600 m in the west-central region (Pillmore et al., 1984). The K-
130 Pg boundary layer is located in the lower part of the Raton Formation, near the top of the lower
131 coal zone beneath a cliff-forming series of predominantly channel sandstones (Pillmore et al.,
132 1984). The depositional environment of this part of the Raton Formation is interpreted to
133 represent an ever-wet environment containing a combination of meandering channels with
134 low-lying floodplains and local backswamps, where crevasse splays periodically interrupted the
135 deposition of organic material (Pillmore and Flores, 1987; Schwartz et al., 2021). These quiet,
136 freshwater, coal swamp conditions extended over a broad alluvial plain and represented an
137 ideal environment to receive and preserve the fallout of impact ejecta related to the Chicxulub
138 impact event (Sharpton et al., 1990).

139



141 **Figure 1. Geological context of the Starkville-South K-Pg boundary site (SVS). (A)**
142 Paleogeographic map of northern America during the latest Cretaceous, highlighting the Raton
143 Basin, the Chicxulub impact structure and other (very proximal to intermediate) K-Pg localities
144 (adapted from Claeys et al., 2002; Scotese, 2004; Schulte et al., 2010; Snedden and Galloway,
145 2019; Goderis et al., 2021). **(B)** Simplified geological map of the Raton Basin in New Mexico and
146 Colorado displaying the location of SVS and other neighboring K-Pg boundary sites. Geological
147 map adapted from Tweto (1979) and Hoffman and Jopnes (2005). K-Pg boundary localities from
148 Pillmore et al. (1984). **(C-D)** Field photos showing a succession of the lower coal zone of the
149 Raton Formation with a prominent channel sandstone, a mudstone evolving into a
150 carbonaceous shale, then a prominent pink kaolinitic tonstein, darkgrey carbonaceous shale,
151 followed by a lignite interval. Camper van and hammer for scale (pictures by Jan Smit). **(E)**
152 Schematic lithology of Starkville with iridium concentrations (black circles) and percentages of
153 fern spores (open circles). The highest iridium concentrations documented in literature for
154 Starkville are found at the boundary between the upper part of the boundary claystone (the
155 'fireball layer' rich in shocked minerals) and the overlying lignite, as indicated schematically also
156 with a yellow arrow in F. The fern spike consists predominantly of the trilete fern species
157 *Cyathidites diaphana* (adapted from Pillmore & Flores, 1987; Tschudy et al., 1984; Schulte et al.,
158 2010). **(F)** Plaster-jacketed block yielding the SVS K-Pg boundary stratigraphy, currently on
159 display at the Rieskrater Museum in Nördlingen, Germany (Berlin et al., 2010), with highest
160 iridium concentrations from literature projected on the sample. **(G)** Mosaic-scan of the small
161 epoxy resin block cut out of the block from F, the yellow rectangle indicates the mapped
162 surface for the μ -XRF analysis. **(H)** Laboratory setup at the Vrije Universiteit Brussel for the μ -
163 XRF analyses on the horizontally oriented SVS K-Pg epoxy slab using a Bruker M4 Tornado μ -XRF
164 instrument.

165

166 This study focuses specifically on the Starkville South K-Pg site (SVS; Fig. 1B-E) in
167 southern Colorado because the lithology at this site is well-described in literature as
168 characteristic for the Raton Basin (Pillmore et al., 1984; Pillmore and Flores, 1987; Pollastro and
169 Pillmore, 1987). The clastic fraction of this section is regarded as relatively unaltered with
170 abundant impact-related minerals (e.g., shocked quartz, feldspar) preserved (Sharpton et al.,
171 1990), and its boundary claystone interval is characterized by highly elevated iridium
172 concentrations between 16-56 ppb (Pillmore et al., 1984; Goderis et al., 2013) (Fig. 1E), which
173 highlights the presence of geochemical traces of Chicxulub impact ejecta products. The basal
174 part of a lignite overlying the boundary clay interval is also characterized by a pronounced fern
175 spike, a clear marker for the K-Pg mass extinction horizon (Tschudy et al., 1984) (Fig. 1E). The
176 section at SVS consists of ~25 m stratigraphy (Pillmore & Flores, 1987), comprising a prominent
177 channel sandstone, a mudstone evolving into a carbonaceous shale, followed by a prominent
178 pinkish kaolinitic tonstein and a darkgrey carbonaceous shale that transitions into a Paleocene
179 lignite interval (Fig. 1C-G). SVS is situated in a roadcut ~4 km south of the exit to the town of
180 Starkville from the Interstate Highway 25, southern Colorado (37°5'57.30"N; 104°31'14.50"W).
181 Geochemical and petrographic analysis of SVS has so far mostly focused on the pink kaolinite
182 layer and on bulk samples at a centimeter to millimeter stratigraphic sampling (Sharpton et al.,
183 1990; Pillmore et al., 1984; Pollastro & Pillmore, 1987; Izett, 1990). Here, the main focus lies on

184 the sedimentary sequence directly above the kaolinitic claystone. A multi-proxy workflow at a
185 micrometer-scale resolution is introduced to identify the ejecta components and to unravel the
186 depositional processes of the SVS K-Pg event stratigraphy.

187

188 **3 Materials and Methods**

189 In 2005, researchers from the University of New Mexico excavated a large K-Pg boundary block
190 from the Starkville-South K-Pg site (Berlin et al., 2010), which is currently on display in the
191 Rieskrater Museum (Nördlingen, Germany). This sample was stabilized with plaster of Paris and
192 epoxy, and was subsequently cut and polished, preserving a continuous ~25 cm thick
193 microstratigraphic succession of an Upper Cretaceous silt- to claystone, a K-Pg boundary
194 interval of kaolinitic clay and carbonaceous shale, and a Paleocene lignite (Fig. 1F). A
195 representative subsample from this resin block (~3 x 5 cm in size), comprising the kaolinitic
196 tonstein to lignitic part of the K-Pg boundary complex (Fig. 1G), was selected to study in a non-
197 destructive manner the geochemistry and petrography of SVS at a high spatial resolution of 25
198 μm . As only this embedded block was available for study, no destructive bulk geochemical
199 analysis was possible, and no thin sections were made, for example to study the potential
200 presence of shock features within ejected grains.

201 *3.1 Micro-X-ray fluorescence*

202 The major and trace element composition of the SVS resin slab was determined using energy-
203 dispersive micro-X-ray fluorescence ($\mu\text{-XRF}$) scanning at the Vrije Universiteit Brussel, Belgium
204 (VUB), and at Bruker nano GmbH in Berlin, Germany. For this purpose, an M4 TORNADO
205 benchtop $\mu\text{-XRF}$ surface scanner (Bruker nano GmbH, Berlin, Germany) equipped with a Rh
206 tube as X-ray source, a focusing polycapillary lens (20 μm for Mo $\text{K}\alpha$) and two XFlash 430 Silicon
207 Drift detectors was used (Fig. 1H). This flexible technique allows to obtain high-resolution
208 elemental distributions by scanning flat sample surfaces in a rapid, non-destructive, and cost-
209 efficient manner (de Winter and Claeys, 2017; Wouters et al., 2020; Kaskes et al., 2021). As the
210 sample contained an abundant hydrocarbon fraction which causes a higher attenuation depth
211 for the X-rays (Kaskes et al., 2021), the $\mu\text{-XRF}$ mapping was performed using both detectors at
212 intermediate X-ray source energy settings (50 kV and 300 μA) and with the sample oriented
213 horizontally in the vacuum chamber. This orientation was selected to avoid sample material
214 being excited at a higher stratigraphic interval, invoking potential shifts in the
215 chemostratigraphy caused by the oblique incident X-ray beams and the higher attenuation
216 depth of the organic material. The measurements were carried out under near vacuum
217 conditions (20 mbar) with a pixel size resolution of 25 μm and integration time of 50 ms per
218 pixel. This allowed to generate elemental distribution maps; multi-element maps and single-
219 element heatmaps (Fig. 2) in which the color intensity is directly correlated with the intensity of
220 the characteristic element signal per given pixel. To convert the element distribution maps into
221 continuous line profiles, the data was extracted from two subareas with a continuous
222 stratigraphy. This work was done using the linescan function integrated into the M4 TORNADO
223 Software. The extracted data allowed to produce two partly overlapping chemical profiles of

224 the K-Pg boundary interval at 25 μm stratigraphic resolution, which are hereafter termed
225 integrated-area linescans (IAL; i.e. IAL-1 and IAL-2).

226 To obtain quantitative $\mu\text{-XRF}$ linescans, 15 (nanometer-milled) pressed powder pellets of
227 certified geological reference materials (provided by myStandards GmbH, Kiel, Germany) are
228 measured by applying repeated $\mu\text{-XRF}$ spot analyses ($n = 3$ per standard) using the same
229 settings as the mapping described above, but with an acquisition time of 240 s (real-time) per
230 25 μm pixel size. A similar nano-particulate pressed powder pellet (NIOZ Foraminifera House
231 Standard-2-Nano-Pellet) has recently been developed and tested for a satisfactory
232 homogeneous composition, which resulted in various paleoclimate research applications (Boer
233 et al., 2022). The 15 nano-pellets consist of certified reference materials that range from felsic
234 and mafic igneous rocks to carbonates (details in Table 1; values obtained from the GeoREM
235 database (Jochum et al., 2005)). The first quantification of the XRF spectra is performed by
236 using the fundamental parameters algorithm (Rousseau, 1984a; Rousseau, 1984b; Rousseau
237 and Bouchard, 1986) integrated in the M4 TORNADO $\mu\text{-XRF}$ software, which is based on the
238 theoretical Sherman equation (Sherman, 1955) that allows to infer from the measured spectra
239 the respective elemental concentrations (De Winter et al., 2017). The quantification results (in
240 non-normalized oxides wt%) of the 15 reference samples were plotted versus the reference
241 values, to obtain the sensitivity correction (slope) and potential intercept value required to
242 quantify the unknown sample. The results in Fig. 4, show all a very good linear correlation ($R^2 >$
243 0.99). This method of standard supported fundamental parameter quantification not only
244 allows to improve the analytical results but serves also as a validation. All chemical profile data
245 of IAL-1 and IAL-2 can be found in Supplementary Table S1, and is plotted in Fig. 5 and Fig. S3.

246 *3.2 Scanning electron microscopy*

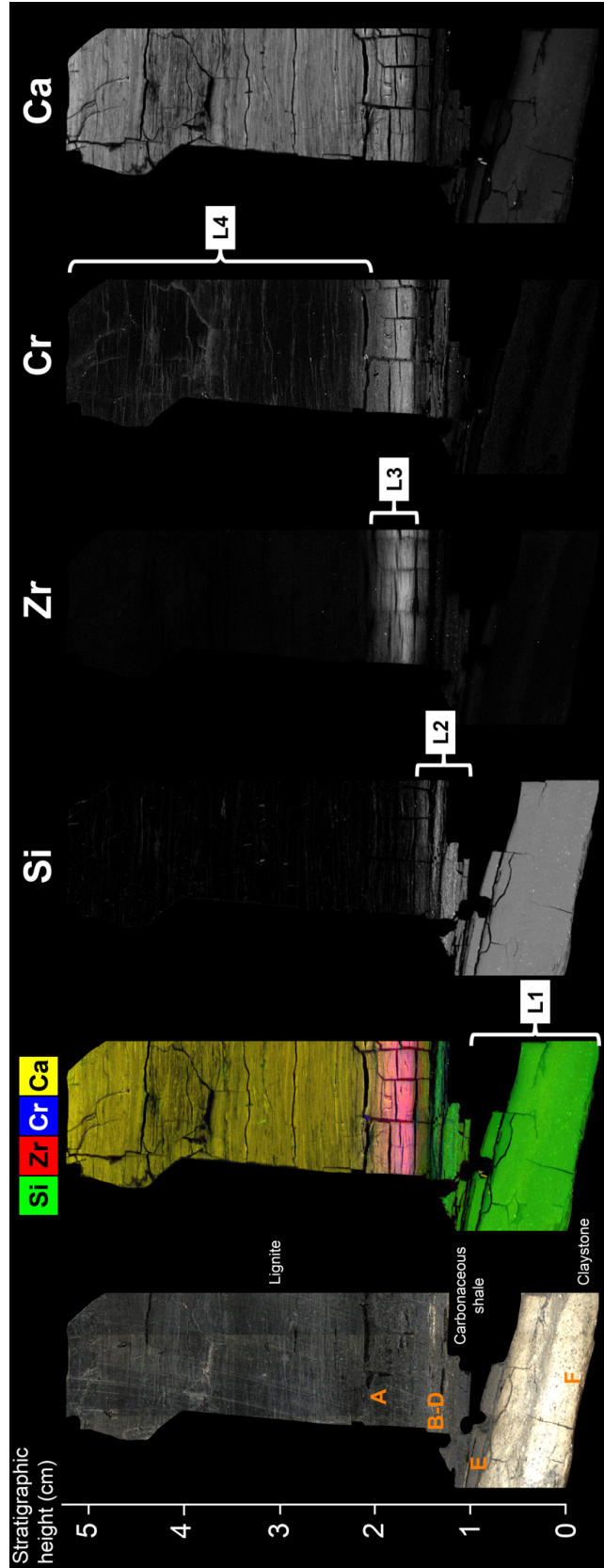
247 Detailed petrographic examination of the SVS slab (Fig. 3) is carried out using a Quanta 20 (FEI
248 Company, Hillsboro, USA) low vacuum environmental scanning electron microscope (ESEM) at
249 the Institute of Natural Sciences, Brussels, Belgium (RBINS) (Decr e et al., 2020). For this study,
250 environmental SEM is preferred because of the large mounting stage (L x W x H: 14 x 6 x 4 cm)
251 and the ability to visualize microtextures in an organic-rich sample while avoiding the use of a
252 carbon coating. The ESEM at the RBINS is equipped with an Apollo 10 Silicon Drift energy
253 dispersive X-ray spectrometer (EDS; EDAX, Pleasanton, USA), which is used to verify the
254 composition of mineral and ejecta phases using qualitative spot analyses on backscatter
255 electron (BSE) images.

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261 **Figure 2. Overview of elemental distributions.** Multi-element μ -XRF map of Si (green), Zr (red),
262 Cr (blue) and Ca (yellow), and semi-quantitative, single-element heat-maps for Si, Zr, Cr, and Ca,
263 displayed with deconvoluted settings. Based on geochemical and petrographic observations,
264 the SVS microstratigraphy is divided into four sublayers (L1-L4) marked on the maps. The letters
265 A-F on the optical image refer to the ESEM-BSE images shown in Fig. 3. Additional single-
266 element heatmaps are shown in Fig. S1.

267

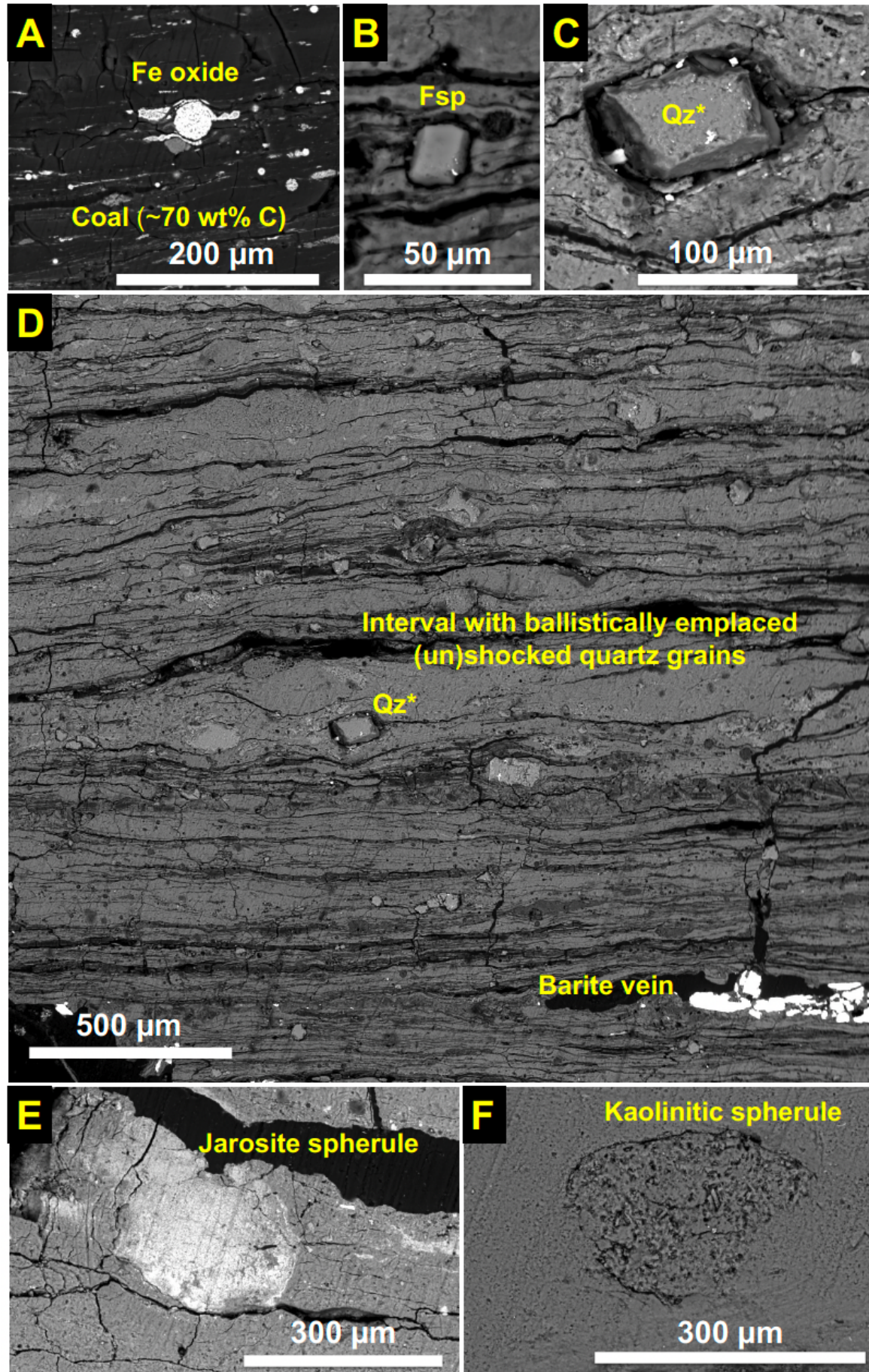
268 **4 Results**

269 The μ -XRF mapping of the SVS slab detected 32 elements (Fig. S1), namely: Na, Mg, Al, Si, P, S,
270 Cl, K, Ca, Ti, V, Cr, Mn, Fe, Co, Ni, Cu, Zn, Ga, Ge, As, Rb, Sr, Y, Zr, Nb, Mo, Ba, Pb, Th, and U, and
271 Rh (linked to Rh tube). Of this list of major and trace elements, P, S, As, Co, Ge, Mo, Pb, Th, U
272 and Rh could not be quantified accurately based on the matrix-matched multi-standard
273 calibrations of nanopowdered pressed pellets of certified reference materials (Figs. 4; S2). The
274 calibration protocol relying on non-normalized mass concentrations resulted in a slightly better
275 coefficient of determination (R^2) and more realistic slope values than the net intensity
276 approach, so the quantifications using the exported mass concentrations from the Bruker M4
277 software based on the fundamental parameters method are used. Using this method, R^2 values
278 are > 0.99 for Na_2O , MgO , Al_2O_3 , SiO_2 , K_2O , CaO , TiO_2 , MnO , Fe_2O_3 , Cl, V, Cr, Ni, Cu, Zn, Ga, Rb,
279 Sr, Y, Zr, Nb, and Ba (Table 2). Key major oxides and trace elements for the identification of
280 impact ejecta products and intervals represent SiO_2 (potentially linked to ejected quartz grains:
281 Bohor et al., 1984), Al_2O_3 and K_2O (clay components: Pollastro and Pillmore, 1987), Zr (e.g.,
282 ejected zircon: Krogh et al., 1993a), Cr and Ni (for Ni-rich spinels (Robin et al., 1992),
283 microkrystites (Belza et al., 2017), and for an admixed meteoritic component (Goderis et al.,
284 2013)). The results of the integrated area linescan (IAL) quantifications are discussed below and
285 are shown in Fig. 5 (for SiO_2 , Zr, Cr and Ni profiles) and Fig. S3 for all major and trace element
286 profiles.

287 By focusing primarily on the distribution and IAL-based chemical profiles of SiO_2 , Zr, Cr and Ni,
288 and using a verification by ESEM-EDS, the K-Pg boundary interval is divided into four distinct
289 layers (Figs. 2 and 5). The lowermost stratum (L1: > 0.98 cm thick) represents a claystone that is
290 relatively homogeneous in composition showing consistently high SiO_2 (~ 53 wt%) and Al_2O_3
291 contents (~ 37 wt%) (Fig. 2), and relatively low K_2O (generally < 1 wt%), Zr and Cr concentrations
292 (both < 100 ppm) and very low Ni concentrations (< 12 ppm). ESEM-EDS analysis showed that
293 this layer yields altered impact spherules (Fig. 3E-F). The overlying carbonaceous shale layer (L2:
294 ~ 5.5 mm thick, from ~ 0.98 – 1.53 cm stratigraphic height) has clear hotspots enriched in both
295 SiO_2 (up to 51 wt%) and Zr (first clear Zr hotspot of 428 ppm at 0.98 cm stratigraphic height;
296 highest Zr hotspot value of 801 ppm at 1.5025 cm height) (Fig. 2), and contains granular
297 petrographic features with ejected quartz and feldspar grains as evidenced by ESEM-EDS (Fig.
298 3B-D). L2 also shows some elevated peaks in Cr values (on average 120 ppm with a maximum
299 value up to 630 ppm at 1.155 cm height) and a very small number of peaks with Ni enrichments
300 (one peak up to 90 ppm). The highest bulk iridium values available in literature for Starkville

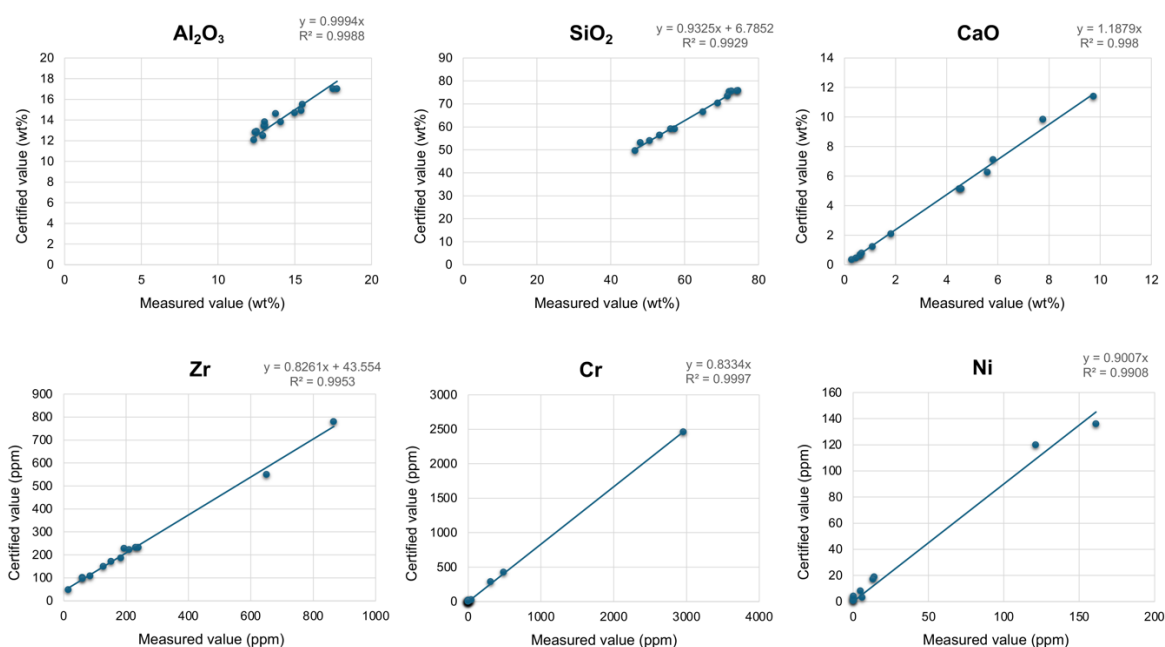
301 South (56 ppb) is observed in this layer with ejected shocked mineral grains (Pillmore et al.,
302 1984), which corresponds to L2 in this study. The third layer (L3: ~8.2 mm thick from ~1.53 -
303 2.35 cm) is a transition interval from carbonaceous shale to lignite (Fig. 3A) and is characterized
304 by low SiO₂ values (< 25 wt% with on average 11.7 wt%), and a sharp peak with both very high
305 concentrations of Zr (maximum value of 3984 ppm at 1.79 cm stratigraphic height) and Cr
306 (maximum value of 839 ppm at 1.71 cm height). Ni also shows clear enrichments in this
307 interval, with a maximum value of 474 ppm Ni located at 2.065 cm stratigraphic height (Fig. 5).
308 Above 2.35 cm stratigraphic height, the Zr concentrations drop consistently below detection
309 limits (> 44 ppm), which we therefore link to the presence of a fourth layer (L4: > 3 cm thick;
310 from ~2.35 cm upwards). This is a lignite that yields, besides a low Zr content, also low SiO₂
311 values (on average 8.7 wt%, with a maximum up to 11.5 wt%), and low Cr concentrations (< 200
312 ppm; on average 11 ppm Cr). In contrast, L4 displays large variations in Ni with high values up
313 to 455 ppm (Fig. 5).

314



316 **Figure 3. Representative backscattered electron (BSE) images of impact ejecta components.**
317 (A) Lignite matrix with iron oxide. (B) Ejected feldspar from L2. (C) Ejected quartz grain in L2
318 showing downwarping of surrounding sediment layers indicating ballistic deposition in soft
319 sediment. (D) Overview of sublayer L2 showing ejected quartz minerals with downwarping
320 strata and a barite vein, the latter likely represents authigenic mineralization during peat
321 formation (Rudmin et al., 2018). (E) Jarosite impact spherule in the top part of sublayer L1. (F)
322 Kaolinitic spherule from lower part of L1. The positions of the BSE images are indicated on the
323 slab photo from Fig. 2 with the four sublayers. Mineral abbreviations based on Whitney and
324 Evans, 2010).

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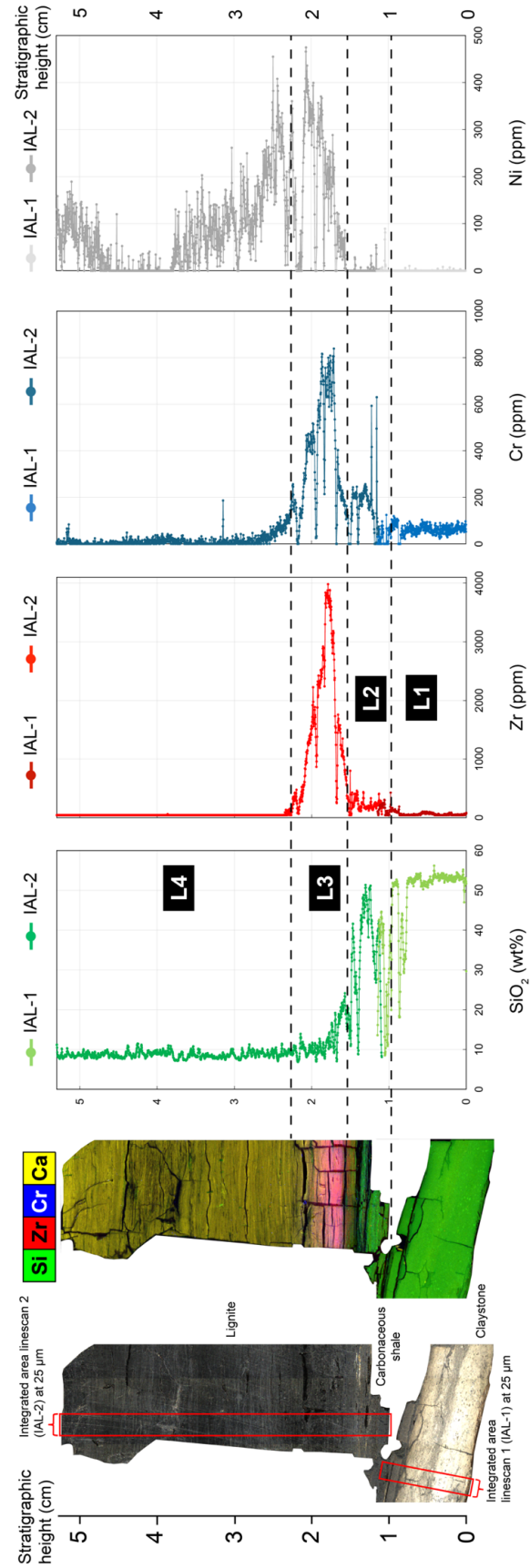


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327 **Figure 4. Calibrations using μ -XRF spot analyses on certified reference materials.**
328 Representative calibration curves using the mass concentrations exported from the M4 Bruker
329 software and reference values for the major oxides Al₂O₃, SiO₂, CaO, and trace elements Zr, Cr,
330 and Ni. All calibration curves are displayed in Supplementary Figure 2 and the linear regression
331 slope (and possible intercept) with associated coefficients of determination (R²) are shown in
332 Table 2.

333

334



336 **Figure 5. Overview of quantitative high-resolution (25 μm) $\mu\text{-XRF}$ linescans of the SVS K-Pg**
337 **boundary interval.** Chemical profiles at 25 μm stratigraphic resolution produced using the
338 integrated area linescan option in the Bruker M4 $\mu\text{-XRF}$ software, showing the vertical
339 variations of SiO_2 (in wt%), Zr, Cr, and Ni (all in ppm). The two red rectangles refer to the two
340 regions of the integrated-area linescans (IAL-1 and IAL-2: see the two colors in the element
341 profiles). The four sublayers (L1-L4), based on geochemistry and petrography, are marked on
342 the profiles.

343

344 5 Discussion

345 5.1 Novel $\mu\text{-XRF}$ linescan workflow

346 The $\mu\text{-XRF}$ calibration curves on a range of nanopowdered press pellets yield excellent
347 regression lines for 22 elements in total ($R^2 > 0.99$, see Table 2), suitable for accurate
348 quantifications. The integrated-area linescan (IAL) quantification reveals to be a successful
349 measurement mode as it provides a representative integration while still preserving a high
350 stratigraphic resolution of 25 μm necessary for detecting subtle changes in depositional
351 processes. This way, with a single $\mu\text{-XRF}$ element mapping run already reliable chemical profiles
352 can be extracted without the need for additional point-linescans. In this case study, an
353 integrated area of ~ 3 mm wide is selected (Fig. 5). Using a wider area in this sample is not
354 feasible due to the presence of cracks filled up by epoxy, resulting in a combined acquisition
355 time of 6 seconds per stratigraphic slice of 25 μm at a measurement time of 50 ms per pixel.

356 To identify the intervals that might have been chemically affected by the presence of a
357 (micro)crack filled with epoxy, the presence of chlorine was used as a proxy. There are some
358 peaks present in Cl (> 0.3 wt%) in the chemical profile, but the influence on the overall
359 chemostratigraphic trends in key elements such as Si, Zr, Cr and Ni is relatively limited as it
360 affects only thin intervals of < 300 μm (Fig. S3). Moreover, some of the cracks have been filled
361 with the mineral barite (BaSO_4 ; as verified by ESEM-EDS, see Fig. 3D), but these barite veins are
362 not thicker than 400 μm , so the effect on the general chemical trends are also considered
363 limited (Fig. S3).

364 The sample selected for this study is also challenging in terms of quantifications as it contains a
365 large organic fraction. The M4 Bruker TORNADO $\mu\text{-XRF}$ instrument cannot detect C, but EDS
366 spot analyses reveal C concentrations of around 70% by weight in sublayer 3 and 4. Here, a
367 universal method for $\mu\text{-XRF}$ is developed suitable for a range of lithologies and applied without
368 sum-normalization. K-Pg boundaries are often characterized by abrupt lithological changes, for
369 example from a pure limestone to a claystone in marine settings, and a 'one fits all calibration'-
370 approach suitable for a range of matrix compositions is most likely the optimal solution to
371 produce continuous chemical profiles across lithological boundaries. Previous laser
372 ablation-inductively coupled plasma-mass spectrometry (LA-ICP-MS) analysis on the K-Pg
373 boundary preserved in drill core samples from the Demarara Rise (235 μm spotsize) and
374 Chicxulub crater site (150 μm spotsize) showed that a nugget effect can strongly influence a

375 chemical enrichment in siderophile elements (Berndt et al., 2011; Goderis et al., 2021). By
376 systematically introducing μ -XRF mapping and IAL into this workflow, the negative influence of
377 the nugget effect can be visualized and avoided as much as possible before detailed trace and
378 platinum group element analysis is being carried out, for example relying on LA-ICP-MS.

379

380 *5.2. Geochemical fingerprinting of impact ejecta phases and intervals*

381 For the identification of impact ejecta phases, SiO₂, Zr, Cr, and Ni are selected. K₂O and Al₂O₃
382 are used to identify different types of clay minerals. Subunit L1 is characterized by a relatively
383 homogeneous composition dominated by SiO₂ (54.2 ± 3.4 wt%) and Al₂O₃ (36.7 ± 3.2 wt%),
384 consistent with EDS spot analyses (Fig. 3F) and bulk XRF values (54.8 ± 0.4 wt% SiO₂; 40.6 ± 0.5
385 wt% Al₂O₃, comparison is based on an anhydrous sum-normalization) of the same K-Pg
386 boundary claystone at Starkville-South from Izett (1990). This corresponds to a mineralogy
387 almost exclusively dominated by kaolinite (Al₂Si₂O₅(OH)₄), as also evidenced by bulk powder X-
388 ray diffraction (XRD) data from Pollastro & Pillmore (1987).

389 The overlying subunit L2 shows more chemical variations with a generally lower and more
390 variable SiO₂ content (47.9 ± 11.9 wt%), lower Al₂O₃ concentrations (24.2 ± 7.9 wt%), but higher
391 K₂O values (1.8 ± 0.6 wt% for L2) compared to L1 (0.8 ± 0.2 wt% K₂O for L1). Pollastro &
392 Pillmore (1987) observed that the shale beds found throughout the K-Pg boundary interval in
393 the Raton Basin consist of a suite of minerals typical of shale, including quartz (20-40 wt%),
394 feldspar, illite-smectite clays, mica (usually illite), and kaolinite. This mineral assemblage
395 explains the more variable composition and the incorporation of K in the presence of the
396 mineral phases K-feldspar (KAlSi₃O₈) and illite (K,H₃O)(Al,Mg,Fe)₂(Si,Al)₄O₁₀[(OH)₂, (H₂O)]).

397 The L2 subunit contains evidence of ejected quartz and feldspar grains based on μ -XRF and
398 ESEM-EDS analyses. Hotspots of zirconium on the high-resolution μ -XRF maps of the scanned
399 epoxy slab likely also reveal the presence of ejected zircon. However, an ESEM-EDS study did
400 not confirm the presence of zircon yet, which could be explained by the attenuation depth of Zr
401 for X-rays (i.e., 0.2-1.1 mm as shown in Kaskes et al., 2021) and hence the presence of a
402 potential zircon grain below the sample surface. More sample material, thin sections and a
403 more detailed petrographic investigation are needed to identify these zircons throughout their
404 microstratigraphy to use them for example for U-Pb dating relying on depth-profiling LA-ICP-MS
405 (Rasmussen et al., 2019) or to characterize their shock state using Electron Backscatter
406 Diffraction (EBSD) (Cavosie et al., 2018).

407 Subunit L3, on top of the spherule (L1) and ejected quartz bed (L2), represents the transition
408 zone between a carbonaceous shale and lignite. L3 is strongly enriched in Zr, Cr, and Ni. The
409 enrichment in Zr is interpreted to be linked to the presence of ultrafine dust composed of
410 pulverized Yucatán granitic basement. Based on previous geochemical analyses on K-Pg
411 boundaries (de la Parra et al., 2022) and impactites from the Chicxulub crater (Goderis et al.,
412 2021; Feignon et al., 2022), Cr and Ni have successfully been used as a marker for meteoritic

413 contribution as these siderophile elements behave chemically similarly to platinum group
414 elements such as iridium (Schmitz, 1992; Goderis et al., 2013).

415 Unit L4 consists purely of a lignite with an estimated carbon content value of 70 wt%. The
416 remaining inorganic matter in the lignite shows relatively low values of SiO₂ (2.1 ± 0.9 wt% after
417 recalculation based on 30 wt% mineral fraction) and Al₂O₃ (1.4 ± 0.6 wt%), but higher
418 concentrations in CaO (3.2 ± 1.3 wt%), and Fe₂O₃ (19.8 ± 3.4 wt%). A similar trend with
419 relatively high CaO and Fe₂O₃ concentration values is found in the Neogene Velenje lignite in
420 Slovenia (Markič, 2006), showing that these elements are mostly organically bound. The
421 Velenje lignite corresponds to a fine detrital coal, interpreted to be deposited at a swamp-lake
422 border closer to open water settings, which might also be the case for the SVS lignite (Markič,
423 2006).

424

425 *5.3. Impact ejecta chronology and emplacement mechanisms*

426 Figure 6 summarizes the chronology and emplacement mechanisms of the Chicxulub impact
427 ejecta sequence. The Starkville K-Pg impact ejecta sequence started with the deposition of
428 (now replaced) glassy impact spherules (microtektites), which have been detected as ~300 μm
429 sized homogeneous, spherulitic objects in the basal claystone. These spherules subsequently
430 altered to kaolinite, goyazite, or jarosite in the acidic reducing environment of the coal swamps,
431 and now form the thick kaolinitic boundary claystone (L1) (Izett, 1990; Pillmore et al., 1984;
432 Pollastro & Pillmore, 1987). These formerly glassy melt spherules are interpreted to have
433 formed by near-instantaneous shock-melting (Belza et al., 2015) and launched in a turbulent
434 ejecta curtain with angles of 30-45° relative to the surface outside the crater (Alvarez et al.,
435 1995), followed by rapid quenching in flight during the ballistic transport to the place of
436 deposition. The latter likely occurred for US Western Interior sites within ~10-15 minutes (Fig.
437 6; Alvarez et al., 1995; Artemieva & Morgan, 2020).

438 The overlying carbonaceous shale layer (L2) consists of multiple ejected quartz and feldspar
439 grains that are thought to have been ballistically expelled out of the crater as part of the 'warm
440 'fireball phase' with a trajectory most likely steeper than ~65°. Compared to the microtektites,
441 this would result in a slighter slower deposition rate in the order of ~20 minutes to 2 hours,
442 depending on the size of the minerals, the launch angle and the final site of deposition (Fig. 6)
443 (Alvarez et al., 1995; Kring and Durda, 2002). This layer is known as the 'fireball layer'.
444 Chronometric dating using U-Pb of highly shocked zircons from this layer in the Raton Basin of
445 Colorado have revealed resetting ages around ~66 Ma, consistent with their association to the
446 Chicxulub impact structure (Krogh et al., 1993b)

447 Alvarez et al. (1995) proposed that the iridium, as part of an extremely hot fireball consisting of
448 vaporized meteorite and target rock carbonate and silicate, likely deposited as a separate
449 veneer on top of the ejected quartz layer. This proposed 'third layer' has likely been identified
450 in this study using μ-XRF at the Starkville K-Pg site on the basis of a transitional interval (L3)
451 between a carbonaceous shale and a lignite that is enriched in Zr, Cr, and Ni. This coaly interval

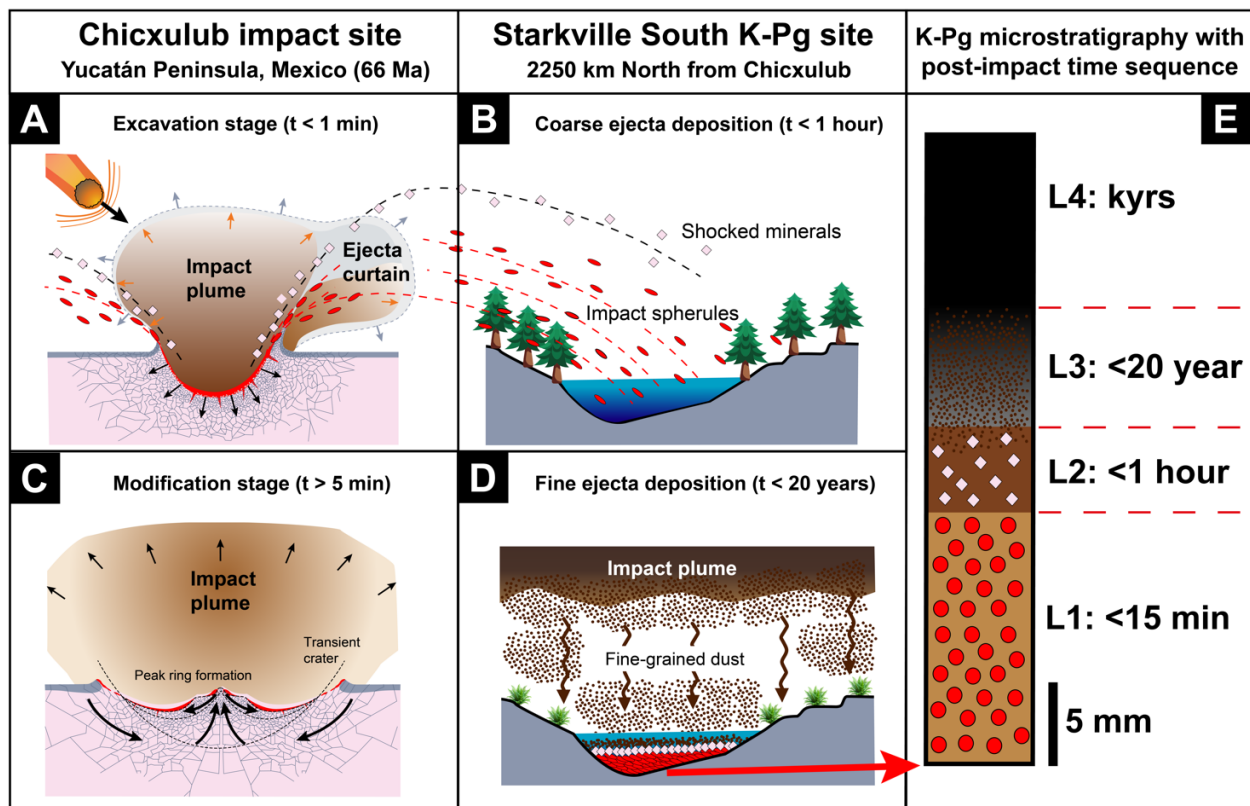
452 is characterized by the highest Zr concentrations of the identified intervals (up to 4725 ppm;
453 Fig. 5), which is interpreted here as being related to a pulverized fraction of the unmelted felsic
454 crystalline basement incorporated within the impact dust plume. Geochemical analyses of
455 granitoids from the recent IODP-ICDP Expedition 364 drilling within the Chicxulub peak ring
456 have shown average Zr values of ~100 ppm (Feignon et al., 2021). The enriched Zr values in L3
457 are most probably related to the refractory behavior and high condensation temperature of Zr
458 (T_{50} of 1722 K (Wood et al., 2019)) within the impact dust plume (Belza et al., 2017), likely
459 steering its slow atmospheric deposition.

460 The bulge with high Zr values (> 2500 ppm) from Starkville, between ~1.7-1.9 cm stratigraphic
461 height, corresponds also to a broad peak in high Cr values (> 600 ppm) (Fig. 5). These high Cr
462 values in the coaly interval (L3) most likely correspond to the Ir and PGE anomaly, as can also be
463 deduced from a similar linear relationship between these two elements as found in the
464 Chicxulub crater itself (Goderis et al., 2021) and other terrestrial and marine K-Pg boundaries
465 around the world (Schmitz, 1992; Goderis et al., 2013; Feignon et al., 2022). The highest Ir
466 values for Starkville described in literature are found in the shocked mineral-bearing
467 carbonaceous shale ("fireball layer": ~56 ppb; Pillmore et al., 1984) or from the interval
468 between this layer and the lignite above (~16 ppb; Tschudy et al., 1984; Fig. 1E). In this study,
469 this corresponds to L2 and the transition between L2 to L3, whereas the highest Cr values in this
470 study are found in L3. Elevated Ir values have been found in a coaly interval above the K-Pg
471 boundary claystone at the site of Madrid East and Clear Creek North based on previous low-
472 resolution studies with up to 10.9 and 14.6 ppb Ir, respectively (Izett, 1990). However, these
473 bulk powder-based studies are relatively low-resolution and future work should focus on more
474 high-resolution PGE datasets. Considering a direct relationship between Cr and Ir, this would
475 mean that the return to background values for Cr marks the final atmospheric settling of fine
476 meteoritic and granitic dust, which has been estimated to have taken less than 20 years post-
477 impact based on atmospheric modelling (Kring and Durda, 2002; Goderis et al., 2021; Senel et
478 al., 2023). Also, peaks of Ni and Cr on the μ -XRF maps might hint to the possible presence of
479 microkrystites containing Ni-Cr-rich and Mg-poor spinel group minerals (Smit, 1999; Kyte &
480 Smit, 1986). So far, these microkrystites have petrographically not been verified during the
481 ESEM-EDS survey, but their presence has been recorded from the uppermost claystone bed
482 from the Tanis K-Pg site (DePalma et al., 2019). An in-depth petrographic analysis on more
483 material from the Raton Basin is preferred to confirm or deny the presence of microkrystites.
484 The highest Ni values in L3 (> 300 ppm) are seen stratigraphically slightly higher (~1.85-2.09 cm
485 height) compared to Zr and Cr (~1.7-1.9 cm) and this upward shift might be related to Ni
486 remobilization in the organic rich environment.

487 This pattern for Ni is also seen in the overlying L4. This sublayer still displays high values of Ni
488 (up to 455 ppm), compared to very low Zr (>50 ppm) and low Cr values (>200 ppm). Ni is more
489 mobile than Cr and Zr (Agnieszka and Barbara, 2012) and therefore likely explains such a
490 contrasting profile. L4 becomes more depleted in Al_2O_3 and SiO_2 upwards and more CaO rich
491 (Fig. S3), suggesting that the lignite become more organic rich with a smaller influence of
492 detrital input towards the top in a paludal/lacustrine paleoenvironment. This lignite is
493 approximately 5 cm thick throughout the Raton Basin and shows the presence of a fern spike:

494 an anomalous abundance of fossil fern spores relative to other palynoflora in the aftermath of
 495 the Chicxulub impact (Tschudy et al., 1984). The duration of this period of fern dominance in
 496 the earliest Danian has been estimated to have lasted for (several) millennia (kyr) (Berry, 2019).
 497 This age indication likely marks the upper boundary of the Ni remobilization. However, more
 498 detailed work on fine-grained ejecta components in the lignite, in conjunction with ejecta
 499 modelling and high-resolution palynology provide more constraints on this timeframe.
 500 Generally, applying a similar μ -XRF-IAL method on other K-Pg boundaries in the Raton Basin and
 501 beyond may constrain potential post-impact alteration in more detail with the aim to further
 502 disentangle the complex ejecta dynamics related to the Chicxulub impact event.

503



504

505 **Figure 6. Schematic visualization of key time intervals in the emplacement of Chicxulub**
 506 **impact ejecta and an idealized K-Pg microstratigraphy of the Raton Basin.** This simplified
 507 model (not to scale, based on Pierazzo and Artemieva, 2012; Belza, 2015; Collins et al., 2020)
 508 displays the results of an 45° oblique impact from the NE. The Yucatán target stratigraphy and
 509 the different types of coarse impact ejecta (glassy impact melt spherules and shocked minerals
 510 derived from the granitic basement) and fine-grained impact ejecta (including silicate dust) are
 511 highlighted with their interpreted timing post-impact (Artemieva & Morgan, 2020; Alvarez et
 512 al., 1995; Goderis et al., 2021, Senel et al., 2023). A) Excavation stage at the Chicxulub crater
 513 region showing the initiation of the ejecta curtain (< 1 minute post-impact). B) Coarse-grained
 514 ejecta deposition at the Starkville-South K-Pg boundary site showing first the deposition of
 515 impact spherules (< 15 minutes) followed by the ballistic emplacement of (shocked) basement

516 mineral grains. C) The modification stage in the crater region showing peak ring formation and
517 the rise of the impact plume into the stratosphere (> 5 minutes). D) Fine-grained ejecta
518 deposition at the Starkville-South K-Pg boundary site derived from an impact dust plume rich in
519 pulverized granitic material and an admixed meteoritic component (< 20 years after impact). E)
520 An idealized K-Pg microstratigraphy of the Raton Basin showing the interpreted chronology
521 displaying a kaolinitic claystone interval with glassy impact spherules (L1), followed by a
522 carbonaceous shale interval with shocked minerals (L2), which is capped by a coaly interval
523 enriched in fine Zr and Cr rich dust (L3), that transitions into a lignite without major
524 atmospheric dust input (L4).

525

526 **6 Conclusions**

527 The μ -XRF workflow approach developed in this study generates high-resolution
528 chemostratigraphic records of key stratigraphic intervals. This is exemplified here with a
529 continuous sediment slab from a terrestrial Cretaceous-Paleogene (K-Pg) boundary. This
530 workflow is powerful in geochemically fingerprinting specific impact ejecta products and
531 thereby elucidating the timeline of the complex ejecta processes involved. High-resolution μ -
532 XRF mapping and quantitative linescans (25 μ m) in combination with ESEM-EDS petrography
533 revealed that the Starkville South sequence has a more complex microstratigraphy than the
534 'dual-layer' succession described in literature for US Western Interior K-Pg sites. In this study,
535 the sample stratigraphy encompasses four layers (L1 – L4). First, a basal claystone is identified
536 with abundant glassy impact spherules that have now been altered to kaolinite and jarosite
537 (L1), due to the acidic and reducing conditions in a local swamp environment. This first lithology
538 is followed by a carbonaceous shale interval rich in ejected quartz grains (L2). Using ESEM-EDS
539 analyses, these two ejecta intervals, which are interpreted to have formed by ballistic transport
540 from the Chicxulub crater region and were likely emplaced within ~1 hour after impact at
541 Starkville. Subsequently, pronounced enrichments in both zirconium and chromium are
542 detected in the overlying lignite layer (L3), which are attributed to fine-grained impact dust
543 comprising pulverized granitoid basement with admixture of meteoritic material, likely
544 deposited <20 years after impact through slow atmospheric settling. Based on the previously
545 observed linear relationship between chromium and iridium, this suggests that a large part of
546 the iridium-rich particles settled during the deposition of the basal part of the coaly layer
547 instead of during the underlying 'fireball layer'. Based on low zirconium contents, a final
548 interval in the lignite layer (L4) is identified, which is still enriched in nickel likely linked to post-
549 impact mobilization in a detrital coal swamp environment during the millennia post-impact.
550 These geochemical observations can serve as input parameters for ejecta models, which can
551 shed more light on the timing and mechanisms of impact ejecta processes for Chicxulub and
552 other impact events in the Solar System.

553

554 **As Applicable – Inclusion in Global Research Statement**

555 The Authorship: Inclusion in Global Research policy aims to promote greater equity and transparency in
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566 received, and/or description of end-users of the research. You can learn more about the policy in this
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568 [Gutiérrez et al.](#), [Tully et al.](#) Please note that these statements are titled as “Global Research
569 Collaboration Statements” from a previous pilot requirement in *JGR Biogeosciences*. The pilot has ended
570 and statements should now be titled “Inclusion in Global Research Statement”.

571

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581

582 **Open Research**

583 All μ -XRF data from this study is available via the Zenodo online repository:

584 <https://zenodo.org/records/14552419>.

585

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839 **Tables**

840

#	Reference name	Details of institute	Material
1	BHVO-2	U.S. Geological Survey (USGS)	Basalt
2	RGM-2	U.S. Geological Survey (USGS)	Rhyolite
3	OREAS-20a	Ore Research and Exploration Pty.Ltd., Bayswater North, Australia (OREAS)	Granodiorite
4	GSP-2	U.S. Geological Survey (USGS)	Granodiorite
5	AGV-2	U.S. Geological Survey (USGS)	Andesite
6	AC-E	Centre de Recherches Petrographiques et Geochimiques (CRPG), Vandoeuvre-lès-Nancy, France	Granite
7	BCR-2	U.S. Geological Survey (USGS)	Basalt
8	OKUM	International Association of Geoanalysts, Nottingham, UK	Komatiite
9	GH	Centre de Recherches Petrographiques et Geochimiques (CRPG), Vandoeuvre-lès-Nancy, France	Granite
10	JA-1	Geological Survey of Japan, Higashi, Tsukuba, Ibaraki	Andesite
11	SARM-1	Council for Mineral Technology, Randburg, South Africa, (MINTEK)	Granite
12	RMG-1	U.S. Geological Survey (USGS)	Rhyolite
13	JR-2	Geological Survey of Japan, Higashi, Tsukuba, Ibaraki	Rhyolite

14	JR-1	Geological Survey of Japan, Higashi, Tsukuba, Ibaraki	Rhyolite
15	JB-2	Geological Survey of Japan, Higashi, Tsukuba, Ibaraki	Basalt

841 **Table 1:** List of certified reference materials used in this study.

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#	Element or oxide	Slope value	Intercept	R ²
1	Na ₂ O (wt%)	0.7985	0.8022	0.9902
2	MgO (wt%)	1.0201	0.3914	0.9986
3	Al ₂ O ₃ (wt%)	0.9994	0	0.9988
4	SiO ₂ (wt%)	0.9325	6.7852	0.9929
5	K ₂ O (wt%)	0.9701	0.1741	0.9977
6	CaO (wt%)	1.1879	0	0.9980
7	TiO ₂ (wt%)	1.0736	0	0.9993
8	MnO (wt%)	0.9750	0	0.9996
9	Fe ₂ O ₃ (wt%)	0.8968	0	0.9997
10	Cl (ppm)	1.6902	0	0.9986
11	V (ppm)	0.9439	0	0.9923
12	Cr (ppm)	0.8334	0	0.9997
13	Ni (ppm)	0.9007	0	0.9908
14	Cu (ppm)	0.7171	0	0.9936
15	Zn (ppm)	0.7361	0	0.9961
16	Ga (ppm)	0.6268	0	0.9906
17	Rb (ppm)	0.9236	0	0.9982
18	Sr (ppm)	0.8182	0	0.9996
19	Y (ppm)	1.0142	12.12	0.9914
20	Zr (ppm)	0.8261	43.55	0.9953
21	Nb (ppm)	0.7595	0	0.9921
22	Ba (ppm)	1.0773	0	0.9924

843 **Table 2:** Slope, intercept and R² values for the μ -XRF regression lines for a selection of elements
844 or oxides based on a calibration with the certified reference materials listed in Table 1 (see Fig.
845 S2 for the calibration curves).

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Supplementary Material

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Disentangling impact ejecta dynamics using micro-X-ray fluorescence (μ -XRF): a case study from the terrestrial Cretaceous-Paleogene (K-Pg) boundary

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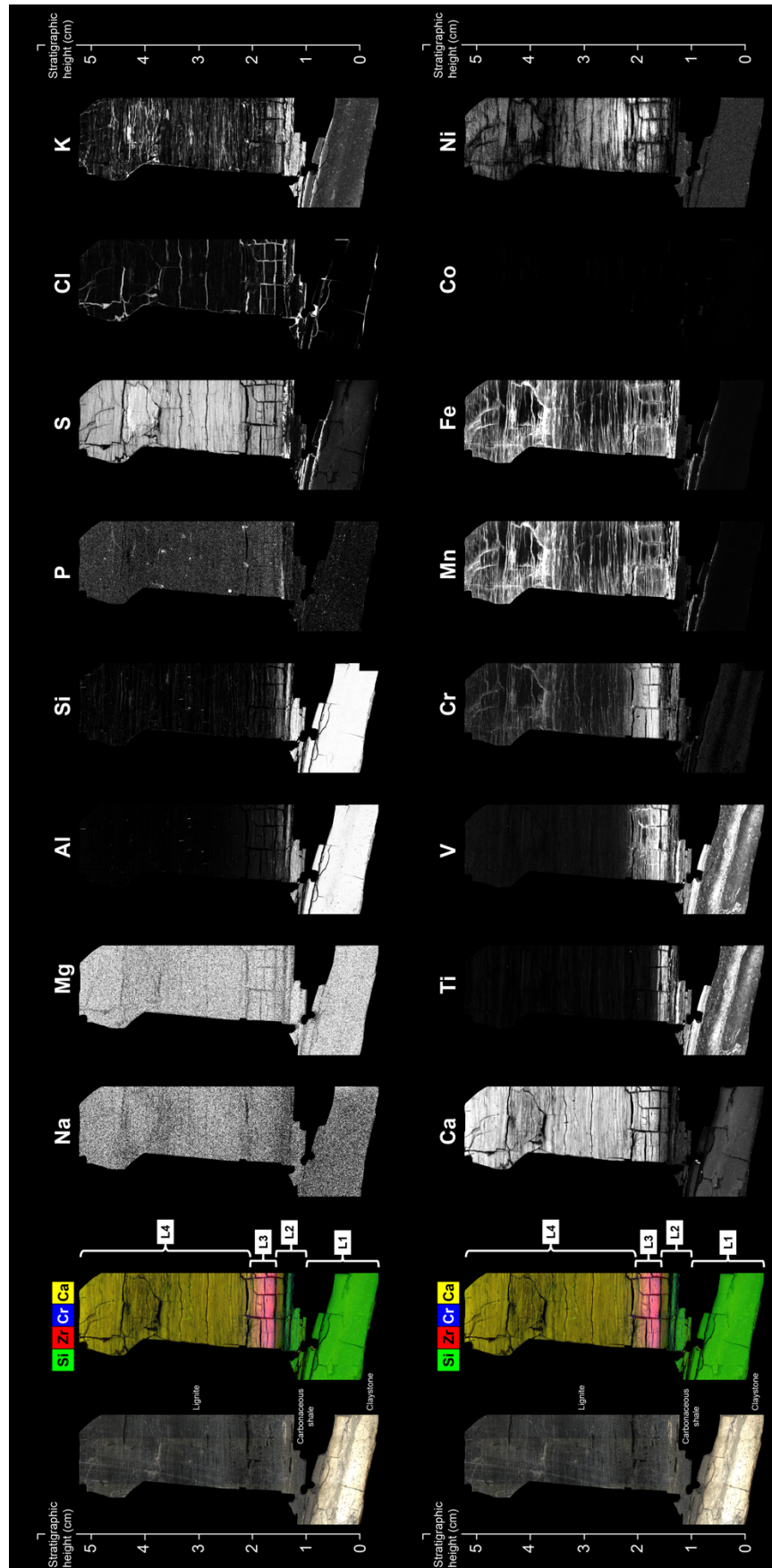
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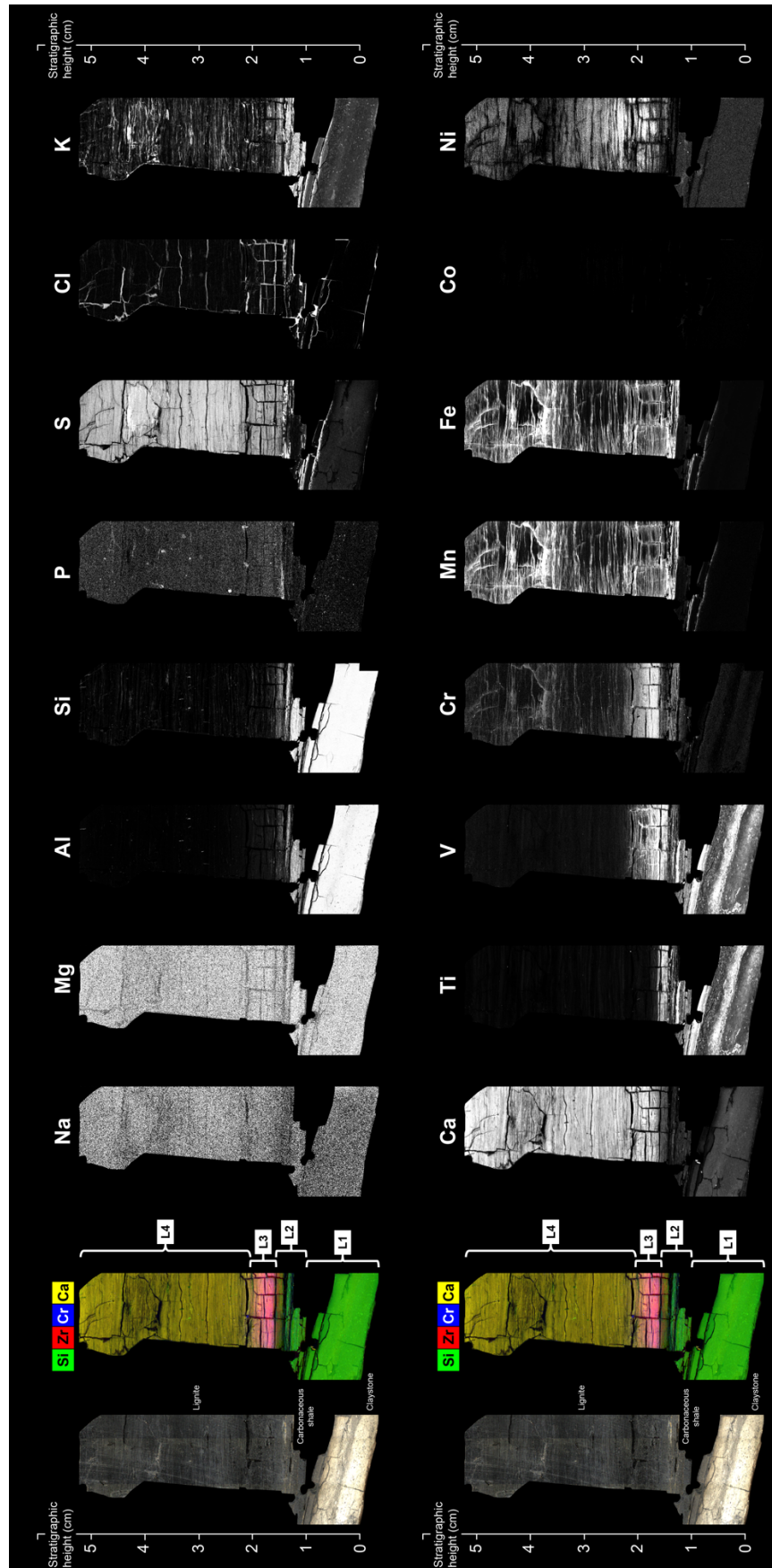
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Table S1
Figures S1 to S3

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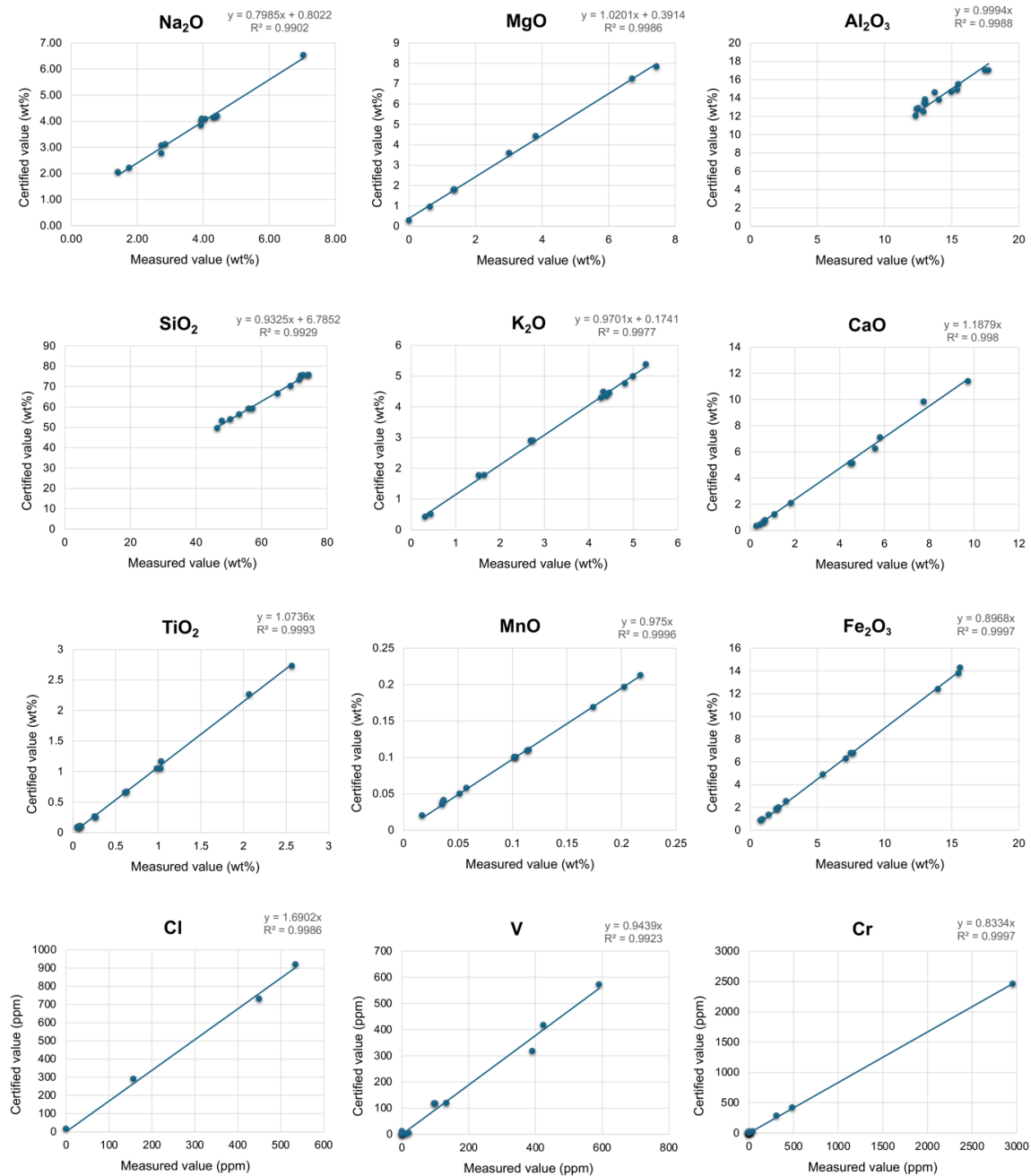
Table S1. Concentration data of major and trace elements of the Starkville-South (svs) K-Pg boundary slab based on the two integrated-area linescan (IAL) profiles (at 25 μ m stratigraphic resolution), following the μ -XRF calibration values displayed in Table 2. Data is available through the Zenodo online repository: <https://zenodo.org/records/14552419>.



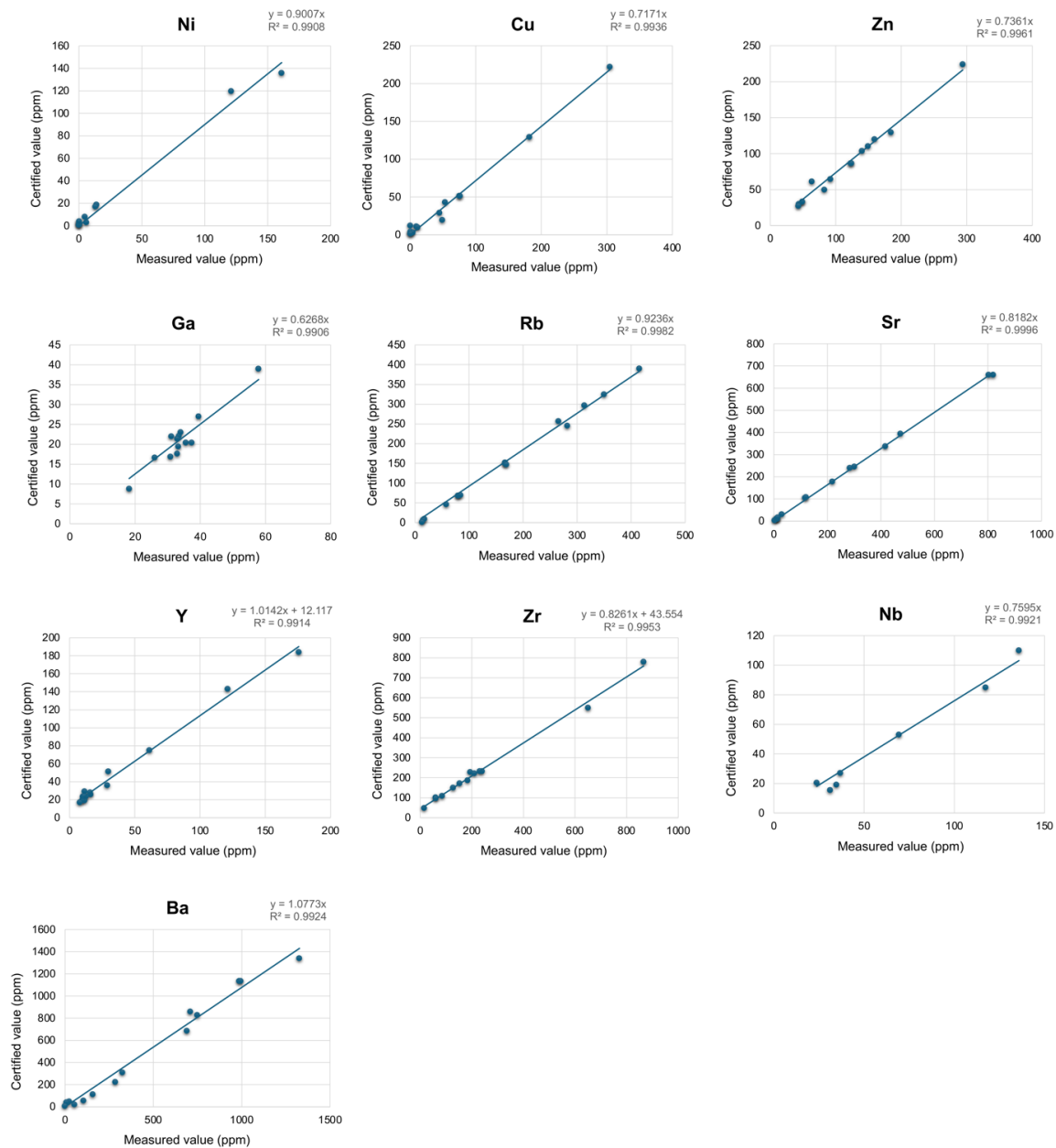


876 **Figure S1.** All μ -XRF single-element heatmaps of the SVS sample, showing the semi-quantitative
877 distribution of 32 elements displayed as a black and white heatmap using deconvoluted settings in
878 the Bruker TORNADO M4 software.

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882 **Figure S2.** Calibration curves of all 22 elements studied in this μ -XRF project, based on extracted mass
883 concentrations from the CRMs listed in Table 1.

884
885 **See below:**

886 **Figure S3.** Selected integrated-area linescan (IAL) profiles of major and trace elements. See Fig. 5 for
887 the two IAL positions of the SVS sample.

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