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3	Magnetic stripes preserved in the continent-ocean transition zone offshore
4	NW Australia
5	
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17	Key Points
18	
19	1) We identify seaward-dipping reflectors across the Cuvier Abyssal Plain and use
20	chemical data to argue it is not oceanic crust as thought.
21	2) Magnetic stripes are thought unique to oceanic crust, but we show they can form in
22	non-oceanic crust in continent-ocean transition zones.
23	3) We move the continent-ocean boundary of NW Australia oceanwards by >500 km.
24	
25	

26 Abstract

27 Magnetic stripes have long been used to define oceanic crust and the onset of submarine sea-28 floor spreading. Identifying magnetic stripes is crucial for plate kinematic reconstructions, 29 and for estimating spreading rates and timing of continental break-up: the most landward 30 magnetic stripes are assumed to mark the oldest oceanic crust. However, continental crust 31 heavily intruded by magma during active sub-aerial rifting in Ethiopia, records magnetic 32 reversals akin to those observed in oceanic crust, implying magnetic stripes may not be diagnostic of seafloor spreading. Recent work on South Atlantic volcanic passive margins has 33 34 identified linear magnetic anomalies within the continent-ocean transition zone (COTZ), which is presently overlain by seaward dipping reflector (SDR) sequences of stacked lavas. 35 36 Perhaps surprisingly, magnetic stripes have rarely been observed in COTZ crust along 37 passive margins elsewhere. We use magnetic, 2D seismic reflection, and geochemical data to 38 examine the lithospheric structure of the Cuvier margin, offshore Northwest Australia, which 39 hosts well-defined, ≤220 km long magnetic stripe anomalies M10N–M5. Although 40 interpreted previously as oceanic crust, we suggest these magnetic stripes occur on a >500km wide, Early Cretaceous COTZ. We suggest the magnetic anomalies record magnetic 41 42 reversal signatures originating from sheeted dykes and gabbroic intrusions within the sub-SDR crust. Our interpretation implies the outer limit of Australia's COTZ may thus be up to 43 44 >500 km further oceanward than previously recognised. Furthermore, this work strengthens a growing consensus that magnetic reversals at rifted margins may not be truly diagnostic of 45 the onset of oceanic crust formation. 46

47

48 Plain Text Summary

Earth's ocean floors are characterised by a barcode-like pattern of magnetic stripes parallel to
mid-ocean ridges. These stripes record reversals in Earth's magnetic North and South field

51 when new rock was forming and moving away from submarine mid-ocean ridges; this is 52 fundamental evidence for continental drift and plate tectonic theory. Magnetic stripes have 53 long been considered unique to oceanic crust, but recent studies suggest they may form on 54 continental crust. We study the Cuvier Abyssal Plain offshore NW Australia, which displays clear magnetic stripes and is thought comprise oceanic crust. We use seismic reflection data, 55 56 which provides ultrasound-like images of Earth's subsurface, to study the crusts structure and show it contains thick sequences of lavas that were erupted in a sub-aerial environment. We 57 58 also reinterpret the chemistry of lavas from the proposed mid-ocean ridge that formed the 59 crust, and demonstrate the magma creating these rocks interacted with continental material. Our evidence suggests the Cuvier Abyssal Plain is not formed of oceanic crust, but rather is a 60 61 hybrid of continental and oceanic crust. Our results mean that: (1) the continent-ocean 62 boundary of NW Australia could be >500 km further offshore; and (2) magnetic stripes may 63 not be diagnostic of oceanic crust. 64 65 **Index Terms** 1020; 1021; 1517; 8105; 8178 66 67 **Keywords** 68 Magnetic stripe; oceanic crust; continental crust; continent-ocean transition zone; seaward-69 70 dipping reflectors; Australia 71 72 73 **1. Introduction** Development of magnetic reversal anomalies (stripes) during oceanic crust formation is 74 fundamental to modern plate tectonic theory (e.g., Vine & Matthews, 1963). Magnetic stripes 75

76 immediately adjacent to passive continental margins are commonly interpreted to mark a 77 basin's oldest oceanic crust, and have historically been used to define the continent-ocean boundary (COB) (e.g., Eagles et al., 2015; Rabinowitz & LaBrecque, 1979; Talwani & 78 79 Eldholm, 1973; Veevers, 1986). Recognition and accurate mapping of COBs are critical to 80 palinspastic and plate kinematic reconstructions (e.g., Eagles et al., 2015). However, 81 improved long-offset seismic reflection and refraction data reveal continental break-up can 82 produce complex crustal domains up to several hundred kilometres wide, i.e. the continentocean transition zone (COTZ), characterised by localized lithospheric thinning and 83 84 occasional mantle exhumation (see Eagles et al., 2015 and references therein). Uncertainties in data resolution and interpretation make ascertaining the nature of COTZ crust challenging, 85 86 meaning it is difficult to uniquely define COBs (e.g., Eagles et al., 2015). Furthermore, linear 87 magnetic stripes have recently been identified: (i) in heavily intruded, transitional continental 88 crust of the onshore Afar Rift, Ethiopia (Bridges et al., 2012), which is expected to form part 89 of a COTZ assuming full lithospheric rupture occurs; (ii) along the COTZ offshore Iberia-90 Newfoundland non-volcanic passive margin, recorded by intrusion of magma into exhumed 91 and serpentinised mantle prior to break-up (Bronner et al., 2011); (iii) within part of the 92 volcanic passive margin COTZ offshore NW Australia (i.e. the Gascoyne margin; Direen et 93 al., 2008); and (iv) in magmatic crust, which is wholly comprised of new igneous material 94 but formed during sub-aerial spreading (i.e. not oceanic crust), offshore South America 95 (Collier et al., 2017; McDermott et al., 2018). Given magnetic stripes can seemingly develop within non-oceanic crust, these observations question whether magnetic stripes can be used to 96 diagnose seafloor spreading and COBs (Rooney et al., 2014). 97 98 The Cuvier margin, offshore NW Australia is an ideal location to explore the

99 development of magnetic stripes during the transition to seafloor spreading: it has well-

100 developed magnetic anomalies, calibrated to the global geomagnetic timescale (Robb et al.,

101 2005), and reflection and refraction seismic data capable of imaging the margin's deep 102 crustal structure. Previous seismic refraction and magnetic studies suggest the Cuvier 103 Abyssal Plain (CAP) is underlain by unambiguous oceanic crust, formed by submarine 104 spreading; this interpretation is based on the presence of magnetic stripes distributed about an 105 inferred submarine spreading centre along the Sonne Ridge (Fig. 1) (e.g., Falvey & Veevers, 106 1974; Larson et al., 1979; MacLeod et al., 2017; Robb et al., 2005). Basalts dredged from the 107 Sonne Ridge are interpreted to have an enriched MORB-like chemistry, supporting the 108 inference that the CAP comprises oceanic crust (Crawford & von Rad, 1994; Dadd et al., 109 2015). Our concurrent analysis of magnetic and 2D seismic reflection data, coupled with an examination of published chemical data, suggest the well-defined magnetic stripes of the 110 111 CAP actually occur within the COTZ. In particular, we recognise multiple packages of up to 112 ~3 km thick seaward-dipping reflector (SDR) sequences, which can extend across several 113 defined linear magnetic anomalies, imaged in seismic reflection data spanning ~200 km of 114 the CAP. Furthermore, reinterpretation of chemical data indicates basalts from the Sonne 115 Ridge, interpreted to be an abandoned submarine spreading centre (e.g., MacLeod et al., 116 2017; Robb et al., 2005), contain a continental signature. Based on these observations and 117 data, we propose the CAP may not be oceanic crust and is instead representative of a COTZ 118 with an outer-limit >500 km oceanwards from where it is currently defined. Our 119 interpretation has important implications for plate reconstructions and heat flow and basin 120 modelling of the NW Australian margin. Furthermore, we argue that magnetic stripes may 121 not be a unique feature of oceanic crust. 122



123

124 Figure 1: (A) Location map of the study area highlighting key tectonic elements, including areas of recognised seaward-dipping reflectors (SDRs) (Holford et al., 2013; Symonds et al., 125 1998) and previously interpreted approximate (approx.) limits of the continent-ocean 126 127 boundary (COB; Eagles et al., 2015) and continent-ocean transition zones (COTZs; Direen et al., 2008; Symonds et al., 1998). Inset: study area location offshore NW Australia. AAP -128 Argo Abyssal Plain, CAP - Cuvier Abyssal Plain, CRFZ - Cape Range Fracture Zone, GAP 129 - Gascoyne Abyssal Plain, GP - Gallah Province, NCB - North Carnarvon Basin, SCB -130 South Carnarvon Basin, Cu – Cuvier margin COTZ, SR – Sonne Ridge, SjR – Sonja Ridge, 131 132 WP – Wallaby Plateau, WS – Wallaby Saddle, WZFZ – Wallaby-Zenith Fracture Zone. Dredge sites 55 (samples 055BS004A and 055BS004B) and 57 (sample 057DR051A) are 133 also shown (Dadd et al., 2015). (B) Total magnetic intensity grid (EMAG2v2), interpreted 134

magnetic chrons (based on Robb et al., 2005). See Supplementary Figure S1 for anuninterpreted version.

137

138 2. Definitions

We define a COTZ as the crustal domain separating unambiguous continental crust and 139 140 unambiguous normal oceanic crust (Eagles et al., 2015). Along volcanic or magma-rich passive margins, COTZ's commonly comprise seismically isotropic, fast (>7 km s⁻¹) crust, 141 142 overlain by SDR sequences of sub-aerially erupted, mafic volcanic products (e.g. lava flows) 143 interbedded with terrestrial sedimentary rocks (e.g., Eldholm et al., 1989; Larsen & Saunders, 1998; Menzies et al., 2002; Symonds et al., 1998). These SDR-bearing domains have been 144 145 further sub-divided into: (i) 'transitional crust', i.e. stretched and heavily intruded continental 146 crust (e.g., Eldholm et al., 1989); (ii) or 'magmatic crust' (also termed 'igneous' crust), which 147 effectively is structurally the same as oceanic crust but instead formed by sub-aerial, rather 148 than submarine, spreading (e.g., Collier et al., 2017; McDermott et al., 2018; Paton et al., 149 2017). Where we can study active continental break-up and development of a possible future COTZ in Ethiopia, the latter phases of rifting (i.e. prior to rupture) involve localisation of 150 151 extension into narrow zones (~20 km wide) dominated by dyke intrusion (e.g., Ebinger & Casey, 2001; Keranen et al., 2004; Mackenzie et al., 2005; Maguire et al., 2006). 152 153 Observations from rifted margins and active rifts therefore suggest the COTZ at volcanic 154 passive margins is marked by a compositional and structural spectrum some way between 155 continental and oceanic end-members (Fig. 2). From their landward limit, at the age which corresponds to onset of COTZ development, we expect the proportion of intruded material to 156 157 continental crust (i.e. transitional crust) to increase oceanwards, until no or very little continental crust remains and the COTZ comprises igneous intrusions and extrusions formed 158 159 in sub-aerial spreading centres (i.e. magmatic crust) (Fig. 2). Given the localisation of dyking

within narrow zones we may expect an oceanwards reduction in the continental signature of
magma chemistry as they become more MORB-like (Fig. 2). Because uncertainties in data
resolution and interpretation mean we cannot identify a distinct boundary between
transitional or magmatic crust (cf. Eagles et al., 2015), we henceforth combine these domains
and refer to them both simply as 'COTZ crust' (Fig. 2).

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Figure 2: Schematic model (not to scale) of a continent-ocean transition zone along a 168 169 volcanic passive margin; for simplicity the lithospheric mantle is not shown. As magma 170 intrudes continental crust, likely as dykes at mid- to upper-crustal levels and larger gabbroic 171 bodies in the lower crust, it becomes 'transitional crust' (e.g., Eldholm et al., 1989). Continued intrusion and dyking leads to localisation of magmatism within narrow zones 172 173 where sub-aerial spreading dominates extension and there is little, if any, continental crust remaining (i.e. magmatic crust) (i.e. 'magmatic crust'; e.g., Collier et al., 2017; Paton et al., 174 175 2017). We categorize transitional and magmatic crust as 'COTZ crust'. Sub-aerial spreading 176 centres may feed extensive lava flows that later, through subsidence, become seaward-177 dipping reflectors (SDRs); SDR subsidence leads to rotation of underlying dykes

(Abdelmalak et al., 2015). A similar rotation of lavas and dykes is observed in oceanic crust(Karson, 2019).

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181 **3. Geological Setting**

The North Carnarvon Basin forms the southernmost part of the NW Australian volcanic 182 183 passive margin, bound by the Argo Abyssal Plain to the north, and the Gascoyne Abyssal Plain and CAP to the west (Fig. 1) (Longley et al., 2002; Stagg et al., 2004). Basin formation 184 185 occurred during multiple phases of Permian-to-Late Jurassic rifting, culminating in Early 186 Cretaceous break-up of the Gascoyne and Cuvier margin rift segments (Longley et al., 2002). We sub-divide the basin into the ~400 km wide Gascoyne and 180 km-wide Cuvier margin 187 188 sectors, separated by the NW-trending Cape Range Fracture Zone (Fig. 1). A 200-250 km 189 wide COTZ adjacent to the Gascoyne margin (i.e. the Gallah Province), comprises 2–5.5 km 190 thick SDRs and high-velocity lower crust formed during magnetic anomalies M10N-M5n 191 (i.e. ~136–131 Ma), with interpreted normal oceanic crust emplaced from chron M3r (~130 192 Ma) onwards (Fig. 1) (Direen et al., 2008). Seismic reflection data reveal SDR sequences 193 along the Cuvier margin, beneath the continental slope, within a previously interpreted 50–70 km wide COTZ (Fig. 1) (Hopper et al., 1992; Symonds et al., 1998). The CAP is partly 194 195 bound to the SW by the Wallaby Plateau (Fig. 1), a large bathymetric high likely comprising thinned continental crust overlain by volcanic and sedimentary rocks (e.g., Colwell et al., 196 197 1994; Daniell et al., 2009; Olierook et al., 2015; Stilwell et al., 2012). The CAP lies \sim 5 km below sea level and comprises a 1–3.3 km thick sedimentary 198 sequence overlying 6–10.5 km thick crystalline crust, inferred to be oceanic in origin due to 199 200 the recognition of chrons M10N-M5 (Fig. 1B) (Falvey & Veevers, 1974; Larson et al., 1979). The ~175 km long Sonne Ridge and ~100 km long Sonja Ridge, which bisect the CAP 201 and extend into the Wallaby Plateau (Fig. 1), are interpreted as probable extinct oceanic 202

203 spreading ridges (e.g., MacLeod et al., 2017; Mihut & Müller, 1998; Robb et al., 2005). The 204 Sonne Ridge has also been interpreted as a pseudofault (i.e. an apparent offset in magnetic 205 stripes formed by ridge jumps; Hey, 1977), based on changes in gravity intensity across it and 206 termination of the Cape Range Fracture Zone directly north of the ridge (Gibbons et al., 207 2012). In their model, Gibbons et al. (2012) define a new spreading centre ~100 km to the SE 208 and parallel to the Sonne Ridge (Fig. 1). However, Gibbons et al. (2012) use the COB of 209 Heine and Müller (2005) to define the termination of the Cape Range Fracture Zone, but this 210 model does not account for the Gallah Province being part of the Gascoyne COTZ and, 211 hence, the required north-westward extension of the fracture zone (Direen et al., 2008; Robb 212 et al., 2005). We therefore favour and use the interpreted magnetic chron configuration of 213 Robb et al. (2005), centred on the Sonne Ridge, which predicts half-spreading rates adjacent 214 to both Gascoyne and Cuvier segments were similar during chrons M10-M5 (~4.5 cm/yr), 215 decreasing to ~ 3 cm/yr by chron M3.

216

217 4. Dataset and methodology

218 4.1. Seismic reflection data

219 We constrain the lithospheric structure of Cuvier margin COTZ using pre-stack time-220 migrated seismic reflection and OBS-derived velocity data from the EW0113 survey. The 221 seismic reflection survey was acquired in 2001 by the R/V Maurice Ewing using an 8445 in³ 222 airgun array and a 6 km streamer length (Tischer, 2006). Due to extreme amplitude contrasts 223 between the shallow and deep sections of the original migrated data, a time-dependent gain filter and root filter were applied to improve amplitude balance and enhance deep reflectivity 224 225 (see supplementary information for details). Depth conversion of the seismic data was performed using a velocity model based on velocities interpreted from OBS data by Tischer 226 227 (2006). The OBS array was co-located with seismic line EW0113-6, which is located ~70 km

along-strike from line EW0113-5 (Fig. 1). Each line is >300 km long and extends up to ~200
km out into the CAP from the currently defined COTZ, nearly reaching the Sonne Ridge
(Fig. 1).

231

4.2. Magnetic data

233 To examine the regional magnetic anomalies, we utilise the EMAG2v2 and EMAG2v3 Earth Magnetic Anomaly Grids (Maus et al., 2009; Meyer et al., 2017). EMAG2v2 is a 2 arc min 234 235 resolution grid derived from marine, airborne, and satellite magnetic data, but uses a priori 236 information to interpolate magnetic anomalies in areas where data gaps are present (Fig. 3) 237 (Maus et al., 2009; Meyer et al., 2017). In contrast, EMAG2v3 uses more data points to 238 derive magnetic anomaly maps but assumes no a priori information (Fig. 3) (Meyer et al., 239 2017). In ocean basins with a relatively poor coverage of magnetic data available, such as the 240 CAP, clear linear magnetic anomalies in EMAG2v2 thus typically appear poorly developed or are absent in EMAG2v3 (Fig. 3) (Meyer et al., 2017). This difference in the presence and 241 242 appearance of linear magnetic anomalies between the EMAG2v2 and EMAG2v3 grids is 243 because knowledge of seafloor spreading processes was incorporated into, and therefore 244 influenced, interpolation during construction of the EMAG2v2 grid (Maus et al., 2009; 245 Meyer et al., 2017). Importantly, the apparent reduction in magnetic stripes observed in 246 EMAG2v3, compared to EMAG2v2 (Fig. 3), does not necessarily mean these features are 247 absent, but rather that the available data is insufficient to unambiguously confirm their presence (Meyer et al., 2017). Comparing the EMAG2v2 and EMAG2v3 grids, coupled with 248 shiptrack magnetic data (Robb et al., 2005), allows us to interrogate the magnetic architecture 249 250 of the CAP (cf. Meyer et al., 2017). In particular, we use EMAG2v2 to interpret possible 251 magnetic chrons, picked on positive peaks (Fig. 1), and attempt to correlate them to the 252 EMAG2v3 grid and shiptrack magnetic data. From these comparisons, we tied interpreted

- 253 magnetic stripes to seismic line EW0113-5 using the synthetic profiles of Robb et al. (2005).
- 254 We use the magnetic polarity reversal sequence and absolute ages of Gradstein and Ogg

255 (2012).

256



Figure 3: Total magnetic intensity grids EMAG2v2 and EMAG2v3 (Maus et al., 2009; Meyer
et al., 2017), compared with shiptrack magnetic data (Robb et al., 2005). Key tectonic
elements also shown (see Fig. 1 for legend).

262

263 4.3. Geochemical data

264 To evaluate whether the Sonne Ridge is an extinct seafloor spreading centre (e.g., Mihut & Müller, 1998; Robb et al., 2005), which implies it consists of oceanic crust with a MORB or 265 266 MORB-like affinity along its length, we examine chemical data for a Sonne Ridge basalt lava 267 sample (i.e. sample 057DR051A; Dadd et al., 2015). The analysed sample was dredged from the south-western section of the Sonne Ridge, where it bounds the south-eastern extent of the 268 Wallaby Plateau (Dadd et al., 2015; Daniell et al., 2009; Olierook et al., 2015). We compare 269 270 the Sonne Ridge sample to two collected from within the Wallaby Plateau (i.e. samples 055BS004A and 055BS004B), towards its south-western margin (Dadd et al., 2015). The two 271

- 272 Wallaby Plateau basalts yield plagioclase ⁴⁰Ar/³⁹Ar plateau ages of 125.12±0.9 Ma and
- 273 123.80±1.0 Ma, whilst two analyses of the Sonne Ridge sample yielded less precise ages of
- 274 120±14 Ma and 123±11 Ma (Olierook et al., 2015).
- 275

5. Observations and interpretation of the Cuvier Margin structure

- 277 5.1. Reflection seismology
- 278 5.1.1. Observations
- 279 The geological structure imaged in line EW0113-6 is very similar to that of EW0113-5 (Fig.
- 4), supporting the use of velocities from EW0113-6 in depth conversion. A prominent,
- continuous, high-amplitude seismic reflection is recognised across the CAP and interpreted
- as the interface between crystalline crust and overlying sedimentary rocks (Fig. 4). The Moho
- was picked at the base of a flat-lying zone of moderate to high amplitude, discontinuous
- seismic reflections, and is ~17 km deep in the SE, shallowing oceanward on EW0113-5 to
- $\leq 14 \text{ km}$ (Fig. 4A). On EW0113-6, the Moho remains relatively flat at depths of ~16–17 km
- 286 (Fig. 4B). Overall, the crystalline crust is ~8–10 km thick (Fig. 4).
- 287



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289

Figure 4: Interpreted and uninterpreted, depth-converted seismic lines (A) EW0113-5 and (B)
EW0113-6 showing crustal structure of the Cuvier margin. A co-located magnetic anomaly
profile showing interpreted magnetic chrons is presented for (A) (after Robb et al., 2005).
Seismic data are displayed with reverse polarity, where a downward increase in acoustic
impedance is represented by a negative (red) reflection event and a downward decrease in
acoustic impedance is a positive (black) reflection event. For details of the velocity model
used, see supplemental information.

297

298 The 1–3 km-thick, uppermost crystalline crustal layer comprises a layered, moderate-299 to high-amplitude seismic facies, which locally contains ≤ 20 km wide, ≤ 3 km thick wedges of coherent, high-amplitude, seaward-diverging reflections (Figs 4 and 5). Where welldeveloped wedges are absent, this crustal layer contains discontinuous, horizontal to gently
seaward-dipping reflections (Fig. 4). Seismic velocities for this layer are 4–5 km/s (Fig. 4;
Tischer, 2006).

304



305

306 Figure 5: Zoomed in view of EW0113-5 highlighting the seismic character of interpreted

307 SDR packages (see Fig. 4A for location).

309 The uppermost crystalline layer is underlain by a low-amplitude, weakly-reflective, 310 1-2.8 km thick, seismic facies layer (i.e. 'opaque crust', Figs 4 and 5) and a deeper, low-311 amplitude, 3.5–6 km-thick, seismic facies layer that locally contains prominent, high-312 amplitude, dipping reflections (i.e. 'reflective crust', Figs 4 and 5). The few reflections that 313 are observed within the opaque crust typically have moderate- to high amplitudes, although 314 low-amplitude also occur, and have variable dips (Figs 4 and 5). Dipping reflections within 315 the reflective crust terminate at the Moho and primarily dip oceanward at $20-30^{\circ}$, although 316 several reflections, particularly on EW0113-6, dip landwards (Figs 4 and 5). Seismic 317 velocities of the opaque and reflective layers are 6.8–7.2 km/s (Fig. 4; Tischer, 2006).

318

319 *5.1.2. Interpretation of crustal structure*

320 Based on their geometry and seismic velocity, we interpret the divergent reflection wedges in 321 the uppermost crystalline crust layer as SDRs, likely comprising interbedded subaerial 322 igneous rocks deposited in lava flows and terrestrial sedimentary strata (Figs 4 and 5) (e.g., 323 McDermott et al., 2018; Menzies et al., 2002; Planke & Eldholm, 1994). We infer flatterlying reflections are gently-dipping lavas. The observed structure and velocities of the opaque 324 325 and reflective crust, defined by transparent seismic facies and discordant high-amplitude 326 reflections respectively (Figs 3 and 4), is consistent with the seismic character of sheeted 327 dyke and lower crustal gabbro intrusions in oceanic crust (e.g., Eittreim et al., 1994; Paton et 328 al., 2017). However, the likelihood that SDR emplacement occurred sub-aerially suggests the 329 CAP could instead comprise COTZ crust. In both the oceanic and COTZ crust scenarios, observed moderate- to high-amplitude reflections within these crustal packages may be 330 331 interpreted as igneous intrusion contacts, primary layering within gabbros, or lower crustal strain markers (e.g., shear zones) (e.g., Abdelmalak et al., 2015; Paton et al., 2017; Phipps-332 333 Morgan & Chen, 1993).

334

335 5.2. Magnetic anomalies

336 Assessment of the EMAG2v2 and ship-track magnetic data reveal 10 km wide, continuous 337 magnetic stripes with lengths of ≤ 220 km across much of the CAP (Figs 1, 3, and 4). Although magnetic anomalies in the EMAG2v3 grid are suppressed relative to EMAG2v2, 338 339 linear anomalies can still be distinguished across the CAP and in the Gallah Province (Fig. 3). Due to the lower resolution of magnetic anomalies in the EMAG2v3 grid, magnetic chrons 340 341 cannot be confidently attributed and we thus rely on EMAG2v2 and shiptrack data to define 342 probable chrons (Fig. 3). Proximal to the Australian continent, long-wavelength magnetic 343 anomalies can only be broadly assigned to chron M10N (Figs 1, 3, and 4) (Robb et al., 2005); 344 within the distal 100 km of seismic line EW0113-5, chrons M10n–M5r (~135.3–131.4 Ma) 345 are clearly defined and have amplitudes of $\leq \pm 100$ nT (Figs 1, 3, and 4). Chrons M8r–M7n coincide with a prominent ≤ 3 km thick, ~ 25 km long, package of seaward-dipping reflectors 346 347 in the uppermost crystalline crust layer (Fig. 4A). 348 Robb et al. (2005) interpret the magnetic anomalies M10N–M6 southeast of the 349 Sonne Ridge as conjugate to a more poorly developed set of anomalies northwest of the ridge 350 (Fig. 1). These poorly defined chrons (i.e. M10N–M6) northwest of the Sonne Ridge 351 terminate abruptly against the Cape Range Fracture Zone, abutting younger magnetic chrons 352 on the Gascoyne Abyssal Plain (Fig. 1). Chron M5n is the first to occur continuously alongstrike across both the Cuvier and Gascoyne margin segments (Fig. 1). 353 354 5.3. Geochemistry of basalts dredged from the Sonne Ridge 355 356 The only basalt collected from the Sonne Ridge displays a relatively flat, Rare Earth Element

357 (REE) pattern (Fig. 6A). Based on this observation, Dadd et al. (2015) interpret the basalt to

358 have a slightly enriched MORB-like source, supporting the inference that the CAP comprises

359 oceanic crust (e.g., Hopper et al., 1992; Larson et al., 1979; Mihut & Müller, 1998). Although 360 a flat REE pattern can be indicative of the shallow melting regime of MORB generation, it 361 does not preclude other settings. By replotting the trace element and radiogenic isotopic 362 compositions of the Sonne Ridge sample, we show the sample has characteristics that could 363 be suggestive of a more continental source (Fig. 5). It must be kept in mind that the Sonne 364 Ridge sample is heavily altered (Dadd et al., 2015), which likely explains the elemental enrichment in Pb, Ba, and Rb, as well as elevated ⁸⁷Sr/⁸⁶Sr. However, the negative Nb 365 366 anomaly, in part defined by neighbouring element Th that is unlikely to be affected by 367 contamination, and unradiogenic ε_{Nd} , cannot be ascribed to alteration (Fig. 6). Instead, the negative Nb anomaly and unradiogenic ε_{Nd} can be interpreted as a chemically evolved, 368 369 continental or sedimentary contribution to the magmas. Finally, the chemical similarity of the 370 Sonne Ridge basalt to two ~124 Ma samples from the Wallaby Plateau (Fig. 6) is notable 371 because this area is interpreted to comprise intruded continental crust, supported by the 372 recovery of pre-breakup sedimentary rocks from the plateau during dredging (Daniell et al., 373 2009; Olierook et al., 2015; Stilwell et al., 2012). Therefore, the chemical characteristics of 374 the Sonne Ridge lava suggest the basalt could originate from either melting of sub-375 continental lithospheric mantle (SCLM) or contamination of a MORB-like magma by 376 assimilation as it ascended through continental crust (cf. Dadd et al., 2015). Importantly, the 377 Sonne Ridge likely represents the rift/spreading axis of the CAP (Robb et al., 2005), such that 378 the basalt sample represents one of the youngest magmas within the system. We thus consider it plausible that older magmas emanating from the ridge system, including the SDRs 379 380 imaged in the seismic reflection data, could have a more pronounced continental signature. 381





Figure 6: (A) Primitive mantle normalized incompatible element diagram comparing the
dredged Sonne Ridge and Wallaby Plateau basalt lava samples with average (ave.)
compositions of MORB variants (Hofmann, 2014), Globally Subducting Sediment (GLOSS)
(Plank & Langmuir, 1998), and continental crust (Rudnick & Fountain, 1995). (B) Plot of
ɛ(Nd) versus ⁸⁷Sr/⁸⁶Sr, illustrating that the Sonne Ridge and Wallaby Plateau samples are
distinct from MORB (based on data collated in Hofmann, 2014).

391 6. Discussion

392 6.1. Nature of the Cuvier margin crust

393 Subsidence profiles for normal oceanic crust predict mid-ocean ridge axes at water depths of

~3 km and subsidence rates of ~90 m/Myr for the first 10 Ma after break-up (e.g., Menard,

1969; Parsons & Sclater, 1977; Stein & Stein, 1992). We identify a previously unrecognised

upper-crustal crystalline basement layer within the CAP that contains well-developed, ≤ 3

397 km-thick, SDRs and extends >200 km seaward from the previously-interpreted COTZ (Figs

- 398 1, 4, and 5). Packages of SDRs observed and drilled on other continental margins indicate the
- comprising lavas extruded sub-aerially and are typically overlain by sub-aerial to shallow-
- 400 marine strata deposited during subsequent margin subsidence (e.g., Larsen et al., 1994).
- 401 Although the observed SDRs across the CAP are not drilled, linking their origin to sub-aerial

402 volcanism conforms with observations from DSDP site 263 (Fig. 1), which intersected Early 403 Cretaceous (≤130.8 Ma), shallow-marine sedimentary rocks immediately deposited 404 immediately on top of the crystalline basement (Veevers & Johnstone, 1974; Willcox & 405 Exon, 1976). The presence of SDRs and overlying shallow-marine sedimentary rocks 406 suggests sub-aerial to shallow-marine conditions were prevalent across the Cuvier margin 407 during formation of the CAP, which is inconsistent with subsidence profiles for normal 408 oceanic crust (e.g., Menard, 1969; Parsons & Sclater, 1977; Stein & Stein, 1992). One 409 possible explanation for the observed sub-aerial to shallow-marine conditions is that the CAP 410 actually comprises COTZ crust rather than oceanic crust, having formed through sub-aerial 411 rifting and spreading (Fig. 7D) (cf. Collier et al., 2017; McDermott et al., 2018; Paton et al., 412 2017). The chemistry of a basalt sample from the Sonne Ridge, particularly its Nd isotopic 413 composition and refractory trace element abundances (Fig. 6), allows for a contribution of 414 sub-continental lithospheric mantle (SCLM) or interaction with continental crust, supporting 415 the hypothesis that full continental rupture may not have occurred even during the final 416 phases of CAP spreading.

417 We interpret the previously recognised SDRs beneath the continental slope on the 418 Cuvier margin (Hopper et al., 1992; Symonds et al., 1998) as marking the transition from 419 mechanical- to magma-dominated rifting and the landward limit of the COTZ, which likely 420 comprises heavily intruded continental crust (i.e. 'transitional' crust) (Fig. 7A). A progressive 421 localisation in dyking as rifting continued led to the eventual generation of the sub-aerial 422 Sonne Ridge, which was perhaps akin to 20 km wide Quaternary magmatic segments in the Ethiopian Rift (e.g., Ebinger & Casey, 2001; Hayward & Ebinger, 1996; Kendall et al., 423 424 2005), where the crust likely comprises >50% new mafic material (Daniels et al., 2014). Our proposed crustal structure for the CAP therefore involves a gradual north-westwards change 425 426 from continental crust into COTZ crust, which we suggest is likely characterised by







Figure 7: (A) Map showing the potential limits of the COTZ based on interpreting the Cap
and Gallah Province as transitional and/or magmatic crust. (B-D) Schematic maps showing
the development of COTZ crust and the onset of oceanic crust accretion adjacent to the
Gascoyne and Cuvier margins, during formation of chrons (B) M10, (C) M6 and (D) M3r.
See Figure 1 for chron ages. Location of present day coastline shown for reference.

443

444 Linear magnetic stripes recorded across the CAP reveal an apparent conjugate set of chrons M10N-M6 centred on the Sonne Ridge (Robb et al., 2005), suggesting our inferred 445 446 COTZ crust may extend out to chron M5n, broadly coincident with the north-western limit of 447 the Gallah Province on the Gascovne margin (Figs 1 and 7B). Magnetic reversals in the 448 Gallah Province are similar, albeit less developed, to those in the CAP (Direen et al., 2008; 449 Robb et al., 2005). Direen et al. (2008) argue these magnetic anomalies resulted from the 450 stacking of wedges of positively magnetised basalts (including SDRs) and variably 451 magnetised volcaniclastic conglomerates atop transitional magmatic and modified continental 452 crust. We thus suggest that magnetic anomalies along the Gascovne margin sector are also likely related to geomagnetic polarity reversals recorded by COTZ crust underlying the 453 454 SDRs. Chemical signatures indicating continental lithosphere contamination, similar to those 455 identified in the CAP, are recorded in basalts from the Gascoyne margin (Direen et al., 2008). 456

457 6.1.1. Tectonic-magmatic evolution of the Cuvier and Gascoyne margins

During chrons M10N-M5r (i.e. ~136–131 Ma), we suggest stretching and intrusion of
continental crust adjacent to both the Gascoyne and Cuvier margins produced COTZ crust,
likely comprising transitional crust, which evolved into magmatic crust as intrusion became
localised and the proportion of remnant continental crust decreased (Figs 7B-D). Robb et al.
(2005) proposed that, during chron M5r-M5n (~131 Ma), a ridge jump in the Cuvier sector

463 led to abandonment of the Sonne Ridge, with a short-lived period of spreading about the 464 Sonja Ridge preceding development of a new ridge ~250 km northwest of the Sonne Ridge (Fig. 7D). We consider that this jump linked the Cuvier and Gascoyne ridges, promoting the 465 466 onset of true seafloor spreading from chron M3r (~130 Ma) onwards in both sectors (Fig. 467 7D). This initiation of oceanic crust formation transferred the conjugate region of COTZ 468 crust recording M10N-M6 anomalies to the Australian plate (Fig. 7D). With the recognition 469 of a >500 km wide zone of COTZ crust, our interpretation implies that the outer limit of NW 470 Australia's COTZ in the Cuvier margin sector may be positioned substantially further 471 oceanward (>500 km) than previously interpreted (Fig. 7A). Extension of the COTZ across the CAP has implications for the timing and kinematics of plate reconstructions of the NW 472 473 Australian margin, with the onset of true seafloor spreading potentially ~3 Myr later than 474 suggested by previous studies (e.g., Robb et al., 2005).

475

476 **6.2. Development of magnetic stripes during breakup**

477 Stacked SDR wedges on the CAP overlie COTZ crust and span several chrons (e.g. M8n-M7r, Fig. 3), indicating the magnetic stripes likely record magnetic reversal signatures 478 479 originating from sub-SDR rocks. Recent forward magnetic modelling of conjugate, ship-track 480 magnetic profiles by Collier et al. (2017) suggest magnetic signals over SDRs arise from a 481 combination of stacked and rotated lavas, producing a long-wavelength positive anomaly that 482 can sometimes mask reversals, and linear magnetic anomalies caused by dyke intrusion in the underlying crust. This is consistent with studies of onshore incipient spreading centres (e.g. 483 Ethiopia), where magnetic anomalies likely originate from axial intrusion by dykes in heavily 484 485 intruded upper continental crust (Bridges et al., 2012). We suggest that the well-developed linear anomalies over the CAP may therefore originate from intrusions in the COTZ crust 486 487 underlying the SDRs.

488 The lack of broad positive magnetic anomalies over the transitional to magmatic crust 489 of the CAP may be due to the relatively low SDR thickness (≤ 3 km) compared to the South 490 Atlantic margins (<8 km thick; Collier et al., 2017). The less-clearly developed and higher 491 amplitude magnetic reversals on the Gascoyne Abyssal Plain may relate to interference from the greater SDR thicknesses (\leq 5.5 km) relative to the Cuvier margin (Direen et al., 2008). We 492 493 suggest that SDR thickness and, thereby, preservation of magnetic anomalies within 494 transitional or magmatic crust can partly be attributed to extension rate. For example, the 495 extension rate during SDR eruption offshore NW Australia (~4.5 cm/yr half rate; Robb et al., 496 2005) is substantially faster than the inferred extension rates for the South Atlantic during 497 magmatic crust formation (~1.1 cm/yr half-rate; Paton et al., 2017). Slower extension rates 498 (e.g. South Atlantic) likely promote stacking of lava flows (Eagles et al., 2015), leading to 499 interference and development of the long-wavelength positive magnetic anomalies (e.g., 500 Moulin et al., 2010). Extension rate may also influence magnetic anomaly development by 501 affecting the width of magnetic stripes: reversal anomalies will be narrowest at slow 502 spreading ridges (Vine, 1966). The narrower anomalies, combined with the greater potential 503 for vertical stacking of lavas, will tend to suppress magnetic anomaly preservation.

504

505 **6.** Conclusions

We present evidence for a previously unrecognised, >500 km-wide region of continent-ocean transition zone (COTZ) crust adjacent to the Northwest Australian continental margin. We suggest the COTZ crust likely comprises a spectrum of 'transitional', i.e. heavily intruded continental crust, and 'magmatic' crust formed by addition of igneous material at a sub-aerial ridge. This COTZ crust records well-developed linear magnetic reversal anomalies. Seawarddipping reflectors (SDRs) overlie the 5–8.5 km thick COTZ crust, which hosts intrusions that are the primary source of the observed magnetic signal. The recognition of this region of

513 COTZ crust implies that the outer limit of Australia's continent-ocean transition zone may be
514 >500 km further oceanward than previously interpreted. That COTZ crust on volcanic
515 margins can record the development of clear seafloor spreading-type magnetic stripes,
516 implies magnetic stripes alone are not a reliable proxy for the onset of seafloor spreading and
517 the extent of oceanic crust.

518

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527

528 Figure Captions

529 Figure 1: (A) Location map of the study area highlighting key tectonic elements, including

areas of recognised seaward-dipping reflectors (SDRs) (Holford et al., 2013; Symonds et al.,

531 1998) and previously interpreted approximate (approx.) limits of the continent-ocean

boundary (COB; Eagles et al., 2015) and continent-ocean transition zones (COTZs; Direen et

533 al., 2008; Symonds et al., 1998). Inset: study area location offshore NW Australia. AAP –

534 Argo Abyssal Plain, CAP – Cuvier Abyssal Plain, CRFZ – Cape Range Fracture Zone, GAP

535 – Gascoyne Abyssal Plain, GP – Gallah Province, NCB – North Carnarvon Basin, SCB –

536 South Carnarvon Basin, Cu – Cuvier margin COTZ, SR – Sonne Ridge, SjR – Sonja Ridge,

537 WP – Wallaby Plateau, WS – Wallaby Saddle, WZFZ – Wallaby-Zenith Fracture Zone.

Dredge sites 55 (samples 055BS004A and 055BS004B) and 57 (sample 057DR051A) are
also shown (Dadd et al., 2015). (B) Total magnetic intensity grid (EMAG2v2), interpreted
magnetic chrons (based on Robb et al., 2005). See Supplementary Figure S1 for an
uninterpreted version.

542

543 Figure 2: Schematic model (not to scale) of a continent-ocean transition zone along a volcanic passive margin; for simplicity the lithospheric mantle is not shown. As magma 544 545 intrudes continental crust, likely as dykes at mid- to upper-crustal levels and larger gabbroic 546 bodies in the lower crust, it becomes 'transitional crust' (e.g., Eldholm et al., 1989). Continued intrusion and dyking leads to localisation of magmatism within narrow zones 547 548 where sub-aerial spreading dominates extension and there is little, if any, continental crust 549 remaining (i.e. magmatic crust) (i.e. 'magmatic crust'; e.g., Collier et al., 2017; Paton et al., 550 2017). We categorize transitional and magmatic crust as 'COTZ crust'. Sub-aerial spreading 551 centres may feed extensive lava flows that later, through subsidence, become seaward-552 dipping reflectors (SDRs); SDR subsidence leads to rotation of underlying dykes 553 (Abdelmalak et al., 2015). A similar rotation of lavas and dykes is observed in oceanic crust 554 (Karson, 2019). 555

Figure 3: Total magnetic intensity grids EMAG2v2 and EMAG2v3 (Maus et al., 2009; Meyer
et al., 2017), compared with shiptrack magnetic data (Robb et al., 2005). Key tectonic
elements also shown (see Fig. 1 for legend).

559

560 Figure 4: Interpreted and uninterpreted, depth-converted seismic lines (A) EW0113-5 and (B)

561 EW0113-6 showing crustal structure of the Cuvier margin. A co-located magnetic anomaly

562 profile showing interpreted magnetic chrons is presented for (A) (after Robb et al., 2005).

acoustic impedance is a positive (black) reflection event. For details of the velocity model 565 566 used, see supplemental information. 567 Figure 5: Zoomed in view of EW0113-5 highlighting the seismic character of interpreted 568 569 SDR packages (see Fig. 4A for location). 570 571 Figure 6: (A) Primitive mantle normalized incompatible element diagram comparing the dredged Sonne Ridge and Wallaby Plateau basalt lava samples with average (ave.) 572 573 compositions of MORB variants (Hofmann, 2014), Globally Subducting Sediment (GLOSS) 574 (Plank & Langmuir, 1998), and continental crust (Rudnick & Fountain, 1995). (B) Plot of ϵ (Nd) versus ⁸⁷Sr/⁸⁶Sr, illustrating that the Sonne Ridge and Wallaby Plateau samples are 575 576 distinct from MORB (based on data collated in Hofmann, 2014). 577 Figure 7: (A) Map showing the potential limits of the COTZ based on interpreting the Cap 578 579 and Gallah Province as transitional and/or magmatic crust. (B-D) Schematic maps showing the development of COTZ crust and the onset of oceanic crust accretion adjacent to the 580 581 Gascoyne and Cuvier margins, during formation of chrons (B) M10, (C) M6 and (D) M3r. 582 See Figure 1 for chron ages. Location of present day coastline shown for reference. 583 References 584 585 Abdelmalak, M. M., Andersen, T. B., Planke, S., Faleide, J. I., Corfu, F., Tegner, C., et al. (2015). The ocean-continent transition in the mid-Norwegian margin: Insight from 586 seismic data and an onshore Caledonian field analogue. Geology, 43(11), 1011-1014. 587

Seismic data are displayed with reverse polarity, where a downward increase in acoustic

impedance is represented by a negative (red) reflection event and a downward decrease in

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