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1 **Are magnetic stripes on the Cuvier Abyssal Plain (offshore NW Australia)**
2 **diagnostic of oceanic crust?**

3

4 Running title: Origin of the Cuvier Abyssal Plain

5

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15

16 **Abstract**

17 Magnetic stripes have long been used to define the presence and age of oceanic crust.

18 However, continental crust heavily intruded by magma can record magnetic reversals akin to
19 those observed in oceanic crust. We re-evaluate the nature of the Cuvier Abyssal Plain

20 (CAP), offshore NW Australia, which hosts magnetic stripes and has previously been defined

21 as oceanic crust. We use magnetic, 2D seismic reflection, and geochemical data to test

22 whether the CAP structure and composition is consistent with unambiguous oceanic crust.

23 We show chemical data from a basalt within the CAP, previously described as displaying an

24 enriched MORB-like signature, actually contains evidence of contamination by continental

25 material. We also recognise seaward-dipping reflector (SDR) sequences across the CAP.

26 Borehole data from overlying sedimentary rocks suggests these SDRs were emplaced in a
27 shallow-water (<200 m depths) or sub-aerial environment. Our results indicate the CAP may
28 not be unambiguous oceanic crust. Instead, we suggest the CAP could comprise a spectrum
29 of heavily intruded continental crust (akin to present-day Ethiopia) through to fully oceanic
30 crust, recording the evolution from continental rifting to progressively magma-dominated,
31 sub-aerial to shallow-water extension. Our work supports suggestions that magnetic reversals
32 may not be truly diagnostic of oceanic crust.

33

34 Supplementary material: Enlarged and uninterpreted versions of the magnetic data and
35 seismic reflection lines are available at.

36

37 Development of magnetic reversal anomalies (stripes) during oceanic crust formation is
38 fundamental to modern plate tectonic theory (e.g., Vine & Matthews 1963). Where magnetic
39 stripes occur adjacent to passive continental margins, they are commonly interpreted to mark
40 a basin's oldest, unambiguous oceanic crust. Such stripes have also historically been used to
41 define the continent-ocean boundary (e.g., Talwani & Eldholm 1973; Rabinowitz &
42 LaBrecque 1979; Veevers 1986; Eagles et al. 2015): here we define unambiguous oceanic
43 crust as comprising layers of pillow basalts, sheeted dikes, and gabbro formed during deep-
44 marine (≥ 2 km water depths) seafloor spreading (e.g., McDermott et al. 2018). However,
45 continental break-up can produce broad, complex zones that have structural and geochemical
46 traits lying somewhere between unambiguous continental crust and unambiguous oceanic
47 crust (e.g., Skogseid et al. 1992; Symonds et al. 1998; Planke et al. 2000; Skogseid et al.
48 2000; Direen et al. 2007; Bastow & Keir 2011). The crustal affinity of such break-up zones is
49 accordingly difficult to distinguish (e.g., Eagles et al. 2015). Where these so-called continent-
50 ocean transition zones (COTZs) occur, as opposed to an abrupt continent-ocean boundary,

51 magnetic stripes within the adjacent unambiguous oceanic crust are expected to mark the
52 outer limit of the COTZ (e.g., Pickup et al. 1996; Direen et al. 2007; Eagles et al. 2015; Paton
53 et al. 2017; Peron-Pinvidic et al. 2019). Using magnetic stripes to recognise and accurately
54 map the extent and age of oceanic crust has proved critical to palinspastic and plate kinematic
55 reconstructions (e.g., Heine & Müller 2005; Eagles et al. 2015; Causer et al. 2020). Yet linear
56 magnetic stripes akin to those hosted by unambiguous oceanic crust have recently been
57 identified along: (i) the onshore Afar Rift, Ethiopia in heavily intruded continental crust
58 (Bridges et al. 2012), an area expected to eventually become a COTZ assuming full seafloor
59 spreading develops; (ii) magma-poor passive margin COTZs offshore Iberia and
60 Newfoundland, where magnetic anomalies are recorded by magmatic intrusion into exhumed
61 and serpentinised mantle prior to break-up (Bronner et al. 2011); (iii) part of the magma-rich
62 passive margin COTZ offshore NW Australia (i.e. the Gascoyne margin; Direen et al. 2008);
63 and (iv) offshore South America where the margin comprises ‘magmatic crust’ wholly
64 comprised of new igneous material, which differs from normal oceanic crust in that it formed
65 via sub-aerial and/or shallow-water extension, not deep-marine spreading (Collier et al. 2017;
66 McDermott et al. 2018). These observations suggest magnetic stripes could develop within
67 non-oceanic crust, which questions whether they can be confidently used to diagnose seafloor
68 spreading, detect continent-ocean boundaries, and be used as hard constraints on palinspastic
69 and plate kinematic reconstructions (Rooney et al. 2014).

70 Given recent studies have shown magnetic stripes may not be diagnostic of
71 unambiguous oceanic crust formed during deep-marine seafloor spreading (e.g., Direen et al.
72 2008; Bridges et al. 2012; Collier et al. 2017; McDermott et al. 2018), it is worth re-
73 evaluating the nature of previously defined oceanic crust adjacent to passive margins. Here,
74 we test the origin of the Cuvier Abyssal Plain (CAP), offshore NW Australia through an
75 integrated analysis of 2D seismic reflection data from the CAP, coupled with a re-

76 examination of published chemical data (Fig. 1). The CAP hosts well-developed magnetic
77 anomalies distributed about an inferred spreading centre along the Sonne Ridge (Fig. 1B)
78 (e.g., Robb et al. 2005). These anomalies, coupled with observations from seismic refraction
79 data, have previously been used to suggest that the CAP comprises unambiguous oceanic
80 crust (Fig. 1) (e.g., Falvey & Veevers 1974; Larson et al. 1979; Robb et al. 2005; MacLeod et
81 al. 2017). Basalts dredged from the Sonne Ridge have been interpreted to reflect an enriched
82 MORB-like chemistry, supporting the inference that the CAP crust is oceanic (Crawford &
83 von Rad 1994; Dadd et al. 2015). The CAP has thus previously been defined as oceanic crust
84 in all regional and global palinspastic and palaeogeographic reconstructions involving NW
85 Australia (e.g., Heine & Müller, 2005; Gibbons et al., 2012).

86 Within our seismic reflection data we recognise multiple packages of up to ~3 km
87 thick seaward-dipping reflector (SDR) sequences, which cumulatively extend at least ~300
88 km oceanwards from the continental shelf. These SDR packages are likely dominated by
89 lavas and occasionally span several well-defined, broadly linear magnetic anomalies.
90 Lithological and biostratigraphic data from the Deep Sea Drilling Project (DSDP) Site 263
91 borehole indicate sedimentary strata above these SDRs were deposited in neritic
92 environments (<200 m water depths), implying lava emplacement was sub-aerial or shallow-
93 water. Our reinterpretation of chemical data also indicates basalts from the Sonne Ridge
94 contain a continental signature (e.g., Robb et al. 2005; MacLeod et al. 2017). We suggest that
95 the CAP may not be unambiguous oceanic crust and could instead comprise a spectrum of
96 crustal types, ranging from heavily intruded continental crust to fully magmatic crust
97 generated during sub-aerial or shallow-water extension. We consider the CAP may define a
98 COTZ with an outer-limit >500 km oceanwards from the previously defined passive margin
99 boundary. Our interpretation has important implications for palinspastic and plate kinematic
100 reconstructions involving the NW Australian margin, as well as for heat flow and basin

101 modelling studies. More generally, our study supports the arising hypothesis that magnetic
102 stripes may not be a unique feature of oceanic crust.

103

104 **Continent-Ocean Transition Zones**

105 Where we can study active, magma-rich continental break-up onshore in Ethiopia, the
106 processes driving development of a possible future COTZ during the latter stage of rifting
107 (i.e. prior to rupture) involve the localisation of extension into narrow (~20 km wide) zones
108 that progressively become dominated by dyke intrusion (e.g., Ebinger & Casey 2001;
109 Keranen et al. 2004; Mackenzie et al. 2005; Maguire et al. 2006; Daniels et al. 2014). In these
110 sub-aerial and so-called ‘magmatic segments’, $\geq 50\%$ of the crust may comprise new,
111 intruded mafic material (Daniels et al., 2014). Along volcanic or magma-rich passive
112 margins, COTZs commonly contain seismically isotropic, fast ($>7 \text{ km s}^{-1}$) crust, overlain by
113 SDR sequences of mafic volcanic products (e.g. lava flows) interbedded with sedimentary
114 rocks emplaced within sub-aerial or shallow-water environments (e.g., Eldholm et al. 1989;
115 Larsen & Saunders 1998; Symonds et al. 1998; Menzies et al. 2002). The COTZ crust
116 underlying these SDR-bearing domains comprises a spectrum of stretched and heavily
117 intruded continental crust (e.g., Eldholm et al. 1989) to ‘magmatic crust’ (also termed
118 ‘igneous’ crust), which record processes similar to those active within Ethiopian magmatic
119 segments (e.g., Collier et al. 2017; Paton et al. 2017; McDermott et al. 2018). Magmatic crust
120 is structurally similar to oceanic crust but instead formed during sub-aerial or shallow-water
121 extension along a magmatic segment, rather than deep-marine spreading at a mid-ocean ridge
122 (e.g., Collier et al. 2017; Paton et al. 2017; McDermott et al. 2018); i.e. although the
123 formation of magmatic crust requires the prior break-up of continental crust, we note that this
124 event may not coincide with full continental lithospheric rupture, which necessarily involves
125 the mantle lithosphere.

126 Observations from rifted margins and active rifts suggest that COTZs at magma-rich
127 passive margins are marked by a compositional and structural spectrum, bounded by
128 unambiguous continental and oceanic crust end-members (Fig. 2). From the landward limit of
129 COTZs, we expect the proportion of magma intruded into continental crust to increase
130 oceanwards (Fig. 2). As dyking localises, eventually no continental crust will remain (i.e.
131 break-up of continental crust), and the COTZ will solely comprise igneous intrusions and
132 extrusions emplaced along magmatic segments during sub-aerial or shallow-water extension
133 (i.e. magmatic crust; Fig. 2) (e.g., Paton et al. 2017; McDermott et al. 2018). Because of
134 uncertainties in data resolution and interpretation mean, it is often difficult to constrain the
135 progression from intrusion of continental crust to the onset of magmatic crust emplacement.
136 We therefore combine these domains and refer to them both simply as a COTZ (Fig. 2).
137 Diminishment of the buoyant support maintaining these dense, sub-aerial or shallow-water
138 magmatic segments will promote their subsidence (e.g., Corti et al. 2015; McDermott et al.
139 2018). As these magmatic segments subside to water depths of ≥ 2 km, plate-spreading drives
140 the generation of unambiguous oceanic crust, comprising layers of pillow basalts, sheeted
141 dykes, gabbro, and oceanic mantle lithosphere (e.g., McDermott et al. 2018); i.e. full
142 continental lithospheric rupture has occurred by this point. Across COTZs and into
143 unambiguous oceanic crust, we may therefore expect an oceanwards reduction in the
144 continental signature of magma chemistry as they become more MORB-like (Fig. 2).

145

146 **Geological Setting**

147 *Crustal Structure and Age*

148 The North Carnarvon Basin and South Carnarvon Basin form part of the NW Australian
149 magma-rich passive margin, bound by the Argo Abyssal Plain to the north, and the Gascoyne
150 Abyssal Plain and Cuvier Abyssal Plain (CAP) to the west (Fig. 1A) (Longley et al. 2002;

151 Stagg et al. 2004). Basin formation occurred during multiple phases of Permian-to-Late
152 Jurassic rifting, culminating in Early Cretaceous break-up of the Gascoyne and Cuvier
153 margin rift segments (Fig. 3A) (Longley et al. 2002). We sub-divide the study area into the
154 ~400 km wide Gascoyne and 180 km-wide Cuvier margin sectors, separated by the NW-
155 trending Cape Range Fracture Zone (Fig. 1A). The CAP is bound to the SW by the Wallaby
156 Plateau and Wallaby Saddle (Fig. 1A).

157

158 *Previously interpreted continent-ocean transition zones (COTZs)*

159 A 200–250 km wide COTZ lies northwest of the continental Exmouth Plateau (i.e. the Gallah
160 Province; Fig. 1A). This COTZ comprises 2–5.5 km thick SDRs and high-velocity lower
161 crust (Direen et al. 2008). The Gallah Province COTZ preserves magnetic stripes formed
162 during chrons M10N–M5n (~136–131 Ma; Valanginian-to-Hauterivian), with interpreted
163 oceanic crust (Gascoyne Abyssal Plain) emplaced onwards from chron M3r (~130 Ma,
164 Hauterivian; Figs 1B and 3A) (Direen et al. 2008). Along the Cuvier Margin, beneath the
165 modern continental slope, seismically imaged SDR sequences are interpreted to mark a 50–
166 70 km wide COTZ (e.g., Figs 1A and C) (Hopper et al. 1992; Symonds et al. 1998).

167

168 *Cuvier Abyssal Plain*

169 Adjacent to the Cuvier Margin, the CAP lies ~5 km below sea level and comprises a ~1–3.3
170 km thick sedimentary sequence overlying ~6–10.5 km thick crystalline crust (e.g., Fig. 1C)
171 (Hopper et al., 1992). Based on recognition of magnetic chrons M10N–M5 within the CAP, it
172 has been interpreted as oceanic crust emplaced initially at ~136 Ma (i.e. Valanginian; Figs 1B
173 and 3A) (Falvey & Veevers 1974; Larson et al. 1979). The presence of chrons M10N–M5
174 within the Gallah Province COTZ (Direen et al. 2008) and the inference that seafloor
175 spreading began in the CAP at ~136 Ma, suggest that continental rupture and seafloor

176 spreading adjacent to the Cuvier Margin occurred ~5–6 Myr before the Gascoyne Margin
177 (Fig. 3A) (Falvey & Veevers 1974; Larson et al. 1979). Critically, the occurrence of magnetic
178 stripes in the non-oceanic (i.e. COTZ) crust of the Gallah Province questions whether the
179 presence magnetic stripes in the CAP should be used to classify it as oceanic crust (e.g.,
180 Direen et al. 2008; Bridges et al. 2012; Collier et al. 2017; McDermott et al. 2018).

181

182 *Origin of the Sonne and Sonja ridges*

183 Under the assumption that the CAP comprises oceanic crust, the ~175 km long Sonne Ridge
184 and ~100 km long Sonja Ridge, which extend into the Wallaby Plateau (Fig. 1A), have been
185 interpreted as probable extinct oceanic spreading ridges (e.g., Mihut & Müller 1998; Robb et
186 al. 2005; MacLeod et al. 2017); in this model, spreading is interpreted to have jumped from
187 the Sonne Ridge to the Sonja Ridge at ~131.7 Ma (Hauterivian). Geochemical analyses of a
188 basalt dredged from the Sonne Ridge along its extension into the Wallaby Plateau, suggest it
189 has a slightly enriched MORB-like signature, supporting the inference that the Sonne Ridge
190 is an oceanic spreading centre (Dadd et al. 2015). An alternative interpretation forwarded for
191 the Sonne Ridge is that it represents a ‘pseudofault’ (i.e. an apparent offset in magnetic
192 stripes formed by ridge jumps; Hey 1977); this interpretation is based on changes in gravity
193 intensity across the structure and the possible termination of the Cape Range Fracture Zone
194 directly north of the ridge (Gibbons et al. 2012). In their model, Gibbons et al. (2012) define
195 a different oceanic spreading centre ~100 km to the SE and parallel to the Sonne Ridge (red
196 dashed line in Figs 1A and B). However, we note that Gibbons et al. (2012) use the COB of
197 Heine & Müller (2005) to define the termination of the Cape Range Fracture Zone. This COB
198 interpretation does not consider the Gallah Province is a COTZ and, hence, does not include
199 the required north-westward extension of the fracture zone beyond the seaward limit of the
200 Exmouth Plateau (Heine & Müller 2005; Robb et al. 2005; Direen et al. 2008). Overall, given

201 the repetition of magnetic stripe sequences either side of the Sonne Ridge, and to a lesser
202 extent the Sonja Ridge, we favour their interpretation as spreading ridges and thus use the
203 interpreted magnetic chron configuration of Robb et al. (2005). This magnetic chron
204 configuration is centred on the Sonne Ridge and predicts half-spreading rates adjacent to both
205 Gascoyne and Cuvier margins were similar during chrons M10–M5 (~4.5 cm/yr), decreasing
206 to ~3 cm/yr by chron M3 (Robb et al. 2005).

207

208 *The Wallaby Plateau and Wallaby Saddle*

209 The Wallaby Plateau is a large bathymetric high (Fig. 1A), containing up to ~7.5 km thick
210 sequences of volcanic and sedimentary rocks, which are typically expressed in seismic
211 reflection data as packages of diverging and dipping reflections that appear similar to SDRs
212 (e.g., Colwell et al. 1994; Daniell et al. 2009; Stilwell et al. 2012; Olierook et al. 2015).

213 Interpretation of seismic reflection and magnetic data, coupled with chemical,
214 geochronological, and biostratigraphic analyses of dredge samples, suggests the Wallaby
215 Plateau probably comprises ~124 Myr old, continental flood basalts and interbedded
216 sedimentary strata emplaced on a fragment of extended continental crust (see Olierook et al.
217 2015 and references therein). Between the Wallaby Plateau and the Australian continent is
218 the Wallaby Saddle, a bathymetric low containing SDRs but no magnetic stripes, interpreted
219 by Symonds et al. (1998) to comprise ‘transitional’ crust (Figs 1A and B). The Wallaby
220 Plateau and Wallaby Saddle seemingly preserve a range of crustal types typical of a COTZ,
221 but not unambiguous continental crust or unambiguous oceanic crust. Similarities between
222 their structure and that of the CAP may therefore imply the latter is not truly oceanic crust.

223

224 *Sedimentary Cover on the Cuvier Abyssal Plain*

225 The top of the crystalline basement within the CAP corresponds to a high-amplitude
226 reflection in seismic data, which is overlain by a ~1–3.3 km thick, sedimentary succession
227 broadly comprising sub-horizontal reflections (e.g., Fig. 1C) (e.g., Veevers & Johnstone
228 1974; Hopper et al. 1992). Biostratigraphic and lithological data for the sedimentary cover
229 are available from the DSDP Site 263 borehole, which was drilled in 1972 and terminates
230 ~100–200 m above the basement (Figs 1A and 3B) (e.g., Bolli 1974; Scheibnerová 1974;
231 Wiseman & Williams 1974; Holbourn & Kaminski 1995). From a depth of 200 m below
232 seabed to the base of the borehole (746 m), the sedimentary strata intersected at DSDP Site
233 263 comprise black claystones; between depths of ~475–746 m these claystones are silty and
234 contain abundant kaolinite (Fig. 3B) (Robinson et al. 1974; Compton et al. 1992). Within the
235 lowermost ~20 m of the drilled sequence, and in thin units at around ~500 m depth, the
236 siltstones occur that are poorly sorted, and contain angular quartz grains (Fig. 3B) (Robinson
237 et al. 1974). Analyses of benthic foraminifera from DSDP Site 263 suggest the black,
238 kaolinitic claystones spanning ~475–746 m are likely Hauterivian-to-Middle Barremian,
239 passing upwards into Albian-to-Aptian black claystones (Fig. 3B) (Holbourn & Kaminski
240 1995); these age ranges are supported by dinoflagellate distributions and carbon isotope
241 stratigraphy (Wiseman & Williams 1974; Oosting et al. 2006). A gradual upwards transition
242 from coarsely to finely agglutinated foraminifera species, coupled with an upwards increase
243 in the scarcity of shallow-water taxa (e.g., *Hyperamina* spp.) and a corresponding decrease in
244 grain size, suggests that the Hauterivian-to-Middle Barremian strata record deepening neritic
245 (i.e. <200 m water depth) conditions (e.g., Fig. 3B) (Robinson et al. 1974; Veevers &
246 Johnstone 1974; Holbourn & Kaminski 1995; Oosting et al. 2006). Sedimentary rocks
247 recovered from the Pendock-1 borehole, which is located on the current continental shelf, are
248 sedimentologically similar and of comparable age to those penetrated in DSDP 263 (Veevers
249 & Johnstone 1974; Holbourn & Kaminski 1995). These similarities to Pendock-1 suggest that

250 the Hauterivian-to-Middle Barremian strata sampled by DSDP Site 263 can broadly be
251 correlated to the Winning Group of the North and South Carnarvon basins (Figs 1A and 3)
252 (Veevers & Johnstone 1974; Holbourn & Kaminski 1995).

253

254 **Dataset and methodology**

255 *Seismic reflection data*

256 To assess the crustal structure of the CAP and surrounding areas, we interpret seven 2D
257 seismic lines from four, pre-stack time-migrated reflection surveys (Fig. 1A) (see
258 Supplementary Table 1 for acquisition and processing details for each survey). Seismic lines
259 EW0113-5, EW0113-6, and repro-n303 are each >400 km long and extend from the
260 continental shelf up to ~300 km into the CAP from the currently defined COTZ (i.e. they are
261 orthogonal to the structural trend of the margin); EW0113-5 and EW0113-6 span the mapped
262 area of SDRs in the Cuvier COTZ and the location of the extinct spreading centre predicted
263 by Gibbons et al., (2012), whereas repro-n303 images the Sonne Ridge (Fig. 1A). Due to
264 extreme amplitude contrasts between the shallow and deep sections of the original migrated,
265 EW0113 data, we applied a time-dependent gain filter and root filter to improve amplitude
266 balance and enhance deep reflectivity (see supplementary information for details). Lines
267 s135-05, s135-08, and s310-59 image the inferred ‘transitional’ crust of the Wallaby Saddle
268 and the intruded continental crust of the Wallaby Plateau (e.g., Symonds et al., 1998;
269 Goncharov and Nelson, 2012; Olierook et al., 2015). The NE-trending seismic line s135-11
270 was also interpreted as ties together the margin-orthogonal seismic lines and provides a
271 margin-parallel image of the southernmost Exmouth Plateau continental crust, the CAP, and
272 the Wallaby Plateau (Fig. 1A).

273 Although time-migrated seismic reflection data allows us to qualitatively and
274 quantitatively characterise crustal structure, seismic velocity information is required to

275 convert depth information from seconds two-way time (TWT) to metres. To provide context
276 for the thicknesses and depths of some discussed structures, we depth-converted the
277 EW0113-5 and EW0113-6 seismic data using interval velocities derived from ocean-bottom
278 seismometer (OBS) data (Table 1) (Tischer 2006). The OBS array was co-located with
279 seismic line EW0113-6, which is located ~70 km along-strike from line EW0113-5 (Fig. 1A);
280 the geological structure imaged in line EW0113-6 is very similar to that of EW0113-5,
281 supporting the use of velocities from EW0113-6 to depth convert both lines. As velocity data
282 from across the Wallaby Plateau and Wallaby Saddle is limited (Goncharov & Nelson 2012),
283 and because along-strike variation in geology will likely promote changes in the velocity
284 structure, lines s135-05, s135-08, and s310-59 are presented in time. For easier comparison
285 between seismic data from the CAP and the Wallaby Plateau and Wallaby Saddle, we do not
286 depth-convert repro-n303 or s135-11. Interpretation of reflection configurations (e.g., dip
287 values) in time-migrated data are only qualitative, and may change if depth-converted.

288

289 ***Magnetic data***

290 To examine the regional magnetic anomalies, we utilise the EMAG2v2 and EMAG2v3 Earth
291 Magnetic Anomaly Grids (Maus et al. 2009; Meyer et al. 2017). EMAG2v2 is a 2 arc min
292 resolution grid derived from marine, airborne, and satellite magnetic data, but uses *a priori*
293 information to interpolate magnetic anomalies in areas where data gaps are present (Fig. 4A)
294 (Maus et al. 2009; Meyer et al. 2017). In contrast, EMAG2v3 uses more data points to derive
295 magnetic anomaly maps but assumes no *a priori* information (Fig. 4B) (Meyer et al. 2017).
296 In ocean basins with a relatively poor coverage of magnetic data available, such as the CAP,
297 clear linear magnetic anomalies in EMAG2v2 thus typically appear poorly developed or are
298 absent in EMAG2v3 (cf. Figs 4A and B) (Meyer et al. 2017). This difference in the presence
299 and appearance of linear magnetic anomalies between grids is because (assumed) knowledge

300 of seafloor spreading processes was incorporated into, and therefore influenced, interpolation
301 during construction of the EMAG2v2 grid (Maus et al. 2009; Meyer et al. 2017). Importantly,
302 the apparent reduction in magnetic stripes observed in EMAG2v3, compared to EMAG2v2
303 (Figs 4A and B), does not necessarily mean these features are absent, but rather that the
304 available data is insufficient to unambiguously confirm their presence in non-directionally
305 gridded data such as EMAG2v3 (Meyer et al. 2017). Comparing the EMAG2v2 and
306 EMAG2v3 grids, coupled with shiptrack magnetic data (Robb et al. 2005), allows us to
307 interrogate the magnetic architecture of the CAP (cf. Meyer et al. 2017). In particular, we use
308 EMAG2v2 to interpret possible magnetic chrons, picked on positive peaks (Fig. 1B), and
309 attempt to correlate them to the EMAG2v3 grid and shiptrack magnetic data. From these
310 comparisons, we tied interpreted magnetic stripes to seismic line EW0113-5, EW0113-6, and
311 repro-n303 using the synthetic profiles of Robb et al. (2005). To update the absolute ages of
312 the interpreted magnetic anomalies (Robb et al. 2005), we use the time-calibrated, magnetic
313 polarity reversal sequence of Gradstein & Ogg (2012).

314

315 *Geochemical data*

316 To evaluate whether the Sonne Ridge is an extinct seafloor spreading centre (e.g., Mihut &
317 Müller 1998; Robb et al. 2005) consisting of oceanic crust with a MORB or MORB-like
318 affinity along its length, we examine chemical data from a dredged basalt lava sample
319 collected along its extension into the Wallaby Plateau (i.e. Site 57 - sample 057DR051A;
320 diamond 57 in Fig. 1A) (Daniell et al. 2009; Dadd et al. 2015; Olierook et al. 2015). We
321 compare the Sonne Ridge sample to two samples collected from near the south-western
322 margin of the Wallaby Plateau (diamonds 55 and 52 in Fig. 1A) (i.e. Site 55 - samples
323 055BS004A and 055BS004B) (Dadd et al. 2015). Two Wallaby Plateau basalts dated from
324 Site 52 (Fig 1A) yield plagioclase $^{40}\text{Ar}/^{39}\text{Ar}$ plateau ages of 125.12 ± 0.9 Ma and 123.80 ± 1.0

325 Ma, whereas two analyses of the Sonne Ridge sample yielded less precise ages of 120 ± 14 Ma
326 and 123 ± 11 Ma (Olierook et al. 2015).

327

328 **Results**

329 *Reflection seismology*

330 *Cuvier Abyssal Plain*

331 We interpret a prominent, continuous, high-amplitude seismic reflection across the CAP; this
332 represents the interface between crystalline rock and overlying sedimentary strata (e.g., Figs
333 1C and 5). The Moho was picked at the base of a sub-horizontal zone of moderate-to-high-
334 amplitude, discontinuous seismic reflections, and is broadly flat-lying at ~ 16 – 17 km or ~ 10 s
335 TWT (Fig. 5). On EW0113-5, the Moho appears to become shallower oceanwards (to ≤ 14
336 km), although our interpretation of repro-n303 suggests it may deepen again beneath the
337 Sonne Ridge (Figs 5A and D). Overall, the crystalline crust is ~ 8 – 10 km (~ 3 – 5 s TWT) thick
338 (Figs 1C and 5). On EW0113-5 and EW0113-6, there is no clear evidence for the spreading
339 ridge interpreted by Gibbons et al., (2012), which is expected to occur within magnetic chron
340 M9n (Figs 1A and 5).

341 Across the CAP, the ~ 1 – 3 km-thick, uppermost crystalline crustal layer comprises a
342 layered, moderate- to high-amplitude seismic facies (SF1; Fig. 5). On NW-trending seismic
343 lines orthogonal to the margin (i.e. EW0113-5, EW0113-6, and repro-n303), SF1 locally
344 contains ≤ 40 km wide, ≤ 4.5 km thick wedges of coherent, high-amplitude, dipping
345 reflections that predominantly diverge seaward (Figs 5 and 6); adjacent to the Sonne Ridge
346 on its NW side, a package of dipping reflections diverge landwards (e.g., Fig. 5D). There is
347 no correlation between the location and width of these wedges relative to the magnetic
348 chrons; e.g., some packages of seaward-diverging reflections span several chrons (Fig. 5).
349 Where well-developed wedges are absent, SF1 contains discontinuous, horizontal to gently

350 seaward-dipping reflections (Fig. 5). On line s135-11, which is oriented parallel to the
351 margin, most reflections within SF1 are either sub-horizontal or dip gently north-eastwards
352 (Fig. 7). Seismic velocities for SF1 are estimated to be ~4–5 km/s (Fig. 5; Tischer 2006).

353 In places, the uppermost crystalline layer (SF1) is underlain by a low-amplitude, near
354 transparent seismic facies (i.e. SF2), which is particularly clear on lines EW0113-5 and
355 EW0113-6 (Figs 5 and 6). SF2 is up to ~2.8 km thick, being thinnest and occasionally absent
356 beneath wedges of dipping reflections within SF1 (Figs 5 and 6). The few reflections that
357 occur within SF2 typically have low-to-moderate to amplitudes and variable dips (Figs 5 and
358 6). On repro-n303, at the seaward termination of an overlying wedge in SF1, a ~15 km wide
359 swarm of landward-dipping reflections are present in SF2 (Fig. 5D). There is no clear SF2
360 observed on line s135-11, even in areas where it is encountered on the intersecting margin-
361 orthogonal lines (Fig. 7).

362 Beneath SF2 we recognise a ~3.5–6 km (<2 s TWT) thick, low-amplitude layer that
363 locally contains prominent, high-amplitude, dipping reflections and discontinuous, moderate
364 amplitude, sub-horizontal reflections (SF3; Figs 5–7). On line EW0113-5, the inclined
365 reflections within SF3 terminate at the Moho and primarily dip oceanwards at 20–30° (Fig.
366 5A). On lines EW0113-6 and repro-n303, however, reflections within SF3 dip both
367 oceanwards and landwards (Figs 5B, C, and 6). Mapped reflections within SF3 on s135-11
368 primarily dip towards the SE, extending from the top of the layer down into the mantle,
369 cross-cutting but not offsetting NE-dipping, gently inclined reflections (Fig. 7). Similar mid-
370 and lower-crustal reflection configurations to SF2 and SF3, respectively, occur in the seismic
371 data presented by Hopper et al. (1992) (Fig. 1C). Seismic velocities of SF2 and SF3 are 6.8–
372 7.2 km/s (Fig. 4; Tischer 2006).

373

374 *Wallaby Plateau and Wallaby Saddle*

375 Building on previous investigations of seismic data across the continental-to-COTZ crust of
376 the Wallaby Plateau and Wallaby Saddle, here we (re)interpret several 2D seismic lines and
377 compare their structure to that of the CAP. Similar to the CAP, a prominent, continuous,
378 high-amplitude seismic reflection marks the interface between crystalline rock and overlying
379 sedimentary strata across the Wallaby Plateau and Wallaby Saddle (Figs 7 and 8). Within the
380 Wallaby Saddle, the crust appears to be ~5–6 s TWT thick, although the Moho can only
381 tentatively be interpreted, and can also be sub-divided into: (i) SF1, itself containing up to ~4
382 s TWT thick, 12 km wide wedges of diverging reflections that typically dip seawards; (ii)
383 restricted zones that seismically appear similar to SF2 described from the CAP; and (iii) a
384 1.5–3 s TWT thick SF3 unit that contains reflections with variable dips, including prominent
385 swarms of landward-dipping reflections that cross-cut but do not offset other reflections and
386 that typically occur at the oceanward termination of SF1 wedges (Fig. 8). Derivation of
387 interval velocities from seismic reflection stacking velocities suggests rocks comprising SF1
388 have velocities of ~2.5–5.3 km s⁻¹ (see insets in Fig. 8C) (Goncharov & Nelson 2012). It is
389 difficult to determine whether SF1-SF3 continue across the full extent of the Wallaby Saddle
390 in s310-59 because there appears to be a distinct change in reflection configuration (Fig. 8C).
391 In particular, we observe that although reflectivity in west of this change is decreased,
392 reflections towards the top of the crust are broadly sub-parallel to the basement reflection and
393 those within the mid- to lower-crustal areas are either gently inclined landwards, or
394 moderately inclined oceanwards (Fig. 8C).

395 Seismic reflection imaging of the Wallaby Plateau reveals the crust is up to ~7 s TWT
396 thick (e.g., at the Sonne Ridge), thicker than that of the Wallaby Saddle (~5–6 s TWT thick)
397 but that there is no apparent significant change in Moho depth between the two crustal
398 domains (Figs 7 and 8); we note these observations are based on time-migrated data and may
399 thus be invalidated if there any differences in velocity structure between the two areas not

400 previously recognised. The crust of the Wallaby Plateau is also thicker than that of the CAP,
401 and its underlying Moho is located at deeper levels (~ 12 s TWT; Fig. 7). Towards the SW
402 margin of the Wallaby Plateau, a ~ 40 km wide, apparently NE-trending rift system occurs,
403 comprising normal faults with throws of up to ~ 1 s TWT that bound and dissect a graben
404 (Figs 8B–D). Reflections within the upper section of the Wallaby Plateau crust are typically
405 moderate-to-high amplitude and form layered packages, which are either conformable to the
406 top basement horizon or that diverge (Fig. 8). The diverging packages of dipping reflections
407 appear similar to SF1 observed in the CAP and Wallaby Saddle (Figs 5 and 8). Derivation of
408 interval velocities from seismic reflection stacking velocities suggests rocks comprising these
409 diverging reflector sequences have velocities of ~ 2.5 – 5.3 km s⁻¹ (Fig. 8C) (Goncharov &
410 Nelson 2012). Due to uncertainties regarding the reliability of seismic processing within the
411 middle and lower crustal sections of the Wallaby Plateau, e.g., where imaging is hindered by
412 seabed multiples, it is difficult to confidently interpret reflections as real geological features
413 and not artefacts. However, we note that in these middle and lower crustal sections,
414 reflections are low-to-moderate amplitude and broadly dip gently in various directions; in
415 places, steeply inclined reflections are observed that appear to cross-cut but not offset gently
416 dipping reflections (Fig. 8). These steeply inclined mid- to lower-crustal reflections typically
417 appear to be located beneath diverging reflection packages, or beyond their down-dip
418 termination (Fig. 8).

419

420 *Comparison of magnetic anomalies to seismic reflection data*

421 EMAG2v2 and ship-track magnetic data reveal that 10 km wide, ≤ 220 km long magnetic
422 stripes cover much of the CAP (Figs 1B, 4, and 5). No magnetic stripes can confidently be
423 identified and dated within the Wallaby Plateau and none are observed within the Wallaby
424 Saddle (Figs 1B and 4). Although magnetic anomalies in the EMAG2v3 grid are suppressed

425 relative to EMAG2v2, subtle, linear anomalies can still be distinguished across the CAP and
426 in the Gallah Province (cf. Figs 4A and B). Due to the lower resolution of magnetic
427 anomalies in the EMAG2v3 grid, magnetic chrons cannot be confidently attributed and we
428 thus rely on EMAG2v2 and shiptrack data to define possible chrons (Fig. 4). Proximal to the
429 Australian continent, long-wavelength magnetic anomalies can only be broadly assigned to
430 chron M10N (~135.9–134.2 Ma; Figs 1B, 4, and 5) (Robb et al. 2005); across parts of
431 seismic lines EW0113-5, EW0113-6, and repro-n303, chrons M10n–M5r (~135.3–131.4 Ma)
432 are clearly defined and have amplitudes of $\leq \pm 100$ nT (Figs 1B, 4, and 5). On all three seismic
433 lines, chrons M8r–M7n (~133–132 Ma) coincide with a package of seaward-dipping
434 reflectors observed in SF1, which on EW0113-5 is ≤ 3 km thick and ~25 km long (Fig. 5).

435 Robb et al. (2005) interpret the magnetic anomalies M10N–M6 southeast of the
436 Sonne Ridge as conjugate to a more poorly developed set of anomalies northwest of the ridge
437 (Fig. 1B). These chrons NW of the Sonne Ridge (i.e. M10N–M6; ~135.9–131.7 Ma)
438 terminate abruptly north-eastwards against the Cape Range Fracture Zone, abutting magnetic
439 chrons on the Gallah Province, and to the SW are cross-cut by chrons (M5n?; 131.7–130.6
440 Ma) mirrored either side of the Sonja Ridge (Fig. 1B). Beyond the outermost chron (M10N)
441 interpreted in the CAP, chron M5n is the first to occur continuously along-strike across both
442 the Cuvier and Gascoyne margin segments, extending into the Wallaby Plateau (Fig. 1B).

443

444 *Geochemistry of basalts dredged from the Sonne Ridge*

445 The only basalt collected from the Sonne Ridge displays a relatively flat Rare Earth Element
446 (REE) pattern (Fig. 9A). Based on this observation, Dadd et al. (2015) interpret the basalt to
447 have a slightly enriched MORB-like source, supporting the inference that the CAP comprises
448 oceanic crust (e.g., Larson et al. 1979; Hopper et al. 1992; Mihut & Müller 1998). Although a
449 flat REE pattern can be indicative of a shallow melting regime related to MORB generation,

450 it does not preclude other settings. By replotting the trace element and radiogenic isotopic
451 compositions of the Sonne Ridge sample, we show the sample has characteristics that could
452 suggest a more continental source (Fig. 9A). It should be noted that the Sonne Ridge sample
453 is heavily altered (Dadd et al. 2015) which could explain the elemental enrichment in Pb, Ba,
454 and Rb, as well as elevated $^{87}\text{Sr}/^{86}\text{Sr}$. However, the sample exhibits unradiogenic ϵ_{Nd} and a
455 negative Nb anomaly that is in part defined by a relative enrichment in the neighbouring
456 element Th, which likely cannot be attributed to contamination; the combination of these
457 features likely cannot be ascribed to alteration (Fig. 9). Instead, the negative Nb anomaly and
458 unradiogenic ϵ_{Nd} may indicate a chemically evolved, continental or sedimentary contribution
459 to the magmas. The chemical similarity of the Sonne Ridge basalt to two ~124 Ma samples
460 from the Wallaby Plateau (Fig. 9), which is interpreted to comprise intruded continental crust
461 (Daniell et al. 2009; Stilwell et al. 2012; Olierook et al. 2015), is also consistent with a
462 continental or sedimentary contribution to the Sonne Ridge magmas.

463

464 **Interpretation and Discussion**

465 Since the recognition that it contains magnetic stripes, the CAP has been considered to
466 comprise unambiguous oceanic crust that formed at ~136 Ma (Valanginian) in response to
467 seafloor spreading at the Sonne Ridge (e.g., Falvey & Veevers 1974; Larson et al. 1979;
468 Hopper et al. 1992; Robb et al. 2005). An oceanic origin for the CAP has been supported by
469 seismic reflection-based observations that it has a thin crust relative to adjacent continental
470 blocks (e.g., Fig. 1C) (e.g., Hopper et al. 1992), and chemical data, which suggest it has a
471 MORB-like signature (Dadd et al. 2015). The apparent certainty that the CAP is oceanic
472 means it has been unquestionably treated as such in all geological models of the evolution of
473 NW Australia, including regional and global palinspastic and plate kinematic reconstructions
474 (Heine & Müller 2005; Gibbons et al. 2012). However, the identification of linear magnetic

475 anomalies within non-oceanic crust, in areas such as Ethiopia and the Atlantic margins,
476 requires us to re-evaluate the origin of crustal domains previously classified as oceanic crust
477 based on the presence of magnetic stripes (e.g., Bronner et al. 2011; Bridges et al. 2012;
478 Collier et al. 2017; McDermott et al. 2018).

479

480 *Seismic facies interpretation*

481 Beneath the sedimentary cover across the CAP, we recognise three distinct layers (SF1–SF3;
482 Figs 5–8). We particularly identify a newly recognised upper-crustal layer (SF1) in the CAP
483 that comprises well-developed wedges of divergent, seaward-dipping reflectors (SDRs) (Figs
484 5 and 6). These SDRs are ≤ 4.5 km thick, likely have seismic velocities of $\sim 4\text{--}5$ km s⁻¹
485 (Tischer 2006), and that collectively extend >300 km seaward from the previously interpreted
486 COTZ (Figs 5 and 6). Diverging SDRs are also observed within the: (i) the previously
487 defined, 50–70 km wide COTZ along the Cuvier Margin beneath the continental slope, where
488 they are up to ~ 5 km thick (e.g., Fig. 1C) (e.g., Hopper et al. 1992; Symonds et al. 1998); and
489 (ii) across the Wallaby Saddle and Wallaby Plateau, where they are $\sim 5\text{--}10$ km thick and have
490 similar seismic velocities ($2.5\text{--}5.3$ km s⁻¹) to those in the CAP (Figs 5, 6, and 8) (e.g.,
491 Symonds et al. 1998; Sayers et al. 2002; Goncharov & Nelson 2012). The lack of boreholes
492 penetrating these SDR sequences offshore NW Australia means we cannot determine their
493 composition or the nature of underlying crust. However, SDR sequences that are
494 geometrically and geophysically (e.g., seismic velocities typically range from $\sim 3\text{--}5$ km s⁻¹)
495 similar to those from offshore NW Australia (e.g., SF1) have been recognised along other
496 passive margins, developed on both heavily intruded continental crust and thickened oceanic
497 crust (e.g., Hinz 1981; Larsen & Saunders 1998; Harkin et al. 2020). Where these SDRs have
498 been drilled, or are exposed onshore (e.g., Iceland and Greenland), they comprise interbedded
499 basaltic lavas, tuffs, and sedimentary rocks formed during sub-aerial, or perhaps shallow-

500 water, continental breakup and crustal spreading (e.g., Bodvarsson & Walker 1964; Mutter et
501 al. 1982; Roberts et al. 1984; Eldholm et al. 1987; Larsen et al. 1994a; Geoffroy et al. 2001;
502 Harkin et al. 2020). Based on similarities in structure and seismic velocities to SDRs studied
503 elsewhere, we suggest that SF1 comprises spreading-related volcanic rocks interbedded with
504 sedimentary layers (Figs 1C, 5, 6, and 8) (e.g., Mutter et al. 1982; Hopper et al. 1992;
505 Symonds et al. 1998; Planke et al. 2000; McDermott et al. 2019; Harkin et al. 2020).

506 SDR sequences (SF1) can develop on heavily intruded continental crust or thickened
507 oceanic crust (e.g., Larsen & Saunders 1998). The observed structure and seismic velocities
508 ($6.8\text{--}7.2\text{ km s}^{-1}$) of SF2 and SF3 in the CAP, defined by transparent seismic facies and
509 discordant high-amplitude reflections, respectively (Figs 5–8), are consistent with the typical
510 seismic character of sheeted dykes and lower crustal gabbro intrusions in oceanic crust (e.g.,
511 Eittreim et al. 1994; Paton et al. 2017). However, we note that these seismic facies are not
512 uniquely diagnostic of oceanic crust but can also occur in COTZs, where moderate- to high-
513 amplitude reflections may represent igneous intrusions (e.g., dykes), primary layering within
514 gabbros, or texturally distinct lower crustal shear zones within otherwise homogenous
515 crystalline rocks (e.g., Phipps-Morgan & Chen 1993; Abdelmalak et al. 2015; Paton et al.
516 2017). For example, the swarm of landward dipping reflections within SF2 and SF3 at the
517 down-dip termination of an SDR sequence may correspond to dykes; i.e. they cross-cut but
518 do not offset background reflections and are thus not faults or shear zones (e.g., Figs 5 and 8)
519 (e.g., Abdelmalak et al., 2015; Phillips et al., 2018).

520

521 ***Implications of SDR recognition for the CAP***

522 *Origin of SDR lavas*

523 Lavas within SDR wedges are inferred to emanate from and be thickest at axial magmatic
524 segments, where they were likely fed by sub-vertical dykes. With continued plate divergence,

525 these lavas subside and rotate to dip inwards towards their eruption site (e.g., Planke &
526 Eldholm 1994; Paton et al. 2017; Norcliffe et al. 2018; Tian & Buck 2019); this subsidence
527 also rotates underlying feeder dykes, which will dip away from the magmatic segment (e.g.,
528 Lenoir et al. 2003; Abdelmalak et al. 2015). SDRs across the CAP appear to dip and diverge
529 north-westwards, except one SDR-like package of concave-upwards reflections that borders
530 and diverges south-eastwards towards the Sonne Ridge; i.e. we define a conjugate set of
531 SDRs that occur either side of and dip towards the Sonne Ridge (Fig. 5). Although only one
532 SDR package to the NW of the Sonne Ridge dips south-eastwards towards the ridge, we
533 suggest that the other SDR packages, which dip north-westwards, relate to and were formed
534 at the Sonja Ridge (Fig. 5). From these SDR geometries and distribution, coupled with the
535 previously inferred conjugate sets of magnetic chrons (Fig. 1B), our results are consistent
536 with suggestions that (Falvey & Veevers 1974; Larson et al. 1979; Robb et al. 2005;
537 MacLeod et al. 2017): (i) extension within the CAP was predominantly centred on the Sonne
538 Ridge during chrons M10N–M5r (~136–131 Ma); before (ii) briefly jumping to the Sonja
539 Ridge at ~131 Ma (chron M5n), which interrupted subsidence and rotation of the SDR wedge
540 immediately to the NW of the Sonne Ridge and instead produced north-westwards diverging
541 SDRs. There are no changes in SDR divergence direction either side of chron M9n in
542 EW0113-5 or EW0113-6, suggesting no spreading centre existed here (Figs 5A and B) (cf.
543 Gibbons et al. 2012).

544 The chemistry of a basalt sample from the Sonne Ridge, particularly its Nd isotopic
545 composition and refractory trace element abundances (Fig. 9), indicates it could originate
546 from either: (i) melting of sub-continental lithospheric mantle (SCLM); or (ii) contamination
547 of a MORB-like magma by assimilation as it ascended through continental crust (cf. Dadd et
548 al. 2015). Because these chemical data provide evidence that the Sonne Ridge basalt
549 interacted with continental material, these same data cannot thus be used as definitive

550 evidence that the CAP comprises oceanic crust (cf. Dadd et al. 2015). The basalt sample
551 dredged from its present-day bathymetric expression can be considered a product of one of
552 the youngest magmas within the CAP system (e.g., Robb et al. 2005). We thus think it
553 plausible that older magmas emanating from the Sonne Ridge, including the SDRs imaged in
554 the seismic reflection data, could have a more pronounced continental signature.

555

556 *Environment of SF1 lava emplacement*

557 Borehole and field data reveal SDR lavas typically erupt sub-aerially, but can develop sub-
558 aqueously (e.g., Bodvarsson & Walker 1964; Mutter et al. 1982; Roberts et al. 1984; Eldholm
559 et al. 1987; Larsen et al. 1994b; Symonds et al. 1998; Planke et al. 2000; Geoffroy et al.
560 2001; Harkin et al. 2020). Determining the environment and age of SDR deposition can help
561 establish whether they likely formed via: (i) seafloor spreading at a mid-ocean ridge,
562 consistent with previous interpretations that the CAP comprises unambiguous oceanic crust
563 (e.g., Falvey & Veevers 1974; Larson et al. 1979; Hopper et al. 1992; Robb et al. 2005); or
564 (ii) magmatic addition along a sub-aerial or shallow-water axis during the transition from
565 continental rifting to full plate separation (i.e. the CAP does not comprise oceanic crust) (e.g.,
566 McDermott et al. 2018). However, from their seismic character alone it can be difficult to
567 determine whether SDRs formed in sub-aerial, shallow-water, or deep-marine environments
568 (e.g., compare inner and outer SDR character and inferred emplacement conditions; Symonds
569 et al. 1998; Planke et al. 2000).

570 Observations from the DSDP Site 263 borehole, which terminates ~100–200 m above
571 the crystalline crust, indicate the sedimentary cover deposited above the SDRs: (i) comprises
572 poorly sorted silty claystones that include angular quartz grains and abundant kaolinite,
573 consistent with a neritic (i.e. <200 m water depth) depositional environment (Fig. 3B) (e.g.,
574 Robinson et al. 1974; Veevers & Johnstone 1974; Compton et al. 1992; Holbourn &

575 Kaminski 1995; Oosting et al. 2006); (ii) contain coarsely agglutinated foraminifera species
576 and taxa such as *Hyperamina* spp. within the lowermost intersected strata, which are typical
577 of shallow-marine conditions (Holbourn & Kaminski 1995); and (iii) based on
578 biostratigraphic data were deposited at least in the middle Barremian (e.g., ~127 Ma), but are
579 perhaps as old as Hauterivian (~132.6–129.4 Ma) (Oosting et al. 2006). Deposition of the
580 lowermost sedimentary cover intersected by DSDP Site 263 thus occurred up to ~9 Myr (i.e.
581 ~135.9–127 Ma) after formation of the CAP crust they rest upon, which records chron M10N
582 (135.9–134.2 Ma), during development of crust hosting chrons M7–M1n (132.5–126.3 Ma;
583 Fig. 1B). Critically, SDR-bearing crust cools and subsides as it is transported away from its
584 emplacement site (e.g., Planke & Eldholm 1994; Paton et al. 2017; Norcliffe et al. 2018; Tian
585 & Buck 2019). The presence of strata deposited in the neritic zone (<200 m water depth)
586 above SF1 in DSDP 263, after ~9 Myr of crustal cooling and subsidence, thus implies lava
587 eruption during the early stages of CAP formation occurred: (i) in a sub-aerial or shallow-
588 water environment (i.e. comparable to the inner SDRs of Symonds et al. 1998; Planke et al.
589 2000), if we assume the underlying crust only subsided in the ~9 Myr between SDR
590 emplacement and sediment deposition; or (ii) at a moderately deep-marine spreading centre
591 (i.e. comparable to the outer SDRs of Symonds et al. 1998; Planke et al. 2000), but localised
592 uplift elevated the DSDP 263 area to bathymetric depths equivalent to the neritic zone prior
593 to deposition of overlying strata. We lack the data from strata directly overlying or
594 interbedded with the SDRs to test these two interpretations regarding lava emplacement
595 depth, but note that the relatively flat-lying crystalline crust across the interpreted CAP
596 seismic lines (except for the Sonne Ridge) provides no evidence of post-spreading uplift,
597 perhaps suggesting a sub-aerial or shallow-water environment of emplacement is most
598 plausible (Fig. 5).

599

600 *Nature of CAP crust*

601 Seismic and magnetic data alone are insufficient to determine the origin of the CAP crust
602 because the SDRs, seismic facies (SF1–SF3), and magnetic stripes these data illuminate can
603 be recognised in both oceanic crust and COTZs (e.g., Larsen & Saunders 1998; Symonds et
604 al. 1998; Planke et al. 2000; Bridges et al. 2012; Collier et al. 2017; McDermott et al. 2018).
605 We also show that the chemical data available for a basalt at the Sonne Ridge possesses a
606 continental signature and are thus inconclusive regarding whether or not the crust is oceanic
607 (Fig. 9) (cf. Dadd et al. 2015). However, based on lithological and biostratigraphic data from
608 the sedimentary cover intersected by DSDP Site 263, we suggest: (i) the inferred lavas within
609 SF1, at least during the early stages of CAP formation (i.e. chron M10N), likely erupted in a
610 sub-aerial, or perhaps shallow-water (<200 m water depth), environment; and (ii), assuming
611 the underlying crystalline crust had since subsided relative its position during formation, that
612 the syn-depositional, ~9 Myr old Sonne Ridge was elevated above at least the base of the
613 neritic zone. These constraints on SDR emplacement depth are inconsistent with the CAP
614 being oceanic crust since mid-ocean ridges in such a setting are expected to occur at water
615 depths of ~3 km after 5–10 Myr of spreading (e.g., Menard 1969; Parsons & Sclater 1977;
616 Stein & Stein 1992).

617 We envisage that crustal structure of the CAP could involve a gradual north-
618 westwards change from the continental crust of the Cuvier Margin into a COTZ, which is
619 likely characterised at its landward edge by heavily intruded continental crust and
620 progressively becomes increasingly magma-dominated towards the Sonne Ridge (Figs 2 and
621 10). Our data are insufficient to determine where, or if, there is a transition from heavily
622 intruded continental crust to magmatic crust, which would mark break-up of the continental
623 crust within the CAP. Repetition of the M10N-M6 chrons centred on the Sonne Ridge
624 suggests the possible COTZ of the CAP may extend at least out to chron M5n, which is: (i)

625 >500 km oceanwards of the outer- limit of the previously defined Cuvier COTZ (e.g., Hopper
626 et al. 1992; Symonds et al. 1998); and (ii) broadly coincident with the north-western limit of
627 the Gallah Province on the Gascoyne margin (Direen et al. 2008) (Figs 1B and 10B). From
628 the distribution of the magnetic chrons (Fig. 1B), our recognition of a continental chemical
629 signature within the Sonne Ridge basalt (Fig. 9), and the probable sub-aerial or shallow-water
630 elevation of the ridge during extension, make it plausible that full continental lithospheric
631 rupture may not have occurred in the CAP (Fig. 10). Instead, we suggest rupture of the
632 continental lithosphere and onset of seafloor spreading occurred simultaneously offshore the
633 Cuvier and Gascoyne margins at ~131 Ma (Hauterivian), following an oceanwards ridge
634 jump from the Sonja Ridge, producing unambiguous oceanic crust recording chron M5 (Figs
635 1 and 10) (e.g., Robb et al. 2005; Direen et al. 2008). Continuation of the COTZ across the
636 CAP has implications for the timing and kinematics of plate reconstructions of the NW
637 Australian margin, with the onset of deep-marine seafloor spreading potentially ~3 Myr later
638 than suggested by previous studies (e.g., Robb et al. 2005).

639 Interpreting the CAP as a COTZ developed through sub-aerial, or at least shallow-
640 water, extension implies its crust was: (i) thicker during SDR emplacement, but concurrently
641 and/or subsequently thinned during continued magmatic extension and late-stage stretching
642 (e.g., Bastow & Keir 2011; Bastow et al. 2018); (ii) less dense and thus more buoyant than
643 100% oceanic crust, because it likely retained a significant proportion of continental material;
644 and (iii) thermally buoyant due to the presence of abundant hot intrusions and underlying,
645 decompressing mantle. That these processes can maintain rift axes at above or near sea-level
646 elevations is demonstrated by the onshore occurrence of active rift zones, characterised by
647 heavily-intruded continental crust, in the Main Ethiopia Rift and Afar (e.g., Hayward &
648 Ebinger 1996; Ebinger & Casey 2001; Mackenzie et al. 2005; Bridges et al. 2012).

649 Because the degree of thermal subsidence is at least partly controlled by crustal
650 density, we would expect oceanic crust to thermally subside more than less dense, heavily-
651 intruded continental crust. Given the Hauterivian-to-Middle Barremian sedimentary strata
652 overlying the SDRs were deposited in neritic conditions (Veevers & Johnstone 1974;
653 Holbourn & Kaminski 1995; Oosting et al. 2006), it is apparent the CAP subsided from near
654 sea-level to a current, unloaded basement depth of ~6.5 km; this total subsidence is greater
655 than predicted for dense, thermally subsiding oceanic crust (Stein & Stein 1992). To interpret
656 the CAP as COTZ crust, our results would require other mechanisms, in addition to thermal
657 subsidence, to influence its subsidence history. For example, post-breakup decay of
658 asthenospheric thermal anomalies may account for some elevation discrepancies via removal
659 of dynamic support of the margin (e.g., Czarnota et al. 2013). Finally, the CAP COTZ may
660 have involved some late-stage stretching prior to terminal breakup and the onset of seafloor
661 spreading, akin to processes observed today in the sub-aerial Red Sea rift in Ethiopia (e.g.,
662 Bastow & Keir 2011; Daniels et al. 2014).

663

664 *Development of magnetic stripes during break-up*

665 Recent forward modelling of conjugate, ship-track magnetic profiles by Collier et al. (2017)
666 suggest magnetic signals over SDRs arise from a combination of stacked and rotated lavas,
667 producing a long-wavelength positive anomaly that can sometimes mask reversals, and linear
668 magnetic anomalies caused by dyke intrusion in the underlying crust. Stacked SDR wedges
669 on the CAP are part of a possible COTZ and span several chrons (e.g. M8n-M7r), but are
670 ≤ 4.5 km thick (Figs 5 and 6). These observations indicate the CAP magnetic stripes likely
671 record magnetic reversal signatures originating from sub-SDR rocks; i.e. the SDRs and flat-
672 lying lavas are too thin to dominate the magnetic signature (cf. Collier et al. 2017). In
673 contrast, the less-clearly developed yet higher amplitude magnetic reversals in the Gallah

674 Province COTZ may relate to interference from the greater SDR thicknesses (≤ 5.5 km)
675 relative to the CAP (Direen et al. 2008). Our inference that the magnetic signature is derived
676 from sub-SDR rocks is consistent with studies of onshore incipient spreading centres (e.g.
677 Ethiopia), where magnetic stripes likely originate from axial intrusion by dykes in heavily
678 intruded, upper continental crust, rather than overlying lavas (Bridges et al. 2012).

679 We suggest that SDR thickness and, thereby, preservation of magnetic anomalies
680 within a COTZ can partly be attributed to extension rate. For example, the extension rate
681 during SDR eruption offshore NW Australia (~ 4.5 cm/yr half rate; Robb et al. 2005) is
682 substantially faster than the inferred extension rates for the South Atlantic during magmatic
683 crust formation (~ 1.1 cm/yr half-rate; Paton et al. 2017). Slower extension rates (e.g. South
684 Atlantic) likely promote stacking of lava flows to produce thicker SDRs (Eagles et al. 2015),
685 leading to interference between the magnetic signal of the SDRs and sub-SDR crust and thus
686 the development of the long-wavelength positive magnetic anomalies (e.g., Moulin et al.
687 2010). Extension rate may also influence magnetic anomaly development by affecting the
688 width of magnetic stripes; reversal anomalies will be narrowest at slow spreading ridges
689 (Vine 1966). The narrower anomalies, combined with the greater potential for vertical
690 stacking of lavas, will tend to suppress magnetic anomaly preservation.

691

692 **Conclusions**

693 The recognition of magnetic stripes within the Cuvier Abyssal Plain (CAP), offshore NW
694 Australia, has led to the assumption that it comprises oceanic crust generated by conventional
695 seafloor spreading at the Sonne Ridge, probably at water depths of ≥ 2 km. We challenge this
696 assumption, in line with the growing consensus that magnetic stripes are not necessarily
697 diagnostic of oceanic crust and can instead form in continent-ocean transition zones
698 (COTZs). Using regional 2D seismic reflection lines we demonstrate that the uppermost layer

699 in the CAP crystalline line crust contains seaward-dipping reflector (SDR) sequences, akin to
700 those observed in the previously defined COTZ of the Cuvier Margin and Wallaby Saddle, as
701 well as on the heavily intruded continental crust of the Wallaby Plateau. Through comparison
702 to SDRs recognised elsewhere, we suggest those observed across the CAP comprise lavas,
703 interbedded with sedimentary strata, erupted from an axial magmatic segment. Lithological
704 and biostratigraphic data from a borehole penetrating the CAP sedimentary cover, which
705 were deposited in neritic (<200 m water depth) conditions, require the underlying crystalline
706 crust to have been at shallow-water depths ~9 Myr after its formation and thus imply SDR
707 emplacement occurred in a shallow water or sub-aerial environment. We also re-interpret
708 chemical data from a basalt dredged along the Sonne Ridge and, contrary to previous work,
709 show that it exhibits a continental chemical signature. Overall, these data and interpretations
710 suggest the CAP may not comprise unambiguous oceanic crust, but could instead represent a
711 >500 km wide COTZ where extension likely became more magma-dominated, producing
712 heavily-intruded continental crust (akin to present-day Ethiopia) through to magmatic crust.
713 In our model, break-up of the continental crust occurred during the formation of the CAP, but
714 full continental lithospheric rupture occurred outboard of the COTZ following a ridge jump at
715 ~130 Ma. Our re-evaluation of the CAP crustal type supports suggestions that COTZs along
716 volcanic passive margins may record the development of magnetic stripes, which thus should
717 not be used alone as a reliable proxy for the onset of seafloor spreading and the extent of
718 oceanic crust.

719

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727 previous versions of this manuscript.

728

729 **Figure captions**

730 Figure 1: (A) Location map of the study area highlighting the seismic lines used in this study
731 and key tectonic elements, including areas of recognised seaward-dipping reflectors (SDRs)
732 (Symonds et al. 1998; Holford et al. 2013) and previously interpreted approximate limits of
733 the continent-ocean boundary (COB; Eagles et al. 2015) and continent-ocean transition zones
734 (COTZs; Symonds et al. 1998; Direen et al. 2008). Inset: study area location offshore NW
735 Australia. AAP – Argo Abyssal Plain, CAP – Cuvier Abyssal Plain, CRFZ – Cape Range
736 Fracture Zone, GAP – Gascoyne Abyssal Plain, GP – Gallah Province, NCB – North
737 Carnarvon Basin, EP – Exmouth Plateau, PB – Perth Basin, SCB – South Carnarvon Basin,
738 Cu – Cuvier margin COTZ, SR – Sonne Ridge, SjR – Sonja Ridge, WP – Wallaby Plateau,
739 WS – Wallaby Saddle, WZFZ – Wallaby-Zenith Fracture Zone. Dredge sites 52, 55 (samples
740 055BS004A and 055BS004B), and 57 (sample 057DR051A) are also shown (Dadd et al.
741 2015). (B) Total magnetic intensity grid (EMAG2v2), interpreted magnetic chrons (based on
742 Robb et al. 2005). See Supplementary Figure S1 for an uninterpreted version. (C)
743 Uninterpreted and interpreted seismic line (i.e. seismic profile 670) across the Cuvier Margin,
744 imaging the crustal structure beneath the continental shelf and the deep abyssal plain
745 (modified from Hopper et al., 1992). Velocity profiles from refraction experiments shown;
746 see Hopper et al., (1992) for details. See Figure 1A for approximate line location and
747 Supplementary Figure S2 for an enlarged version of the uninterpreted seismic line.

748

749 Figure 2: Schematic model (not to scale) of a continent-ocean transition zone along a magma-
750 rich passive margin, which depicts the evolution from unambiguous continental crust to
751 unambiguous oceanic crust; for simplicity the lithospheric mantle is not shown. As magma
752 intrudes continental crust, likely as dykes at mid- to upper-crustal levels and larger gabbroic
753 bodies in the lower crust, it becomes ‘heavily intruded continental crust’ (e.g., Eldholm et al.
754 1989). Continued intrusion and dyking leads to localisation of magmatism within narrow
755 zones where there is little, if any, continental crust remaining (i.e. ‘magmatic crust’; e.g.,
756 Collier et al. 2017; Paton et al. 2017). We categorize heavily intruded continental crust and
757 magmatic crust as ‘COTZ crust’. Sub-aerial, magma-assisted rifting may feed extensive lava
758 flows that later, through subsidence, become seaward-dipping reflectors (SDRs). SDR
759 subsidence leads to rotation of underlying dykes (Abdelmalak et al. 2015); a similar rotation
760 of lavas and dykes is observed in oceanic crust (Karson 2019).

761

762 Figure 3: Tectono-stratigraphic chart for the Exmouth Plateau and Cuvier Margin (after
763 Hocking et al., 1987; Arditto, 1993; Partington et al. 2003; Willis, 2005; Reeve et al. 2016).
764 (B) Comparison between stratigraphic data from DSDP 263 and Pendock-1 boreholes
765 (modified from Veevers & Johnstone, 1974; Holbourn & Kaminski, 1995). See Figure 1A for
766 borehole locations.

767

768 Figure 4: Total magnetic intensity grids EMAG2v2 and EMAG2v3 (Maus et al. 2009; Meyer
769 et al. 2017), compared with shiptrack magnetic data (Robb et al. 2005). Key tectonic
770 elements also shown (see Fig. 1 for legend). In (A), CAP – Cuvier Abyssal Plain, GAP –
771 Gascoyne Abyssal Plain, GP – Gallah Province, NCB – North Carnarvon Basin, PB – Perth
772 Basin, SCB – South Carnarvon Basin, Cu – Cuvier margin COTZ, WP – Wallaby Plateau,
773 WS – Wallaby Saddle.

774

775 Figure 5: Interpreted and uninterpreted, depth-converted seismic lines (A) EW0113-5 and (B)
776 EW0113-6, and the time-migrated line (D) repro n303 showing crustal structure of the Cuvier
777 Margin; see Figures 1A and 5C for line locations. The tie-co-located magnetic anomaly
778 profile showing interpreted magnetic chrons is presented for (A–D) (after Robb et al. 2005).
779 See Supplementary Figure S2 for an enlarged version of the uninterpreted seismic lines.

780

781 Figure 6: Zoomed in view of EW0113-5 highlighting the seismic character of interpreted
782 SDR packages (see Fig. 5A for location).

783

784 Figure 7: Interpreted and uninterpreted, time-migrated seismic line s135-11; see Figure 1A
785 for location. See Supplementary Figure S2 for an enlarged version of the uninterpreted
786 seismic line.

787

788 Figure 8: Interpreted and uninterpreted, time-migrated seismic lines (A) s135-s135_05, (B)
789 s135-08, and (D) s310-59 showing crustal structure of the Wallaby Plateau and Wallaby
790 Saddle; see Figures 1A and 8D for line locations. See Supplementary Figure S2 for an
791 enlarged version of the uninterpreted seismic lines.

792

793 Figure 9: (A) Primitive mantle normalized incompatible element diagram comparing the
794 dredged Sonne Ridge and Wallaby Plateau basalt lava samples with average (ave.)
795 compositions of MORB variants (Hofmann 2014), Globally Subducting Sediment (GLOSS)
796 (Plank & Langmuir 1998), and continental crust (Rudnick & Fountain 1995). (B) Plot of
797 $\epsilon(\text{Nd})$ versus $^{87}\text{Sr}/^{86}\text{Sr}$, illustrating that the Sonne Ridge and Wallaby Plateau samples are
798 distinct from MORB (based on data collated in Hofmann 2014).

799

800 Figure 10: (A) Map showing the potential limits of the COTZ based on interpreting the CAP
801 and Gallah Province as transitional and/or magmatic crust. (B-D) Schematic maps showing
802 the development of COTZ crust and the onset of oceanic crust accretion adjacent to the
803 Gascoyne and Cuvier margins, during formation of chrons (B) M10, (C) M6 and (D) M3r.
804 See Figure 1 for chron ages. Location of present day coastline shown for reference.

805

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Table 1: Interval velocities

Layer	Seismic velocity (km s ⁻¹)
Water column	1.5
Sedimentary strata Seaward-dipping reflectors (SDRs)	2.0–2.8
Sub-SDR crust	4.9
Upper mantle	6.8–7.2
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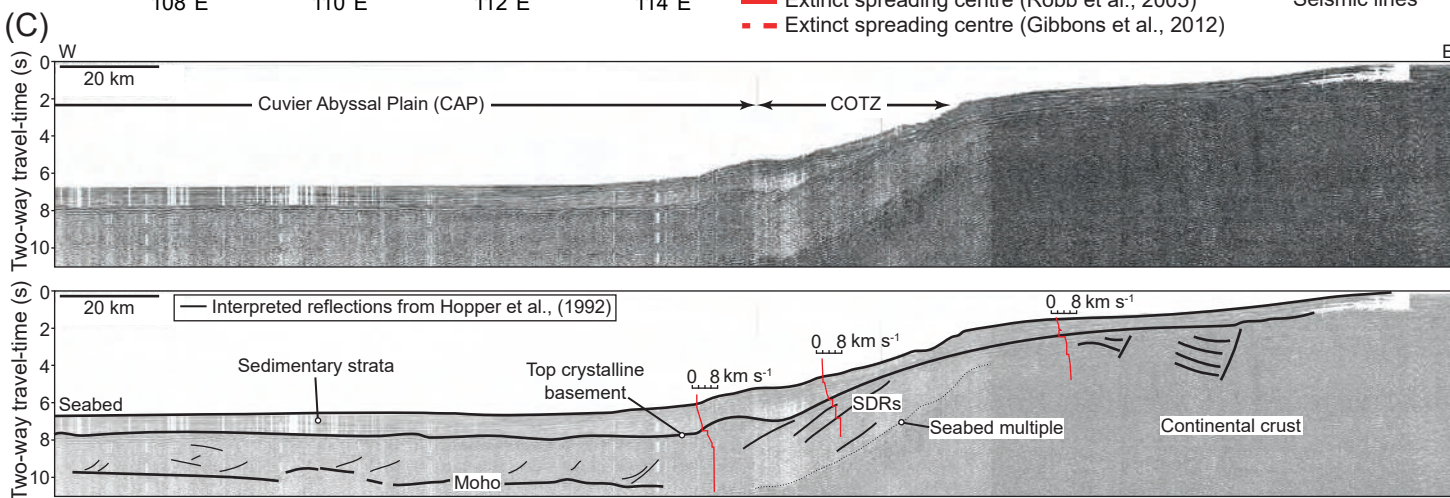
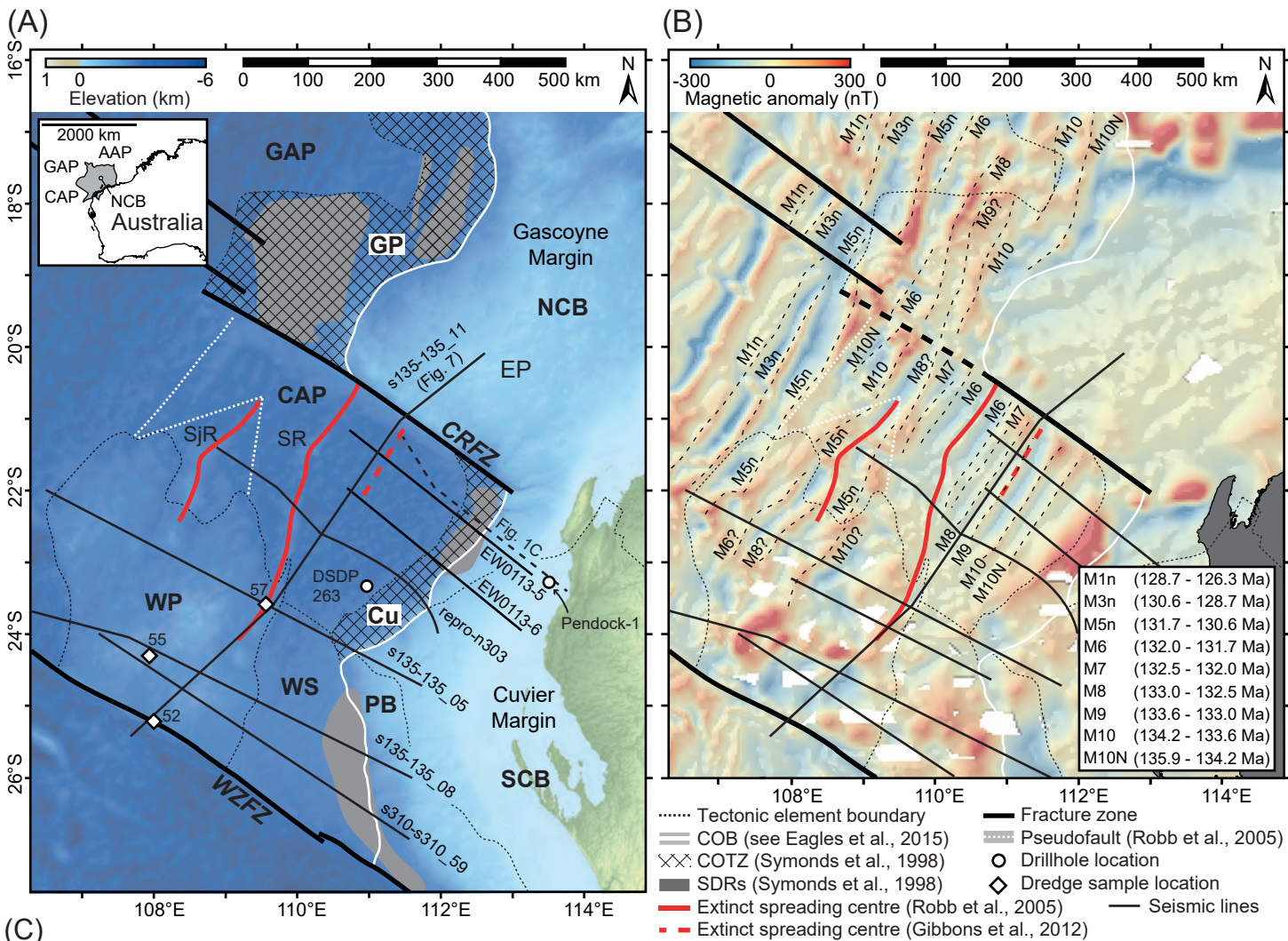


Figure 2

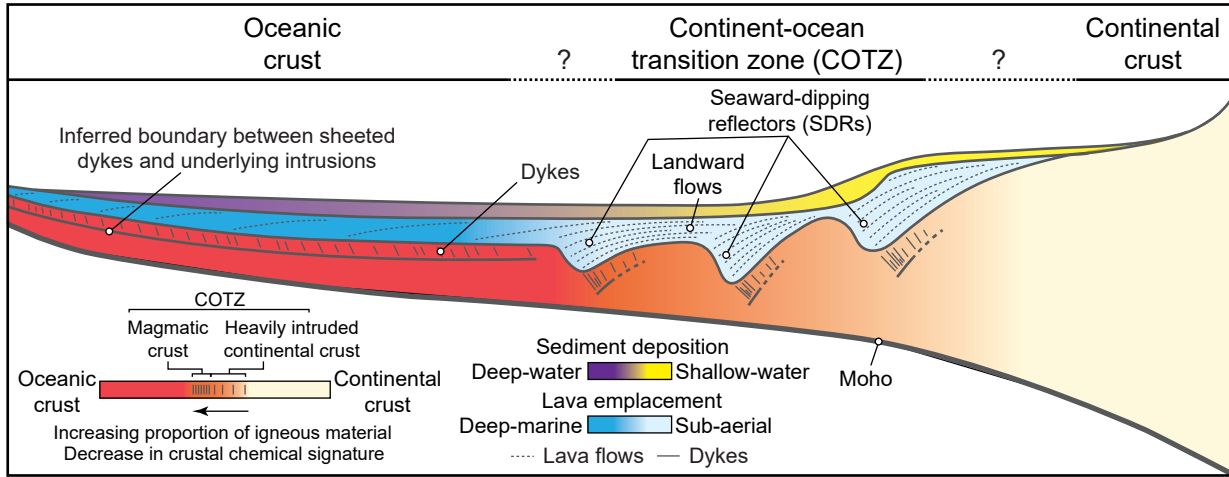


Figure 3

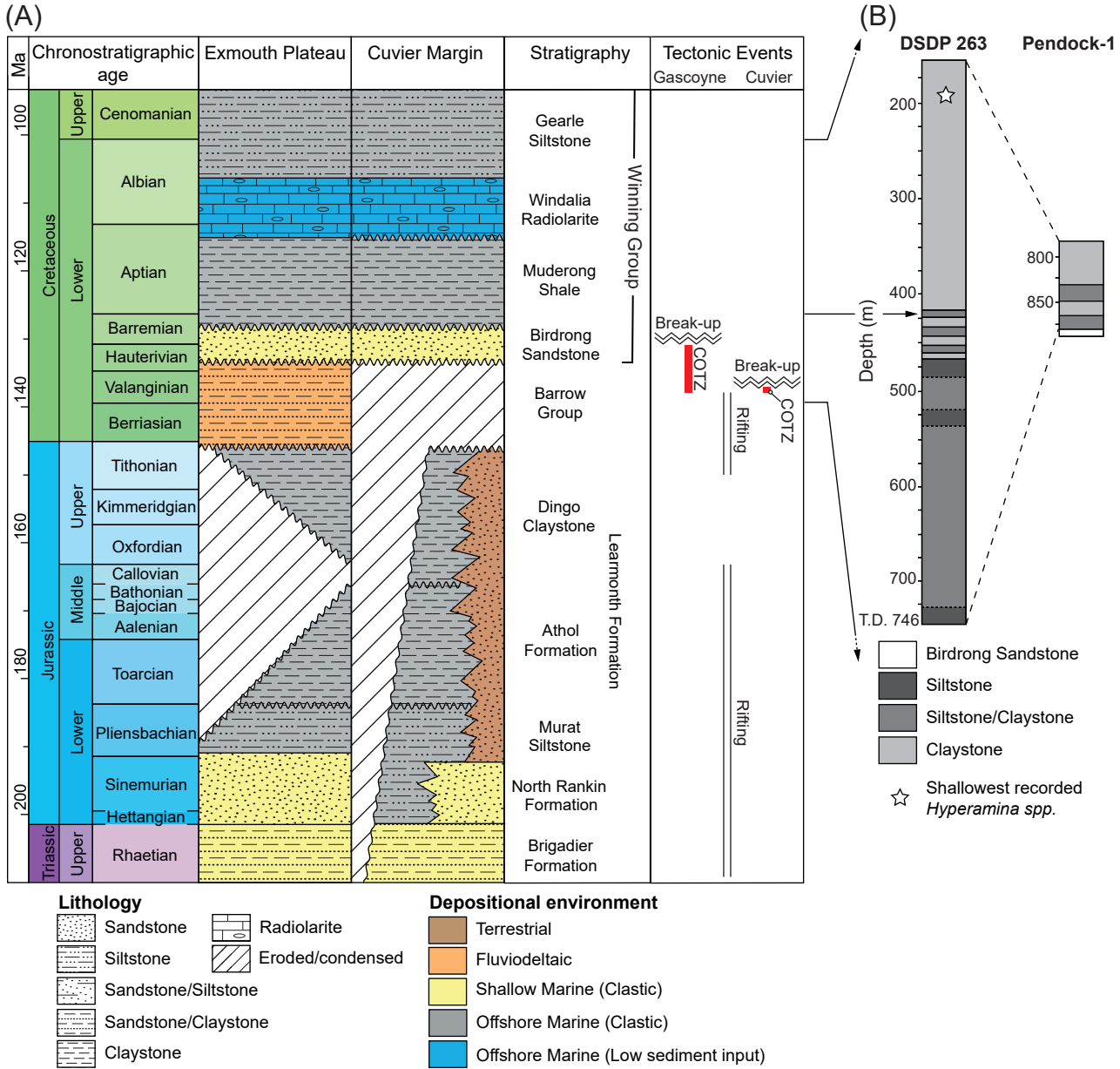


Figure 4

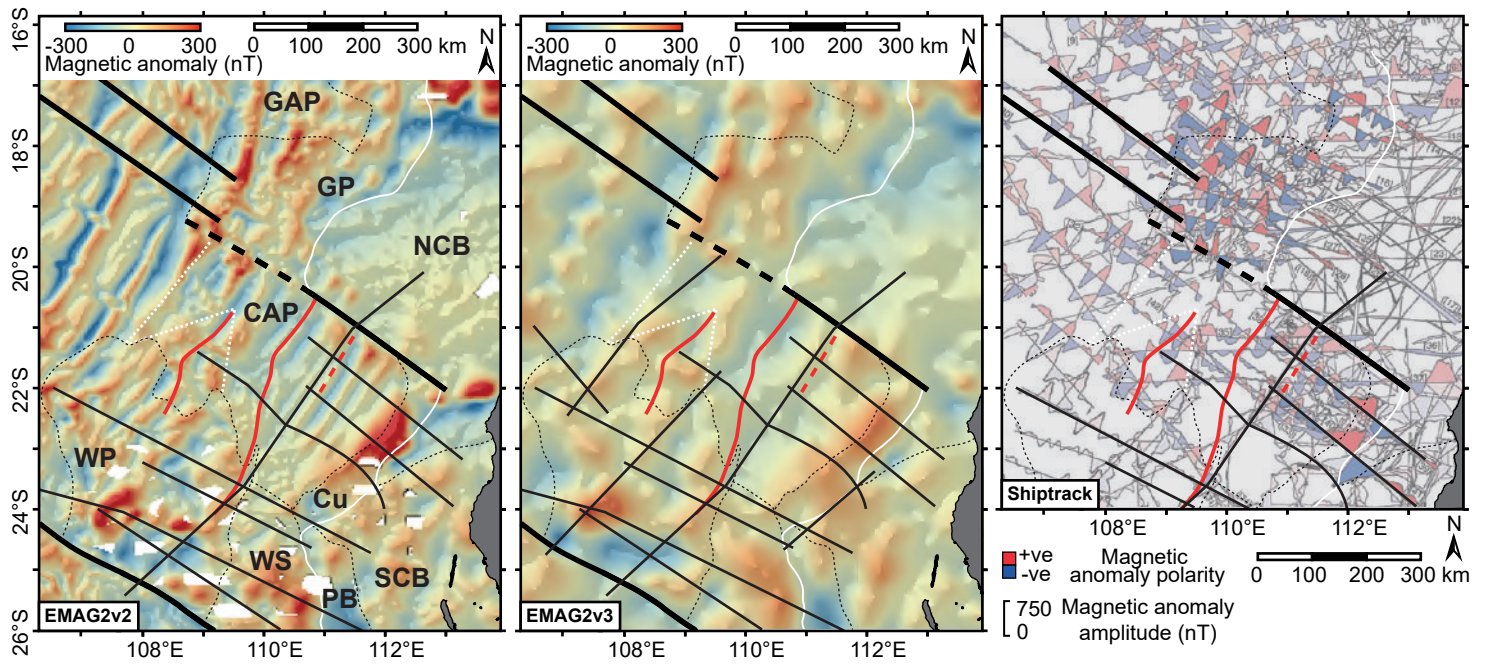


Figure 5

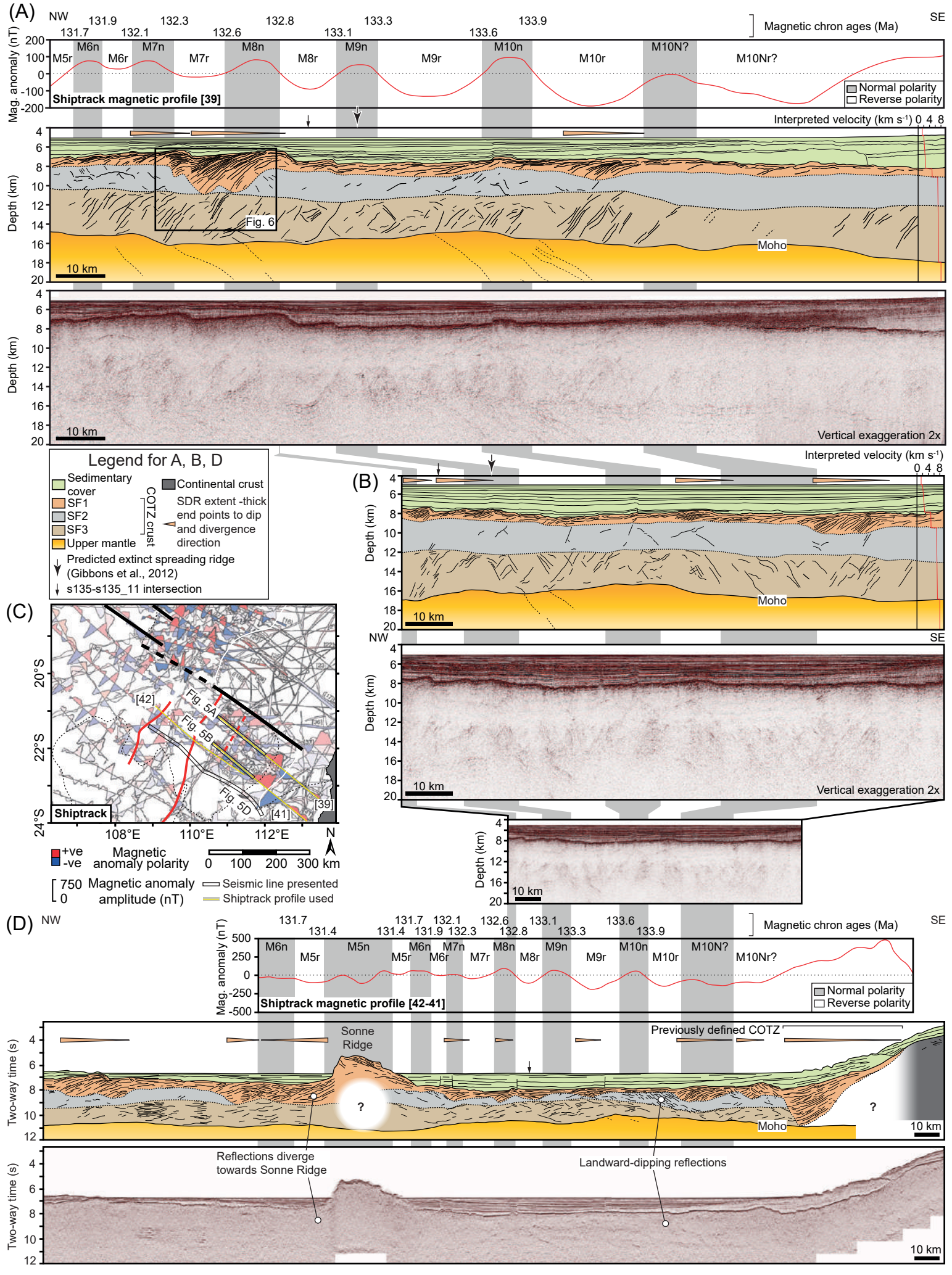


Figure 6

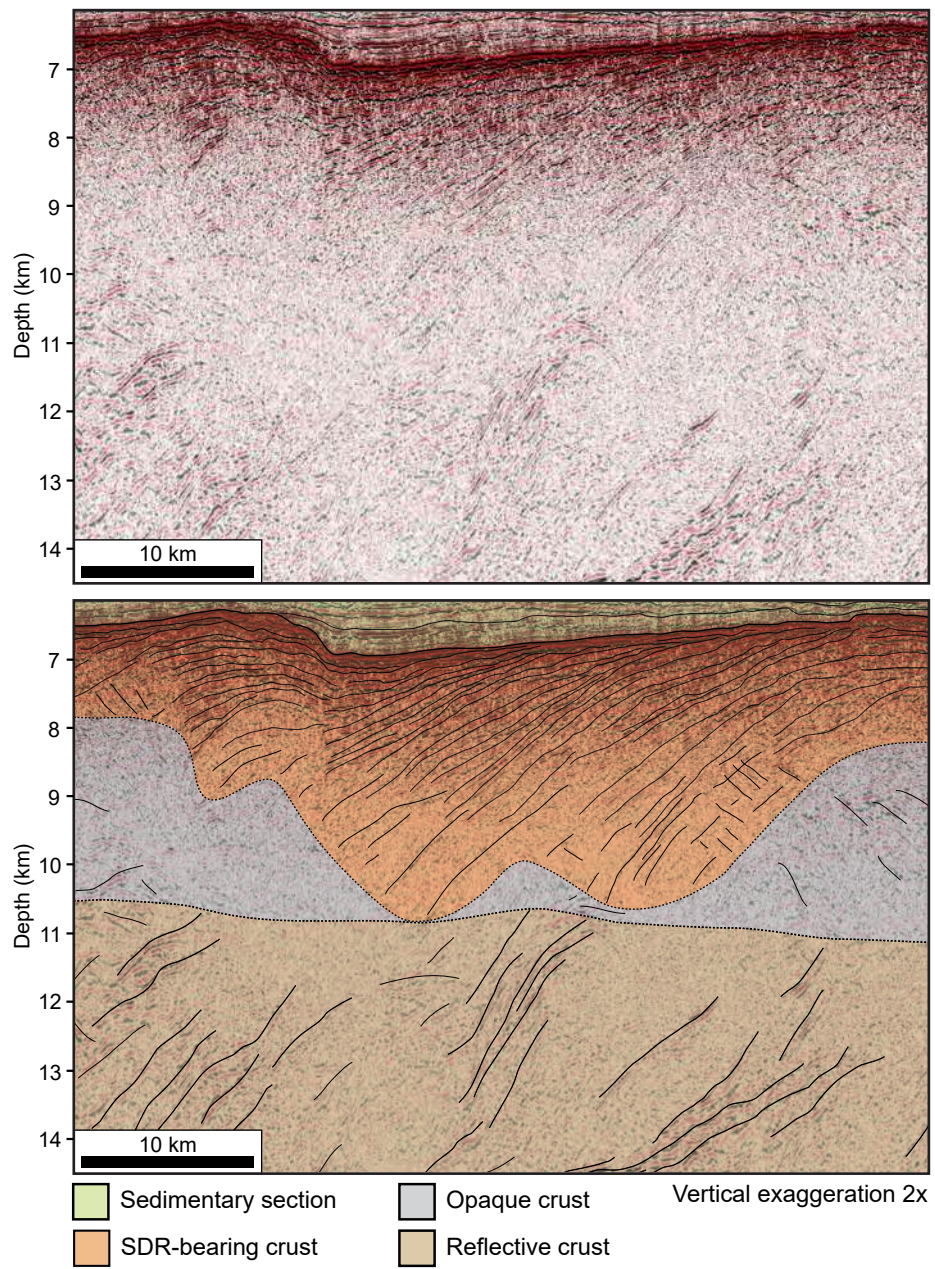


Figure 7

