Deep source regions for Patagonia backarc volcanism imaged by finite frequency body wave tomography

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- 7 Abstract
- 8

9 The subduction of Chile Ridge beneath South America beginning 12-16 Myr ago opened a gap 10 in the subducting slab beneath southern Patagonia, which migrated northward and is located 11 today at 46°S. Geodynamic processes associated with the slab window are poorly understood. 12 Here we apply finite-frequency P and S body wave tomography to seismic data from several 13 temporary arrays as well as regional stations to image seismic heterogeneities down to 650 km 14 depth. The results show strong low velocity anomalies extending to 400 km depth beneath recent 15 back-arc volcanism between 46°S and 48°S, suggesting a link to thermal upwelling in the upper 16 mantle. The southern edge of the Nazca slab extends aseismically down to at least 350 km and 17 has steeper dip than previously suggested. We also image low upper mantle seismic velocities 18 beneath the Patagonia icefields, suggesting low viscosity modulates the patterns of uplift and 19 horizontal deformation observed by GNSS.

20

21 Plain summary

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23 The presence of a tear in the downgoing plate brings hot material from the asthenosphere and 24 modifies locally the mantle flow associated with the subduction zone. Patagonia is a good 25 example where the subduction of the spreading ridge created an opening in the slab. Using arrival 26 times of P and S waves, we develop a new seismic model of the mantle in this region. We show 27 strong back-arc slow velocity anomalies extending to 400 km depth beneath recent back arc 28 volcanism between 46°S and 48°S, suggesting this volcanism is linked to upwelling in the upper 29 mantle. We show that the southern edge of the Nazca slab extends to at least 350 km and has 30 steeper dip than previously suggested. Additionally, our new model suggests lateral variations in 31 mantle viscosity that help to explain the present-day bedrock deformation across the Patagonia 32 icefield.

33 1. Introduction

The subduction of a spreading ridge opens a window in the subducting slab, allowing the 35 36 asthenosphere to upwell and interact with the shallow mantle and crust. Slab windows have a 37 profound influence on geodynamics processes, mantle flow pattern, and thermal structure of 38 subduction zones [Groome and Thorkelson, 2009; Sanhueza et al, 2023]. Magmatism in a slab 39 window is distinctly different from that found in typical volcanic arcs, with adakitic melts associated 40 with slab edges and a volcanic gap in the center of the window [Thorkelson and Breitsprecher, 41 2005]. Magmatism may shift to the backarc region, with tholeiitic lavas presumably due to mantle 42 upwelling observed [Gorring et al, 1997; Thorkelson and Breitsprecher, 2005]. However, it is 43 unclear whether this magmatism results from uppermost mantle processes or is instead due to 44 deeper mantle upwelling.

45

46 The Patagonia slab window (Figure 1a) formed 14 Ma in Southern Patagonia with subduction of 47 the Chilean spreading ridge beneath South America is a good example of a migrated slab window 48 [Breitsprecher and Thorkelson, 2009]. Geochemical studies identified the variation of 49 geochemical signature of magma along the subduction with a transition from adakitic to basaltic 50 signature in the slab melt and presence of an asthenospheric window [Thorkelson and 51 Breitsprecher, 2005]. The back arc region shows widespread recent volcanic activity with the 52 development of massive plateau lavas [Ramos and Kay 1992; Gorring et al. 1997; Guivel et al., 53 2006, Guest et al, 2024]. From analogue modelling and the compilation of available ages of back 54 arc magmatism and kinematic reconstruction of the edge of the Antarctic slab, Guillaume et al. 55 (2010) suggest that recent volcanism could be due to lateral flow of sub-slab mantle.

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57 The Patagonian slab window underlies the Northern and Southern Patagonian Icefields (NPI and 58 SPI), so the thermal and rheological perturbations resulting from the window may be important 59 for the solid earth response to the ice load [Klemann et al, 2007; Lange et al, 2014; Mark et al, 60 2022; Russo et al, 2022]. Currently the NPI and SPI are undergoing rapid ice mass loss [Richter 61 et al. 2019], associated with a fast bedrock uplift [Dietrich et al. 2010; Richter et al. 2016]. The 62 intensity of the solid-earth response to the glacial retreat can be explained by the regional scale 63 rheology with low viscosity in the asthenosphere and a thin lithosphere [Lange et al. 2014; Mark 64 et al. 2022; Russo et al. 2022].

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A recent regional scale Rayleigh wave imaging study [Mark et al. 2022] shows the presence of a slow velocity anomaly in the uppermost mantle which is interpreted as a thermal erosion of the lithospheric mantle over the youngest part of the slab window. From shear wave splitting analysis, Ben-Mansour et al. (2022) found a strong anisotropy with low uppermost mantle shear velocities and an absence of mantle lithosphere. In addition, other studies have imaged the structure immediately beneath the Chile Triple Junction (CTJ) [Russo et al, 2010; Miller et al, 2023] and provided large-scale regional tomography [Portner et al, 2020; Kondo et al, 2024]. 81 Here we use the data-adaptive, multiscale tomographic approach of Hung et al. [2011] and 82 relative sensitivity kernels to combine asynchronous datasets from several different temporary 83 deployments [Maupin 2021] to map the 3-dimensional structure of the Patagonia slab window. 84 The new P and SH seismic velocity models yield broad regional coverage, with better resolution 85 than large-scale models and extend to greater depths than Rayleigh wave models, which are 86 limited to the upper 150-200 km. We provide new constraints on the present-day state of the 87 mantle and how it contributes to back arc magmatism. We also provide additional constraints on 88 the manner in which earth structure contributes to the fast crustal uplift beneath the Patagonia 89 icefields.

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- 91 2. Data and methods
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2.1. Broadband seismic data and processing

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The dataset for this study (**Figure 1**) consists of seismograms from the GUANACO network of twenty-six broadband seismographs deployed from 2018-2021 [Wiens and Magnani, 2018], as well as previous deployments in the Chile triple junction region [Russo et al, 2010; Miller et al, 2023]. Additionally, we use permanent stations from the Chile Seismic Network [Barrientos and CSN team, 2018] and the global network. We used three component seismic waveforms of these stations from teleseismic events between 28° and 98° of epicentral distance and magnitudes greater than 5.4.

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103 Numerous studies show that body wave travel time delays are frequency dependent [Hung et al. 104 2004; Hung et al.; 2011; Kolstrup and Maupin 2015]. To account for frequency-dependent delays, 105 we measure direct P and SH wave travel times on waveforms band-pass-filtered around the 106 secondary seismic noise peak (~0.2 Hz). We use two frequency bands for P-waves (0.03-0.125 107 Hz and 0.3-1.5 Hz) and one frequency band for S-waves (0.03-0.125 Hz). Relative travel times in 108 different frequency bands are estimated from an automated processing procedure [Kolstrup and 109 Maupin, 2015] combining an iterative cross-correlation and stack algorithm (ICCS) [Lou et al., 110 2013] and the multichannel cross-correlation method [Van Decar and Crosson, 1990].

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Data are corrected for crustal delay times using the crustal model from Mark et al. (2022) (Figure S1). To ensure a common reference time and include the effect of the topography, the crustal corrections are computed at each station by considering a model from 50 km below sea level and up to the free surface. In regions of thick sediments, reverberations and their interference introduce a significant frequency dependence of the traveltimes measured on band-passed
waveforms [Yang and Shen, 2006; Ritsema et al., 2009; Kolstrup and Maupin, 2015]. The crustal
correction of finite-frequency travel time residuals should therefore be made frequency
dependent. This is done here following the procedure of Kolstrup and Maupin (2015).

- 120
- 121 2.2. Finite frequency body wave tomography
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To invert the relative travel times to obtain relative seismic velocity anomalies, we use the 3-D Born-Fréchet kernels, which link the influence of velocity heterogeneities and finite frequency travel time shifts [Dahlen et al., 2000; Hung et al. 2000; Schmandt and Humphreys, 2010; Maupin and Kolstrup, 2015]. The kernels are demeaned to form relative kernels to correct for the unevenness in space and time of the data distribution [Maupin, 2021].

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129 We use the data-adaptive, multiscale, finite-frequency tomography method of Hung et al (2011) 130 to take advantage of the distribution of seismic stations and the azimuthal distribution of 131 teleseismic events. The model is parametrized in terms of wavelet basis functions in 3-D, where 132 the calculations assume an isotropic Earth. The model has 65x33 points in the horizontal plane 133 with interspacing of 0.31°. The model extends from 50 to 650km depth with a 37.5 km grid spacing, 134 but shallower and deeper models have also been tested (Figure S2). The inversion includes 135 station terms, but tests with different weights for those show that they do not significantly affect 136 the inverted model. As damping acts on the wavelet coefficients, there is no norm or gradient 137 damping, but an ideal preservation of the resolved elements in the model [Hung et al. 2011, 138 Kolstrup et al. 2015]. The resulting models do not provide absolute velocities but only velocity 139 anomalies relative to an unknown average 1-D model. The optimal models for P and S are 140 determined by selecting the models with the best compromise between model norm and variance 141 on one side and data misfit reduction on the other side. The misfit reduction is high, well above 142 60% even for smooth models, and 88% and 86% for the chosen P and S models respectively 143 (Figure S3).

- 145 3. <u>Results</u>
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147 3.1. P and SH relative velocity models

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149 The new finite-frequency P and SH models provide improved resolution and image the mantle 150 down to 650 km (Figure 2). Overall, the new models show larger velocity contrasts compared to 151 previous results [Portner et al. 2022; Kondo et al. 2024], but the tomographic inversion used here 152 conserves amplitudes quite well compared to other methods. The major patterns of the P and 153 SH models are similar, with the magnitude of the S wave anomalies larger than the P wave 154 anomalies as expected. There are some differences between the P and SH models in the detailed 155 structure of the anomalies. These differences could be due to the unequal effects of the 156 compositional and thermal anomalies on P and S wave velocities. However, because these 157 differences could alternatively result from differences in data volume, signal to noise ratio, and 158 frequency content, we do not extensively describe and interpret these differences and reserve 159 most of our interpretation to the larger patterns reflected in both P and SH models. (Figure 2).

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161 The most prominent velocity anomaly is a large region of slow P and SH velocities anomalies (-162 4/-6%) between 46°S and 49°S (Figures 2 and 3), centered beneath the Meseta del Lago Buenos 163 Aires volcanic region [Guivel et al., 2006, Guest et al, 2024]. Recent surface wave [Mark et al. 164 2022] and body wave [Russo et al. 2010, Portner et al., 2020; Rodriguez et al., 2021; Miller et al. 165 2023; Kondo et al. 2024] seismic models have shown the presence of a similar anomaly in this 166 region but have not resolved its lateral and depth extent. Both P and S wave models show it 167 extending down to at least 400 km, where it is underlain by faster velocities in the P-wave model. 168 The anomaly is truncated in the north by faster velocities likely associated with the subducting 169 Nazca slab. To the east, the lateral extent of the anomaly is limited at shallower depths (< 200 170 km) by the lithospheric mantle of the Deseado Massif (Figures 2 and 3). The E-W cross section 171 C-C' suggests that this anomaly extends eastward beneath the Deseado Massif but the absence 172 of seismic stations in Argentina limits the interpretation of anomalies in this part of the model. 173



176Figure 2. Map views of $\delta V_P/V_P$ (top) and $\delta V_S/V_S$ (bottom) at 120km, 200km, 280km, 350 km and 420 km177depth. The dashed black line represents the surface extension of the Deseado Massif, green dashed line178represents the maximum extension of the slab window and yellow contours are the location of the Northern

179 and Southern Patagonia Icefield.



Figure 3. Cross-sections across P model. Green stars represent earthquakes from the USGS
catalogue. The tectonic features on the map follow the same conventions as in Figure 2.

The northern part of the model region south to 46°S is dominated by fast seismic velocities likely 185 186 associated with the subducting Nazca slab north of the triple junction. This is supported by the 187 nice correlation with the Wadati-Benioff seismicity of the Nazca slab as seen in the E-W cross 188 section A-A' in Figure 3. There is a clear contrast between a fast anomaly (3-4%) north of the 189 CTJ and a slow velocity anomaly (3-4%) in the south. This contrast matches the present-day 190 location of the slab edge from the paleo-reconstruction of the Patagonia slab window by 191 Breisprecher and Thorkelson (2009). It matches also the boundary between a seismically active 192 region to the North and a quiescent region to the South, where the absence of seismicity has 193 been interpreted as the absence of a slab beneath south Patagonia. The P model shows clearly 194 that the fast relative velocity anomalies to the north of the CTJ (green dot on Figure 2) extends 195 eastward and down to 350 km. Due to the nature of S waves and their frequency content, the SH 196 model does not constraint the shape of the fast anomaly as well.

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198 3.2 Resolution tests

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200 Checkerboard tests are used to evaluate the horizontal and vertical (Figures S4-S6) resolution 201 in our datasets. For each synthetic model, we test different sizes of checkers and velocity 202 perturbations (+/- 4%). We add Gaussian noise to the synthetic data prior to inversion with 203 standard deviation of 0.04 s and 0.14 s for P and S waves, respectively, as estimated from the 204 actual data. The inversion is performed using the same parameters and damping as used with 205 real data. The checkerboard tests show good recovery down to 420 km for both P waves (Figures 206 S4 and S5) and SH waves (Figures S4 and S6). Tests also show that the P-waves datasets can 207 resolve anomalies that are about two times smaller than the SH waves datasets.

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209 In addition, we test several scenarios concerning the geometry and extension of the Nazca slab 210 (Figures S7 and S8), the lateral dimensions of the slow velocity perturbation beneath 46°S and 211 48°S (Figures S9-S10), and the presence of two distinct shallow and deeper slow velocity 212 perturbations at the eastern edge of the study (Figures S11). Results show that a deeper and 213 steep Nazca slab is better resolved down to 350km than a one segment slab for both P waves 214 and SH waves (profile A-A' Figures S8). Tests on slow velocity perturbations between 46°S and 215 48°S (Figures S10-S11) show a good recovery in P and SH waves down to 420km and the ability 216 to distinguish two distinct slow velocity anomalies in this region.

A comparison of our Vsh model with the Rayleigh wave Vsv model from Mark et al. (2022)
 (Figures S12) reveals relatively good agreement regarding the spatial distribution of fast and slow
 velocity anomalies across the slab window.

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221 4. Discussion

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4.1. Slow velocity mantle anomalies associated with backarc volcanism

The tomographic models indicate an association between slow velocity mantle anomalies and backarc volcanism (**Figure 2**). Southern Patagonia exhibits extensive back arc volcanism, likely associated with the northward migration of a slab window over the last 14 Myr [Breitsprecher and Thorkelson, 2009; Guillaume et al. 2010]. The source of this magmatism and its distinctive geochemical signatures, as well as its relationship to the slab window are not well understood [Gorring et al, 2003; Guest et al. 2024]. 230 A strong slow velocity mantle anomaly is located beneath the Meseta del Lago Buenos Aires 231 volcanic region, a locus of much of the Pleistocene backarc volcanism (Figure 1). Continental 232 backarc volcanism is often assumed to occur by decompression melting from lithospheric 233 extension, with shallow sources in the upper 100 km [Vasey et al, 2021]. In contrast, the slow 234 velocity seismic anomaly beneath the Mesta del Lago Buenos Aires extends to depths of at least 235 400 km, indicating a deeper geodynamic process. The slow velocity anomalies also indicate a 236 connection eastward towards the Atlantic coast at depths greater than 250 km, although 237 resolution of the eastward extent is limited by the absence of seismic stations in eastern Argentine 238 Patagonia. An eastward mantle connection is contrary to the typical assumption that mantle flows 239 through the slab window from west to east [Russo et al, 2010; Hu et al, 2017; Ben Mansour et al, 240 2022]. However, it is consistent with the geochemical affinity of Meseta del Lago Buenos Aires 241 volcanics with South Atlantic mantle plume sources [Soager et al, 2021; Guest et al, 2024].

242 The slow velocity anomalies beneath the backarc volcanoes are truncated to the east at depths 243 shallower than 200 km by the Deseado Massif, which is incompletely imaged in the new models. 244 The Deseado Massif is part of the Malvinas/Falkland Islands lithospheric block, a Precambrian 245 terrain with strong mantle isotopic affinities with the Namaqua-Natal belt in south Africa [Marshall, 246 1994; Mundl et al., 2015; Schilling et al. 2017]. This lithospheric block is imaged farther to the 247 south (52°-54°) with higher velocities in the upper 200 km, consistent with a thicker continental 248 lithosphere (Mark et al. 2022, and Figure 2). The thicker lithosphere of the Deseado Massif may 249 partially inhibit deep seated volcanism and localize the Meseta del Lago Buenos Aires volcanics 250 along its western boundary. Petrological and geochemical analyses of xenoliths from the 251 lithospheric mantle support the idea of mantle heterogeneity inherited from South Atlantic 252 hotspots (Discovery-Shona-Bouvet) [Soager et al. 2021; Mallick et al. 2023; Jalowitzki et al. 2024] 253 prior to the opening of the Atlantic Ocean. Contamination of melts by the lithospheric mantle with 254 Atlantic geochemical affinity provides an alternative explanation for the Atlantic mantle plume 255 signature in the Mesata del Lago Buenos Aires volcanics. Interestingly, the Pali Aike volcanic field 256 (PAV, Figure 1), with extensive Pleistocene and Holocene activity, is located on the interior of 257 the Malvinas/Falkland lithospheric block [Shilling et al, 2017]. The occurrence of these volcanics 258 in a region of thicker lithosphere, as imaged by high seismic velocities in the upper 200 km (Figure 259 2), demonstrates that backarc volcanics can erupt even on the interior of older lithospheric blocks, 260 possibly aided by fractures or gaps in the lithosphere that are too small to image in this study.

Between 50°S and 52°S, recent magmatism (< 2Myrs) has been recorded in the region of Cerro
del Fraile. This volcanism is located on the top of a slow velocity anomaly in the shallow mantle
(~ 120-160 km depth). Our models are consistent with geochemical analysis of mantle xenoliths

264 showing metasomatism of adakitic basalt and interpreted as the magmatic signature of the 265 subduction of the current oceanic crust of the Antarctica Plate [Killian and Stern 2002; Stern 2004]. 266

- 267

4.2. The geometry of the aseismic southernmost Nazca slab

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269 The present-day geometry of the southern edge of the Nazca slab is poorly constrained. 270 Intermediate-depth earthquake hypocenters extend to depths of about 150 km. The depth extent 271 of seismicity is limited by warm slab temperatures due to the slab's young lithospheric age, which 272 is 7 Ma at the 45°S trench axis and increases northward [Tebbens et al, 1997]. Breitsprecher and 273 Thorkelson (2009) estimated the position of the slab edge using plate reconstructions and 274 assuming a poorly constrained shallow dip angle. The updated Nazca slab model by Portner et 275 al. (2020) suggests an aseismic extension of the Nazca slab down to ~300 km in this region.

276

277 By combining data from all previous temporary seismic deployments with the new Chilean seismic 278 network, we have a larger picture of the lateral and depth extent of the fast anomaly associated 279 with the Nazca slab. The new P-wave tomographic model shows that the slab dip increases 280 substantially at depths of 150-200 km, becoming steeply dipping at about 300 km depth (Figure 281 3). The lower resolution S-wave model shows high velocity slab material north of the triple junction 282 but does not define the geometry as well. Due to the steep dip, the slab does not extend east of 283 69°W within the upper 500 km, in contrast to the reconstruction in Breitsprecher and Thorkelson 284 [2009], which shows the slab underlying the Atlantic coast at about 300 km depth. As a result of 285 the steeper dip, as well as the northward motion of the triple junction over the past 12 Myrs, 286 considerable Nazca slab material may lie in the transition zone beneath the Deseado Massif. 287 Particularly at larger depths of around 300km, the high-velocity region associated with the slab 288 moves eastwards when approaching the slab edge. This is reminiscent of the slab deformation 289 by mantle toroidal flow around slab edges in the numerical experiments of Kiraly et al. (2017).

290

291 4.3. Rheological heterogeneity beneath the Patagonia icefields

292 The NPI and SPI are located on top of the slab window between 46°S and 51°S. Their ice mass 293 has been responding very sensitively to changing climate and atmospheric circulation. Changes 294 of the ice load drive a delayed bedrock deformation due to glacial isostatic adjustment (GIA). 295 According to GIA models, a homogeneous viscoelastic half space responds to an ice-mass loss 296 with an elliptical geographic pattern through a) uplift with a concentrical pattern about the 297 barycenter of the lost ice mass, and b) a symmetrical horizontal extension radially away from that 298 barycenter. The effect of a reduced mantle viscosity of that half space is that the new isostatic

equilibrium be restored in a shorter time, implying larger deformation rates [Weerdesteijn et al,
2022]. A localized low-viscosity anomaly, eccentrical with respect to the ice-load symmetry,
enhances the horizontal deformation rates in the direction from the load center towards that
anomaly [Kaufmann et al. 2005; Klemann et al. 2007].

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304 GNSS observations in the SPI area reveal intense uplift, with rates reaching 4 cm/year in the 305 northern part of the SPI, consistent with a GIA model assuming a homogeneous half space 306 [Dietrich et al. 2010; Lange et al. 2014; Richter et al. 2016]. However, the observed uplift rates 307 are smaller than the modelled ones in the southern part of SPI, south of 50°S, and larger than 308 predicted by the simple GIA model in the northern part of SPI. Furthermore, the observed 309 horizontal deformation is not symmetric about the N-S axis of the icefields, with a stronger 310 eastward extension to the east of the load center compared to the westward extension to the west 311 [Richter et al. 2016].

312

313 Seismic models offer an opportunity to estimate lateral changes in mantle viscosity, since both 314 velocity and viscosity are largely controlled by temperature [lvins et al, 2023]. However, recent 315 detailed Rayleigh wave seismic structure models for southern Patagonia are limited to depths of 316 about 150 km [Mark et al, 2022]. The new models presented here can extend our understanding 317 of rheological variations in this region to deeper depths and can thus provide explanations for the 318 discrepancies between the observed deformation patterns and those predicted by GIA models 319 based on homogenous rheology. The new seismic models show a north-south dichotomy, with 320 the NPI and the northern part of the SPI underlain by low velocities over most of the upper 300 321 km, and the southern SPI underlain by faster velocities (Figure 4). This is in agreement with the 322 Rayleigh wave results of Mark et al (2022) but shows that this dichotomy extends deeper than 323 previously imaged. This dichotomy explains the increase from South to North of the observed 324 uplift compared to the homogeneous model, as a consequence of the decrease in mantle viscosity 325 in that direction, provided that the GIA model assumes a correct ice-load distribution. The new 326 seismic models also show very strong low-velocity anomalies in the backarc near both icefields 327 at 120 km depth, and for the NPI at deeper depths. These velocities anomalies can explain the 328 asymmetric horizontal deformation pattern as a modulation of the GIA-driven radial extension by 329 a localized, low viscosity just east of the ice-load axis.





Figure 4. Enlarged view of the tomographic model in the vicinity of the Patagonia icefields at depths
of 120km, 200km, 280km, 350 km and 420km. The tectonic features on the map follow the same
conventions as in Figure 2.

331

336 5. Conclusions

Ρ

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338 We present new regional scale finite-frequency P and SH waves tomography models of southern 339 Patagonia, using a wavelet-based multiscale parametrization that increases the recovery of 340 amplitude and location of seismic velocity anomalies. The new models show a fast velocity anomaly north of the CTJ associated with the aseismic southernmost Nazca slab and a slow 341 342 velocity anomaly in southern Patagonia related to the present-day location of the slab window. 343 We find that slow velocity anomalies beneath back-arc volcanoes in the northern part of the slab 344 window extend to depths of ~ 400 km, suggesting that the volcanism does not result from shallow 345 mantle processes but is instead linked to upwelling thermal anomalies from deeper regions of the

- 346 upper mantle. We also show that the southern edge of the Nazca slab extends downward to at 347 least 350km and has a steeper slope than in previous models. Finally, the new tomographic 348 models illuminate the structural heterogeneity beneath the Patagonia icefields. Our models show 349 a north-south dichotomy, with the NPI and the northern part of the SPI underlain by low velocities 350 and the southern SPI underlain by faster velocities down to depths of about 300 km. Very strong 351 low-velocity anomalies in the backarc near both icefields support the idea of that more complex 352 rheological models may be able to explain geodetic observations in Patagonia.
- 353
- 354 Open Research
- 355 Data used in this study are from the temporary seismic networks GUANACO, SEARCH and CRSP
- 356 temporary seismic networks respectively under the network code 1P, YJ and XJ. Permanents
- 357 stations from the Chile network and Global network are under the network code C/C1 and GT.
- 358 Data can be obtained from the EarthScope data center:
- 359 1P: <u>https://www.fdsn.org/networks/detail/1P_2018/</u>,
- 360 XJ: <u>https://www.fdsn.org/networks/detail/XJ_2004/</u>,
- 361 YJ: https://www.fdsn.org/networks/detail/YJ_2004/,
- 362 C: https://www.fdsn.org/networks/detail/C/,
- 363 C1: https://www.fdsn.org/networks/detail/C1/
- 364 GT: <u>https://www.fdsn.org/networks/detail/GT/</u>).
- Tomography models showed and discussed in this study can be found in Ben-Mansour and Maupin (2025).
- 367
- 368 Acknowledgements

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- 379 Figures were produced using GMT software [Wessel et al., 2013] using perceptually uniform
- 380 colour maps generated by Fabio Crameri [Crameri, 2018].

382 References 383 384 Barrientos, S, National Seismological Center (CSN) Team (2018), The Seismic Network of Chile. 385 Seismological Research Letters, 89, 2A, 467-474. 386 387 Ben-Mansour, W, Wiens, DA, Mark, HF, Russo, RM, Richter, A, Marderwald, E, Barrientos, S 388 (2022), Mantle flow pattern associated with the Patagonian slab window determined from 389 azimuthal anisotropy, Geophysical Research Letters, 49,18, e2022GL099871. 390 391 Ben-Mansour, W, Maupin, V (2025), Deep source regions for Patagonia backarc volcanism 392 imaged by finite frequency body wave tomography [Dataset & Software]. Zenodo. 393 https://doi.org/10.5281/zenodo.10028616 394 Breitsprecher, K, Thorkelson, DJ (2009), Neogene kinematic history of Nazca-Antarctic - Phoenix 395 slab windows beneath Patagonia and the Antarctic Peninsula, Tectonophysics, 464, 1-4, 10-20. 396 397 Crameri, F (2018). Scientific colour-maps: perceptually uniform and colour blind friendly, 398 www.fabiocrameri.ch/colourmaps. 399 400 Dahlen, FA and Hung, S-H and Nolet, Guust (2000), Frechet kernels for finite-frequency 401 traveltimes: I. Theory, Geophysical Journal International, 141, 1, 157-174. 402 403 Dietrich, R, Ivins, ER, Casassa, G, Lange, H, Wendt, J, Fritsche, M (2010), Rapid crustal uplift 404 in Patagonia due to enhanced ice loss, Earth and Planetary Science Letters, 2898, 1-2, 22-29. 405 406 Gorring, ML, Kay, SM, Zeitler, PK, Ramos, VA, Rubiolo, D, Fernandez, MI, Panza, JL (1997), 407 Neogene Patagonian plateau lavas: continental magmas associated with ridge collision at the 408 Chile Triple Junction, Tectonics, 16, 1, 1--17. 409 410 Gorring, M, Singer, B, Gowers, J, Kay, Suzanne M (2003), Plio--Pleistocene basalts from the 411 Meseta del Lago Buenos Aires, Argentina: evidence for asthenosphere--lithosphere interactions 412 during slab window magmatism, Chemical Geology, 193, 3-4, 215-235. 413

414 Groome, WG and Thorkelson, DJ (2009), The three-dimensional thermo-mechanical signature of 415 ridge subduction and slab window migration, Tectonophysiccs, 464, 1-4, 70-83. 416 417 Guest, IA, Saal, AE, Mallick, S, Gorring, ML, Kay, SM (2024), The volcanism of the Meseta del 418 Lago Buenos Aires, Patagonia: the transition from subduction to slab window, Journal of 419 Petrology, 65, 6. 420 421 Guillaume, B, Moroni, M, Funiciello, F, Martinod, J, Faccenna, C (2010), Mantle flow and dynamic 422 topography associated with slab window opening: Insights from laboratory models, 423 Tectonophysics, 496,1-4,83-98. 424 425 Guivel, C, Morata, D, Pelleter, E, Espinoza, F, Maury, RC, Lagabrielle, Y, Polve M, Bellon, H, 426 Cotten, J, Benoit, M, SUarez, M, De la Cruz, R (2006), Miocene to Late Quaternary Patagonian 427 basalts (46--47 S): geochronometric and geochemical evidence for slab tearing due to active 428 spreading ridge subduction, Journal of Volcanology and Geothermal Research, 149, 3-4, 346--429 370. 430 431 Hu, Jiashun and Faccenda, Manuele and Liu, Lijun (2017), Subduction-controlled mantle flow 432 and seismic anisotropy in South America, Earth and Planetary Science Letters, 470, 13-24. 433 434 Hung, S-H and Dahlen, FA and Nolet, Guust (2000), Frechet kernels for finite-frequency 435 traveltimes: II. Examples, Geophysical Journal International, 141, 1,175-203. 436 437 Hung, SH, Chen, WP, Chiao, LY (2011), A data-adaptive, multiscale approach of finite-frequency, 438 traveltime tomography with special reference to P and S wave data from central Tibet, Journal of 439 Geophysical Research: Solid Earth, 116, B6. 440 441 Ivins, E. R., W. van der Wal, D. A. Wiens, A. J. Lloyd, and L. Caron (2023). Antarctic upper mantle 442 rheology, The Geochemistry and Geophysics of the Antarctic Mantle, Geological Society, London, 443 Memoirs, 56, 267-294. 444 445 Jalowitzki, T. Sumino, H. Conceicao, RV, Schilling, ME, Bertotto, GW, Tassara, A, Gervasoni, 446 F, Orihashi, Y, Nagao, K, Rocha, MP, Antonio de Freitas Rodrigues R (2024), Pristine helium 447 from the Karoo mantle plume within the shallow asthenosphere beneath Patagonia, Nature

448 Communications, 15,1, 6402.

- Kaufmann, G, Wu, P, Ivins, ER (2005), Lateral viscosity variations beneath Antarctica and their
 implications on regional rebound motions and seismotectonics, Journal of Geodynamics, 39, 2,
 165-181.
- 453
- Kilian, R and Stern, CR (2002), Constraints on the interaction between slab melts and the mantle
 wedge from adakitic glass in peridotite xenoliths, European Journal of Mineralogy, 14, 1, 25-36.
- 456
- 457 Kiraly, A, Capitanio, F, Funiciello, F, Faccenna, C (2017), Subduction induced mantle flow:
 458 Length-scales and orientation of the toroidal cell, Earth and Planetary Science Letters, 479, 284459 297.
- 460

Klemann, V., Ivins, E., Martinec, Z., Wolf, D. (2007). Models of active glacial isostasy roofing
warm subduction: case of the South Patagonian Ice field, Journal of Geophysical Research: Solid
Earth, 112, B09405.

464

Kolstrup, ML, Hung, SH, Maupin, V (2015), Multiscale, finite-frequency P and S tomography of
the upper mantle in the southwestern Fennoscandian Shield, Geophysical Journal International,
202, 1, 190-218.

- Kondo, Y, Obayashi, M, Sugioka, H, Shiobara, H, Ito, A, Shinohara, M, Iwamori, H, Kinoshita, M,
 Miller, M, Tassara, C, Ojeda, J (2024), Seismic image of the central to southern Andean
 subduction zone through finite-frequency tomography, Journal of Geophysical Research: Solid
 Earth, 129,11.
- 472
- 473 Lange, H, Casassa, G, Ivins, ER, Schroder, L, Fritsche, M, Richter, A, Groh, A, Dietrich, R (2014),
- 474 Observed crustal uplift near the Southern Patagonian Icefield constrains improved viscoelastic
- 475 Earth models, Geophysical Research Letters, 41, 3, 805-812.
- 476
- 477 Lou, X, Van Der Lee, S, Lloyd, S (2023), AIMBAT: A python/matplotlib tool for measuring
 478 teleseismic arrival time, Seismological Research Letters, 84, 1, 85-93.
- 479
- 480 Mallick, S, Kuhl, SE, Saal, AE, Klein, EM, Bach, W, Monteleone, BD, Boesenberg, JS (2023),
- 481 Evidence of South American lithosphere mantle beneath the Chile mid-ocean ridge, Earth and
- 482 Planetary Science Letters, 620, 118320.
- 483

484	Mark, HF, Wiens, DA, Ivins, ER, Richter, A, Ben Mansour, W, Magnani, MB, Marderwald, E,
485	Adaros, R, Barrientos, S(2022), Lithospheric erosion in the Patagonian slab window, and
486	implications for glacial isostasy, Geophysical Research Letters, 49, 2, e2021GL096863.
487	
488	Marshall, JEA (1994), The Falkland Islands: a key element in Gondwana paleogeography,
489	Tectonics, 13, 2, 165-181.
490	
491	Maupin, V, Kolstrup, ML (2015), Insights in P-and S-wave relative traveltime tomography from
492	analysing finite-frequency Frechet kernels, Geophysical Journal 202, 3,1581-1598.
493	
494	Maupin, V (2021), Combining asynchronous data sets in regional body-wave tomography,
495	Geophysical Journal International, 224, 1, 401-415.
496	
497	Miller, M, Priestley, K, Tilmann, F, Bataille, K, Iwamori, H (2023), P wave teleseismic tomography
498	of the subducted Chile rise, Journal of South American Earth Sciences, 128, 104474.
499	
500	Mundl, A, Ntaflos, T, Ackerman, L, Bizimis, M, Bjerg, EA., Hauzenberger, CA (2015),
501	Mesoproterozoic and Paleoproterozoic subcontinental lithospheric mantle domains beneath
502	southern Patagonia: Isotopic evidence for its connection to Africa and Antarctica, Geology, 43, 1,
503	39-42.
504	
505	Portner, DE, Rodriguez, EE, Beck, S, Zandt, G, Scire, A, Rocha, MP, Bianchi, MB, Ruiz, M,
506	Francca, GS, Condori, C, Alavadaro , P (2020), Detailed structure of the subducted Nazca slab
507	into the lower mantle derived from continent-scale teleseismic P wave tomography, Journal of
508	Geophysical Research: Solid Earth, 125, 5.
509	
510	Ramos, VA and Kay, SM (1992), Southern Patagonian plateau basalts and deformation: backarc
511	testimony of ridge collisions, Tectonophysics,205, 1-3, 261282.
512	
513	Richter, A, Ivins, E, Lange, H, Mendoza, L, Schroder, L, Hormaechea, JL, Casassa, G,
514	Marderwald, E, Fritsche, M, Perdomo, R, Horwath, M, Dietrich R (2016), Crustal deformation
515	across the Southern Patagonian Icefield observed by GNSS, Earth and Planetary Science
516	Letters, 52, 206-215.
517	

519 R (2019), The rapid and steady mass loss of the Patagonian icefields, throughout the GRACE 520 era: 2002--2017, Remote Sensing, 11, 8, 909. 521 522 Ritsema, J, Van Heijst, HJ, Woodhouse, JH, Deuss, A (2009), Long-period body wave traveltimes 523 through the crust: implication for crustal corrections and seismic tomography, Geophysical 524 Journal International, 179, 2, 1255-1261. 525 526 Rodriguez, EE, Portner, DE, Beck, SL, Rocha, MP, Bianchi, MB, Assumpccao, M, Ruiz, M, 527 Alvarado, P, Condori, C, Lynner, C (2021), Mantle dynamics of the Andean Subduction Zone from 528 continent-scale teleseismic S-wave tomography, Geophysical Journal International, 224, 3, 1553-529 1571. 530 531 Russo, RM (2010), Subduction of the Chile Ridge: Upper mantle structure and flow, Gsa Today, 532 20. 533 534 Russo, RM, Luo, H, Wang, K, Ambrosius, B, Mocanu, V, He, J, James, T, Bevis, M, Fernandes, 535 R (2022), Lateral variation in slab window viscosity inferred from global navigation satellite system 536 (GNSS)--observed uplift due to recent mass loss at Patagonia ice fields, Geology, 50, 1, 111–115. 537 538 Sanhueza, J, Yanez, G, Buck, WR, Araya Vargas, J, Veloso, E (2023), Ridge subduction: 539 Unraveling the consequences linked to a slab window development beneath South America at 540 the Chile Triple Junction, Geochemistry, Geophysics, Geosystems, 24, 9, e2023GC010977. 541 542 Schilling, ME, Carlson, RW, Tassara, A, Conceiccao, RV, Bertotto, GW, Vasquez, M, Munoz, 543 Daniel, Jalowitzki, T, Gervasoni, F, Morata, Di (2017), The origin of Patagonia revealed by Re-544 Os systematics of mantle xenoliths, Precambrian Research, 294, 15-32. 545 546 Schmandt, B and Humphreys, E (2010), Seismic heterogeneity and small-scale convection in the 547 southern California upper mantle, Geochemistry, Geophysics, Geosystems, 11,5. 548 549 Soager, N, Holm, PM, Massaferro, GI, Haller, M, Traun, M K (2021), The Patagonian intraplate 550 basalts: A reflection of the South Atlantic convection cell, Gondwana Research, 91, 40-57. 551 Stern, CR (2004), Active Andean volcanism: its geologic and tectonic setting, Revista geologica

Richter, A, Groh, A, Horwath, M, Ivins, E, Marderwald, E, Hormaechea, JL, Perdomo, R, Dietrich,

552 de Chile,31,2, 161-206.

553	
554	Tebbens, S. F., S. C. Cande, L. Kovacs, J. C. Parra, J. L. LaBrecque, and H. Vergara (1997), The
555	Chile ridge: A tectonic framework, Journal of Geophysical Research: Solid Earth, 102, B6, 12035-
556	12059.
557	
558	Thorkelson, DJ and Breitsprecher, K (2005), Partial melting of slab window margins: genesis of
559	adakitic and non-adakitic magmas, Lithos, 79,1-2,25-41.
560	
561	Van Decar, JC and Crosson, RS (1990), Determination of teleseismic relative phase arrival times
562	using multi-channel cross-correlation and least squares, Bulletin of the Seismological Society of
563	America, 80,1, 150-169.
564	
565	Vasey, DA, Cowgill, E, Cooper, KM (2021), A preliminary framework for magmatism in modern
566	continental back-arc basins and its application to the Triassic-Jurassic tectonic evolution of the
567	Caucasus, Geochemistry, Geophysics, Geosystems, 22, 6.
568	
569	Weerdesteijn, MFM, Conrad, CP, Naliboff, JB (2022), Solid Earth Uplift Due To Contemporary
570	Ice Melt Above Low-Viscosity Regions of the Upper Mantle, Geophysical Research Letters, 49,17.
571	
572	Wessel, P., Smith, W., Scharroo, R., Luis, J., Wobbe, F. (2013). Generic Mapping Tools: improved
573	version released, EOS, Trans. Am. geophys. Un., 94, 409-410.
574	
575	Wiens, DA and Magnani, MB (2018), Solid Earth response of the Patagonia Andes to post-Little
576	Ice Age glacial retreat, International Federation of Digital Seismograph Networks.
577	https://doi.org/10.7914/SN/1P_2018.
578	
579	Yang, T and Shen, Y (2006), Frequency-dependent crustal correction for finite-frequency seismic
580	tomography, Bulletin of the Seismological Society of America, 96, 6, 2441-2448.

Supplementary Materials: Deep source regions for Patagonia backarc volcanism imaged by finite frequency body wave tomography

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Figure S1a. Moho depth (left) and sediments thickness (right) from Mark et al. (2022) used for the crustal correction. DM: Deseado Massif, AMB: Austral Magallanes Basin. Yellow dashed lines represent the maximum extension of the slab window and blue contour the Patagonia icefield. Black triangles are the locations of seismic stations used in this study.



Figure S1b. Travel times from 50km below sea-level to the surface, including topography, for, from left to right, low-frequency P waves, high-frequency P waves and SH waves. Note the frequency dependence of the P-wave travel times in regions with thick sediments where low frequencies display reduced travel times.



Figure S2. Test of the effect of the depth extent of the model. Vertical cross sections in P and SH models extending from left to right down to 800, 650 and 500km depth. *The tectonic features on the map follow the same conventions as in Figure 2 in the manuscript.*



Figure S3. Trade-off between model variance and data misfit (left) and Trade-off between model norm and data misfit. The star represents the preferred solutions shown in the manuscript for P and SH.

Checkerboard tests



Figure S4a. Synthetic input model for a checkerboard test with cells of 1.25° in horizontal dimensions, 150km in vertical dimension, $\delta ln V_{\rm P} = \pm 4\%$ (top), and $\delta ln V_{\rm S} = \pm 4\%$ (bottom). Maps at 120, 200, 280, 350 and 420 km depth.





Figures S5a. Vertical cross-sections in the checkerboard synthetic model for P waves presented in Figure S4a. Map view of study area with the location of different profiles. *The tectonic features on the map follow the same conventions as in Figure 2 in the manuscript.*



Figure S5b. Recovery for checkerboard test presented in Figure S5a







Figure S6b. Recovery for checkerboard test presented in Figure S6a

Structural tests



Fast velocity perturbation/Nazca slab

Figure S7. Synthetic and recovery results for P and SH waves associated with a fast velocity perturbation (+3%) between 40°S and 46°S simulating a 100km thick slab penetrating down to 200 km and 350km respectively.



Figure S8. Synthetic and recovery results for P and SH waves in cross section associated with a 100km thick fast velocity perturbation (+3%) representing two possible options for the depth extension of the Nazca slab. Green stars represent earthquake from USGS catalogue.

Slow velocity perturbation/Slab window



Figure S9. Synthetic and recovery results for P and SH waves associated with a low velocity perturbation (-3%) between 46°S and 49°S down to 280km and 420km depth respectively.

Two distinct low velocity perturbations/Slab window and one shallow fast velocity perturbation



Figure S10. The same as Figure S9 with addition to the east of the previous anomaly of a deep low velocity perturbation (-3%) and a shallow fast one (+3%).



Figure S11. Vertical cross sections for the resolution tests presented in Figure S10.



Figure SM12. Comparison of shear wave structure at 87.5km (top) and 162.5km(bottom) derived from body wave tomography (this study) and surface wave tomography (Mark et al. 2022). *The tectonic features on the map follow the same conventions as in Figure 2 in the manuscript.*