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1 **Origin of the Harappan Ernestites: Geochemical Insights into Provenance and**
 2 **Fabrication**

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4 M. K. Mahala¹, Jyotiranjan S. Ray^{1,*}, A. K. Kanungo², G. N. S. Sree Bhuvan³, A. Chatterjee^{1,4},
 5 B. G. George⁵, N. Sorcar^{6,7}, Y. S. Rawat⁸, J. S. Kharakwal⁹, and S. V. Rajesh¹⁰

6 ¹Physical Research Laboratory, Navrangpura, Ahmedabad 380009, India

7 ²Indian Institute of Technology Gandhinagar, Palaj, Gandhinagar 382055, India

8 ³Department of Earth Sciences, Pondicherry University, Puducherry 605014, India

9 ⁴Department of Geology, Presidency University, College Street, Kolkata 700073, India

10 ⁵Department of Earth Sciences, Indian Institute of Technology Bombay, Mumbai 400076, India

11 ⁶National Centre for Earth Science Studies, Akkulam, Thiruvananthapuram 695011, India

12 ⁷Korea Polar Research Institute, Incheon 21990, Republic of Korea

13 ⁸Archaeological Survey of India, Dharohar Bhawan, 24 Tilak Marg, New Delhi 110011, India

14 ⁹JRN Rajasthan Vidyapeeth, Sahitya Sansthan, Udaipur 313001, India

15 ¹⁰Department of Archaeology, University of Kerala, Thiruvananthapuram 695581, India

16

17 *Correspondence: jsray@prl.res.in

18

19 **Abstract**

20 Advancements in stone bead technology, particularly in drilling techniques, emerged during the
 21 Indus Valley (Harappan) civilization. Long-constricted cylindrical drill bits, made from a unique
 22 stone called Ernestite, were a distinctive feature of this culture. The origin of Ernestite remains
 23 uncertain, as it has no known natural analogue and its synthesis process is poorly understood.
 24 This study presents a mineralogical and geochemical investigation of Ernestite stones and drill
 25 bits from multiple Harappan and contemporaneous sites in Gujarat, India, to uncover their origin.
 26 The $^{87}\text{Sr}/^{86}\text{Sr}$ and $\varepsilon_{\text{Nd}}(0)$ of the drill bits overlap with those of the Ernestite stones, confirming
 27 their genetic relationship. The texture and presence of pseudo-mullite ($\text{SiO}_2 > 40\text{wt\%}$) with high
 28 Al-Ti-bearing hematite suggest that Ernestites are synthetic, created through a sintering process
 29 at $\sim 1100^\circ\text{C}$. An abundance of sand to silt-sized detrital quartz, along with Fe-Ti-Zr-rich minerals,
 30 indicates the use of unevenly powdered sandstones and laterites as raw materials, with
 31 geochemical ties to regional sources.

32 Introduction

33 The prehistoric Indus Valley Civilization (IVC), also known as the Harappan civilization, was
 34 one of South Asia's most advanced civilizations of its time, renowned for its sophisticated urban
 35 architecture and material culture¹⁻³. This civilization is famous for its fortified structures,
 36 efficient drainage systems, standardized seals and weights, and advanced technology employed
 37 in the manufacture of a diverse range of artifacts crafted from stone, metal, and shell⁴⁻⁷. Findings
 38 from nearly 2,500 sites across diverse geographic zones reveal that this civilization had broader
 39 spatial coverage compared to the contemporary Mesopotamian and Egyptian civilizations⁵⁻⁸.
 40 Most of the Harappan sites have been discovered along the river valleys of the Indus and
 41 Ghaggar-Hakra systems, distributed across Afghanistan, Pakistan, and northwestern India^{9,10}. It
 42 is generally believed that the Harappan culture began as small agro-pastoral communities in its
 43 Early phase (>5000-2600 BCE), which matured into an urban civilization, recognized as the
 44 Harappan phase (2600-1900 BCE), demonstrating remarkable advancements in town planning,
 45 food production, and the technology of pottery and bead manufacturing. Subsequently, the
 46 society declined through de-urbanization in the Late Harappan phase (1900-1300 BCE)^{10,11}.

47 Stone beads are one of the critical indicators of cultural and trade practices within
 48 prehistoric South Asian civilizations^{6,7}. The manufacture of stone beads began with the
 49 perforation of soft stones (e.g., limestone, steatite, and lapis lazuli) and later with hard stones
 50 (e.g., chert, agate, and jasper). The earliest evidence of stone beads dates back to the Mesolithic
 51 period (e.g., Jwalapuram)¹²; significant developments in bead production technologies, such as
 52 drilling, shaping, coloring, and mounting onto ornaments, occurred in the Neolithic and
 53 Chalcolithic periods² and became a key component of regional and external trades during the
 54 Harappan civilization^{5,7}. Ancient Gujarat was well known for its rich agate resources, which
 55 attracted the Harappans to this region, and bead manufacturing industries/workshops were
 56 established in several urban centres in Kutch and Saurashtra^{7,13}. Although various beads of
 57 different materials were in use, the long cylindrical beads of harder stones, typically jasper and
 58 carnelian, were manufactured through perforation using constricted cylindrical drill bits cut out
 59 from unique chips/stones called Ernestites^{14,15}, since their hardness is higher than agate (~7.5 on
 60 Mohs' scale^{5,13,16}). The beads are characterized by a drill hole section with a stepped profile, as
 61 the drill bits are typically wide at the tip and narrow at the mid-section⁷ (Fig. 2b).

62 The name “Ernestite” was given temporarily by Kenoyer and Vidale¹⁴ after Ernst J.H.
 63 Mackay, but it remains in use. Ernestites are a signature finding of the urban phase of the
 64 Harappan civilization; however, they have been reported in large numbers from the late phase,
 65 single-cultured Harappan and Sorath Harappan sites as well^{17,18}. Many Ernestite stones and drill
 66 bits have been found in close association with bead workshops in several Harappan sites in
 67 Pakistan (e.g., Harappa, Mahenjo-daro, Chahnu-daro)^{7,14} as well as in India (e.g., Dholavira,
 68 Khirsara, Kanmer)^{15,17,19}, and in a few Sorath Harappan sites such as Bhagatrav, Bagasra,
 69 Shikarpur, Nagwada^{18,20-22}. Some important Harappan and Sorath Harappan sites, including
 70 those where Ernestites have been reported, are shown on the map (Fig. 1). Primarily

71 manufactured by the artisans of the Harappan civilization, these unique materials almost became
 72 extinct in subsequent cultural periods⁵.

73 Kenoyer and Vidale¹⁴ described Ernestite at Mohenjo-daro as a rock composed of a
 74 mottled greyish-green to yellow-brown matrix with dark brown to black irregular patches or
 75 dendritic formations. Based on the XRD analysis of samples from Mohenjo-daro, Chanhudaro,
 76 and Harappa, they opined that these are metamorphic rocks composed of quartz, sillimanite,
 77 mullite, hematite, and titanium oxide phases. Law⁷ observed significant quartz, mullite-
 78 sillimanite, and hematite phases in two samples from Harappa, as well as mullite and cristobalite
 79 in the other two. He found from XRD and EMPA analyses that the light and dark matrices
 80 consisted of clay-sized (<2µm) Al-Si bearing phases, compositionally similar to mullite and
 81 sillimanite, in addition to quartz. The dark matrix contained additional phases such as hematite,
 82 titanohematite, rutile, and zircon. He suggested that the Ernestite is likely a highly indurated
 83 tonstein flint clay, sufficiently heat-treated (up to 1100°C) to yield its characteristic hardness,
 84 based on the limited mineralogical and chemical data from his study and earlier experimental
 85 studies on clays. Tonstein is a kaolinitic (flint) claystone formed by diagenesis of volcanic ash in
 86 a swampy or non-marine environment²³. Law⁷ did suggest a few probable locations for the
 87 sources of tonstein closer to the city of Harappa; however, he did not provide experimental proof
 88 of the transformation of any natural rock or mineral to Ernestite by heating. His study attempted
 89 to explain the presence of Ernestite constituent minerals like zircon and rutile by attributing their
 90 likely occurrence in tonstein. Besides, no reasoning was provided for the observation that the
 91 mullites in the Ernestites had significantly different compositions than those that are
 92 stoichiometry mandated²⁴.

93 Because of the sheer number of Ernestite drill bits reported from the Harappan city of
 94 Dholavira in Gujarat (1212), Prabhakar et al.¹⁵ hypothesized that the sources of Ernestite raw
 95 materials were located within the Kutch province of Gujarat. The XRD analyses of two samples
 96 of Ernestites from Dholavira and one sample from Bhagatrav, done by Prabhakar et al.¹⁵ and
 97 Prasad and Prabhakar¹⁷, respectively, showed the presence of quartz, hematite, and
 98 sillimanite/mullite. No cristobalite has been reported in Ernestites from any of the Indian sites.
 99 An ambiguity persists about the provenance (source regions) of the Ernestite raw materials.
 100 Earlier workers^{7,15} speculated both local (Kutch/Ratanpur) and regional (Gujarat) sources, and
 101 there exists no isotopic data to establish the source(s) unambiguously.

102 Despite their ubiquitous presence in the Harappan settlements (Fig. 1), the origin of the
 103 Ernestite stones and drill bits remains uncertain. Hence, deciphering the Ernestite source
 104 materials and their geologic origin is vital to understanding the stone drilling technology and the
 105 inter-regional communication network during the Harappan period. In this study, we have
 106 addressed the following poorly understood aspects of the Ernestites with detailed petrography,
 107 mineralogy, mineral chemistry, geochemical, and isotopic investigations from three Harappan
 108 sites (Dholavira, Khirsara, Kanmer) and one Sorath Harappan site (Bhagatrav) in Gujarat, India:
 109 (1) What are the characteristics of the Ernestites? (2) What raw materials were used for their

110 manufacturing, and (3) What were the geologic sources for these raw materials? In addition, we
111 have attempted to shed some light on the manufacturing process of these stones.

112 Methods

113 Owing to our limited access to the Harappan artifacts, only six samples could be included in this
114 study, consisting of three Ernestite stone samples (Fig. 2a) and three drill bits (Fig. 2b) from four
115 sites (Khirsara, Kanmer, Dholavira, and Bhagatrv; Fig. 1) in Gujarat. The sample from Kanmer
116 is associated with the mature Harappan phase. It comes from the collection of Kharakawal et
117 al.¹⁹. The Bhagatrv sample is related to the Sorath Harappan phase²¹ and comes from the
118 collection of Kanungo²². A sample from Dholavira represents the Mature/Late Harappan phase¹⁵.
119 The stratigraphic contexts of the samples can be found in the references given for each location.
120 The Ernestite from Bhagatrv was subsampled into three parts. There were two drill bits from
121 Kanmer (the first and third from the left in Fig. 2b) and one from Khirsara. Two laterite samples
122 and two sandstone samples from the island of Khadir, on which Dholavira is located, were also
123 studied. These rocks belong to the Paleogene and Jurassic periods, respectively. The sandstones
124 belong to the Khadir Formation, which is equivalent to the Jhurio Formation, and the laterites are
125 correlatable to either the Matanomadh Formation or the Naredi Formation of the Kutch
126 mainland²⁵. A brief description of the geology of Kutch and Khadir is provided in the
127 Supplementary Information. Because of their size and rarity, the drill bits were analyzed only for
128 Sr-Nd isotopic compositions, whereas the stones were powdered for mineralogical, geochemical,
129 and isotopic analyses.

130 Petrographic studies were conducted on thin sections of all three Ernestite samples using
131 transmitted and reflected light. Grain size analysis was done using the inbuilt software (Stream
132 Basic) associated with the petrographic microscope (Olympus® BX-53). The mineralogical
133 compositions of the Dholavira and Bhagatrv Ernestites whole stone powders were determined
134 by X-ray diffraction (XRD) using a Bruker D2 Phaser diffractometer at the Physical Research
135 Laboratory (PRL). Raman spectra for selected phases were obtained using a Renishaw® Invia
136 Reflex Raman microscope with a spectrometer with a 514nm laser excitation wavelength at
137 Pondicherry University. The laser beam, with 0.5 to 5% power, was focused on the sample spot
138 through an X50 microscope objective lens with an exposure time between 60 and 120 seconds.
139 The instrument was calibrated using the Raman silicon band of 520.7 cm⁻¹, with the help of an
140 internal reference material Si(100), before the samples were analyzed. The raw spectra were
141 preprocessed and subsequently corrected for background.

142 In-situ mineral chemical analyses for the Ernestite stones (thin sections) were conducted
143 using a CAMECA SX-5 Tactics Electron Probe Microanalyzer (EPMA) at the National Centre
144 for Earth Science Studies (NCESS)²⁶. Analyses were performed at 20 kV accelerating voltage
145 and 20 nA beam current with a beam diameter of 1μm. Analytical spots were chosen at the center
146 of individual mineral grains that were devoid of any cracks and had uniform Z-contrast. Analyses
147 for the aluminosilicate matrix and titanohematite phases were carried out on spots away from

148 quartz boundaries. X-ray elemental maps were also generated for all three samples
 149 (Supplementary Information). The major element contents of Ernestites were determined by X-
 150 ray Fluorescence (XRF) spectroscopy using a Rigaku® Supermini200 instrument at PRL and the
 151 pressed pellet method²⁷. Multiple international rock standards were used for calibration, and the
 152 reference material OU-6 from the International Association of Geoanalysts (IAG) was used for
 153 accuracy and precision checks. The major element contents of laterites and sandstones were
 154 measured at the NCESS, using an S4 Pioneer sequential wavelength dispersive-XRF²⁸, with
 155 reference materials VL-1 and MAG-1 used for accuracy and precision checks (Supplementary
 156 Data 1).

157 Bulk sample geochemical and isotopic measurements were carried out at PRL. About 50
 158 mg of sample powder each was digested using conventional HF-HNO₃ and HF-HNO₃-HCl
 159 dissolution protocols for trace element and isotopic analyses, respectively. The details of the
 160 analytical procedures are given in George and Ray²⁹. Trace element concentrations were
 161 measured on a Thermo® HR-ICPMS using BHVO-2 (USGS) as a calibration standard. Machine
 162 drift correction was performed using ¹¹⁵In as an internal standard. The accuracy and precision of
 163 our measurements, determined by repeated analyses of BHVO-2 (as unknown), were better than
 164 2% for REE and 5% for other trace elements. Sr and REE were separated from digested solutions
 165 by conventional cation exchange column chromatography using AG 50W-X8 resin (BioRad®),
 166 and Nd was eluted from REE using Ln-specific resin (Eichrom®), using protocols given in
 167 George and Ray²⁹. Sr and Nd isotopic ratio measurements were performed on a TIMS (Thermo®
 168 Triton Plus) in static multicollection mode. Sr isotopes of some samples were measured on an
 169 MC-ICPMS at PRL³⁰. Instrumental mass fractionation for Sr and Nd isotopic ratios was
 170 corrected using exponential fractionation (internal) correction equations of Thirlwall³¹ and
 171 assuming ⁸⁸Sr/⁸⁶Sr = 8.375209 and ¹⁴⁶Nd/¹⁴⁴Nd = 0.7219. Multiple measurements of SRM-987
 172 and JNd-1 over three years yielded an average of ⁸⁷Sr/⁸⁶Sr = 0.710249±0.000009 (2 σ ; n = 14)
 173 and ¹⁴³Nd/¹⁴⁴Nd = 0.512102±0.000010 (2 σ ; n = 14).

174 **Results**

175 **Petrography and Mineralogy**

176 All Ernestite stone chips from Dholavira, Kanmer, and Bhagatav exhibit heterogeneous physical
 177 appearances, unlike other Harappan artifacts, which demonstrate remarkable homogeneity⁷ (Fig.
 178 2a). They are hard (harder than quartz), highly compact, do not produce streaks, and are difficult
 179 to break. Two clear domains, a yellowish-brown or khaki color phase and a black color phase,
 180 can be distinguished by the naked eye (Fig. 2b). Transmitted and reflected light microscopy
 181 reveals that Ernestite stones contain detrital subangular to subrounded quartz grains (sand to silt-
 182 sized) and angular to sub-angular opaque phases like hematite and ilmenite set in a compact,
 183 fine-grained, light-colored (yellowish/khaki) groundmass of unidentifiable mineral(s) (Fig. 3).
 184 Quartz in Dholavira Ernestite occurs as fractured angular to subangular grains (Fig. 3a,b)

185 compared to the sub-angular to sub-rounded grains in Bhagatrav (Fig. 3c,d) and Kanmer (Fig.
 186 3e,f). The opaque phases (hematite, titanohematite, and ilmenite) appear as narrow bands or
 187 irregular patches. They occur in lower proportions in the Dholavira Ernestite than in the Kanmer
 188 and Bhagatrav stones. Hematite appears gray and displays the characteristic reddish internal
 189 reflection under plane and cross-polar view, respectively, in reflected light (Fig. 3g,h), and is
 190 often associated with ilmenite (shows bi-reflectance). All these detrital phases are essentially
 191 larger than clay-sized ($\sim 4\mu\text{m}$) mineral grains that constitute a claystone. Sand-sized (210-736 μm
 192 diameter; Supplementary Figure 1) detrital (clastic) grains of ilmenite and its partial replacement
 193 by hematite are also observed in the Kanmer Ernestite under a cross-polar view in reflected light
 194 (Fig. 3e,f). Zircon and rutile in Kanmer Ernestite have subrounded to irregular grain boundaries,
 195 confirming their detrital nature (Supplementary Figure 3). The size (longest diameter)
 196 distributions of detrital quartz grains (measured in the thin sections) in Dholavira, Kanmer, and
 197 Bhagatrav Ernestites are presented in a box plot (Fig. 4). Their ϕ (= $-\log_2 d$; d = diameter) sizes
 198 (1.84-6.64) vary between medium sand to fine silt, with half of the distributions falling between
 199 very-fine sand to coarse silt fractions. The quartz grains in Kanmer and Dholavira stones are
 200 moderately sorted ($1\sigma = 0.81$ and 0.71 , respectively), whereas those in the Bhagatrav stone are
 201 moderately well-sorted ($1\sigma = 0.68$). Powder XRD patterns of the Dholavira and Bhagatrav
 202 samples (Supplementary Figure 2) reveal that quartz is the most abundant phase in all the
 203 samples, followed by a mullite-like phase (mullite/sillimanite). Hematite was detected only in the
 204 Bhagatrav dark matrix (Supplementary Figure 2), though it is observed in the petrography of all
 205 Ernestites.

206 The background-corrected Raman spectroscopic data for the Ernestite stone samples are given
 207 in Supplementary Data 3, and the spectra are presented in Fig. 5. Each trace in Fig. 5 represents a
 208 Raman (shift) spectrum for one spot analysis. Characteristic peaks for various minerals were
 209 identified by comparing our data with the reference spectra^{32,33}. The laser excitation of quartz
 210 grains in Kanmer and Bhagatrav samples yielded good Raman signals (Fig. 5b-c). However, the
 211 excitation created a high fluorescence effect in the Dholavira sample, obscuring the Raman
 212 signal (Fig. 5a). We could observe the characteristic 468 cm^{-1} (sharp) peak of quartz only for four
 213 spots (Fig. 5b-c); whereas it was small and weaker in others. We did not observe the
 214 characteristic peaks for cristobalite (114; 231; 418 cm^{-1})³³. Due to their intimate association with
 215 disseminated fine-grained quartz grains, the aluminosilicate matrix in the Dholavira and Kanmer
 216 samples did not yield discernible Raman spectra. However, the Bhagatrav sample yielded good
 217 spectra because of its discrete ellipsoidal occurrence as the dark matrix. From these spectra, we
 218 could identify the matrix mineral to be mullite (Fig. 5c-d). We did not find any indication of the
 219 presence of sillimanite (e.g., 235; 311; 872 cm^{-1})³⁴.

220 **Major and Trace elements**

221 The major oxide and trace element contents of two Ernestite samples from Bhagatrav and
 222 Dholavira, as well as two laterite and two sandstone samples from Khadir Island, are presented

223 in Supplementary Data 1. SiO_2 content (47-61 wt%) is the highest among all oxides, with Al_2O_3 ,
 224 FeO^T , and TiO_2 being other major components. MnO , Na_2O , and P_2O_5 are either very low (<
 225 0.1 wt%) or absent, whereas K_2O and MgO concentrations are minor. Bhagatrav Ernestite has
 226 lower SiO_2 and Al_2O_3 , FeO^T , and TiO_2 than Dholavira Ernestite. The major oxide data of the
 227 laterite and sandstone samples from the Khadir Island are also presented in Supplementary Data
 228 1. Laterites have high Fe_2O_3 (36.7-37.6 wt%), moderate SiO_2 (32.31-32.61 wt%) and low Al_2O_3
 229 (8.77-8.89 wt%), TiO_2 (1.32-1.34 wt%) contents, whereas sandstones are characterized by high
 230 SiO_2 (67.57-68.42 wt%), moderate Al_2O_3 (13.26-13.19), K_2O (1.83-1.84 wt%) and low Fe_2O_3
 231 (2.06-2.09 wt%). Various oxides vs. SiO_2 diagrams plotted for Ernestites, sandstones, and
 232 laterites, along with the published data for Mesozoic sandstones^{35,36} and laterites of Kutch³⁷⁻³⁹,
 233 are presented in Fig. 6. Figure 7 presents the primitive mantle (PM) normalized multi-element
 234 patterns for the Ernestite samples and those for Mesozoic rocks⁴⁰, and laterites of Kutch region³⁹.

235 Mineral Chemistry

236 Representative backscattered electron (BSE) images of various phases in a polished thin section
 237 of the Kanmer Ernestite are given in Supplementary Fig. 3. Mineral compositions of different
 238 phases are provided in Supplementary Data 2. X-ray elemental maps for all three Ernestite stones
 239 (i.e., Kanmer, Bhagatrav, and Dholavira), as well as chemical spot analysis data (both by
 240 EPMA), are provided in Supplementary Figs. 3-6 and Supplementary Data 2. Quartz (SiO_2 : 98-
 241 100 wt%) of varying sizes is dispersed within the light-colored (yellow) fine matrix, which is
 242 mainly composed of aluminosilicate phases (SiO_2 : 40-53 wt%; Al_2O_3 : 40-50 wt%). Although
 243 identified as mullites by XRD, the aluminosilicate matrix phases contain much higher SiO_2 than
 244 that mandated by stoichiometry (i.e., ≤ 30 wt%)²⁴, therefore, we identify these phases as
 245 pseudomullites. Fe-Ti bearing phases, such as hematite (FeO : 71-74 wt%) and ilmenite (TiO_2 :
 246 51-56 wt%), often occur as narrow patches or are finely dispersed within the light (yellow)
 247 matrix. Many hematite grains have a high TiO_2 content (29-40 wt%) and can thus be classified as
 248 titanohematite. The titanohematites also contain an appreciable amount of Al_2O_3 (5-21 wt%).

249 Sr-Nd isotopic ratios

250 Results of Sr and Nd isotopic compositions of Ernestite stones and drill bits, laterites, and
 251 sandstones are provided in Table 1. The $^{87}\text{Sr}/^{86}\text{Sr}$ and $\varepsilon_{\text{Nd}}(0)$ of Ernestite stones and drill bits from
 252 Kutch (Dholavira, Kanmer, and Khirsara) vary between 0.71000 and 0.72282, and
 253 -14.7 and -13.9, respectively. In contrast, the Sr-Nd isotopic compositions of Bhagatrav
 254 Ernestites are more radiogenic in Sr and less radiogenic in Nd ($^{87}\text{Sr}/^{86}\text{Sr} = 0.73022$ to 0.730876;
 255 $\varepsilon_{\text{Nd}}(0) = -18.3$ to -18.1). The drill bits from Kanmer have an identical $\varepsilon_{\text{Nd}}(0)$ of -13.9, and one of
 256 the drill bits has almost identical $^{87}\text{Sr}/^{86}\text{Sr}$ as that of the Ernestite stone (0.72282 vs. 0.72207;
 257 Table 1). The laterite samples collected from the Khadir (Dholavira) have $^{87}\text{Sr}/^{86}\text{Sr}$ varying from
 258 0.7089 to 0.7096, and their $\varepsilon_{\text{Nd}}(0)$ ranges from -7.7 to -7.5, whereas the sandstones have more
 259 radiogenic Sr and Nd ($^{87}\text{Sr}/^{86}\text{Sr} = 0.71494$ and 0.74344; $\varepsilon_{\text{Nd}}(0) = -23.5$ and -17.7).

260 **Discussion**

261 In the first-ever detailed characterization, Kenoyer and Vidale¹⁴ suggested a metamorphic origin
 262 for Ernestites based on their identification of the matrix phases as sillimanite and mullite. Mullite
 263 is a rare mineral and has only been reported from specific contact-metamorphic rocks (in
 264 metamorphosed clays) and pseudotachylites^{41,42}. It is also commonly observed in high-
 265 temperature ceramics and has been synthesized by heating various aluminosilicate minerals (e.g.,
 266 kaolinite, kyanite, andalusite, sillimanite) at high temperatures ($>1100^{\circ}\text{C}$)⁴³⁻⁴⁵. However, we
 267 identify these phases, which exhibit identical XRD and Raman spectra to mullite, as
 268 pseudomullites based on their higher SiO_2 contents ($> 40\text{wt\%}$). Since mullite and pseudomullite
 269 are isostructural, all earlier studies, which relied primarily on XRD data, had incorrectly
 270 identified pseudomullite as mullite. "Pseudomullite" refers to a structure or phase that resembles
 271 mullite (Fig. 8) but is not the true, stoichiometric mullite. It is known to form during the thermal
 272 decomposition of kaolinite or other aluminosilicate materials⁴⁴. Mullite refers to an
 273 experimentally observed solid solution series $\text{Al}_{4+2x}\text{Si}_{2-2x}\text{O}_{10-x}$ with $0.2 < x < 0.9$ (Fig. 8)²³.
 274 According to Lenz et al.²⁴, the stoichiometric composition of synthetic mullite commonly varies
 275 between $3\text{Al}_2\text{O}_3.2\text{SiO}_2$ ($\sim 72\text{wt\% Al}_2\text{O}_3$) and $2\text{Al}_2\text{O}_3.\text{SiO}_2$ ($\sim 78\text{wt\% Al}_2\text{O}_3$). In natural mullites,
 276 Fe_2O_3 substitutes Al_2O_3 , producing a wide range of compositions at $\sim 30\text{wt\% SiO}_2$ (Fig. 8)^{23,41}. In
 277 contrast, stoichiometric sillimanite has ~ 61 wt% Al_2O_3 (Fig. 8). The aluminosilicate phase in the
 278 Ernestite matrix is pseudomullite. It has higher SiO_2 and lower Al_2O_3 than those of natural
 279 mullite or sillimanite (Fig. 8). Contrary to an earlier report of the presence of sillimanite in the
 280 Ernestites¹⁷, our Raman spectroscopic data indicate the complete absence of sillimanite in these
 281 stones (Fig. 5). Pseudomullites are not found in nature. However, it has been observed in
 282 synthetic materials produced by heating, such as ancient clinky potteries⁴⁶, that these materials
 283 likely developed at lower temperatures ($\leq 1100^{\circ}\text{C}$) as a precursor to the formation of mullite from
 284 kaolinite (at $\geq 1200^{\circ}\text{C}$)⁴⁵. Therefore, the presence of pseudomullites unambiguously rules out that
 285 Ernestites are natural rocks, indicating their origin by high-temperature processing. Absence of
 286 mullite also suggests that the sintering temperature never reached 1200°C .

287 The Harappans manufactured the Ernestites as the source stones for drill bits through a
 288 high-temperature heating process that could generate the pseudomullites. Further evidence for a
 289 high-temperature process is provided by the chemical composition of Fe-Ti-bearing phases, as
 290 determined by EPMA analyses. The presence of titanohematites with significant TiO_2 (29-40
 291 wt%) and Al_2O_3 (5-21 wt%) suggests an extensive substitution between Fe_2O_3 and TiO_2 and
 292 between Fe_2O_3 and Al_2O_3 . It is known that at temperatures below 800°C , only a limited solid
 293 solution between TiO_2 and Fe_2O_3 is possible⁴¹. Similarly, in the $\text{Fe}_2\text{O}_3\text{-Al}_2\text{O}_3$ system, higher
 294 Al_2O_3 (up to 10wt%) can be substituted into the hematite (Fe_2O_3) structure at high temperatures
 295 only ($\sim 1000^{\circ}\text{C}$)⁴¹. Therefore, higher TiO_2 and Al_2O_3 contents in the titanohematite confirm a
 296 heating process ($>1000^{\circ}\text{C}$) in Ernestite manufacturing. It is thus apparent that the pseudomullite
 297 matrix was produced during high-temperature sintering. This provides the first geochemical
 298 evidence of sintering being used in the manufacture of Ernestites. This also successfully explains

299 the presence of high-temperature craft objects, such as stoneware bangles^{47,48}, steatite beads⁴⁹,
 300 and furnaces¹⁹, at the Harappan sites. The presence of detrital quartz grains, ilmenite, hematite,
 301 zircon, and rutile suggests that the raw materials used to make the Ernestites were natural.

302 Law⁷ suggested tonstein as the only raw material for Harappan Ernestites. He attributed
 303 the coarser (up to 100 μ m) subhedral quartz or cristobalite grains (detected in his BSE images) to
 304 crystallization of free silica (released during heating) and zircon to a magmatic origin. He further
 305 proposed that the raw materials for the Ernestites (i.e., tonsteins) were sourced from
 306 local/regional sources (i.e., Kutch). Tonsteins are hard and compact kaolinite-altered volcanic ash
 307 layers, generally found in coals and associated sediments²³. These often contain magmatic quartz
 308 and zircon²³. However, the microtextural characteristics of zircon, quartz, ilmenite, and rutile in
 309 Ernestites suggest their detrital (clastic) nature, making them unlikely to have been constituents
 310 of tonsteins. Additional evidence against using tonsteins for Ernestites comes from the presence
 311 of non-radiogenic Nd in the Ernestites ($\epsilon_{Nd}(0) < -14$), because all volcanic ash beds in Kutch
 312 have been linked to the Deccan Traps³⁹, which contain more radiogenic Nd ($\epsilon_{Nd}(0) > -11$, Fig. 9).
 313 Besides, our Ernestite samples contain sand-sized detrital quartz and ilmenite grains in contrast
 314 to a claystone/tonstein that usually contains clay-sized grains ($\leq 2\mu$ m).

315 The size and moderately sorted nature of the detrital quartz grains suggest the use of
 316 coarser raw materials, such as sandstones, which were likely pounded into sand/silt-sized
 317 particles before being processed for sintering. It is possible that the Fe-Ti phases (hematite,
 318 titanohematite, ilmenite, and rutile) observed in the Ernestites also originated from the
 319 sandstones, as sandstones generally contain such heavy minerals. However, our Ernestites appear
 320 to exhibit mixing trends between the Mesozoic sandstones and laterites of the Kutch region in
 321 various oxide vs. SiO₂ plots (Fig. 6). Trace element patterns (Fig. 7) also suggest that such a
 322 mixture is necessary to explain the chemistry of the Ernestites. Besides, high contents of Al₂O₃
 323 (> 20wt%) and high field strength elements (e.g., Sc, V, Cr, and Co) in Ernestites cannot be
 324 achieved by heating the sandstones of the Kutch alone. Therefore, a second end-member,
 325 containing Fe-Ti minerals but low in alkali elements, is needed to explain the Ernestite
 326 chemistry. The laterites of Kutch, derived from the mafic volcanic rocks of the Deccan Traps, fit
 327 the bill. The Paleocene to Eocene lateritic deposits in western Kutch (Matanomadh Formation)
 328 and Saurashtra (Jamnagar) contain both Al-rich (gibbsite, kaolinite) and Fe-rich (goethite,
 329 hematite, ilmenite-rich) phases and are depleted in alkalis³⁷⁻³⁹, and have the required
 330 characteristics of this raw material. Although we discard claystone (tonstein) as the sole raw
 331 material, we do not deny its possible use in combination with sandstone and laterite for the
 332 Ernestite manufacturing.

333 Since kaolinite is a common mineral in laterites, it could have decomposed and
 334 undergone subsequent chemical and structural changes to form the pseudomullite matrix during
 335 the sintering process. The free (amorphous) silica released during the heating of pure kaolinite
 336 recrystallizes as cristobalite upon further heating to ~1350°C⁵⁰. However, when kaolinite is

337 heated with alumina-bearing material, no free silica forms^{51,52}. Therefore, it is likely that during
 338 the sintering process carried out by the Harappans, free silica (SiO_2) formation was suppressed
 339 by the presence of gibbsite (Al(OH)_3) in the laterite. Moreover, gibbsite undergoes thermal
 340 decomposition to boehmite (AlO(OH)) at 200°C, which transforms into a transitional alumina (α -
 341 Al_2O_3) phase at 500°C⁵³. The Al_2O_3 in the titanohematite structure was likely sourced from
 342 gibbsite (lateritic) in the mixture during the α - Al_2O_3 stage. Furthermore, β -cristobalite has been
 343 observed at 1200-1300°C in quartz powder heating experiments⁵⁴. The absence of this phase in
 344 the Ernestites, as confirmed by the Raman spectra of silica phases (Fig. 5), is consistent with our
 345 study's results, which indicate a maximum sintering temperature of 1100°C. We suspect that the
 346 cristobalites observed by Law⁷ represent a higher temperature ($\geq 1200^\circ\text{C}$) or prolonged heating
 347 process.

348 The similar $^{87}\text{Sr}/^{86}\text{Sr}$ and $\varepsilon_{\text{Nd}}(0)$ values of Ernestite stone and drill bits from Kanmer
 349 genetically link the drill bits to the stone. Although it has been well established that Ernestite
 350 stones are the raw materials for long and constricted cylindrical drill bits¹⁴, their isotopic
 351 similarity is the first-ever chemical evidence for the same. Because of the sheer number of
 352 Ernestite stones and drill bits from Dholavira, Law⁷ speculated that the raw materials for the
 353 stones came from either the island itself (i.e., Khadir) or the Kutch region of Gujarat. However,
 354 our geochemical data (Figs. 6 and 7) support a regional sourcing of the raw materials. The
 355 sandstones and laterites of Kutch appear to have been the primary sources of the raw materials
 356 for the Ernestites. In search of more robust evidence for this geological provenance hypothesis,
 357 we make use of the Sr-Nd isotopic compositions of Ernestites and their potential source rocks
 358 (Table 1). The ε_{Nd} and $^{87}\text{Sr}/^{86}\text{Sr}$ of the Ernestites plot between the compositional ranges of the
 359 Mesozoic sandstones and the laterites of Kutch, and can be explained by a two-component
 360 mixing involving these two rock types. The isotopic compositions of laterites of Khadir, which
 361 are developed over volcanic ash deposits³⁹, fall well within the compositional field of the Deccan
 362 Basalts. This suggests that other lateritic horizons in the Kutch and Saurashtra region, developed
 363 over the Deccan Trap rocks, could also have served as sources for the raw material for Ernestites.
 364 A binary mixing model (Supplementary Figure 7) indicates a higher contribution (60-80%) from
 365 the sandstone end-member compared to the laterite end-member in most of our samples, with the
 366 contribution reaching as high as 90-95% for the Bhagatrv sample. Results of a Monte-Carlo
 367 simulation considering $\pm 2\sigma$ range for the above two end-member compositions (Fig. 9) strongly
 368 support the inference that the Harappans used laterite from different weathered (Deccan)
 369 horizons and sand from Mesozoic sandstones from Kutch to manufacture Ernestites.

370 We draw the following conclusions from our investigation of Ernestites, the parent material
 371 for the unique constricted drill bits of the Harappan Civilization, using petrographic,
 372 mineralogical, geochemical, and Sr-Nd isotopic techniques.

373 1. Stone drill bits have been isotopically fingerprinted to the Ernestites, confirming their genetic
 374 link.

375 2. The Ernestites consist of medium sand to fine silt detrital quartz, hematite, ilmenite, zircon,
 376 and rutile welded together in a fine-grained aluminosilicate matrix/groundmass.

377 3. Ernestites' texture (larger mineral grains and their detrital nature), and its Nd isotopic
 378 composition ($\varepsilon_{\text{Nd}}(0) > -11$) rule out the use of tonstein flint as a raw material.

379 4. The aluminosilicate matrix/groundmass phase has been chemically identified as
 380 pseudomullite, though its XRD and Raman spectra are similar to those of mullite.

381 5. The presence of pseudomullites, with high SiO_2 contents (> 40wt%), unambiguously makes
 382 Ernestites artificial, with supporting evidence from the significant substitution of Al_2O_3 and
 383 TiO_2 in hematites. These data also suggest a high temperature (reaching 1100°C) synthesis of
 384 Ernestites.

385 6. Mineralogy, texture, and mineral chemistry suggest Ernestites were manufactured through a
 386 high-temperature sintering process involving sand and clay-bearing raw materials.

387 7. Major and trace elements and Sr-Nd isotopic data point to the likelihood of the raw materials'
 388 regional provenance (sandstones and laterites of Kutch).

389 8. All our findings suggest that Ernestites were likely made in the Harappan centres of Gujarat,
 390 India, and the Ernestite-based drilling technology was exclusive to this civilization.

391 **Author contributions**

392 JSR and AC conceived the study. AKK, YSR, JSK, and SVR supplied samples. AC, MKM,
 393 BGG, NS, and GNSSB conducted the analytical work. MKM, GNSSB, AC, and JSR interpreted
 394 the data. JSR secured the project's funding, and all authors contributed to the writing.

395 **Data Availability Statement**

396 All data generated for this study are in the tables in the manuscript and the supplementary files.

397 **Competing interests**

398 The authors declare no competing interests related to this work.

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405 **Appendix A. Supplementary Information**

406 Supplementary materials for this article have been attached as separate files.

407 1) Supplementary Information (Doc file: Text and Supplementary Figures 1-7)
 408 2) Supplementary Data 1 (Excel file: Major and Trace element data)
 409 3) Supplementary Data 2 (Excel file: EPMA mineral chemistry data)

410 4) Supplementary Data 3 (Excel file: Raman spectroscopy data)

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562

563 **Table 1.** Sr-Nd isotopic data for the Ernestite stones/drills and Laterites

| Sample | Type | Location | $^{87}\text{Sr}/^{86}\text{Sr}$ | $^{143}\text{Nd}/^{144}\text{Nd}$ | $\varepsilon_{\text{Nd}}(0)$ |
|-----------------|-----------|-----------|---------------------------------|-----------------------------------|------------------------------|
| E-2 | Ernestite | Bhagatrav | 0.73022* | 0.511708 | -18.1 |
| E-3 | Ernestite | Bhagatrav | 0.73027* | 0.511703 | -18.2 |
| E-4 | Ernestite | Bhagatrav | 0.73087* | 0.511700 | -18.3 |
| 3304 | Drill bit | Kanmer | 0.72282* | 0.511927 | -13.9 |
| 3285 | Drill bit | Kanmer | 0.71778* | 0.511926 | -13.9 |
| ERN-KM | Ernestite | Kanmer | 0.72207* | 0.511884 | -14.7 |
| ERN-DV | Ernestite | Dholavira | 0.712901 | 0.511900 | -14.4 |
| ERN-KU | Drill bit | Khirsara | 0.709991 | 0.511915 | -14.1 |
| KH-3 | Laterite | Khadir | 0.708990 | 0.512256 | -7.5 |
| KH-4 | Laterite | Khadir | 0.709551 | 0.512245 | -7.7 |
| KH-15-6 | Sandstone | Khadir | 0.714940* | 0.511731* | -17.7 |
| KH-15-27 | Sandstone | Khadir | 0.743439* | 0.511433* | -23.5 |

564 Note: All ratios are TIMS data except those marked with *, which are MC-ICPMS data. The average isotopic ratios
 565 and external reproducibilities determined for the international SRM-987 and JNd-1 in TIMS, after repeated analyses
 566 over three years, are $^{87}\text{Sr}/^{86}\text{Sr} = 0.710249 \pm 0.000009$ (2σ ; $n=14$) and $^{143}\text{Nd}/^{144}\text{Nd} = 0.512102 \pm 0.000010$ (2σ ; $n=14$),
 567 respectively. $\varepsilon_{\text{Nd}}(0) = [(^{143}\text{Nd}/^{144}\text{Nd})_{\text{sample}}/^{143}\text{Nd}/^{144}\text{Nd})_{\text{CHUR}} - 1] \times 10^4$, where CHUR = Chondrite Uniform Reservoir
 568 and (0) stands for present-day value.

569

570 **Figure Captions**

571 **Fig. 1.** A schematic geographical map of western India and Pakistan showing important
 572 Harappan urban centers and cities from which Ernestite stones and drill bits have been reported.
 573 The three Harappan sites and one Sorath Harappan site, from which the Ernestite samples for this
 574 study came, are marked.

575 **Fig. 2.** The photographs of (a) Ernestite stones from Dholavira, Bhagatrav, and Kanmer, and (b)
 576 Ernestite drill bits from Kanmer studied in this work. Note the textural variations between
 577 different samples, as reflected in their colors. The first and third drill bits shown in (b) have been
 578 used for isotopic analyses.

579 **Fig. 3.** Photomicrographs of Ernestite thin sections: (a) sample from Dholavira in plane-
 580 polarized transmitted light showing the presence of quartz (Qz) and opaques (Opq); (b) same as
 581 in (a) with hematite (Hem) displaying characteristic red internal reflection under reflected lights;
 582 (c-d) sample from Bhagatrav in plane-polarized and cross-polarized transmitted lights; (e-f)

583 sample from Kanmer showing detrital ilmenite (Ilm) and quartz (Qz) in plane-polarized
 584 transmitted light; (g-h) sand-sized ilmenite (Ilm) grains and hematite (Hem) patches in the same
 585 sample as in (e) in plane and crossed polarized reflected lights, respectively.

586 **Fig. 4.** Boxplot showing grain size distribution (in ϕ scale) in the Ernestites. Relevant statistical
 587 information is given in boxes inside the figure. The mean (circle) and median (dashed line) are
 588 marked in each box. Bhagatrav-1 and Bhagatrav-2 represent the yellow (khaki) and black
 589 colored groundmasses, respectively. Symbols: n = number of observations; μ = mean; σ =
 590 standard deviation; SK = skewness; K = kurtosis.

591 **Fig. 5.** Raman spectra of silica and aluminosilicate phases in the Ernestite samples from
 592 (a) Dholavira; (b) Kanmer, and (c-d) Bhagatrav, identified as alpha-quartz and mullite,
 593 respectively, based on their Raman shifts. One sample (Bhagatrav-2) was also analyzed for its
 594 Raman shift between 1000 and 1500 cm^{-1} .

595 **Fig. 6.** Plots of various oxides vs SiO_2 for Ernestite samples, laterites, and sandstones from
 596 Khadir Island (data in Table 1). Compositions of Mesozoic sandstones^{35,36} and laterites of
 597 Kutch³⁷⁻³⁹ are plotted as fields for comparison.

598 **Fig. 7.** Primitive Mantle normalized spider diagram for Ernestite samples. Also plotted are the
 599 data for laterites³⁷⁻³⁹ and Mesozoic shales⁴⁰ of Kutch. Normalizing values are from McDonough
 600 and Sun⁵⁵.

601 **Fig. 8.** The plot of Al_2O_3 vs. SiO_2 for the aluminosilicate phases in our Ernestite samples
 602 compared with compositions of natural mullite²⁴, stoichiometric (synthetic) mullite, mullite solid
 603 solution, and stoichiometric sillimanite.

604 **Fig. 9.** The plot of $\epsilon_{\text{Nd}}(0)$ vs. $^{87}\text{Sr}/^{86}\text{Sr}$ of the Ernestite stones and drill bits, along with laterites
 605 and Mesozoic sandstones⁵⁶ of Kutch. The compositional field of Deccan Basalts⁵⁷ is shown as a
 606 pale cyan rectangle for comparison. The points represent the results of a Monte-Carlo binary
 607 mixing simulation involving the Mesozoic sandstone (average composition: $^{87}\text{Sr}/^{86}\text{Sr} =$
 608 0.733600 ± 0.030 & $\epsilon_{\text{Nd}} = -18.7 \pm 6.2$) and the laterite (average composition: $^{87}\text{Sr}/^{86}\text{Sr} =$
 609 0.709271 ± 0.0008 & $\epsilon_{\text{Nd}} = -7.6 \pm 0.30$), at 2σ levels of uncertainty, of the Kutch region. f_{ms}
 610 represents the fraction of sandstone end-member in the mixture.

611