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The New Method of Estimation Greenhouse Effect and Climate Change

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Abstract: Global climate change is one of the major concerns of modern society. To estimate this change, the global mean temperature is often used. Measuring and calculating the Earth's average temperature is a complex, multi-step process that combines data from various sources and employs statistical techniques. Today, datasets containing spatial-temporal data on Earth's temperature are readily available. Although scientists claim to achieve an accuracy of a few tenths of a degree, the fundamental question is not accuracy but whether the global mean temperature is a meaningful metric at all.

This paper demonstrates that the current methodology for determining the global mean temperature is inadequate for estimating the greenhouse effect and climate change, potentially leading to misleading scientific conclusions in the long term. A new methodology is introduced, focusing on the energy budget of the Earth's heating and cooling processes. The total influence of the atmosphere on the greenhouse effect and global warming can be estimated by comparing Earth's temperature with that of the Moon, treated as a bare body. The concept of 'potential temperature for cooling' is introduced as a more appropriate parameter for assessing the greenhouse effect, global warming trends, and climate change. Applying this new methodology to temperature averaging is expected to yield surprising results and offer a more accurate insight into climate change.

Key words: climate change, the average planet's temperature, greenhouse effect, modelling

Introduction

The climate system primarily consists of the land, ocean, ice surfaces, atmosphere, and solar radiation that provides energy. These components interact to produce the conditions on and around the Earth's surface that we refer to as climate [1]. Climate is defined by averaging physical quantities that

32 characterize these conditions over space and time, typically over a 30-year period [2]. The physical
33 quantities considered include surface temperature, precipitation, cloud cover, wind patterns, and more.

34 .
35 As Klaus Hasselmann, Nobel Laureate in Physics (2021), noted, a major task in climate science is the
36 detection problem—identifying the most sensitive climate indices that best distinguish climate change
37 signals from natural climate variability. Examples of such indices include global or regional mean
38 surface temperature, vertical temperature differences, sea ice extent, sea level change, and integrated
39 deep ocean temperatures [3]. Nevertheless, measuring and reconstructing global mean temperatures
40 remains a primary focus of leading organizations such as NASA's Goddard Institute for Space Science,
41 the National Oceanic and Atmospheric Administration (NOAA), and the UK Met Office's Hadley
42 Centre [4].

43 Measuring and calculating global mean temperature involves collecting temperature data from various
44 locations worldwide and averaging these values. The globe is divided into numerous spatial cells, and
45 for each grid cell, the temperature anomaly—the difference between the measured and usual
46 temperature for a given day—is calculated. The average of all anomalies is then determined and
47 compared across years. This process is highly complex, requiring sophisticated methods to address data
48 quality, spatial and temporal gaps, and biases. Moreover, different organizations may use slightly
49 different methodologies and datasets, leading to variations in reported global mean temperature values.
50 Recent advances in atmospheric and oceanic science, coupled with new global measurement systems
51 such as remote sensing and numerical computer models, have enabled quantitative studies and
52 predictions of climate change and global warming [1-6]. For instance, the IPCC Sixth Report (Climate
53 Change 2021) was based on over 14,000 scientific publications, while the IPCC Fifth Report stated with
54 high confidence that human-induced warming had reached approximately 1°C (likely between 0.8°C
55 and 1.2°C) above pre-industrial levels by 2017, increasing at a rate of 0.2°C per decade [1]. These
56 results, derived from models and observations, are impressive yet remain subject to scientific debate.

57 This paper examines a fundamental question: Does the global mean temperature provide a meaningful
58 measure of climate change? The findings suggest that local surface temperatures are misused in the
59 current methodology, which is inadequate for assessing climate change. At a minimum, additional
60 temperature metrics should be considered for comparison.

61

62 THE CURRENT METHOD

63 Summing and averaging temperatures from different regions of the Earth's surface yields a quantity
64 with no direct physical meaning. In other words, global mean temperature is a purely statistical indicator
65 that is insensitive to climate change and has been misinterpreted. The following thought experiments
66 illustrate why the current averaging method is flawed.

67

68 Thought Experiments

69 A Simple Proof of the Current Methodology's Inadequacy

70 Imagine the Earth is heated uniformly to a temperature of $t_{Earth} = 15^{\circ}\text{C}$. Now, suppose that due to
71 climate change, one hemisphere's mean temperature drops to $t_1 = 0^{\circ}\text{C}$, while the other rises to $t_2 =$
72 30°C . The calculated global mean temperature remains $t_{Earth} = (t_1 + t_2)/2 = 15^{\circ}\text{C}$. This
73 demonstrates that infinite temperature distributions—each reflecting different climate conditions—can
74 produce the same mean temperature, proving that mean temperature is insensitive to climate change.

75 A more striking example involves a spherical celestial body heated uniformly to 1K. The total power
76 radiated, according to the Stefan-Boltzmann law, is:

$$77 \quad P = A \sigma T^4, \quad (1)$$

78 where A is surface area, σ is the Stefan-Boltzmann constant, and T is temperature. If half of the body's
79 surface is cooled to 0K while the other half is heated to 2K, the global mean temperature remains 1K,
80 but the power needed to maintain this distribution increases eightfold. This illustrates how global mean
81 temperature fails to reflect energy balance.

82

83 Spatial and Temporal Temperature Distribution Matters

84 The Earth's temperature is governed by energy balance—solar shortwave radiation heats the
85 planet, while long-wave infrared radiation cools it. The Earth's effective radiometric temperature,
86 calculated as:

$$87 \quad T_{Earth} = \sqrt[4]{\frac{(1-\alpha)S}{4\sigma}} \approx 255\text{K} = -18^{\circ}\text{C}, \quad (2)$$

88

89 where S is the solar constant, σ is the Stefan-Boltzmann constant, and α is albedo (~ 0.3), differs
90 significantly from the actual mean surface temperature ($\sim 13.9^\circ\text{C}$), highlighting the atmosphere's
91 role [6].

92 Two problems arise immediately, making these comparisons meaningless. First, the albedo differs
93 between a bare planet and a planet with an atmosphere. Second, the temperature calculated using
94 equation (3) assumes a uniformly heated planet, which never actually occurs in reality.

95 To emphasize the importance of spatial temperature distribution, let us consider an extreme case.
96 Imagine, for simplicity, that one side of the planet is uniformly heated while the other side remains
97 cold at $T_2 \approx 0\text{K}$. The temperature of the heated side can be calculated using equation (2), with the
98 denominator adjusted from four to two, yielding $T_1 \approx 303\text{K}$. Consequently, the global mean
99 temperature, given the same energy input, is now $T_{Earth} = (T_1 + T_2)/2 \approx 152\text{K}$!

100 Now, imagine a planet that is heated uniformly for half of the time by two Suns (i.e., $2S$) and
101 remains unheated for the other half. On average, over time, the planet receives the equivalent of
102 one Sun's energy input. If we further assume that the planet has a low heat capacity and cools
103 immediately when not exposed to a heat source, we obtain the same result as in the previous
104 example.

105 Thus, one can conclude that any spatial or temporal temperature variation leads to a decrease in
106 the global mean temperature. However, this does not necessarily mean the planet is cooling—it
107 simply affects the way energy is distributed over time and space.

108

109 **THE PROPOSED METHOD**

110 The core idea is to establish a relationship between energy balance and global temperature. Due to the
111 complexity of atmospheric influences, let us first analyse a bare planet, such as the Moon. Since it is at
112 the same distance from the Sun as Earth, the solar constant remains unchanged. The Moon's albedo is
113 approximately 0.12, allowing us to estimate its average temperature using the zeroth-order model [6]:

114 $T_{Moon} = 270\text{K}$.

115 However, since the Moon’s temperature distribution is highly inhomogeneous, and the cooling at each
116 point on the surface is proportional to the fourth power of its temperature, it is essential to account for
117 temperature variations over time and space to accurately determine the outgoing energy flux.

118 In theory, the total outgoing energy flux can be calculated by integrating the emitted radiation over a
119 given period and across all points on the surface. At equilibrium, the total incoming energy flux over a
120 certain period must equal the total outgoing energy flux, ensuring an energy balance.

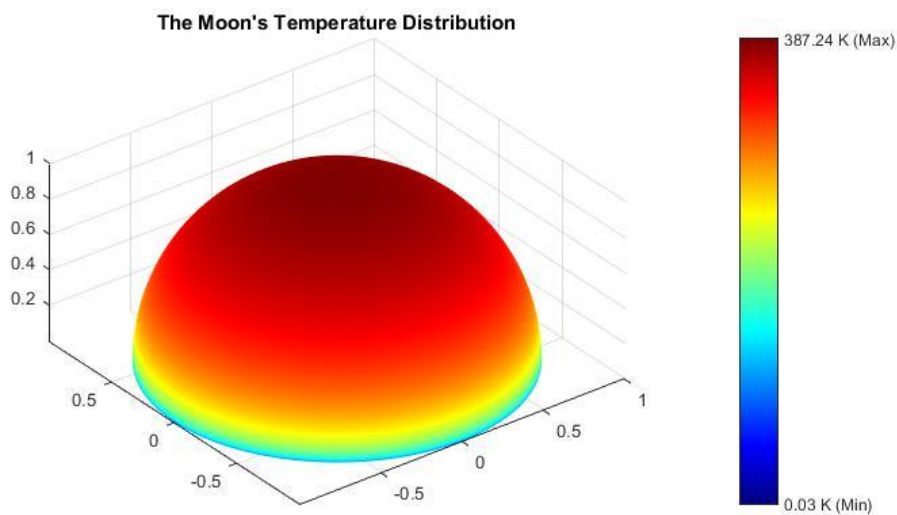
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122 The Simplest Model of the Moon’s Temperature Distribution

123 To demonstrate the application of this methodology, let us assume steady-state radiative equilibrium at
124 each point on the Moon’s sunlit surface, following the cosine law for solar irradiation. Furthermore, we
125 assume that the absorbed energy at any given point is entirely radiated into space as thermal energy:

126
$$S(1 - A)\cos(\theta) = \varepsilon\sigma T_{Moon}^4(\theta), \quad (3)$$

127 where A —the Moon’s albedo (~ 0.11), θ —angle of incidence, ε —the emissivity (~ 0.95) and
128 $T_{Moon}(\theta)$ —represents the surface temperature at a given point. Although this model serves as a first-
129 order approximation of the Moon’s temperature distribution, it ignores rotation, thermal inertia, and heat
130 conduction. Therefore, it should only be used for quick estimations. The resulting “frozen-in-time”
131 temperature distribution is illustrated in Fig. 1.



132

133 Fig. 1. The simplest model of the Moon’s surface temperature distribution

134

(Sun-facing hemisphere)

135 Knowing the surface temperatures in any point, one can now discretize the Moon's surface into N cells
136 with the same temperature and calculate the sum of outgoing flux:

$$137 \quad \varepsilon \sum_{i=1}^N \Delta S_i \sigma T_i^4. \quad (4)$$

138 This flux must be equal to the total absorbed solar energy:

$$139 \quad (1 - A) S \pi R_{Moon}^2. \quad (5)$$

140 To characterize the energy balance with a single parameter (which we call Global Energetic
141 Temperature T_{GET}), we use the following relation:

$$142 \quad \varepsilon \sum_{i=1}^N \Delta S_i \sigma T_i^4 = \varepsilon A \pi R_{Moon}^2 \sigma T_{GET}^4. \quad (6)$$

143 This derived temperature has a physical meaning: it represents the temperature the Moon would have if
144 its entire surface were uniformly heated. In our simplified model, this temperature is approximately
145 274K.

146 In contrast, the traditional approach of computing the global mean temperature using:

$$147 \quad T_{ave} = (\sum_{i=1}^N \Delta S_i T_i) / \sum_{i=1}^N \Delta S_i \quad (7)$$

148 (while assuming that the Moon's dark side is at 0K) yields $T_{ave} \approx 154K$. However, this value has no
149 physical significance; it only serves as an indicator of temperature inhomogeneity rather than a
150 meaningful global metric.

151 The Earth Case

152 The same methodology applies to Earth, though its temperature distribution is significantly more
153 complex due to atmospheric effects. Unlike the Moon, the Earth's cooling mechanism involves both
154 radiation and convection, making it difficult to model temperature evolution purely from radiative
155 balance equations. However, using measured temperature datasets, one can estimate what Earth's
156 cooling would look like in the absence of an atmosphere, keeping the same temperature distribution.
157 The total potential outgoing radiative flux can then be expressed as:

$$158 \quad \varepsilon \sum_{i=1}^N \Delta S_i (\sum_{j=1}^M \sigma T_j^4 \Delta t_j), \quad (8)$$

160 where are ΔS_i – the area of i th cell, N –total number of the grid cells, Δt_j – j time step, M –total
161 number of time steps in the chosen time period, $\tau = \sum_{j=1}^M \Delta t_j$, $S_{Earth} = \sum_{i=1}^N \Delta S_i$ – the total surface
162 area of Earth.

163 The total influence of the atmosphere, including the greenhouse effect, can be easily determined
164 as the difference between the potential cooling energy (given by Equation 8) and the total incoming
165 energy that the Earth would receive during the same period if it had no atmosphere. In this case,
166 the albedo value is adjusted to match that of the Earth's surface, $\alpha_{Earth\ surface} \approx 0.2$ [7], and the
167 total incoming flux is given by:

$$168 \quad (1 - \alpha_{Earth\ surface})S\pi R_{Earth}^2\tau. \quad (9)$$

169 Thus, the correct method for calculating the Earth's temperature over a specific time period—
170 allowing for estimates of climate change (warming or cooling)—is:

$$171 \quad \sum_{i=1}^N \Delta S_i (\sum_{j=1}^M \sigma T_j^4 \Delta t_j) = S_{Earth} \sigma T_{P.C}^4 \tau. \quad (10)$$

172 This temperature $T_{P.C}$, can be referred as the “effective temperature for potential cooling”. Previous
173 parameter—the global mean temperature could serve as an indicator for spatial and temporal
174 temperature redistribution.

175 **CONCLUSION**

176 This leads to the conclusion that the current methodology for estimating global warming and climate
177 change—based on calculating the global mean temperature—is insensitive and inadequate. The primary
178 issue is that the global mean temperature lacks direct physical meaning and has no explicit connection
179 to the energy balance.

180 Mathematically, this inconsistency arises because the global mean temperature is a linear combination
181 of individual temperatures, whereas the energy budget is highly nonlinear, depending on the fourth
182 power of temperature (as dictated by the Stefan-Boltzmann law).

183 To address this limitation, we propose a more physically accurate method, which, when tested, could
184 provide new insights into climate change using existing datasets. The concept of the "effective
185 temperature for potential cooling" offers a more sensitive indicator of global warming and the impact
186 of increasing greenhouse gas concentrations in the atmosphere.

187 An important advantage of this approach is that it can be easily implemented, as it utilizes existing
188 climate datasets. The reference point could be set in the pre-industrial era (1850), allowing for a precise
189 reconstruction of the natural greenhouse effect and a continuous assessment of climate change and
190 global warming based on changes in the effective temperature for potential cooling.

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