

1 Mechanisms of fault mirror formation in carbonate rocks

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12

13 Abstract

14 The development of smooth, mirror-like surfaces provides insight into the mechanical behaviour  
15 of crustal faults during the seismic cycle. To determine the physico-chemical mechanisms of fault  
16 mirror formation, we investigated carbonate fault systems in seismically active areas of central  
17 Greece. Using multi-scale electron microscopy combined with Raman and electron energy loss  
18 spectroscopy we show that fault mirror surfaces do not need to develop from nanogranular  
19 volumes, as suggested by previous investigations. Instead, we demonstrate that decarbonation is  
20 the main transformation process that leads to the formation of smooth surface coatings in these  
21 faults. Piercement structures on top of the fault surface indicate calcite decarbonation, producing  
22 CO<sub>2</sub> and lime (CaO), where lime subsequently reacts to portlandite (Ca(OH)<sub>2</sub>) under hydrous  
23 conditions. Nanoscale imaging and electron diffraction reveals a thin coating of a non-crystalline  
24 material sporadically mixed with nano-clay, forming a complex-composite material that smooths  
25 the slip surface. Spectroscopic analyses reveal that a thin layer of non-crystalline carbon covers

26 the slip surface. We suggest that, during the post-seismic period, ordering (hybridisation) of  
27 amorphous carbon led to the formation of partly-hybridised amorphous carbon but did not reach  
28 full graphitisation. Calcite nanograins, < 50 nm in size, are spatially associated with the carbon  
29 and indicate that the decomposition products acted as a crystallisation medium. Within this  
30 medium portlandite back-reacted with CO<sub>2</sub> to form nanocrystalline calcite. Hence, we suggest  
31 that the nanograins are not the result of comminution during slip but originate from  
32 pseudomorphic replacement of calcite after portlandite. Our results suggest that calcite  
33 decarbonation products may develop across the entire fault surface, controlling the formation of  
34 carbonate fault mirrors, and may facilitate slip on a decarbonation-product glide film.

## 35 [1 Introduction](#)

36 Brittle deformation of upper-crustal rocks can result in high-magnitude seismic events (Scholz,  
37 1998). Mirror slip surfaces (MSSs) along principal slip zones in carbonate rocks provide an  
38 excellent opportunity to investigate the deformation processes that occur in relation to slip events  
39 in the seismogenic zone. The defining feature of MSSs is the high degree of visible light  
40 reflectance resulting from a low surface roughness. MSSs can form at seismic slip velocities and,  
41 therefore, may indicate paleo-seismicity (Siman-Tov et al., 2013; Smith et al., 2013; Fondriest et  
42 al., 2013; Kirkpatrick et al., 2013; Spagnuolo et al., 2015). However, MSS have also been  
43 developed at sub-seismic slip conditions in deformation experiments (Verberne et al., 2014) and,  
44 thus, they are not necessarily diagnostic of seismic slip.

45 Studies suggest that the low surface roughness of MMS may be the result of nano-sized  
46 grains (< 1 µm), which constitute the uppermost layers of principal slip surfaces (Fondriest et al.,  
47 2013; Siman-Tov et al., 2013; Collettini et al., 2014). Therefore, the production of a nanogranular  
48 slip-surface coating has been associated with the development of natural MSSs during high-

49 magnitude seismic events. In addition, nanogranular coatings are also produced in deformation  
50 experiments under seismic conditions (Green et al., 2015; Spagnuolo et al., 2015), supporting this  
51 hypothesis. Current research suggests different formation mechanisms for MSSs related to  
52 nanograins: (1) Formation of long beams by plasticity and subsequent brittle fracturing of the  
53 beams (Siman-Tov et al., 2013), (2) localised dynamic recrystallisation and static recrystallisation  
54 (Smith et al., 2013), (3) dislocation cell formation (Verberne et al., 2013) and (4) sharp  
55 rheological boundaries separating high-, and low-strained regions (Pozzi et al., 2018).

56 One of the main mechanisms of fault gouge formation is grain-size reduction by  
57 comminution. The grain-size dependence of the modified Hall-Petch relation proposed by  
58 (Sammis and Ben-Zion, 2008) illustrates that smaller grain sizes require higher differential  
59 stresses to fracture. Therefore, local stress concentrations during coseismic events may be  
60 favourable sites for generation of nanograins by fracturing micron to millimetre-sized grains.  
61 Ultimately, these nanograins may not only be evidence for paleo-seismic events but may in  
62 addition control the frictional behaviour of the fault (Han et al., 2007; Han et al., 2010; Di Toro et  
63 al., 2011).

64 Less considered effects of fault rock deformation in carbonates are chemo-mechanical  
65 transitions, e.g., decarbonation. Products of decarbonation processes are reported from  
66 experiments (Verberne et al., 2014; Spagnuolo et al., 2015; Delle Piane et al., 2017) and from  
67 natural faults (Vigano et al., 2011; Collettini et al., 2013). The experimental studies report the  
68 presence of amorphous carbon as a deformation product. Some natural faults also exhibit  
69 amorphous glass coatings around euhedral calcite crystals in potential carbonate pseudotachylites  
70 (Vigano et al., 2011). Furthermore, deformation experiments on silicates (Di Toro et al., 2004;  
71 Pec et al., 2012; Hirose et al., 2012; Toy et al., 2015) and on clay-bearing materials (Goldberg et  
72 al., 2016) produce non-crystalline or partly-amorphous silicate materials (Pec et al., 2012).

73 We investigated the surface and the first tens of microns of two natural carbonate fault-  
74 mirror structures to test the hypothesis of a nanograin surface coating. Our high-resolution  
75 electron microscopy and Raman spectroscopy analyses demonstrate that a thin coating of  
76 decarbonation products is able to produce a MMS without the need of a nanogranular surface  
77 coating. In addition, we suggest that the amorphous products may facilitate and lubricate seismic  
78 slip (De Paola et al., 2011; Di Toro et al., 2011) or possibly re-strengthen the fault by  
79 recrystallisation of the decomposition products during the interseismic period.

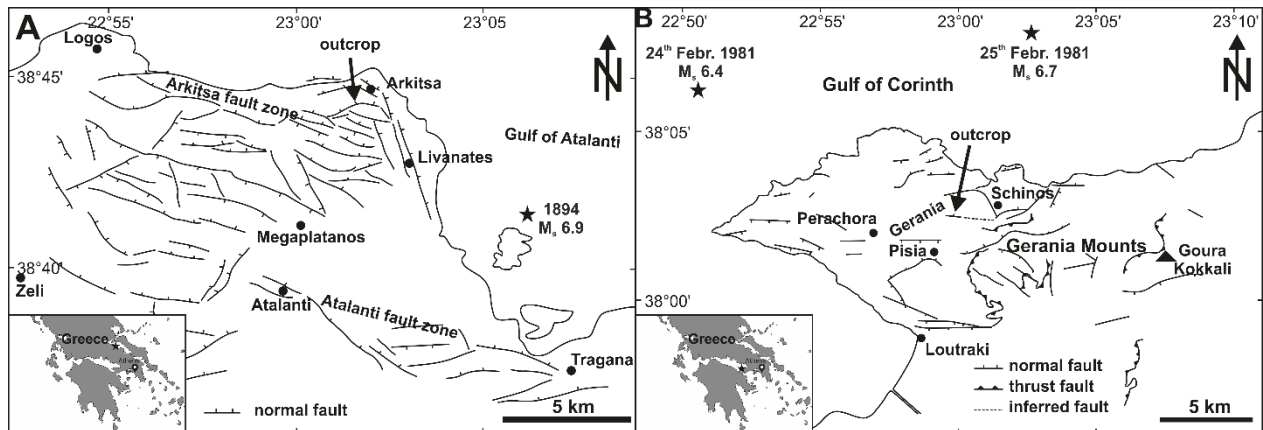
## 80 2. Geological Setting

### 81 2.1 Geology of the Arkitsa fault zone

82 The 700 m long Arkitsa fault (Fig. 1A) belongs to the Kamena Vourla fault system; a  
83 northward-dipping, ESE-WNW striking, active normal fault zone of approximately 50 km length  
84 along the southern coast of the Gulf of Evia. The left-stepping fault system consists of the  
85 Kamena Vourla, the Agios Konstantinos and the Arkitsa fault segments (Roberts and Jackson,  
86 1991; Ganas et al., 1998). The study area (38°43'56.17"N, 23° 0'27.41"E) is situated within the  
87 Pelagonian zone of Central Greece. The Arkitsa fault planes juxtapose Late Triassic to  
88 Middle/Late Jurassic platform carbonates in the footwall with lower Pliocene-Pleistocene to  
89 Quaternary sediments in the hanging wall (Kokkalas et al., 2007). The Arkitsa fault scarp is an  
90 anthropogenically formed fault exposure first described by Jackson and McKenzie (1999).  
91 Quarrying for two decades removed a major part of the hanging-wall colluvium to reveal three  
92 large, smooth fault planes of up to 65 m height (Kokkalas et al., 2007).

93 Historical seismic records from the Gulf of Evia area record about 13 rupture events from  
94 426 BC until the last major event in 1894 (Ganas et al., 1998; Ganas et al., 2006), with a  
95 magnitude  $M_s$  6.9 earthquake hosted within the Atalanti fault zone (Fig. 1A) (Ambraseys and  
96 Jackson, 1990). This number of events contrasts with the approximately 50 slip increments

97 identified by Jackson and McKenzie (1999), where the authors calculated the slip increments  
 98 from fault plane height and average earthquake slip distances. Evidence for Holocene seismic  
 99 activity along the Arkitsa fault planes is recorded by the approximately 1 m unweathered fault  
 100 scarp existing prior to quarrying (Jones et al., 2009).



101  
 102 **Figure 1:** Tectonic maps of the study areas. Insets show the geographical location within central Greece. **A:**  
 103 Location of the studied exposure of the Arkitsa fault zone (modified after Jones et al., 2009; earthquake data:  
 104 NOAA). **B:** Outcrop location of the Schinos fault segment north of Corinth (modified after Kaplanis et al., 2013 and  
 105 Collier et al., 1998).

## 106 2.2 Geology of the Schinos fault zone

107 The Schinos fault zone is located north-east of Corinth within the Gerania mountains with  
 108 the studied exposure ( $38^{\circ} 2'14.40''N$ ,  $23^{\circ} 0'22.33''E$ ) belonging to the Gerania unit of the Internal  
 109 Hellenides (Kaplanis et al., 2013). The stratigraphic succession of the Gerania unit includes from  
 110 bottom to top: Permian sedimentary and volcanic rocks, Upper Triassic neritic limestones and  
 111 dolomites, Lower Jurassic limestones, an Upper Jurassic succession of rift-related sediments  
 112 overlain by either Ammonitico Rosso (Bathonian or Oxfordian age) or radiolarian sediments, and  
 113 ophiolite-derived turbidites of the Beotian flysch (Kaplanis et al., 2013).

114 Historic seismic data record an earthquake sequence in the Corinth region with three main  
 115 shocks: 24<sup>th</sup> February 1981,  $M_s$  6.7; 25<sup>th</sup> February 1981,  $M_s$  6.4 and 4<sup>th</sup> March 1981,  $M_s$  6.4  
 116 (Collier et al., 1998). Based on the epicentral location and a focal mechanism mismatch, Collier  
 117 et al. (1998) interpreted that the 24<sup>th</sup> February event started with displacement on an offshore

118 fault and continued to rupture the Pisia fault, producing a surface break. In addition, Collier et al.  
119 (1998) propose the same succession of events for the 25<sup>th</sup> February event, which caused a  
120 displacement on the Schinos fault, probably forming the most recent event on the fault plane that  
121 we study here.

### 122 3. Methods

123 We analysed drill-core samples from the Arkitsa and Schinos fault exposures using a  
124 range of micro-analytical techniques. In total, 40 drill cores, 2.54 cm in diameter, were collected  
125 from the Arkitsa fault zone and 18 were collected from the Schinos fault zone. The maximum  
126 retrievable drill-core lengths did not exceed 10 cm (Arkitsa) and 5 cm (Schinos) as the porosity  
127 of the fault rock increases drastically and the material loses its cohesion with distance from the  
128 fault surfaces. Seven samples from the Arkitsa fault zone and four samples from the Schinos fault  
129 zone were chosen for further investigation based on the best-preserved mirror surfaces. Sample  
130 cores were sputter coated with an 8-nm thick layer of platinum/palladium for charge deduction.

131 From both faults, two representative samples were selected for electron-transparent foil  
132 preparation using a FEI Helios Nanolab G3 Dualbeam focused ion beam scanning electron  
133 microscope (FIB-SEM). Prior to ion beam deposition of the main platinum strip, a 200-nm layer  
134 of platinum was deposited using the electron beam (2kV, 0.4 nA) to prevent surface  
135 amorphisation by the ion beam as a preparation artefact. Transmission electron microscope  
136 (TEM) investigations of the foils were carried out with a FEI Talos F200X. TEM images were  
137 acquired with 200 kV acceleration voltage and 5 nA or 10 nA beam current depending on final  
138 sample thickness. The FEI Talos SuperEDX detector was used for chemical element mapping via  
139 energy dispersive X-rays (EDX) in scanning transmission electron microscope (STEM) mode.  
140 EDX maps were quantified post-acquisition with Bruker Esprit 1.9 software using the Cliff-  
141 Lorimer approximation and ImageJ.

142 Electron energy loss spectroscopy (EELS) data were acquired with a Zeiss Libra 200FE in  
143 TEM mode at 200 kV with an Omega in-column energy filter. The energy resolution of EELS  
144 analyses was 0.7 eV, measured at the full width half maximum (FWHM) of the zero-loss peak.  
145 Energy loss spectra were obtained at 250,000x magnification with a 100  $\mu\text{m}$  filter-entrance  
146 aperture giving an effective aperture of about 40 nm on the sample. The convergence angle of  
147 EELS was about 0.1 milliradian (mrad), and the acceptance angle was defined by the 60  $\mu\text{m}$   
148 diameter of the objective aperture giving a collection angle of 11.6 mrad. Measurement times of  
149 EEL spectra were set to 1–5 s, with 5–10 frames/spectra, on a slow-scan CCD (model UltraScan  
150 4000), with binning of 4x4 pixels (giving 1024 pixels), and an energy spread of 0.08 eV/pixel,  
151 resulting in an energy range of 82 eV in the acquired spectra. Spectra were deconvolved with the  
152 zero-loss peak to remove the effect of plural scattering and background-subtracted assuming a  
153 power law function with Gatan's Digital Micrograph software.

154 Raman spectroscopy was carried out with a WiTec ALPHA300 R confocal microscope to  
155 identify and characterize possible reaction or decomposition products on the fault surfaces. We  
156 used a laser with a wavelength of 532 nm and a spectral grating of 600 grooves/cm. Data  
157 acquisition, data post-processing and peak fitting were performed with WiTec ProjectFour 4.1  
158 software and Fityk 0.98 (Wojdyr, 2010) peak fitting software. In accordance with (Ferrari and  
159 Robertson, 2000) we determined the intensity ratios of the D to G peak,  $\frac{I(\text{D})}{I(\text{G})}$ , to obtain  
160 information about the degree of crystallinity within carbon phases. We used the approach by  
161 Ferrari and Robertson (2000) based on first principle considerations to determine the peak height  
162 ratios because the information about the less disordered aromatic rings and the clustering of the  
163  $\text{sp}^2$  phase is contained in the intensity maximum of the D peak and not in the width. Calculation

164 of reaction enthalpy values were executed with SUPCRT92 (Johnson et al., 1992) at standard  
165 conditions (1 bar, 298 K).

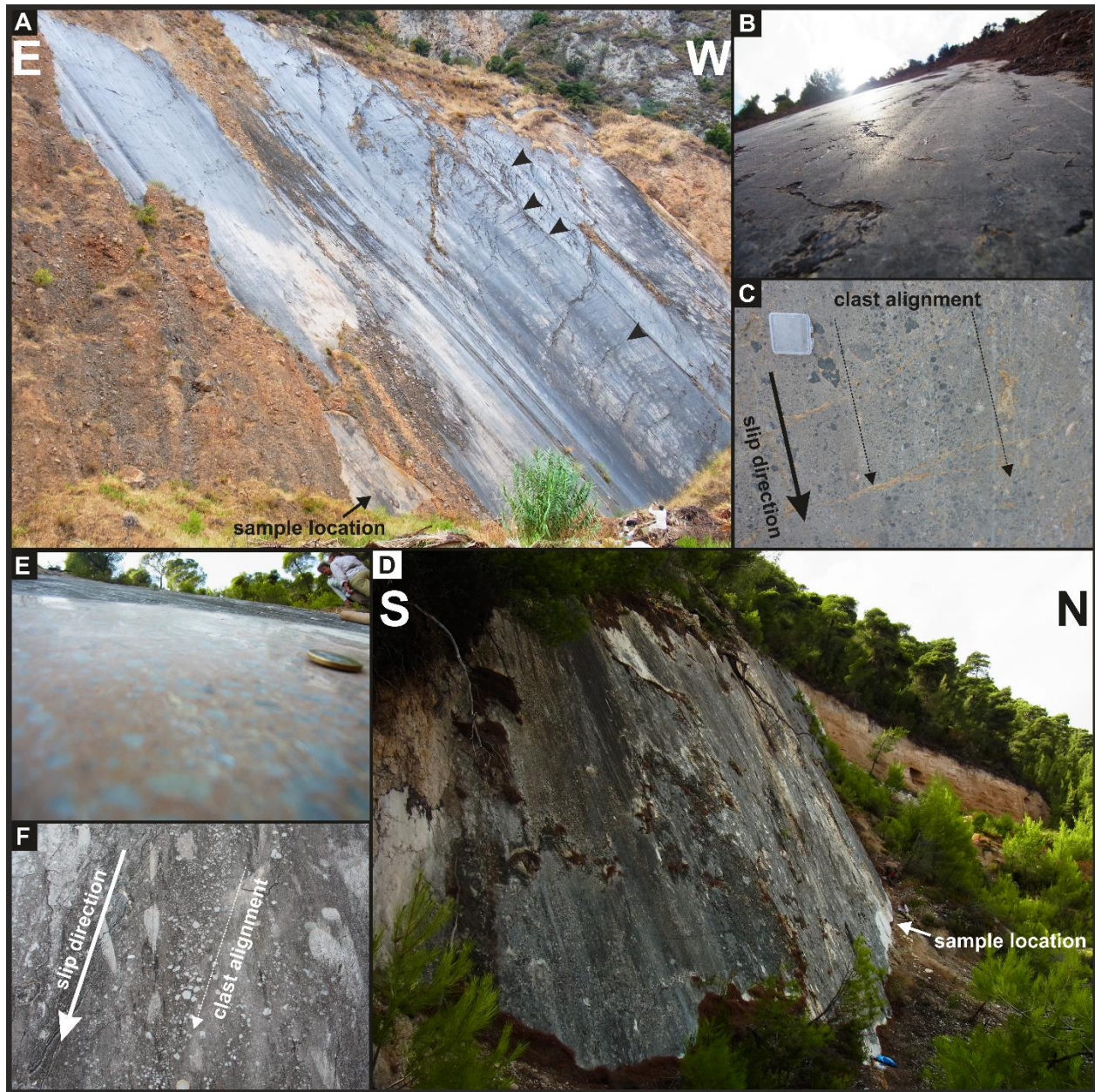
## 166 4. Results

### 167 4.1 Field results

168 The exposed fault surface of the Arkitsa fault (Fig. 2A) exhibits a range of slip-related  
169 structures. On the outcrop scale, the fault exposure steps over, suggesting the presence of several  
170 slip planes inside the fault damage zone. The fault surface contains pronounced slip grooves  
171 parallel to the direction of oblique slip. Fractures with a spacing on the order of one metre are  
172 oriented approximately perpendicular to the grooves (Fig. 2A). Parts of the slip plane are covered  
173 with residual reddish-brown hanging-wall breccia, incorporating fragments of dark host-rock  
174 carbonate up to several decimetres in size (Fig. 2A). The most prominent feature is the low  
175 roughness of the fault surface, which enables the reflection of sunlight (Fig. 2B). The fault rock is  
176 a matrix-supported greyish, cohesive material, with dark carbonate clasts up to several  
177 centimetres in size (Fig. 2C). Slip-parallel alignment of clasts on the fault surfaces can be traced  
178 over several meters (Fig. 2C).

179 The Schinos fault surface (Fig. 2D) exhibits a range of structures formed during slip. We  
180 observe stepovers, indicating the presence of several fault planes inside the fault damage zone.  
181 The fault surface is extremely smooth with areas not only reflecting light but also mirroring the  
182 surrounding vegetation (Fig. 2E). Fragments of grey carbonate host rock are incorporated into the  
183 red-orange footwall fault rock. The clasts are strongly aligned in the slip direction and usually do  
184 not exceed 1–2 cm in size (Fig. 2F). The fault surface contains pronounced slip grooves and a  
185 wavy surface morphology (Fig. 2D). The mirror surfaces of both faults are hosted inside a zone  
186 of high competence with average thicknesses of 10 cm and 5 cm for Arkitsa and Schinos,  
187 respectively.



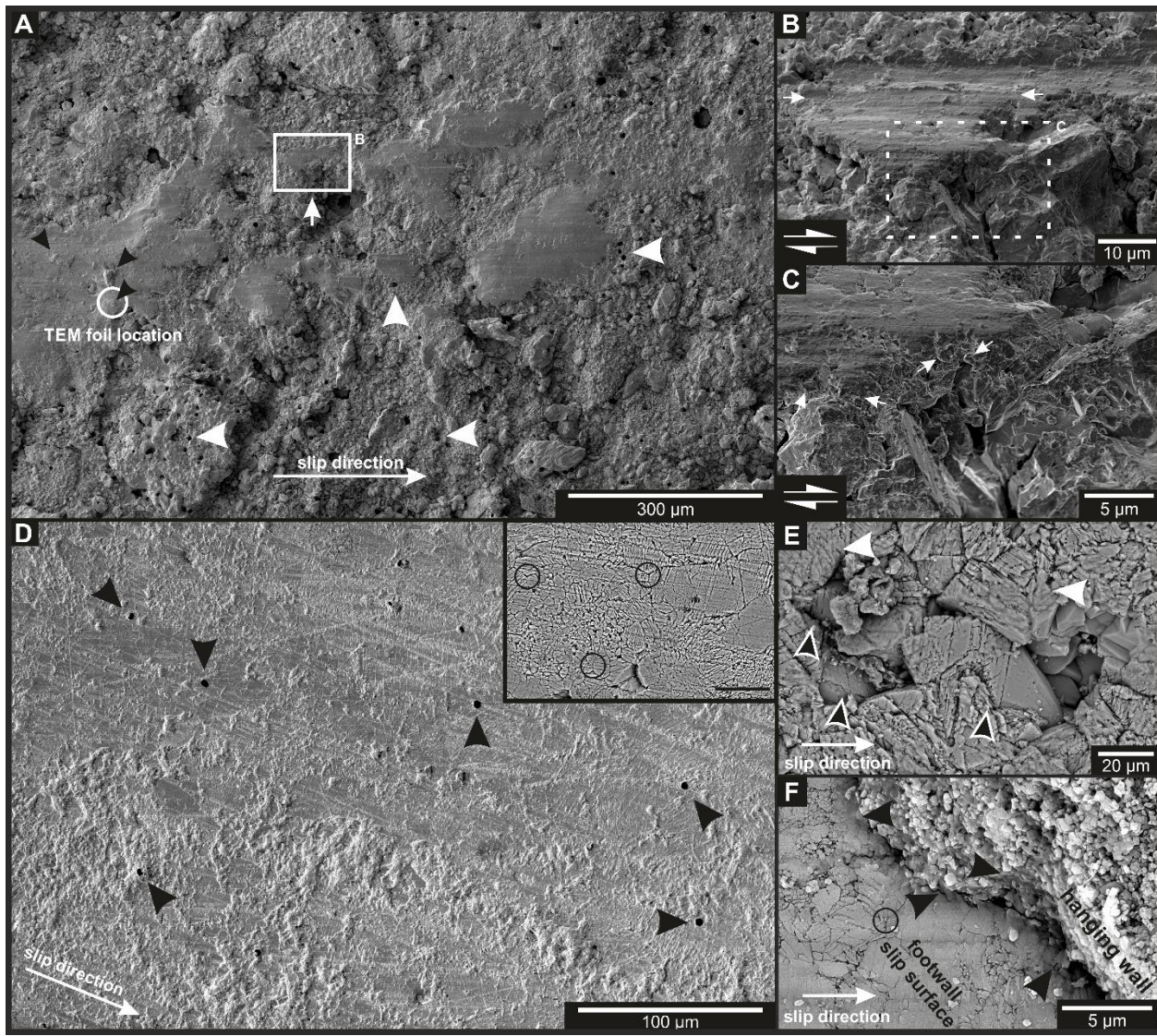


188  
 189 **Figure 2:** Representative field structures from Arkitsa (A – C) and Schinos (D – F) fault exposures. **A:** Overview of  
 190 Arkitsa fault exposure showing one of the three main slip planes with reddish-brown hanging-wall breccia cover on  
 191 part of the slip surface. Black arrows indicate fractures perpendicular to slip direction. Person for scale, bottom  
 192 right corner. **B:** Fault surface reflecting sunlight. **C:** Alignment of host rock fragments parallel to slip direction. Box  
 193 approx. 5 x 4 cm. **D:** Schinos fault surface exposure shows grey weathering of the slip plane. The slip plane is curved  
 194 with the salient pointing out of the figure. Fault scarp to the north is now covered by alluvium. **E:** Image  
 195 demonstrating fault surface reflectivity. Reflection of vegetation along the top edge and above the one-euro coin. **F:**  
 196 Grey, weathered fault plane showing a strong alignment of light grey host-rock clasts parallel to slip direction (one-  
 197 euro coin for scale, bottom edge).

#### 198 4.2 Slip surface microstructures

199 SEM analyses of the Arkitsa and Schinos slip surfaces reveal a low surface roughness on  
 200 the microscopic level. The surfaces are well polished and, in some places, preserve parts of the

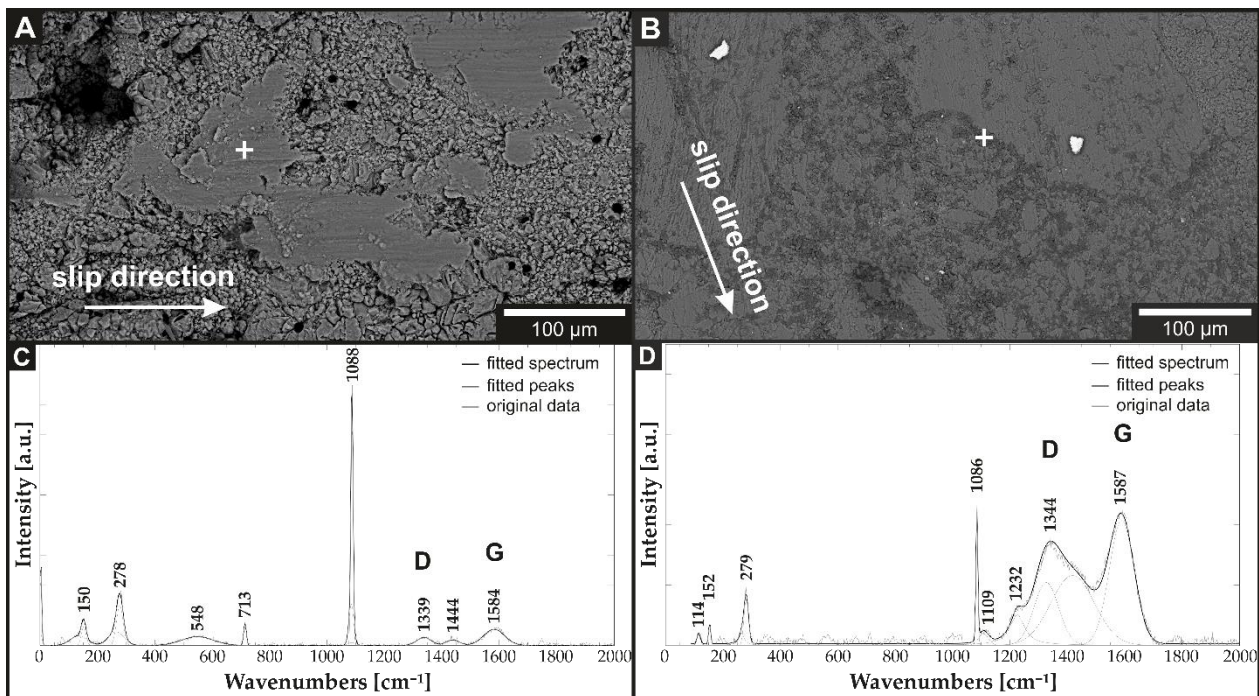
201 hanging wall (Fig. 3A and F). Microgrooves are aligned parallel to the slip direction (Fig. 3B)  
 202 and holes with a diameter of 6–8  $\mu\text{m}$  pierce the slip surfaces (Fig. 3A and D). Localized erosion  
 203 of the Arkitsa slip surface exposes the underlying fault rock, with a grain size of 2–5  $\mu\text{m}$  (Fig.  
 204 3A, B and C). The first 10–20  $\mu\text{m}$  of the fault rock below the Arkitsa slip surface consist of a  
 205 fine-grained deformation product, which is situated on top of a comparatively coarse-grained,  
 206 less deformed material (Fig. 3B and C). The contact between the fine- and coarse-grained  
 207 domains is established by a coating phase (Fig. 3B and C).



208

209 *Figure 3: Representative slip surface microstructures of the Arkitsa (A -C) and Schinos (D - F) exposures. A:*  
 210 *Secondary electron overview image of the Arkitsa slip surface. Smooth slip patches with microgrooves parallel to*  
 211 *slip direction. Uniform holes (white arrows) penetrate the slip surface. Residue of hanging wall on top of the slip*  
 212 *surface patch (black arrows and inside the white circle) with TEM sample location. B:* *Close-up of slip surface from*  
 213 *A. Fine-grained slip surface material situated on top of coarse fault rock grains. Microscope stage tilted to 52°. C:*  
 214 *Close-up of section from B. Roundish grains are covered by coating phase (white arrows). D:* *Secondary electron*  
 215 *image of Schinos slip surface with microgrooves parallel to slip direction. Uniform holes penetrate the principal slip*  
 216 *surface (black arrows). Inset: Backscatter electron image illustrating triple junction grain boundary contacts. Scale*  
 217 *bar = 50 µm. E:* *Backscatter electron image top view onto the slip surface. Cracks in idiomorphic calcite crystals*  
 218 *are visible a few micrometres into the crystal (black-white arrows). Coating phase with low backscatter contrast*  
 219 *covers parts of the slip surface and reduces surface roughness (white arrows) F:* *Backscatter electron image of the*  
 220 *contact between footwall and hanging-wall. Black arrows indicate locations of a coating phase with lower*  
 221 *backscatter contrast.*

222 In contrast, the Schinos fault surface consists of large, truncated grains with boundaries  
 223 that meet in triple junctions (Fig. 3D and inset). A fragmented layer lies on top of undeformed  
 224 calcite crystals (Fig. 3E) and the damage extends about 10–20 µm into the fault rock. Here also a  
 225 coating phase is present on the slip plane (Fig. 3E and F). On both faults, Arkitsa and Schinos,  
 226 the coating phase has a low contrast in backscattered-electron images (Fig 4A and B). The phase  
 227 also coats grain edges and reduces the slip surface roughness (Fig. 3C, E and F). In addition, the  
 228 coating phase infiltrates interstitial regions of the hanging-wall breccia and holds residual pieces  
 229 of hanging wall in place (Fig. 3F).

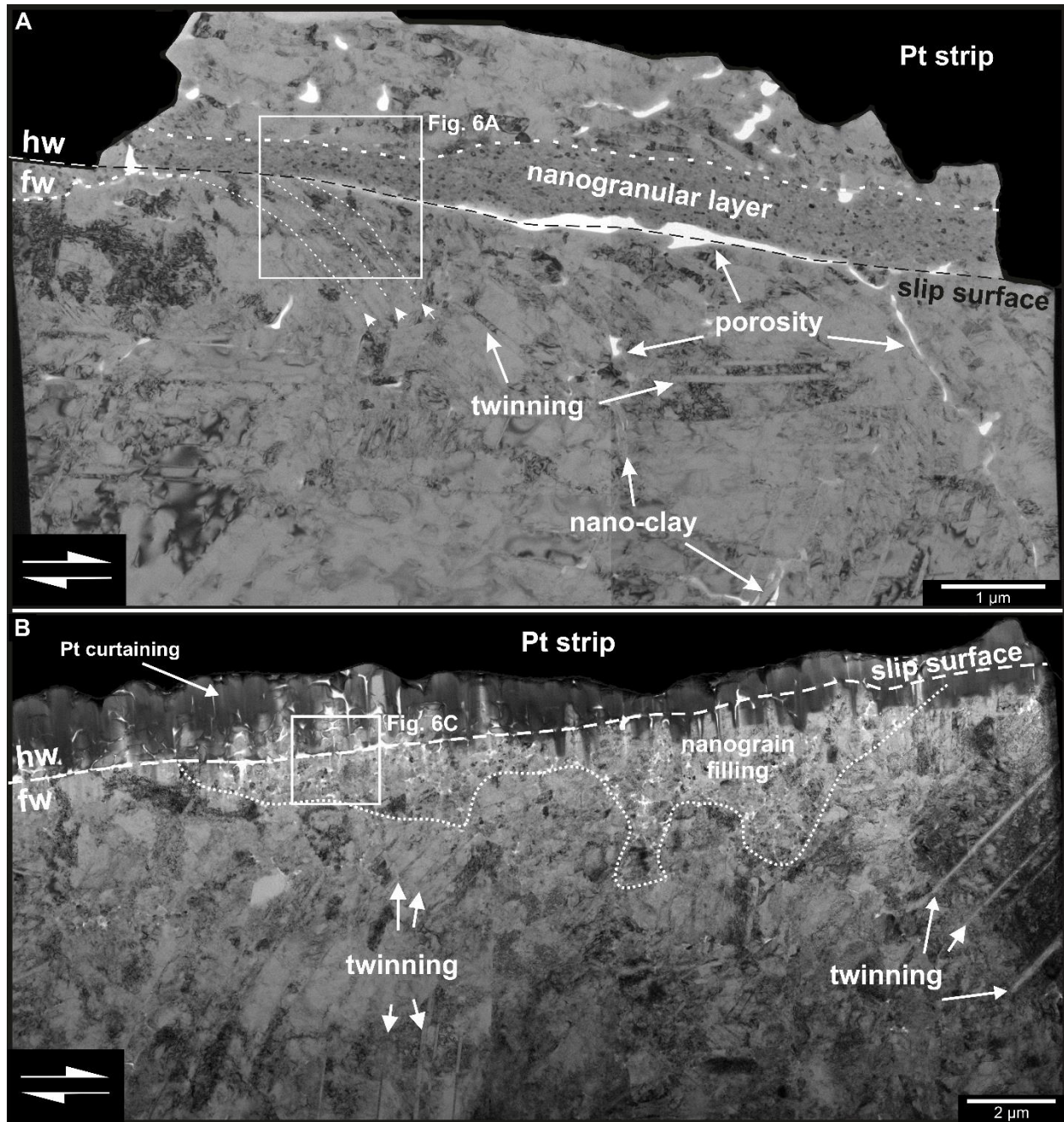


230 *Figure 4: Backscatter electron images with locations of Raman spectroscopy measurements for Arkitsa (A and C)*  
 231 *and Schinos (B and D) slip surfaces. **A and B:** Backscatter electron images of the fault surfaces. A phase with low*  
 232 *backscatter contrast (darker grey) is visible in B. **C and D:** Raman spectra showing the D and G peak position of a*  
 233 *disordered carbon phase.*

234 Figure 4 presents Raman spectra from the coating phase. Two broad peaks between 1200–  
 235 1700  $\text{cm}^{-1}$  indicate the presence of a disordered carbon phase, whilst a peak at 1086  $\text{cm}^{-1}$   
 236 demonstrates the presence of crystalline calcite. The fitted spectrum from the Arkitsa fault plane  
 237 gives a  $\frac{I(D)}{I(G)}$  ratio of 0.543 at a D-peak position of 1339  $\text{cm}^{-1}$  and a G-peak position of 1584  $\text{cm}^{-1}$ .  
 238 The fitted spectrum from the Schinos fault plane gives a  $\frac{I(D)}{I(G)}$  ratio of 0.789 at a D-peak position  
 239 of 1344  $\text{cm}^{-1}$  and a G peak position of 1587  $\text{cm}^{-1}$ .

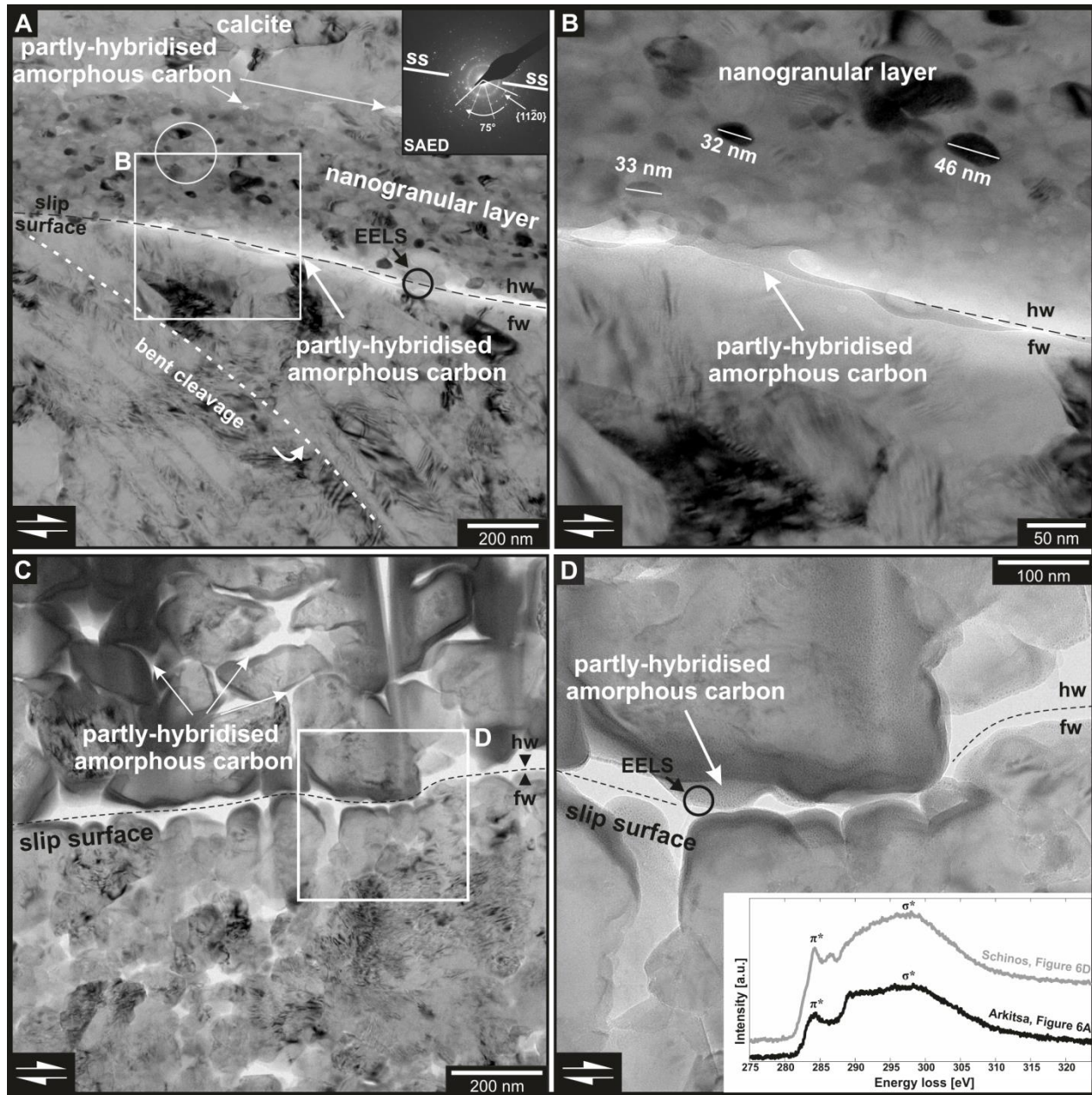
#### 240 4.3 Nanostructures

241 TEM analyses on FIB-SEM foils of both fault-surface exposures reveal a thin surface  
 242 coating. The uppermost layer of both fault-rock exposures is defined by a principal slip surface  
 243 with a thin, non-crystalline coating between hanging wall and footwall (Fig. 5A and B). The  
 244 coating has a homogenous diffraction contrast, shows no lattice fringes (Fig. 6A to D, 8A and F)  
 245 and EDX analyses indicate the presence of carbon (C), iron (Fe), silicon (Si) and aluminium (Al)  
 246 (Fig. 7, Si content of Arkitsa  $\sim$  40 area % estimated with ImageJ). The coating connects hanging-  
 247 wall breccia with the footwall fault surface (Fig. 6B and D) but also extends into the hanging-  
 248 wall breccia (Fig. 6A and C). EELS measurements of the coating phase are given in the inset of  
 249 Fig. 6D. The onset of the  $\pi^*$  peak represents the measured carbon K-edge at 284 eV which is  
 250 followed by a ‘whaleback-shaped’  $\sigma^*$  region. In addition, the measurement of the Schinos sample  
 251 exhibits an additional peak at 286.5 eV possibly indicating C-O bonding.



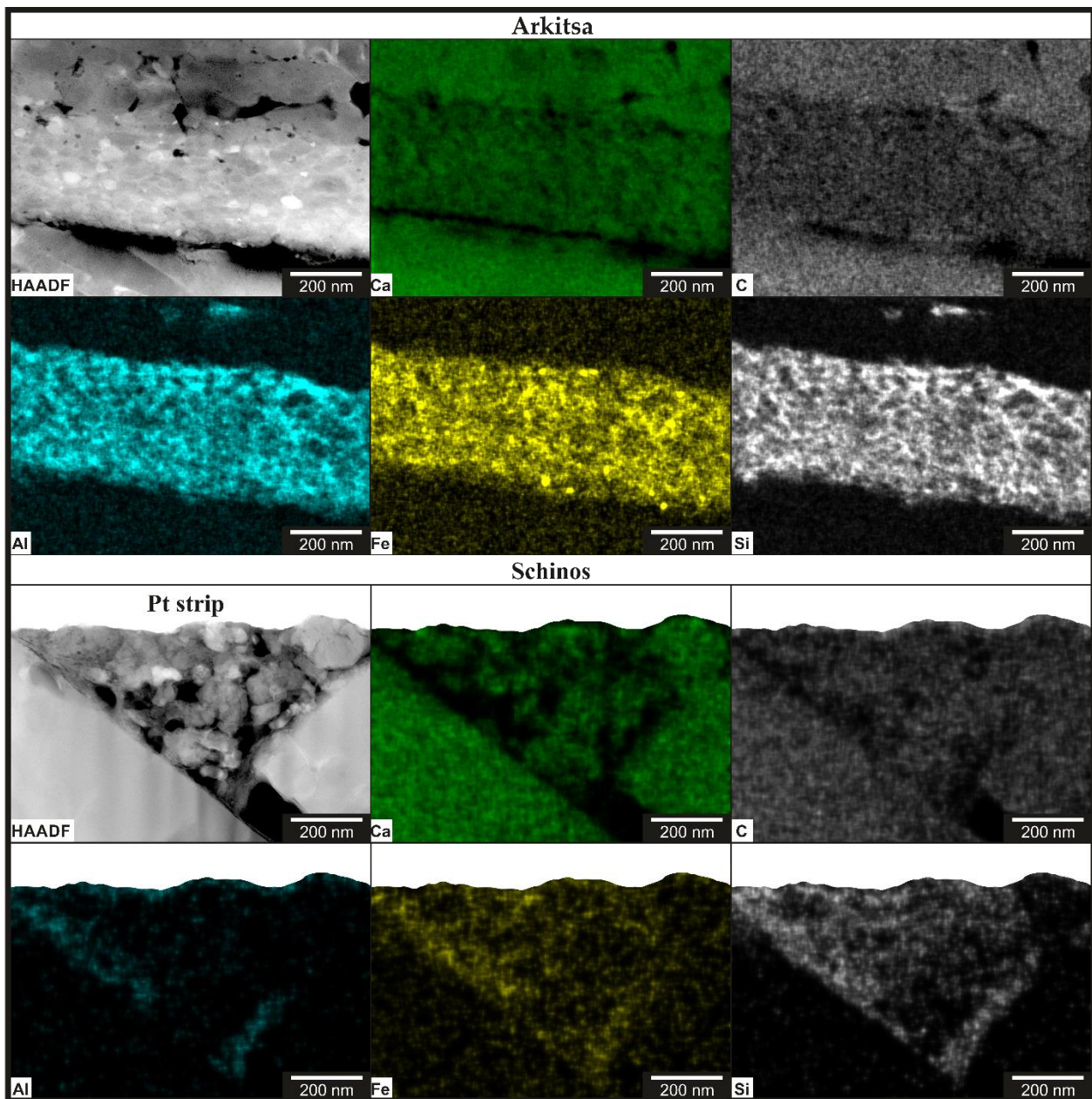
252  
 253 *Figure 5: Overview of TEM cross sections into the principal slip surfaces of Arkitsa and Schinos, including*  
 254 *preserved hanging-wall residue. **A:** Bright-field TEM image overview of representative Arkitsa nanostructures. The*  
 255 *hanging-wall breccia consists of small, elongated calcite fragments held in place by non-crystalline carbon. The*  
 256 *nanogranular layer is part of the hanging-wall breccia and lies on top of deformed calcite crystals. Fractures*  
 257 *(dashed lines) bend down from the principal slip surface into the deformed calcite crystals. Fractures and holes form*  
 258 *porosity inside the mirror slip volume and are filled either with non-crystalline carbon and/or nano-clay. **B:** Bright-*  
 259 *field TEM image with representative Schinos nanostructural hanging-wall to footwall relation. The hanging wall*  
 260 *consists of small grains of several hundred nanometres. Individual grains are surrounded by non-crystalline carbon.*  
 261 *The footwall hosts a nanograin filling, the grain size of which increases abruptly with distance from the slip surface*  
 262 *to grains of several micrometres in size. Larger grains contain twins. hw = hanging wall, fw = footwall.*

263 We observe a direct relation between the spatial distribution of amorphous material and  
264 calcite nanograins. For the Arkitsa fault exposure, a nanogranular layer is present with rounded,  
265 elliptical calcite nanograins of ~50 nm size between the footwall and the hanging-wall breccia  
266 (Fig. 5A, 6A, B and 7). Selected area electron diffraction (SAED) patterns of the Arkitsa  
267 nanogranular layer in Figure 5A demonstrate the polycrystalline nature of the material with  
268 clustered diffraction spots forming two arcs in an angle of about 75° (inset Fig. 6A). Figure 8A  
269 reveals that the thickness of the coating varies along the slip surface and can locally incorporate  
270 calcite grains from the underlying nanostructure. We find that calcite nanograins are not  
271 restricted to the principal slip surface but also occur away from the displacement zone inside the  
272 fault rock (Fig. 8B). The nanograins do not contain any dislocations (Fig. 6B and 8B) and are not  
273 always in grain-to-grain contact in the plane of the section. Calcite nanograins of the Schinos  
274 exposure are mostly located between large grains along the principal slip surface (Fig. 8C and D).  
275 The gaps between large grains are filled with an amorphous material that increases the overall  
276 smoothness of the fault surface. Inside the amorphous material, most nanograins have elongated  
277 shapes, with no dislocations structure and nano-clay minerals often wrap around the grains (Fig.  
278 8D).



279 *Figure 6: TEM cross sections illustrating the relationship between hanging-wall (hw) breccia and footwall (fw)*  
 280 *contact for Arkitsa (A – B) and Schinos (C – D). **A:** TEM bright field close-up from Fig. 5A. Non-crystalline carbon*  
 281 *establishes the contact between hanging wall and footwall. The footwall consists of large, stressed calcite crystals*  
 282 *with bent fractures. Nanogranular layer with non-crystalline carbon and calcite nanograins. White circle indicates*  
 283 *SAED location; black circle depicts EELS measurement. Inset: SAED showing two sets of crystal orientation. ss =*  
 284 *slip surface. **B:** TEM bright field close-up from A. Non-crystalline carbon connects hanging wall and footwall. The*  
 285 *nanogranular layer produces Moiré fringes indicative of overlapping crystal lattices. Nanograins are not always in*  
 286 *grain-to-grain contact. **C:** TEM bright field close-up image from Fig. 5B. The grain size of the hanging wall is 200*  
 287 *nm and larger. The grain size of the footwall directly at slip surface is about 100 nm. **D:** TEM bright field close-up*  
 288 *image from C. The hanging wall and footwall are connected via non-crystalline carbon. Black circle depicts EELS*  
 289 *measurement. Inset shows EELS spectra from black circles in A and D.*

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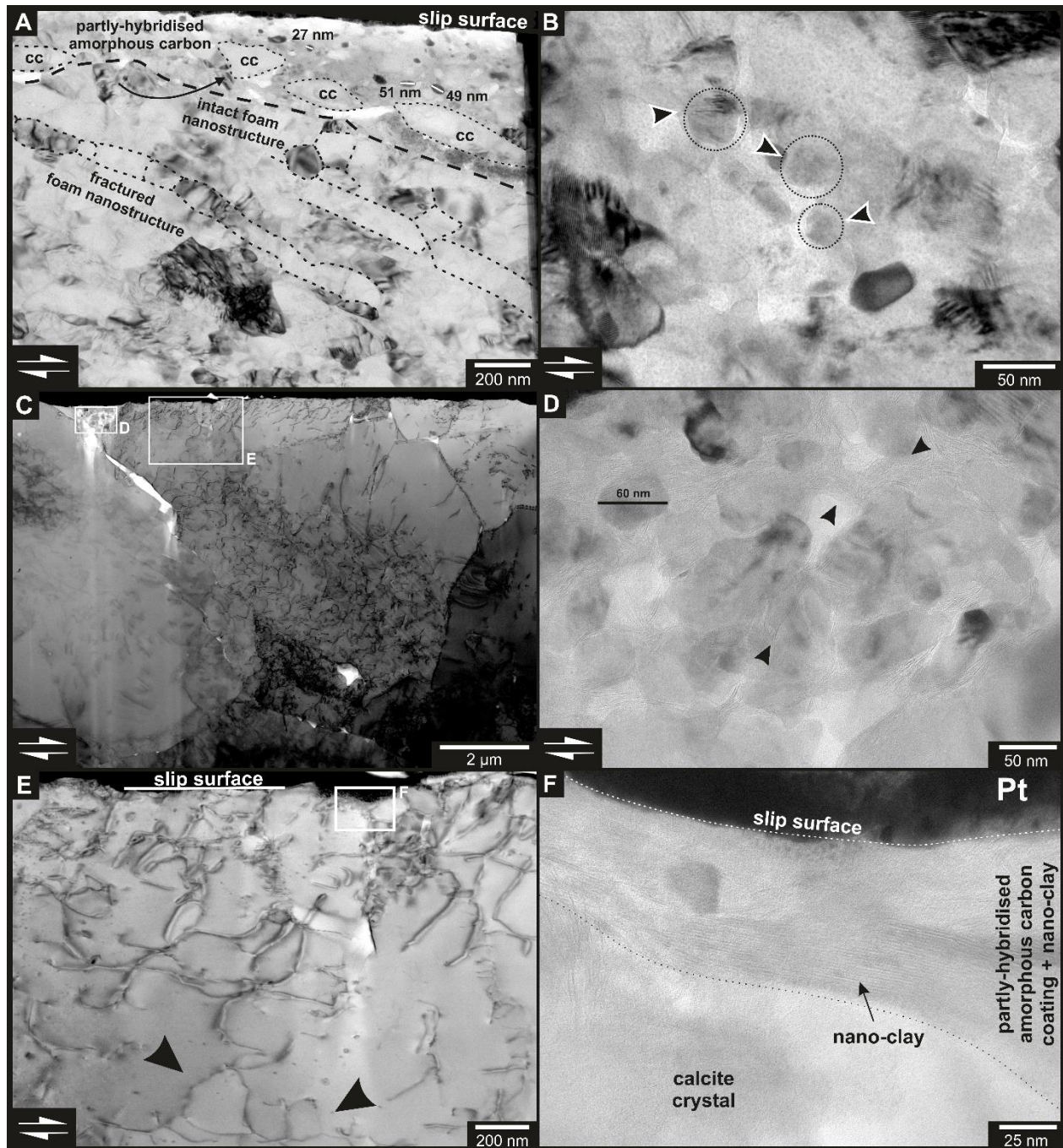


291 *Figure 7: TEM-EDX maps of calcite nanograin and non-crystalline carbon from Fig. 5A (Arkitsa) and Fig. 8C*  
 292 *(Schinos) visualizing the element distribution. Overlap of calcium (Ca) and carbon (C) distribution confirms that the*  
 293 *grains inside the non-crystalline carbon are calcite. Aluminium (Al), iron (Fe) and silicon (Si) are concentrated in*  
 294 *the interstitial region around the calcite grains.*

295 Further TEM investigation of the slip surfaces reveals a variety of slip-derived  
 296 deformation nanostructures. Neither fault exposure exhibits a gradual decrease in grain size  
 297 towards the slip surface but show an abrupt change, instead. Adjacent to the principal slip  
 298 surface, we observe twinned calcite crystals about 2–5  $\mu\text{m}$  in size (Fig. 5A and B). Dislocations  
 299 are homogeneously distributed in the larger crystals from the Arkitsa exposure (Fig. 5A). For the  
 300 Schinos case, the dislocation distribution appears to be less ordered, with dislocations



301 concentrated towards grain contacts in the slip direction (Fig. 8C). Bent cleavage fractures dip  
302 down from the Arkitsa slip surface and exhibit minor displacements together with slight increases  
303 in dislocation density along the fracture planes (Fig. 5A and 6A). Along the slip surface, the  
304 nanostructure of the Arkitsa exposure can vary in intensity and develop a foam nanostructure  
305 (Fig. 8A). The layer thickness of the foam nanostructure is about 1  $\mu\text{m}$  and the grain size ranges  
306 from 100–200 nm with the grain boundaries exhibiting triple junctions at  $\sim 120^\circ$  angles (Fig. 8A).  
307 This nanostructure is not observed in the samples analysed from the Schinos exposure, which  
308 instead contain dislocation arrays (Fig. 8C and E). In general, the foam nanostructure at Arkitsa  
309 appears to be less developed adjacent to the slip surface, especially where the fracture density is  
310 highest.



311 *Figure 8: Nanostructures of fault surface coating and nanograin abundance not associated with the principal slip*  
 312 *surface. **A:** Bright-field TEM image from the Arkitsa exposure. Amorphous slip surface coating with calcite*  
 313 *nanograins and crystal fragments from underlying foam nanostructure (black arrow). A sharp contact separates the*  
 314 *underlying nanostructure from the slip surface coating. Nanostructure with former fracture planes (dashed lines),*  
 315 *which originate from the principal slip surface, and overprint the existing nanostructure. **B:** Calcite nanograins*  
 316 *away from the principal slip surface hosted inside non-crystalline carbon. Tip direction of black-white arrows*  
 317 *indicate lattice fringe orientation of new grains. **C:** Bright-field STEM image overview. Large grains (5 μm) with a*  
 318 *heterogeneous dislocation density distribution. **D:** Calcite nanograins in non-crystalline carbon as a filling between*  
 319 *two grain boundaries directly at the principal slip surface. Clay minerals wrap around the nanograins (black*  
 320 *arrows). **E:** Bright-field STEM close-up from C showing the dislocation structure directly at the principal slip*  
 321 *surface. Black arrows point to dislocation-pinned Frank-Read sources frozen during dislocation bow-out. **F:** Bright-*

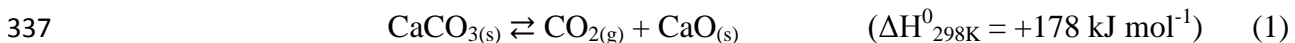
322 *field TEM image from the Schinos exposure. Amorphous slip surface coating mixed with nano-clay on top of calcite*  
323 *crystal. cc = calcite.*

## 324 5 Discussion

### 325 5.1 Processes revealed by nano/microstructures

#### 326 5.1.1 Decarbonation and carbon reduction

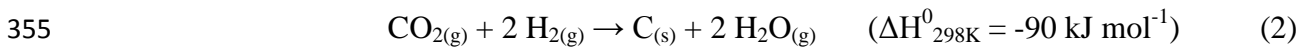
327 Nanostructural investigation of the Arkitsa and Schinos fault exposures reveal the  
328 importance of calcite decarbonation products for the formation of a smooth fault surface. TEM  
329 images reveal a coating of an amorphous solid that we termed “coating phase” or “amorphous  
330 material” above. We suggest that the amorphous solid formed following calcite decarbonation, as  
331 a result of shear heating at asperity contacts or the 'severe' introduction of dislocations during  
332 fault slip. The dislocation production takes place via (re-)activation of Frank-Read sources, which  
333 can multiply dislocations and emit dislocation loops into the crystal after dislocation bow-out.  
334 The Schinos fault exposure preserves an example of two activated Frank-Read sources close to  
335 the principal slip surface (Fig. 8E). In our natural samples, holes piercing the principal slip  
336 surface (Fig. 3A and D) indicate CO<sub>2</sub> degassing as a product of the decarbonation reaction



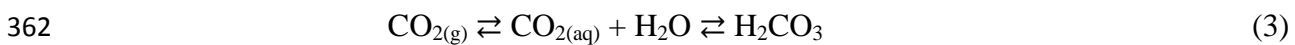
338 reported to occur due to thermal decomposition of calcite starting at 600 °C (Rodriguez-Navarro  
339 et al., 2009). However, experiments by Martinelli and Plescia (2004) recorded CO<sub>2</sub> emission  
340 during ball milling experiments of calcite without increasing the macroscopic temperature. In  
341 addition to the CO<sub>2</sub> emission, the authors inferred the presence of amorphous lime (CaO) in their  
342 milling product as a result of a process termed ‘mechanical liming’. Subsequent heating of the  
343 milling product resulted in a decrease of amorphous lime and the formation of calcite. The  
344 experiments of Martinelli and Plescia (2004) demonstrate that reaction (1) can also take place just  
345 by deforming calcite without reaching macroscopic temperatures of > 600 °C. Based on these  
346 experiments, we suggest that the lime produced during coseismic decarbonation may be partly

347 amorphous. The experiments also suggest that decarbonation can occur at macroscopic  
 348 temperatures lower than 600 °C, although the temperature at asperity contacts could be higher.  
 349 We suggest that 'mechanical liming' is important for wet deformation conditions where fluids are  
 350 present and will buffer the temperature increase.

351 The acquired electron energy loss and Raman spectra indicate the presence of a carbon  
 352 phase. Hence, the presence of this phase implies the occurrence of carbon-reducing processes  
 353 because reaction (1) does not produce elemental carbon. Precipitation of carbon in deformation  
 354 experiments of Oohashi et al. (2014) under a hydrogen atmosphere suggests that the reaction



356 can reduce CO<sub>2</sub> to form carbon. Similarly, Spagnuolo et al. (2015) detected H<sub>2(g)</sub> and CO<sub>2(g)</sub>  
 357 degassing during their experiments. Although reaction (2) is expressed as a chemical vapour  
 358 deposition reaction some of the hydrogen produced can go into solution to reduce dissolved CO<sub>2</sub>.  
 359 This process is known to occur in hydrothermal systems where dissolved hydrogen can interact  
 360 with CO<sub>2(aq)</sub> to form reduced carbon phases (e.g., (Milesi et al., 2015)). Under wet conditions, the  
 361 CO<sub>2</sub> from reaction (1) goes into solution starting the hydrolysis



365 to give carbonic acid, hydrogencarbonate, carbonate ions, and protons (Ruiz-Agudo et al., 2013).

366 The mole fraction solubility (X<sub>1</sub>) of CO<sub>2</sub> in water at room temperature is, with X<sub>1</sub> = 6.15 x 10<sup>-4</sup>,  
 367 one order of magnitude larger than H<sub>2</sub> with X<sub>1</sub> = 1.411 x 10<sup>-5</sup> (Gevantman, 2000). It is, therefore,  
 368 more likely that a larger amount of CO<sub>2</sub> will dissolve in water and form carbonic acid, being the  
 369 more probable proton donor, for reducing carbon from the residual CO<sub>2</sub>. As an alternative, CaO

370 surfaces can act as potential catalytic sites and promote the reaction with H<sub>2</sub>O at kinks, corners  
371 and/or steps. These surfaces can split H<sub>2</sub>O leading to one hydroxyl ion (OH<sup>-</sup>) and one proton (H<sup>+</sup>)  
372 which in turn may also be able to reduce carbon from CO<sub>2</sub> (Kudłacz and Rodriguez-Navarro,  
373 2014). The main requirement for this reaction to occur is the presence of crystalline CaO. This  
374 would either require prior crystallisation of the suspected amorphous lime from reaction (1) or  
375 the lime resulting from reaction (1) was already crystalline. Calcium ions (Ca<sup>2+</sup>) decrease the pH  
376 of the water facilitating CO<sub>2</sub> solubility. Earth-alkali metal reactions with water are commonly  
377 known to produce hydrogen by reaction of the metal with water to form H<sub>2</sub> and the corresponding  
378 metal hydroxide. As the reduction of CO<sub>2</sub> by hydrogen does not require the participation of an  
379 additional reducing agent, it is likely to occur also under closed-system experimental conditions.  
380 With a variety of possible production reactions, it remains challenging to identify the main  
381 reaction responsible for the source of hydrogen for CO<sub>2</sub> reduction. However, a combination of  
382 several processes is likely, with the simplest one being the hydrolysis of CO<sub>2</sub> in H<sub>2</sub>O. The CaO or  
383 Ca<sup>2+</sup> interaction with water can increase the pH up to 12.4, as shown in experiments by Ruiz-  
384 Agudo et al. (2013), which increases the solubility of CO<sub>2</sub> in water and would facilitate reaction  
385 (3). In general, the result would be the precipitation of carbon from the fluid, which would not be  
386 restricted to the principal slip surface.

### 387 5.1.2 Carbon-bond sp<sup>2</sup> hybridisation

388 The obtained  $\frac{I(D)}{I(G)}$  ratios from Raman spectroscopy indicate that the slip surfaces are  
389 coated with nanocrystalline graphite according to the interpretation of the Raman modes by  
390 Ferrari and Robertson (2000). With the obtained  $\frac{I(D)}{I(G)}$  ratios we can calculate the size of the  
391 nanocrystallites according to the equation from (Pimenta et al., 2007)

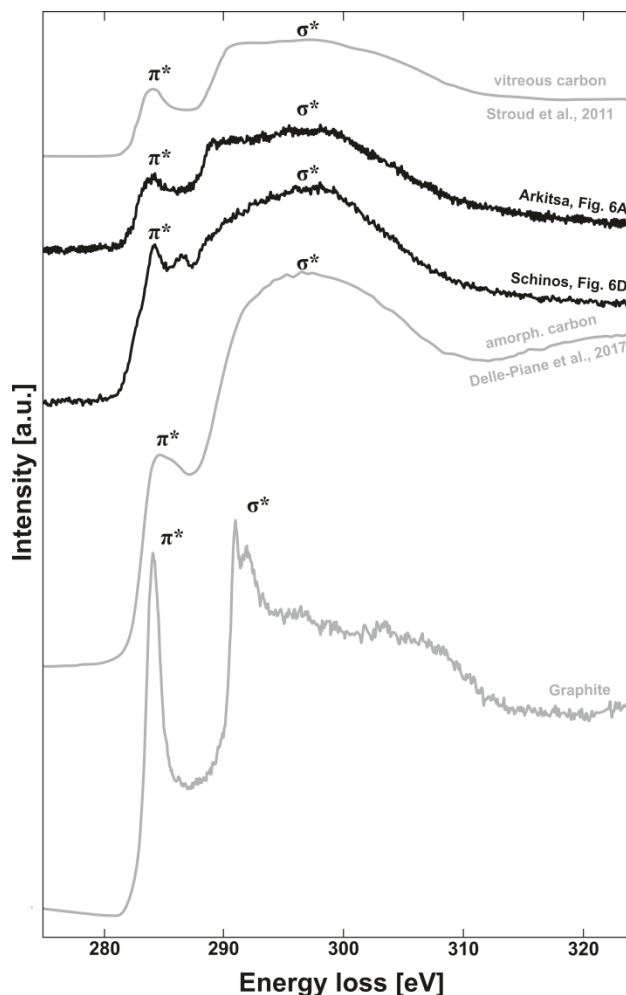
$$L_a \text{ (nm)} = \frac{560}{E_{laser}^4} \left( \frac{I(D)}{I(G)} \right)^{-1} \quad (4),$$

392 where  $L_a$  is the cluster/crystallite size in nm,  $E_{laser}$  is the energy (wavelength) of the laser with  
 393 2.33 eV (532 nm), and  $\frac{I(D)}{I(G)}$  the intensity ratio of D and G peaks. Equation (4) is based on results  
 394 from Pimenta et al. (2007), which revealed an inversely proportional relationship between the  
 395 fourth power of  $E_{laser}$  and the  $\frac{I(D)}{I(G)}$  ratio. For the Arkitsa exposure we calculate a size of  $L_a = 35$   
 396 nm ( $\frac{I(D)}{I(G)} = 0.543$ ) and for the Schinos exposure  $L_a = 24$  nm ( $\frac{I(D)}{I(G)} = 0.789$ ). While Pimenta et al.  
 397 (2007) calculate with  $L_a$  the graphite nanocrystallite size, Ferrari and Robertson (2000) mention  
 398  $L_a$  as the cluster diameter of carbon aromatic rings. The terminology used by Ferrari and  
 399 Robertson (2000) for nanocrystalline graphite and amorphous carbon is based on the starting  
 400 material experiencing disorder. When disordering graphite,  $\frac{I(D)}{I(G)}$  will increase with increasing  
 401 disorder and development of a D peak. With amorphous carbon as the starting material, the  
 402 development of a D peak and the increase in  $\frac{I(D)}{I(G)}$  indicates clustering. Because the D peak arises  
 403 from aromatic rings, for small  $L_a$  the D-mode strength is proportional to the cluster area or  
 404 diameter Ferrari and Robertson (2000).

406 According to our discussion of possible options for carbon reduction in Section 5.1.1,  
 407 where carbon is most likely produced as the outcome of post-seismic, chemical reduction  
 408 processes, the initially precipitated carbon may be amorphous. Nathan et al. (1974) reported for  
 409 glass-like carbon a D-peak position at  $1340 \text{ cm}^{-1}$  and a G-peak position at  $1590 \text{ cm}^{-1}$ , while Wang  
 410 et al. (1990) reported a D-peak position at  $1347 \text{ cm}^{-1}$  and a G-peak position at  $1588 \text{ cm}^{-1}$ . The  
 411 band positions from Wang et al. (1990) were obtained with a 515 nm laser on glass-like carbon,  
 412 which was heat treated at  $2000 \text{ }^\circ\text{C}$ . Wang et al. (1990) also noted that the D-peak position is

413 sensitive to the wavelength of the laser used, exhibiting a band-position shift towards lower  
414 wavenumbers with increasing laser wavelength. The reported values from Nathan et al. (1974)  
415 and Wang et al. (1990) are in agreement with our measured band positions for Arkitsa (D peak:  
416  $1339\text{ cm}^{-1}$ , G peak:  $1584\text{ cm}^{-1}$ ) and Schinos (D peak:  $1344\text{ cm}^{-1}$ , G peak:  $1587\text{ cm}^{-1}$ ), suggesting  
417 that the slip surface coating resembles glass-like carbon. Therefore, instead of the development of  
418 nanocrystalline graphite, our results are consistent with the presence of a partly-ordered carbon  
419 exhibiting an electron-bond environment akin to glass-like carbon. This may be evidence for the  
420 beginning of a clustering process of the aromatic carbon rings ( $\text{sp}^2$ -clustering), which is likely to  
421 have occurred during post-seismic annealing.

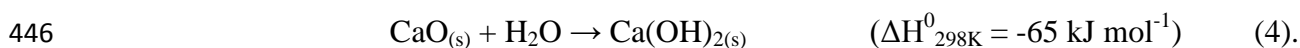
422 High-resolution TEM imaging coupled with electron diffraction suggests that the surface  
423 coating is non-crystalline. This interpretation is supported by our EELS analysis which results in  
424 spectra resembling amorphous carbon (Fig. 9). However, although both have a  $\pi^*$  peak, the phase  
425 identified here has a more pronounced  $\sigma^*$  region indicating a stronger graphitic order, yet lacking  
426 full long-range order as observed in crystalline graphite (Rosenberg et al., 1986). Our spectra  
427 resemble those of Stroud et al. (2011), suggested to be indicative of glass-like carbon with  $\text{sp}^2$   
428 hybridisation. Although the EEL spectra resemble glass-like carbon, we will refer to the carbon  
429 phase from the fault exposures as partly-hybridised amorphous carbon (PHAC). Impurities of Al,  
430 Fe, and Si (Fig. 7) in the PHAC suggest either the concomitant decomposition of silicate phases  
431 (e.g., clays) during slip or the influence of hydrothermal fluids percolating along the fault zone,  
432 precipitating clay. We suggest that the presence of PHAC demonstrates that amorphous carbon is  
433 precipitated during or immediately after slip but undergoes ordering ( $\text{sp}^2$  hybridisation) via post-  
434 seismic annealing, as demonstrated by Raman spectroscopy and EELS.



435  
 436 *Figure 9: Comparison of electron energy loss spectra. Calcite decomposition products presented here have a small*  
 437 *intensity difference of  $\pi^*$  to  $\sigma^*$  peak. The EEL spectrum from vitreous (glass-like) carbon (Stroud et al., 2011)*  
 438 *exhibits a similar  $\sigma^*$ -peak shape compared to the acquired spectra of this study. Reference spectrum of graphite has*  
 439 *three weak  $\sigma^*$  peak features which are absent in all other carbon spectra.*

### 440 5.1.3 Nanograin nucleation

441 Our nanostructural investigations reveal a close spatial relationship between PHAC and  
 442 newly formed calcite nanograins (e.g., Fig. 6A). Nucleation of nanograins inside the  
 443 decomposition product implies a similar chemistry of reactant and product. As discussed in  
 444 Section 5.1.1, reaction (1) also produces lime, which is highly reactive and can form portlandite  
 445 ( $\text{Ca}(\text{OH})_2$ ) under hydrous conditions via the hydration reaction



447 In the presence of  $\text{CO}_2$ , portlandite back-reacts to calcite and water by the carbonation reaction





449 Kudłacz and Rodriguez-Navarro (2014) observed a nucleation-related crystallographic preferred  
450 orientation in two sets of portlandite crystals. The angle between the  $\{11-20\}_{\text{portlandite}}$  planes was  
451  $75^\circ$ , implying that portlandite developed with a topotactic relationship of the  $\{11-20\}_{\text{portlandite}}$   
452 planes parallel to the  $\{110\}_{\text{CaO}}$  planes of lime. The SAED inset of Fig. 6A indicates two preferred  
453 orientations with an angle of about  $75^\circ$  (centre to centre) between the  $\{11-20\}$  planes of calcite in  
454 our natural samples, supporting our described back-reaction steps from lime, through portlandite,  
455 to calcite. In addition, we also observe dislocation-free nanograins away from the slip surface  
456 (Fig. 8B). Hence, we suggest that calcite nanograins are not the outcome of comminution or  
457 plastic grain-size reduction during slip but originate from pseudomorphic growth of calcite after  
458 portlandite and portlandite after lime from decarbonation.

## 459 5.2 Fault surface evolution during the seismic cycle

460 The onset of fault slip triggers a variety of crystal-plastic deformation processes. Cleavage  
461 fractures in Figure 3E demonstrate that slip localises within the first 20  $\mu\text{m}$  inside the fault rock.  
462 This shows that the transfer of strain from the slip surface is low because the cleavage fractures  
463 are concentrated close to the slip surface. Fractures also cross-cut the foam nanostructure and  
464 produce a nanofabric where grains are trapped between sheet-like fracture planes (Fig 8A). Shear  
465 heating may cause a considerable temperature rise at asperity contacts. The temperature increase  
466 or grain fracturing trigger the decarbonation reaction (Eq. 1), which leads to  $\text{CO}_2$  degassing and  
467 the production of lime. It is not fully resolved whether the lime would be amorphous, as reported  
468 by Martinelli and Plescia (2004), or crystalline.

469 During and immediately after fault slip, calcium oxide infiltrates cracks and cavities and  
470 reacts with the fluids present to crystalline portlandite (Eq. 4). The resulting portlandite crystals  
471 back-react with the released  $\text{CO}_2$  to form calcite nanograins (Eq. 5). During the pseudomorphic  
472 replacement of portlandite by calcite, the calcite nanograins preserve the crystallographic

473 nucleation relationship from portlandite. Nano-clay precipitates inside the reaction medium either  
474 from the fluid or from previously decomposed clay minerals and may contribute to the fault-  
475 mirror appearance. Reduction of carbon can occur by reducing CO<sub>2</sub> with dissolved hydrogen (Eq.  
476 2) either from the hydrolysis reaction (Eq. 3) or by water splitting with lime as a catalyst.  
477 Subsequent sp<sup>2</sup>-hybridisation of the amorphous carbon results in aromatic ring clustering and  
478 hence in a stronger near-field bonding of the carbon. As a result, PHAC is formed.

479 Graphitisation of carbonaceous materials in active faults may be an important factor for  
480 fault reactivation because graphite is characterised by a low friction coefficient of  $\mu_{ss} = 0.08$   
481 (Oohashi et al., 2011). Glass-like carbon is a stable, non-graphitising type of carbon which  
482 requires an activation energy of +215 kJ mol<sup>-1</sup> to initiate graphitisation (Saxena and Bragg, 1978;  
483 Hokao et al., 2000). This would imply that on natural faults graphite can only be found after a  
484 considerable energy input to induce graphitisation of glass-like carbon species. Saxena and Bragg  
485 (1978) identified three stages for graphitisation of glass-like carbon: (1) < 1500 °C  
486 dehydrogenation, (2) 1500 °C ≤ 2300 °C stress relief of the structure and (3) > 2300 °C onset of  
487 graphitisation. A high temperature treatment of synthetic glass-like carbon can cause  
488 graphitisation suggesting a thermodynamic control of the graphitisation process. Molecular  
489 dynamics simulations of sp<sup>2</sup> carbon-bond self-assembly illustrate that sp<sup>2</sup>-clustering from  
490 amorphous precursors can commence after 200 picoseconds at ~3200 °C (Powles et al., 2009).  
491 This short time span suggests that the sp<sup>2</sup>-clustering we observe here takes place immediately  
492 after carbon reduction. In addition, the simulations suggest that once sp<sup>2</sup>-bonded stable clusters  
493 form, breaking these bonds to further rearrange and increase sp<sup>2</sup>-ordering may require the need to  
494 overcome an additional energy barrier. This could explain why the PHAC observed here never  
495 reached a higher degree of graphitisation because the energy input was not enough to break the  
496 initial stable carbon clusters. Also, synthetic glass-like carbon displays the often described,

497 entangled, lath-like structure (e.g., (Saxena and Bragg, 1978)). Although the PHAC observed  
498 here lacks the aforementioned nanostructure, it does have the vibrational and electronic properties  
499 of glass-like carbon (Fig. 4 C and D, Fig. 9). Because PHAC is formed by chemical precipitation  
500 and hybridisation, in contrast to pyrolysis synthesis of glass-like carbon (Powles et al., 2009), we  
501 propose that the way the carbon is formed influences the structure of the material while  
502 maintaining a similar interatomic bond structure. This suggestion is supported by Powles et al.  
503 (2009), who point out that the properties of glass-like carbon depend on the formation conditions,  
504 the precursor material, and on the annealing conditions. It is not clear whether the consideration  
505 of geological time scales for the annealing of PHAC at lower temperatures may yield a similar  
506 result in graphitisation.

### 507 5.3 Advances in understanding fault surface evolution

508 Amorphous carbon has been observed in deformation experiments (Verberne et al., 2014;  
509 Spagnuolo et al., 2015; Delle Piane et al., 2017) and its mechanical role has been assessed by  
510 Oohashi et al. (2011). The authors measured a friction coefficient for amorphous carbon of  $\mu =$   
511 0.54 at the initiation of slip and a low steady-state friction coefficient of  $\mu_{ss} = 0.15$  at seismic slip  
512 velocities. Di Toro et al. (2011) speculated about fault lubrication by carbonate decomposition  
513 products but, for the first time, we can demonstrate the extent of amorphous coatings.  
514 Decarbonation is the predominant process for producing a thin fault surface coating and is mainly  
515 responsible for the low surface roughness by smoothing out surface corrugations. Because the  
516 fault surfaces show little variation in the field, our results suggest that the fault planes of both  
517 fault exposures are at present covered with the decarbonation and carbon hybridisation products.

518 Our EEL and Raman spectroscopy results indicate that PHAC has a similar carbon-  
519 bonding environment compared to glass-like carbon. Synthetic glass-like carbon has an entangled  
520 lath-like structure consisting of ordered carbon sheets while the natural PHAC has no

521 nanostructure. The mechanical properties of glass-like carbon, heat treated at 1000 °C, show a  
522 Youngs modulus of  $E = 29$  GPa and a shear modulus of  $G = 12.5$  GPa (Robertson, 1991). Hokao  
523 et al. (2000) report a friction coefficient for glass-like carbon of  $\mu = 0.21$  and  $\mu = 0.13$  for  
524 mixtures of glass-like carbon with graphite. However, as the structure of the PHAC is different  
525 from synthetic glass-like carbon we suggest that mechanical properties of the synthetic material  
526 may not fully reflect the mechanical behaviour of the carbon phase reported here. Because little is  
527 known about the deformation of various crystalline and non-crystalline carbon phases further  
528 experimental investigations are warranted.

529         While previous research focused mainly on grain size reduction mechanisms (Sammis and  
530 Ben-Zion, 2008; Siman-Tov et al., 2013), our results highlight the importance of decarbonation  
531 and back-reactions as part of the overall fault deformation and healing behaviour. In contrast to  
532 (Siman-Tov et al., 2013) we do not identify small grains evolving from fracturing long beams of  
533 calcite. However, our results indicate that, for natural carbonate faults, a nanogranular coating is  
534 not a key feature for producing a fault mirror surface. For example, the Schinos fault exposure  
535 illustrates that MMSs can contain no coating of calcite nanograins (Fig. 7C, E and F). Instead,  
536 MSS are produced by the interplay of grain-size reduction, decarbonation, back-reactions, and  
537 annealing. The present study does not exclude crystal-plastic deformation as an important grain-  
538 size reduction mechanism in fault gouges but demonstrates that it is not the main mechanism to  
539 form a fault mirror surface. The formation of an amorphous phase with low viscosity during  
540 deformation could explain the low friction, stable, near steady state behaviour after the onset of  
541 weakening in Han et al. (2007) as well as the increase in friction coefficient (Spagnuolo et al.,  
542 2015) via re-carbonation and carbon hybridisation after displacement stopped. The proposed  
543 chemical reactions demonstrate that a succession of healing reactions take place after fault slip  
544 ceases.

545 **6 Conclusions**

546 We conclude that decarbonation of calcite and the subsequent reaction of the  
547 decarbonation products produces fault mirror surfaces. The decarbonation process itself may be a  
548 major coseismic fault weakening factor and fault slip is facilitated on a decarbonation-product  
549 glide film. Frictional behaviour during slip may be dictated by lubrication of low viscosity  
550 (amorphous) calcium oxide and (amorphous) carbon. Occasional precipitation of clay inside the  
551 decarbonation products. Post-seismic hybridisation produces partly-hybridised amorphous carbon  
552 and connects footwall with hanging wall. Ultimately, post-seismic carbonation of portlandite  
553 produces new nanograin calcite crystals by pseudomorphic replacement and illustrate the fault  
554 healing.

555

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