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Evaluating Turbulence Parameterizations at Gray Zone Resolutions for the Ocean Surface Boundary Layer

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7	Key Points:
8	• Ocean simulations with partially resolved boundary layer turbulence exhibit resolution-
9	dependent sensitivity to choice of turbulence closure
10	• In the gray zone resolutions, k - ϵ can accurately represent mean-state properties
11	but excessively damp small-scale turbulence
12	• Using Smagorinsky or the implicit method in the gray zone allows more active tur-
13	bulence despite sacrificing mean state fidelity

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14 Abstract

Turbulent mixing in ocean boundary layers is often fully parameterized as a subgrid-15 scale process in realistic ocean simulations. However, recent submesoscale modeling stud-16 ies have advanced to a horizontal grid spacing of $\mathcal{O}(10 \text{ m})$ that is comparable to, or even 17 smaller than, the typical depth of the turbulent surface boundary layer. Meanwhile, ef-18 forts toward realistic large-eddy simulations (LES) nested within regional models require 19 subdomains with similar grid spacings, where turbulent eddies are partially resolved in 20 the mixed layer. The range of intermediate gird resolution, often known as the "gray zone", 21 22 presents challenges for model configuration and analysis, including uncertainties regarding the behavior of common turbulence closures outside of their ideal use cases. In this 23 study, we evaluate three common configurations for subgrid turbulence—k- ϵ , Smagorin-24 sky, and an implicit no-closure method—in the gray zone resolutions for the ocean sur-25 face mixed layer. Results indicate the $k - \epsilon$ closure shows less sensitivity to grid spacing, 26 producing accurate mean mixed-layer profiles even with partially resolved turbulence. 27 However, it overly damps turbulent motions, significantly reducing small-scale variabil-28 ity that could otherwise be captured. The Smagorinsky closure and the implicit method, 29 in contrast, exhibit higher sensitivity to grid spacing, initially performing poorly but con-30 verging toward baseline solutions at finer grids. Our findings provide guidance for sub-31 mesoscale and turbulent-scale modeling, recommending Smagorinsky or implicit meth-32 ods for nesting scenarios at turbulence-permitting resolutions. The k- ϵ closure is more 33 suitable for high-resolution models primarily focused on accurate mean-state represen-34 tations rather than explicitly resolving detailed three-dimensional turbulence. 35

³⁶ Plain Language Summary

Turbulence is an important small-scale process that mixes heat, momentum, and 37 nutrients near the ocean surface and bottom. Typically, ocean models cannot fully re-38 solve turbulent motions due to limited grid resolution and instead rely on parameter-39 izations to approximate the effects of turbulent mixing. However, as models improve and 40 begin to partially resolve turbulence at finer resolutions, choosing appropriate param-41 eterizations becomes increasingly important. Here, we evaluate three common model con-42 figurations—the k- ϵ closure, the Smagorinsky closure, and an implicit no-closure method—across 43 a range of intermediate grid spacings known as the "gray zone". The k- ϵ closure accu-44 rately represents boundary mixing regardless of resolution but suppresses smaller-scale 45 turbulence. In contrast, the Smagorinsky closure and the implicit methods are sensitive 46 to resolution, performing poorly at coarse grids but substantially improving at higher 47 resolutions. Our findings suggest using Smagorinsky or implicit methods for simulations 48 aimed at explicitly resolving turbulent motions (e.g., nested large eddy simulations), while 49 the k- ϵ closure is better suited for scenarios prioritizing accurate mean conditions over 50 detailed turbulence features. 51

52 1 Introduction

Realistic regional ocean simulations now regularly resolve submesoscale eddies at 53 a grid spacing of $\mathcal{O}(1 \text{ km})$ (Gula et al., 2021; Taylor & Thompson, 2023). Through mul-54 tiple nesting steps, many studies have also successfully further refined the model grid spac-55 ing to \mathcal{O} (10 m) to investigate mixed-layer processes, such as Langmuir circulation (Hypolite 56 et al., 2022), symmetric instability (Dong et al., 2022), and submesoscale shelf current 57 (Dauhajre et al., 2019). Relative to the typical value of mixed layer depth at $\mathcal{O}(100 \text{ m})$, 58 realistic simulations at the horizontal resolution of $\mathcal{O}(10 \text{ m})$ can represent, albeit coarsely, 59 large turbulent eddies in the mixed layer. However, despite being turbulence-permitting, 60 these simulations generally employ vertical mixing parameterizations from the Reynolds-61 averaged Navier-Stokes equations (RANS) framework, for example, K-profile parame-62 terization (KPP, Large et al., 1994) and k- ϵ (Jones & Launder, 1972), which are designed 63

to model boundary turbulence at all scales. This leads to concerns about double-counting 64 of turbulence from directly resolved eddies and subgrid processes, both in terms of en-65 ergetics and effects on dynamics. Likewise, the premature use of LES closures (e.g., Smagorin-66 sky) at these intermediate scales is also problematic without representing the energy cas-67 cade in the inertial subrange. This dilemma posed by the lack of scale separation is com-68 monly known as the "gray zone" problem, first introduced by the atmospheric sciences 69 literature (Wyngaard, 2004). In this study, we aim to better understand the sensitiv-70 ity of common turbulence closures in the gray zone resolutions for the ocean surface bound-71 ary layer, with the goal to inform strategies on closure selection that better fit model-72 ing objectives. 73

Representing turbulent flows at the gray zone resolution has been a subject of ex-74 tensive research for atmospheric modeling (F. Chow et al., 2019; Honnert et al., 2020). 75 For example, Zhou et al. (2014) performed simulations for the atmospheric convective 76 boundary layer and found that the gray zone resolution resulted in overly large convec-77 tion cells and a delayed onset of turbulence. Mirocha et al. (2013) and Goodfriend et al. 78 (2015) focused on nested simulations and found that the choice of subgrid closure affected 79 the turbulence flow transition at the coarse-fine domain interface. Treatments at the lat-80 eral boundaries, such as increasing forcing frequency (Brisson et al., 2016), adding syn-81 thetic perturbations (Mazzaro et al., 2017), and optimizing grid aspect ratios (Daniels 82 et al., 2016), can help improve turbulence statistics in the inner domain. The gray zone 83 problem has also led to the active development of many scale-aware turbulence closures 84 for atmospheric modeling (F. K. Chow et al., 2005; Bhattacharya & Stevens, 2016; Kurowski 85 & Teixeira, 2018), which enable the transition from bulk closures to three-dimensional 86 turbulence for the planetary boundary layer. While scale-aware parameterizations (built 87 specifically to span the gray zone) are desirable, these have not yet been studied in depth 88 for the ocean and are not available in most common ocean models (e.g., ROMS, CROCO, 89 and MITgcm). 90

For ocean modeling, the gray zone problem has already been a practical concern 91 and will become more prevalent in the future. Our study thus aims to evaluate the per-92 formance of commonly used turbulence closures (k- ϵ , Smagorinsky, and the implicit method) 93 in the gray zone resolutions that allow the partial representation of mixed layer eddies. 94 Here, we orient our analysis to two common modeling scenarios that may face the gray 95 zone problem. First, for stand-alone submesoscale-resolving simulations—where the ob-96 jective is to highly resolve processes at $\mathcal{O}(100 \text{ m})$ —we focus on the representation of av-97 eraged mixed layer profiles and the evolution of mixed layer depth. Second, for nested 98 LES—where the objective is the best possible representation of turbulence statistics as the gray zone is traversed with telescoping resolution grids—we evaluate the represen-100 tation of turbulent coherent structures, fluxes, and kinetic energy. In the following sec-101 tions, we first define the problem setup and numerical simulations in Section 2. Section 102 3 describes the results and makes recommendations based on the modeling objectives. 103 Finally, results are summarized in Section 4, along with a discussion of the limitations 104 and future research directions. 105

106 2 Methods

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2.1 Model Description

The Coastal and Regional Ocean Community model, CROCO Version 1.3.0 (Auclair et al., 2022), is used to simulate turbulence in the upper ocean boundary layer. In this study, we broadly define the gray zone as a range of horizontal grid spacings extending from the scale of mixed layer depth to well-resolved LES, for which turbulent eddies are allowed to appear (coarsely resolved) in the mixed layer (further details on grid spacing can be found in Section 2.2). The same non-hydrostatic, non-Boussinesq solver is applied to all simulations with the fifth-order weighted essentially non-oscillatory (WENO5) scheme for both momentum and tracer advection. This setup follows CROCO's documentation on LES applications by Jullien et al. (2022) as well as the best practices according to Pressel et al. (2017), although we note below some results that indicate that a higher-order solver may be preferable for some LES applications. Recently, CROCO's non-hydrostatic solver and the WENO5 scheme have been compared to the US National Center for Atmospheric Research's LES model (NCAR-LES) and demonstrated good accuracy for surface boundary layer turbulence problems (Fan et al., 2023).

Two turbulence closures in CROCO, the k- ϵ model (Jones & Launder, 1972) and 122 123 the Smagorinsky-Lilly model (Lilly, 1962), are evaluated for their performance in the gray zone. While KPP is another common closure for ocean modeling, it underperformed in 124 our preliminary tests relative to two-equations models like $k \cdot \epsilon$ for the gray zone resolu-125 tions with large turbulent eddies in the mixed layer. Therefore, we limit our scope to k-126 ϵ as the representative RANS closure in this study. As an additional point of compar-127 ison, we run simulations without an explicit turbulence closure: an alternative approach 128 relying instead on the dissipative nature of the monotonic advection scheme (WENO5) 129 to provide an implicit model of subgrid turbulence. In the LES framework, this is known 130 as implicit LES (ILES) or monotonically integrated LES (MILES) (Grinstein et al., 2007), 131 and it has shown good accuracy for turbulent flows in a wide range of atmospheric and 132 oceanic contexts (Smolarkiewicz et al., 2007; Pressel et al., 2017; Silvestri et al., 2024). 133 Readers should note that implementations of the closures used here may vary across dif-134 ferent models. It is possible that the closure performance in the gray zone could be im-135 proved by targeted re-tuning. However, such ad hoc efforts are likely to be challenging 136 and may not generalize across cases. Hence, we do not include such efforts here (see Sec-137 tion 4). 138

The $k - \epsilon$ model is a common RANS closure used for ocean modeling, for example, 139 in the basin-scale submesoscale-permitting simulations of the Atlantic Ocean (GIGATL) 140 (Gula et al., 2021) and the idealized simulations of ice dynamics in the surface mixed 141 layer (Herman et al., 2020). In CROCO, $k - \epsilon$ is implemented as part of the Generic Length 142 Scale mixing parameterization (Umlauf & Burchard, 2003), which solves the transport 143 equations for turbulent kinetic energy (TKE, k) and its dissipation rate (ϵ) to compute 144 eddy viscosity, K_{gls} . As a vertical mixing parameterization in CROCO, the shear pro-145 duction term for TKE, $P = K_{gls} \left[(\partial_z u)^2 + (\partial_z v)^2 \right]$ accounts only for the vertical shear 146 of the horizontal velocity rather than the full deformation rate. Meanwhile, the advec-147 tion terms in the total derivatives of k and ϵ include both horizontal and vertical com-148 ponents. The eddy viscosity and diffusivity are computed with the stability function from 149 Canuto et al. (2001). The minimum TKE parameter is set to $10^{-10} \text{ m}^2 \text{s}^{-2}$ while the back-150 ground viscosity and diffusivity are prescribed as a constant $10^{-6} \text{ m}^2 \text{s}^{-1}$. 151

Similar to the wide applicability of $k - \epsilon$, the Smagorinsky model is a common clo-152 sure in the LES framework where comparatively large 3D turbulent eddies are explic-153 itly resolved while smaller-scale motions are modeled. Typically, LES requires the grid 154 configuration to be fine enough to resolve at least 80% of the energy in turbulent mo-155 tions (Pope, 2000). For optimal performance, the filter width should be placed within 156 the inertial subrange to ensure a proper representation of the energy cascade. In CROCO, 157 the Smagorinsky model follows the formulation by Lilly (1962) to calculate eddy viscos-158 ity but separates the horizontal and vertical directions, namely 159

$$K_{smag} = \begin{cases} C_s^2 \Delta_x \Delta_y D & \text{(horizontal)}, \\ C_s^2 \Delta_z^2 D & \text{(vertical)}, \end{cases}$$
(1)

where $D = \sqrt{2S_{ij}S_{ij}}$ and $S_{ij} = 0.5(\partial_{x_j}u_i + \partial_{x_i}u_j)$ is the strain rate. The Smagorinsky coefficient C_s is fixed at the canonical value of 0.16, and the filter width is taken as the horizontal grid spacings (Δ_x and Δ_y) and the vertical grid spacing (Δ_z). Note that this anisotropic treatment of the eddy viscosity differs from traditional implementations of the Smagorinsky model, which assume isotropic eddy viscosity (Chamecki et al., 2019).

Additionally, CROCO applies a buoyancy adjustment such that when a form of the crit-165 ical Richardson number (N^2/D^2) is larger than 0.25 (strongly stratified regions), the eddy 166 viscosity is set to a constant background value of $10^{-6} \text{ m}^2 \text{s}^{-1}$. The turbulent diffusiv-167 ity is set to K_{smaq} with a constant turbulence Prandtl number (Pr) equal to 1. Since 168 LES closures are designed for grid resolutions that permit the resolution of isotropic ed-169 dies, their ideal usage excludes the gray zone when the grid has a high aspect ratio (i.e., 170 large horizontal grid spacing). Nevertheless, we believe it is necessary to evaluate the per-171 formance of the Smagorinsky scheme in the gray zone as part of a nesting-down strat-172 egy to achieve realistic LES, despite this application falling outside its intended and val-173 idated use. 174

175 2.2 Simulation Configuration

This study focuses on canonical turbulence regimes relevant to the ocean surface boundary layer, including idealized simulations forced by constant, spatially homogeneous surface buoyancy flux, wind drag, and surface gravity waves in a doubly periodic domain. We note that spatially inhomogeneous mean flows, such as submesoscale fronts, can modify the parameter space of turbulence production and dissipation (Dong et al., 2024; Zheng et al., 2025); however, these cases are left for future work.

A total of five cases are designed to represent different types of turbulent flow (Table 1). To differentiate the relative contribution to turbulence kinetic energy, we map the surface forcings on $La_t - \Lambda$ parameter space (Figure 1). The turbulent Langmuir number (La_t) quantifies the competition between the wind-driven shear turbulence and the vortex forcing associated with the Stokes-drift velocity (McWilliams et al., 1997) and is defined as

$$La_t = (u_*/u_s)^{1/2}, (2)$$

where $u_* = (\tau/\rho_0)^{1/2}$ is the friction velocity and $u_s = \omega^3 a^2 g^{-1}$ is the surface Stokes drift for deep water waves. The surface forcings for wind drag (τ) , wave frequency (ω) , and wave amplitude (a) are applied only in the x-direction. The reference density (ρ_0) is chosen to be 1010 kg m⁻³ (neglecting salinity) along with a standard value of gravity $(g = 9.81 \text{ m s}^{-2})$.

¹⁹³ The stability parameter (Λ) characterizes the relative impacts of surface wind stress ¹⁹⁴ and buoyancy fluxes and is defined as

$$\Lambda = \kappa w_*^3 / u_*^3,\tag{3}$$

where $\kappa = 0.41$ is the von Kármán constant and $w_* = (B_s|h|)^{1/3}$ is turbulent convective velocity. The surface buoyancy flux, B_s , is related to the heat flux F by

$$B_s = \frac{\alpha g}{\rho_0 c_p} F,\tag{4}$$

where $\alpha = 2 \times 10^{-4} \circ \text{C}^{-1}$ is the thermal expansion coefficient used in the linear equation of state and $c_p = 3985 \text{ J} \text{ (kg }^{\circ}\text{C})^{-1}$ is the specific heat capacity. A boundary layer depth, h, of 30 m (set from the initial condition) is used to calculate w_* , even though the evolution of h varies depending on the surface forcings.

To represent the surface wave effect, the wave-current interaction was activated in CROCO (MRL-WCI) for the wave-averaged equations. We prescribe monochromatic wave forcings ($\omega = 1.01 \text{ rad } s^{-1}$ and a = 0.8 m) to generate a uniform Stokes drift profile aligned with the wind forcing ($\tau_x = 0.037 \text{ N m}^{-2}$), resulting in a wind-wave equilibrium at $La_t = 0.3$ for both the wind-wave and convection-wind-wave cases. The same set of forcings was used in the LES study on Langmuir turbulence by McWilliams et al. (1997) and was revisited with non-hydrostatic CROCO by Herman et al. (2020). The other cases have no wave forcing such that $La_t \to \infty$. For the stability parameter, we choose $\Lambda = 0.61$ for the convection-wind case to reflect the shear-dominated turbulence generation with a weak convection. In the realistic convection-wind-wave case, $\Lambda = 1.76$ falls between the 60% and 90% contours of the joint probability density function reported by Li et al. (2019), indicating surface forcing conditions representative of the real ocean. The convection-only case has no wind forcing for $\Lambda \to \infty$ and the shear-only case has no surface buoyancy flux giving $\Lambda = 0$.

Forcing Case	F (W m ⁻²)	$B_s (m^2 s^{-3})$	w_{*} (m s ⁻¹)	$ au_x$ (N m ⁻²)	u_* (m s ⁻¹)	u_s (m s ⁻¹)	La_t	Λ
Convection-only	-100	4.87×10^{-8}	1.14×10^{-2}	0.00	0.00	0.00	∞	∞
Shear-only	0.00	0.00	0.00	0.10	9.95×10^{-3}	0.00	∞	0.00
Convection-wind	-100	4.87×10^{-8}	1.14×10^{-2}	0.10	$9.95 imes 10^{-3}$	0.00	∞	0.61
Wind-wave	0.00	0.00	0.00	0.037	$6.05 imes 10^{-3}$	6.79×10^{-2}	0.30	0.00
Convection-wind-wave	-65	4.87×10^{-8}	9.83×10^{-3}	0.037	6.05×10^{-3}	6.79×10^{-2}	0.30	1.76

 Table 1.
 Surface forcing conditions for the five Scenarios evaluated in the gray zone resolution using the CROCO model.



Figure 1. Five forcing cases, shown as gray circles in the $La_t - \Lambda$ space (see also Chor et al., 2021). The solid blue, gray, and red lines denote regions where more than 25% of the turbulence kinetic energy is produced by Langmuir forcing, wind stress, or buoyancy fluxes, respectively. The dashed white lines represent the joint probability distribution function of the realistic ocean from Li et al. (2019).

Given that the precise definition of the gray zone range remains an area of ongoing research (Beare, 2014; F. Chow et al., 2019; Honnert et al., 2020), we do not attempt to provide a rigorous framing here. Relative to an initial mixed layer depth of 30 m, we choose horizontal grid spacings to be 4, 12, 24, and 48 m ($\Delta = \Delta_x = \Delta_y$, identical in

both the x and y directions). Our post hoc analysis suggests that these grid spacings ef-219 fectively span the gray zone in our simulations, as indicated by the emergence of turbu-220 lent eddies. To evaluate the gray zone results, each forcing case includes a 1.25 m run 221 with the Smagorinsky closure, serving as a well-resolved baseline for comparison. The 222 computational domain is horizontally periodic, with 256×256 grid points for the 1.25 223 m and 4 m runs and 128×128 for the rest. All runs share the same 100-point vertical 224 grid configuration (using CROCO's grid parameters: $\theta_s = 11.97$, $\theta_b = 0$, hc = 401.69, 225 and h = 131.46), which maintains an approximately constant 1 m spacing above 60 m 226 depth and gradually stretches to 4 m near the bottom. A sponge layer is applied to the 227 bottom 20 grid points, where the velocity field is nudged to zero and the temperature 228 is relaxed toward the initial stratification. The combination of the stretched vertical grid 229 and sponge layer damps internal waves and minimizes their reflection at the bottom, help-230 ing to isolate surface mixing dynamics. 231

The initial condition for all simulations consists of a resting, stratified ocean. A 30 m mixed layer with a uniform temperature of 20° C is positioned on top of a stratified interior with $N^2 = 1.96 \times 10^{-5} \text{ s}^{-2}$ (equivalent to 0.01 °C m⁻¹). To accelerate the transition to turbulence in the mixed layer, small Gaussian noise with a zero mean and a standard deviation of 5×10^{-7} °C is added to the temperature field. All runs are integrated for four inertial periods with a constant Coriolis frequency of 10^{-4} s^{-1} .

The non-hydrostatic solver of CROCO uses a time-splitting method with two user-238 defined time steps for the fast and slow modes. For the baseline run at $\Delta = 1.25$ m, we 239 use a slow mode time step of 0.5 s and a fast mode time step of 0.017 s, ensuring the Courant 240 number remains below 0.68 for the Courant-Friedrichs-Lewy (CFL) condition. For gray 241 zone runs at coarser grid spacings, longer time steps are used (1 s for slow mode; 0.05242 s for fast mode). A wide range of time steps was tested, and these values were selected 243 to balance numerical stability and computation efficiency. To further relax the sound-244 related CFL constraint, we set the speed of sound $C_s = 3 \text{ m s}^{-1}$ and the second viscos-245 ity $\lambda = 10 \text{ kg s}^{-1} \text{ m}^{-1}$ in the fast mode. Despite these non-physical values, Fan et al. 246 (2023) shows that they have minimal impact on turbulence statistics. 247

248 3 Results

To evaluate closure performance in the gray zone, we first describe the instanta-249 neous coherent turbulence structures under different forcing scenarios (Section 3.1). Next, 250 spatio-temporally averaged mixed-layer profiles (i.e., temperature, velocity, and fluxes 251 of temperature and momentum) are analyzed to quantify the impact of grid spacing on 252 turbulent mixing, followed by a discussion on the evolution of mixed layer depth. Finally, 253 we compare kinetic energy spectra to assess the closure effect on effective resolution. In 254 this section, we define the gray zone as the range of horizontal grid spacing (Δ) that al-255 lows turbulent eddies to emerge. The vertical grid is fixed in all simulations. 256

257

3.1 Coherent Turbulence Structure

Analyzing flow structures provides a valuable qualitative metric for evaluating marginally resolved eddies in the gray zone. Accurately representing turbulent structures is crucial for nested simulations, as grid-dependent flow artifacts can propagate from the parent domain and contaminate the final solutions (F. Chow et al., 2019). Here, we present coherent turbulence structures as instantaneous flow fields in a horizontal plane near the surface (z=10 m), four inertial periods after model initialization.

²⁶⁴ Convection-driven turbulence exhibits characteristic cell-like patterns, which are ²⁶⁵ visible in the high-resolution simulations shown in Figure 2. For example, at $\Delta = 4$ m ²⁶⁶ (MLD ~ 14 Δ), Smagorinsky, k- ϵ , and implicit runs all display the classic signature of ²⁶⁷ convection cells with strong downwelling boundaries (blue) and broad centers of weak

upwelling (red). Driven by surface cooling, the cellular patterns are similar to those found 268 in other simulations of the surface ocean mixed layer (Chor et al., 2018; Souza et al., 2020) 269 and the atmospheric boundary layer heated from below (Honnert et al., 2011; Zhou et 270 al., 2014). At larger grid spacings, the cellular pattern persists at $\Delta = 12$ m but becomes 271 barely visible at $\Delta = 24$ m and $\Delta = 48$ m. Among the closures, the $k - \epsilon$ runs produce 272 the weakest vertical velocities, suggesting damping of the resolved turbulence. Meanwhile, 273 in the Smagorinsky runs, the size of the convection cells are resolution-dependent, in-274 creasing with the horizontal grid spacing at a rate faster than seen in the implicit case. 275 This behavior may indicate a limitation of the Smagorinsky model in the gray zone, where 276 it struggles to represent subgrid dissipation and impacts kinetic energy across different 277 scales (see Section 3.2.1 for further discussion on subgrid fluxes). 278



Figure 2. Snapshots of the vertical velocity (w) from the convection-only case at 10 m depth after four inertial periods. Each column is organized to show the results of the same closure at different grid spacings.

The shear-driven turbulence case exhibits diverse fine-scale features including streaks 279 and rolls (Figure 3). Unlike convection, the turbulence patterns are highly sensitive to 280 both closure and grid spacing. In particular, $k - \epsilon$ consistently produces a horizontally ho-281 mogeneous flow field with near-zero vertical velocity regardless of grid spacing. Despite 282 the lack of resolved turbulence, $k - \epsilon$ still drives vertical mixing through subgrid processes, 283 as shown later in the flux profile (Section 3.2.2). For Smagorinsky runs, the turbulence 284 structure changes at different grid spacings. The flow fields display eddies at $\Delta = 4$ m 285 and elongated streaks at $\Delta = 12$ m and 24 m. At $\Delta = 48$ m (comparable to the MLD 286 at 40 m), no turbulence appears. The implicit runs exhibit turbulence patterns across 287 the gray zone, but the scale of eddies depends on grid spacing. As shown later in Sec-288 tion 3.2.2, although both the Smagorinsky and implicit runs allow turbulence to form, 289 they struggle to generate sufficient vertical mixing to deepen the mixed layer at coarse 290 grid spacings ($\Delta = 24$ m and 48 m). 291



Figure 3. Snapshots of the vertical velocity (w) from the shear-only case at 10 m depth after four inertial periods. Each column is organized to show the results of the same closure at different grid spacings. Note that no turbulence appears in the k- ϵ runs and the 48 m Smagorinsky run.

The wind-wave case is designed to generate Langmuir turbulence under strong wind 292 and wave conditions typical of the surface ocean (Belcher et al., 2012). With $La_t = 0.3$, 293 the roll-cell pattern of Langmuir circulations (concentrated horizontal convergence zones 294 and strong downwelling lines) exhibits strong sensitivities to grid spacing and subgrid 295 closures (Figure 4). Unlike the homogeneous flow in the shear-only case, $k - \epsilon$ generates 296 the roll-cell pattern with the additional forcing of Stokes drift. However, compared to 297 Smagorinsky and the implicit method, k- ϵ produces relatively smooth velocity gradients, 298 again indicating excessive damping of resolved turbulence (as discussed in Section 3.4 299 when analyzing flow spectra). The relative width of roll cells scales with grid spacing for 300 both Smagorinsky and the implicit method, increasing from about 100 m at $\Delta = 4$ m 301 to over 400 m at $\Delta = 24$ m. In contrast, k- ϵ shows less sensitivity to grid spacing, con-302 sistently producing roll cells with a characteristic width of about 100 m across all grid 303 spacings. At $\Delta = 48$ m, the Smagorinsky solution does not show turbulent eddies, whereas 304 $k - \epsilon$ and the implicit method still generate weak but coherent downwelling lines. This aligns 305 with previous modeling studies by Hypolite et al. (2021, 2022), which reported a sim-306 ilar sensitivity of roll cell width to model resolution in the realistic simulations of the Cal-307 ifornia Current system. 308

The other forcing cases exhibit similar sensitivities in their coherent structure to closure and grid spacing, with additional figures provided in the supporting information. While this section only offers a qualitative discussion of flow patterns, the representation of coherent turbulence structures has important implications for model nesting in the gray zone. We find that $k \cdot \epsilon$ tends to damp turbulent eddies more than Smagorinsky and the implicit method, with the shear-only case being the most pronounced example. When an LES is nested within a regional simulation, excessively laminar flow at



Figure 4. Snapshots of the vertical velocity (w) from the wind-wave case at 10 m depth after four inertial periods. Each column is organized to show the results of the same closure at different grid spacings. Note that no turbulence appears in the 48 m Smagorinsky run.

the boundary may suppress turbulence development in the inner domain. Compared to $k-\epsilon$ in the gray zone, using Smagorinsky and the implicit method produces different turbulence patterns, which can propagate into the inner domains and influence the final solutions. In the next section on mixed-layer profiles, we move toward the quantitative evaluation of boundary mixing in the gray zone.

321 3.2 Mixed Layer Profiles

This section examines the horizontally-averaged properties of the mixed layer—including temperature, velocity, and stratification profiles—for the convection-only (3.2.1), shearonly (3.2.2), and convection-wind-wave cases (3.2.3). Flux profiles are separated into resolved and subgrid components. All profiles are horizontally averaged at each depth and temporally over the last inertial period.

327 3.2.1 Convection-only

Figure 5 presents the temperature and stratification profiles (N^2) from the convection-328 only case. Each panel displays the gray zone profiles of different grid spacings, compared 329 to the baseline solution ($\Delta = 1.25$ m, Smagorinsky) in red. Qualitatively, the averaged 330 profiles under convection-driven turbulence show little sensitivity to changes in grid spac-331 ing and closure, relative to other forcing cases below. For the implicit and Smagorinsky 332 solutions, the most notable deviations from the baseline solution occur near the surface 333 above a depth of 5 m. In contrast, the k- ϵ solutions exhibit greater differences at the bot-334 tom of the mixed layer, between 45 m and 55 m. A clear convergence of results is ob-335 served as the grid spacing is refined. 336



Figure 5. Mean-state temperature (solid lines) and stratification (dotted lines) profiles from the convection case. Along with the temperature profiles (solid lines) in the top panels, the accompanying N^2 profiles are shown as dotted lines. The 1.25 m Smagorinsky run is used as the baseline of comparison (red lines).

Figure 6 presents the vertical velocity variance and temperature flux profiles in the 337 convection case. Regardless of closure, the velocity variances increase with smaller grid 338 spacings, corresponding to more active turbulent eddies (Section 3.1). The damping ef-339 fect of $k - \epsilon$ on resolved turbulence, previously observed in the flow visualization, is also 340 evident here, with the smallest variances relative to those in the Smagorinsky and im-341 plicit runs. For temperature fluxes, all runs show similar total flux magnitudes, regard-342 less of closures or grid spacing. This suggests that larger turbulent temperature anoma-343 lies at lower resolutions compensate for reduced vertical velocities. 344

For the temperature profiles of $k - \epsilon$ in Figure 6.f, the total fluxes remain consistent 345 despite varying degrees of resolved turbulence at different grid spacings. The relative frac-346 tion of subgrid component $(-\overline{K_s\partial_z T})$, where K_s is eddy diffusivity) compensates for changes 347 in the resolved component (w'T'). Small differences emerge in the entrainment layer (40–60 348 m), where the higher-resolution runs produce slightly enhanced negative temperature 349 flux. Previously, Umlauf and Burchard (2005) have also evaluated the performance of 350 $k - \epsilon$ for free convection and reported good approximations of buoyancy flux and entrain-351 ment depth. 352

In Figure 6.d and e, the temperature flux profiles from the Smagorinsky and the 353 implicit runs show striking similarities. The subgrid fluxes by Smagorinsky are negligi-354 ble, suggesting that numerical dissipation from the advection scheme alone provides most 355 of the diffusion. While this behavior is acceptable from a model fidelity perspective, it 356 is not entirely desirable, as the model effectively functions as an implicit LES despite the 357 prescription of an explicit closure. A higher-order advection scheme (not currently avail-358 able in CROCO) could potentially mitigate this issue. However, since this behavior is 359 observed only in free convection, it may not be a concern in more realistic configurations 360 (Section 3.2.3). 361



Figure 6. Averaged vertical velocity variance and temperature fluxes in the convection case, where the total flux is denoted as the solid lines and the subgrid component as dotted lines. The 1.25 m Smagorinsky run is used as the baseline of comparison in red lines.

3.2.2 Shear-only

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Compared to free convection, the shear-only case is far more sensitive to closure 363 choice. In the k- ϵ runs, turbulent mixing is entirely represented by subgrid processes. 364 Figure 7.c shows that $k - \epsilon$ produces identical, well-mixed temperature and velocity pro-365 files at different grid spacings, with zero velocity variance (Figure 8.c) and only subgrid 366 fluxes $(-K_m \partial_z u$ where K_m is eddy viscosity; See Figure 8.f). The lack of resolution sen-367 sitivity reflects the expected behavior of $k \cdot \epsilon$ as a RANS turbulence closure, which is well 368 validated for free-shear flows. Compared to Smagorinsky and the implicit solutions (Fig-369 ure 7.a, b), $k - \epsilon$ generates a deeper mixed layer, with strong stratification in the entrain-370 ment layer at 45 m. The maximum N^2 is almost twice as large as the baseline solution 371 (Figure 7.c). 372

The sensitivity to grid spacing is particularly pronounced for Smagorinsky and the implicit method. At $\Delta = 48$ m, the Smagorinsky solution shows no turbulent eddies (Figure 3). The temperature profile remains at 30 m from the initial condition, while the velocity profile shows strong surface shear above 10 m (Figure 7.e); At $\Delta = 48$ m, the momentum flux is entirely subgrid above 10 m in response to wind drag, while the tem-



Figure 7. Averaged temperature and u-velocity profiles for the shear-only case. Along with the temperature profiles (solid lines) in the top panels, the accompanying N^2 profiles are shown as dotted lines. The 1.25 m Smagorinsky run is used as the baseline of comparison in red lines. Note the differences in the x-axis scale in the bottom row.

perature flux is negligible at all depths (Figure 8.e, h). These profiles suggest that the mixed layer is not being deepened effectively, as further reflected in the MLD time series (Figure 12.b). Similarly, the implicit method at $\Delta = 48$ m also struggles to drive mixed-layer deepening, despite some turbulence appearing in the velocity field in Figure 3 (so is the non-zero variance in Figure 8.a).

Despite difficulties in representing mixed-layer deepening at coarse grid spacings, 383 both Smagorinsky and the implicit method demonstrate better performance at finer grids. 384 From $\Delta = 48$ m to 4 m, the temperature and velocity profiles (Figure 7.a, b, d, and e), 385 as well as the temperature and momentum fluxes (Figure 8.d, e, g, and h), coverage to 386 the baseline solution. The momentum fluxes increase within the mixed-layer interior (10-30)387 m), indicating stronger downward mixing of momentum input from surface wind stress. 388 Near the surface (above 10 m), the Smagorinsky subgrid flux for momentum contributes 389 less to the total momentum flux with smaller grid spacing, suggesting that vertical mix-390 ing is increasingly dominated by resolved turbulence rather than the subgrid closure, as 391 expected for LES. However, the opposite trend is observed in temperature flux, where 392



Figure 8. Averaged vertical velocity variance, together with the momentum and temperature fluxes from the shear case. In the flux profiles, the solid lines show the total flux and the dotted lines show the subgrid component. The 1.25 m Smagorinsky run is used as the baseline of comparison in red lines. Note that all k- ϵ runs produce the same profile consisting only of the subgrid component.

the subgrid contribution in the Smagorinsky runs increases with resolution near the entrainment layer at about 40 m.

³⁹⁵ When comparing the performance of different closures, the shear-only case high-³⁹⁶ lights k- ϵ 's ability to represent mixed-layer deepening with minimal resolution sensitivity in the gray zone. However, for nested simulations, $k \cdot \epsilon$ may suppress the emergence of turbulent eddies, making it problematic for intermediate nesting steps. Instead, transitioning from bulk closures to Smagorinsky or the implicit method may improve turbulence representation. Nevertheless, this approach introduces a trade-off with meanstate accuracy, which must be considered despite the benefits at higher resolutions. While shear-only forcing is uncommon at the ocean surface (Belcher et al., 2012), these findings may have implications for bottom boundary layer simulations (Umlauf et al., 2015; Wenegrat & Thomas, 2020).

3.2.3 Convection-wind-wave

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Beyond the single forcing cases, we extend our analysis to a more oceanographically relevant case driven by combined convection-wind-wave forcing (Figure 1). Similar figures for the convection-shear and wind-wave cases are available in the supporting information.

When comparing the mixed layer profiles of different closures for this case in Fig-410 ure 9, the k- ϵ solutions show the smallest deviation from the baseline, highlighting their 411 accuracy in representing mean mixed-layer profiles even with partially resolved turbu-412 lence (see the instantaneous flow field in Figure S2 from the supporting information). 413 Further, $k - \epsilon$ exhibits minimal sensitivity to grid spacing and maintains good accuracy 414 throughout the gray zone. In comparison, the Smagorinsky and implicit solutions show 415 strong sensitivity to grid spacing. For example, at $\Delta = 48$ m, instead of producing a 416 well-mixed boundary layer, both the Smagorinsky and the implicit solutions exhibit strong 417 temperature inversions and velocity shear near the surface (Figure 9.a, b, d, and e), in-418 dicating poor performance when turbulent eddies are barely resolved. When Δ is refined 419 to 4 m, both Smagorinsky and the implicit method generate a well-mixed boundary layer, 420 with nearly identical profiles converging to the baseline solution. 421

The variance and flux profiles in Figure 10 highlight the role of subgrid closure in 422 driving mixing relative to resolved motions. Similar to the convection-only case, the ver-423 tical velocity variance, which corresponds to the intensity of turbulent eddies, increases 424 with resolution for all three closures (Figure 10.a, b, and c). k- ϵ again stands out for its 425 consistency in representing total fluxes across the gray zone range. k- ϵ also maintains 426 a reasonable partition between resolved and subgrid fluxes, with a smaller relative con-427 tribution from the subgrid closure at higher resolution, suggesting its adaptability to grid 428 refinement (Figure 10.f and i). For Smagorinsky and the implicit method, despite poor 429 performance at coarse resolutions, the total flux converges with smaller grid spacings (Fig-430 ure 10.d, e, g, and h). Notably, the subgrid fluxes by Smagorinsky are considerably smaller 431 than those by k- ϵ at the same grid. At $\Delta = 4$ m, the Smagorinsky solutions and the 432 implicit solutions achieve a similar accuracy relative to the baseline. Here, the small sub-433 grid contribution in the Smagorinsky runs suggests that vertical mixing is well represented 434 by resolved eddies rather than the subgrid closure. 435

Overall, among different closures for the realistic forcings, $k - \epsilon$ best represents mixed-436 layer profiles but damps turbulent eddies. This may be advantageous for submesoscale-437 resolving simulations that require an accurate mixed-layer representation without explic-438 itly resolving turbulent eddies. Despite poor performance at coarse grids, the Smagorin-439 sky and the implicit solutions quickly converge, with a greater fraction of total flux car-440 ried by resolved turbulence. These properties may be useful for nesting applications tran-441 sitioning through gray zone resolutions toward well-resolved LES, where parent solutions 442 can influence turbulence in the nested domain (F. Chow et al., 2019). 443



Figure 9. Averaged temperature and u-velocity profiles for the convection-wind-wave case. Along with the temperature profiles (solid lines) in the top panels, the accompanying N^2 profiles are shown as dotted lines. The 1.25 m Smagorinsky run is used as the baseline of comparison in red lines. Note the differences in the x-axis scale in the bottom row.

3.3 Time dependence

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While the previous analysis focuses on instantaneous and temporally-averaged prop-445 erties, in this section we compare the MLD time series to highlight the joint impacts of 446 grid spacing and closure on the rate of mixed-layer deepening. There are many differ-447 ent methods to diagnose MLD, and the criteria based on a density difference from the 448 surface is commonly used in regional models (Courtois et al., 2017). In this study, the 449 surface value is not used, given that some gray zone simulations are not well mixed near 450 the surface (see Figure 9 for the temperature inversion in the 48 m Smagorinsky run from 451 the convection-wind-wave case). Instead, we define MLD as the depth where the den-452 sity exceeds the mixed layer average (taken between 15 m and 25 m) by 0.005 kg m⁻³. 453 A range of density thresholds $(0.01-0.001 \text{ kg m}^{-3})$ was tested, and the results were found 454 to be qualitatively robust to any reasonable choice of thresholds. We have tested another 455 common MLD criterion based on a density gradient (i.e., depth of N^2 maximum) fol-456 lowing Fan et al. (2023). However, we find the coarse resolution runs tend to generate 457 less pronounced N^2 maxima near the mixed layer bottom (see the stratification profiles 458 in Figure 7.a and b), leading to large fluctuations in the MLD time series at the initial 459



Figure 10. Averaged vertical velocity variance, together with the momentum and temperature fluxes from the convection-wind-wave case. In the flux profiles, the solid lines show the total flux and the dotted lines show the subgrid component. The 1.25 m Smagorinsky run is used as the baseline of comparison in red lines.

time steps. Therefore, to allow a better comparison of MLD evolution across different
 gray zone simulations, we choose the criteria of density threshold.

Figure 11 shows the MLD time series for the convection-only case comparing different closures (panel a-c) and summarizing the mean errors over the last inertial period (panel d). Similarly to the mixed-layer profiles discussed in Section 3.2.1, the mixed-layer



Figure 11. Time series of MLD in the convection-only case by different closures and grid resolutions in panel a-c. The mean MLD errors relative to the baseline run in the last inertial period are shown in panel d.

deepening rate shows minimal sensitivity to grid spacing for all closures. Shortly after model initialization, MLD deepens in all gray zone simulations due to surface buoyancy loss, reaching approximately 55 m after four inertial periods. While all closure solutions converge at smaller grid spacing, the k- ϵ solution at $\Delta = 48$ m exhibits the largest error from the baseline, though still relatively small (1-2 m). This error is consistent with earlier observations in the temperature profiles (figure 6.c), where k- ϵ struggles in the entrainment layer between 45 m and 55 m.

In contrast to convection, the mixed-layer evolution under wind forcing (Figure 12) 472 is more sensitive to closure and grid spacing. As discussed earlier, the k- ϵ runs in the shear-473 only case do not produce turbulent eddies, modeling turbulence mixing entirely as a sub-474 grid process. This results in an identical solution regardless of grid spacing. In Figure 475 12, the k- ϵ solutions exhibit a deeper mixed layer after four inertial periods, consistent 476 with the stratification seen in the temperature profiles (Figure 7.c). For Smagorinsky and 477 the implicit solutions, their inefficiency in driving vertical mixing at coarse grid spacings 478 (e.g., $\Delta = 48, 24, 12$ m) is evident for the lack of mixed-layer deepening in the MLD time 479 series (Figure 12.b, c), although we note these simulations also had a thick transition layer 480 between the mixed-layer and interior with reduced stratification (Figure 7.a, b). Over-481 all, MLD evolution in the shear-only case highlights the accumulation of mean-state er-482 ror in gray zone simulations, which could become particularly problematic for long-time 483 integrations or applications where the fidelity of large-scale conditions is critical for LES. 484

Figure 13 shows mixed-layer evolution under more realistic convection-wind-wave forcing conditions. In addition to mean-state drift, the MLD time series reveals notable



Figure 12. Time series of MLD in the shear-only case by different closures and grid resolutions in panel a-c. The mean MLD errors relative to the baseline run in the last inertial period are shown in panel d.

differences in the onset of mixed-layer deepening. For $k - \epsilon$, the mixed layer begins to deepen 487 immediately, following the baseline solution. However, for Smagorinsky and the implicit 488 method, the deepening of MLD is delayed at coarse resolutions. For example, in the Smagorin-489 sky solution at $\Delta = 48$ m, MLD remains at its initial value for 1.5 inertial periods, lead-490 ing to a shallower MLD for about 5 m ($\sim 10\%$ MLD) after four inertial periods. This 491 suggests that the transition to turbulence also depends on grid spacing and closure (see 492 Figure S9 for the time series of vertical velocity variance), which can be an important 493 consideration for strongly time-dependent turbulence problems such as the diurnal cy-494 cle (Wenegrat & McPhaden, 2015; Sutherland et al., 2016). At coarse grid spacings that 495 barely resolve turbulent eddies, $k - \epsilon$ may offer a better mean-state representation of the 496 mixed layer. When using Smagorinsky or the implicit method in the intermediate nest-497 ing steps toward LES, introducing additional perturbations to initial and boundary con-498 ditions may help spin up turbulent eddies and improve inner-domain solutions. 499

3.4 Effective Resolution

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A model's effective resolution is commonly defined as the horizontal length scale where dissipation—due to both the numerical discretization and turbulence parameterizations begin to significantly influence the solution, such that the underlying dynamics are no longer properly represented. It is often diagnosed by where the kinetic energy spectrum departs from the expected slope at small scales, for example, $k^{-5/3}$ for atmospheric mesoscale dynamics (Skamarock, 2004) and k^{-2} for oceanic submesoscale (Soufflet et al., 2016). Prior work indicates the effective model resolution is often reasonably approximated as



Figure 13. Time series of MLD in the convection-wind-wave case by different closures and grid resolutions (panel a-c). The mean errors relative to the baseline run is calculated for the last inertial period (panel d).

multiples of grid spacing ($\approx 7-10\Delta$), although the precise value can be a complicated 508 factor of many choices of numerical setup, including advection schemes for momentum 509 and tracers, as well as time-stepping routine (Soufflet et al., 2016). Here, we highlight 510 how gray zone simulations can incorrectly represent the dynamics near the turbulence 511 injection scale (e.g., section 3.1), such that even scales much larger than those expected 512 to be affected directly by model dissipation may be contaminated. In this section, we 513 highlight this behavior focusing on the convection-wind-wave case. The results of other 514 forcing cases display similar characteristics and can be found in Figure S10-14 from the 515 supporting information. 516

The effective resolution of the baseline run—due to explicit dissipation from the 517 eddy viscosity and the implicit numerical dissipation from the WENO5 scheme—can be 518 identified by the high-wavenumber roll-off of the spectral kinetic energy (Figure 14). The 519 baseline KE spectra, labeled as 1.25 m (SMAG) in all panels, follow the expected slope 520 of the inertial subrange $(k^{-5/3})$ starting from 50 m until wavelengths of approximately 521 10 m at which point there is a faster roll-off. The clear separation between the inertial 522 subrange and the numerical range indicates an effective numerical resolution of about 523 7Δ , similar to the range found previously for much larger-scale simulations (Soufflet et 524 al., 2016). However, only the baseline run has an unambiguous effective resolution scale 525 that can be identified from the spectral slopes. For the gray-zone runs of larger grid spac-526 ings (e.g., the implicit runs $\Delta > 4$ m), the inertial subrange is not well resolved while 527 still impacted by the numerical dissipation, making it impossible to identify the effec-528 tive resolution based on the expected spectral slope. This highlights the presence of under-529 resolved large eddies and indicates the gray zone nature of these simulations. The mis-530



Figure 14. Turbulent kinetic energy spectra from the convection-wind-wave case at 20 m depth, derived from the 3D velocity fields and averaged over the last inertial period. The dashed vertical lines mark the effective numerical resolution of the 1.25 m Smagorinsky run at 7Δ , and the characteristic -5/3 slope of the inertial subrange is shown in gray.

representation of large eddies and model dissipation is likely to contribute to errors in the mean-state profiles (see Section 3.2.3).

In Figure 14, the k- ϵ spectra display lower levels of energy than the baseline and 533 the implicit and Smagorinsky spectra at the same grid spacings, suggesting that the use 534 of $k - \epsilon$ in the gray zone substantially damps turbulent eddies. This is consistent with the 535 smoothed velocity gradients found in the flow structure (Section 3.1). At $\Delta = 48$ m, 536 the k- ϵ spectrum becomes overly flat with few turbulent eddies explicitly resolved. The 537 damping issue of k- ϵ is the most pronounced in the shear-only case where the closure com-538 pletely suppresses the emergence of turbulent eddies regardless of grid spacings (Figure 539 S11). For the $k - \epsilon$ simulations, the effective resolution cannot be meaningfully defined here, 540 as the dynamics of all scales up to the domain size are not properly represented. For nested 541 simulations across the gray zone, the use of $k - \epsilon$ may be problematic for supplying overly 542 laminar flows for boundary conditions, affecting the turbulence statistics in inner solu-543 tions. The importance of this will vary depending on the role of turbulence advection 544 between parent and child domains. 545

⁵⁴⁶ Compared to $k - \epsilon$, the Smagorinsky closure and the implicit method generate sim-⁵⁴⁷ ilar spectra of higher energy levels for the resolved turbulent motions (Figure 14). This ⁵⁴⁸ is valuable during intermediate nesting steps in the gray zone to better represent tur-⁵⁴⁹ bulence in the inner domain despite the trade-off in the mean-state accuracy (Section ⁵⁵⁰ 3.2.3). However, instead of aligning with the baseline solution, the spectral peaks in the ⁵⁵¹ Smagorinsky and the implicit runs shift towards longer wavelengths with larger grid spac-⁵⁵² ings, representing the model's attempt at representing turbulence even on the coarse grids.

The spectrum shift highlights how the spectral definition of effective resolution (based 553 on roll-off of the slope) may at times be misleading, as there can be shifts in wavenum-554 ber space or energy levels that are independent of the spectral shape (see also Figure S11). 555 For instance, the 24 m grid solution using Smagorinsky has a spectral shape broadly sim-556 ilar to the 1.25 m baseline run, but being shifted such that the spectral peak is at ap-557 proximately 1200 m wavelength rather than 150 m. It would be incorrect to conclude 558 that the effective resolution of this simulation is near 200 m (where there is a roll-off in 559 the spectrum), as comparison with the baseline simulation indicates inaccurate repre-560 sentation of turbulent motions at scales well exceeding 10Δ (in the sense of both the vari-561 ance contained at a given spatial scale and in terms of the coherent structures as dis-562 cussed in section 3.1). This should be considered for nesting applications, where inac-563 curate representations of boundary layer turbulence from the outer domain can be ad-564 vected into the inner domain through the boundary conditions, contaminating the final 565 solution, an issue identified previously for nested LES of the atmospheric boundary layer 566 (Mirocha et al., 2013; Mazzaro et al., 2017). 567

568 4 Summary and Discussion

This study evaluates the gray zone performance of two common turbulence closures, 569 $k \cdot \epsilon$ and Smagorinsky, as well as the implicit method using the WENO5 advection scheme, 570 for ocean surface mixed layer under different forcing conditions. With marginally resolved 571 turbulent eddies, we find that coherent turbulent structures, mean mixed-layer profiles, 572 and kinetic energy spectra exhibit closure-dependent resolution sensitivities. In general, 573 $k \cdot \epsilon$ solutions provide the highest accuracy in representing bulk properties such as mixed 574 layer depth, maintaining well-mixed temperature and velocity profiles across the gray 575 zone range. The total fluxes (subgrid plus resolved) by $k - \epsilon$ also remain mostly consis-576 tent, with reasonable adjustment to resolution in the partitioning between the subgrid 577 and resolved fluxes. However, the instantaneous flow fields and the kinetic energy spec-578 tra suggest $k - \epsilon$ damps the turbulence motions, representing mixing primarily as a sub-579 grid process, even when the grid is fine enough to begin resolving turbulence. The most 580 striking example is the shear-only case, where $k - \epsilon$ suppresses turbulence entirely, pro-581 ducing an identical laminar flow field at all grid spacings tested. In comparison, the Smagorin-582 sky and the implicit solutions show greater sensitivity to grid spacing. With coarse grids, 583 they fail to drive sufficient vertical mixing to deepen the mixed layer, resulting in strong 584 velocity shear and temperature inversions near the surface. However, with smaller grid 585 spacing, both Smagorinsky and the implicit method demonstrate improvement as their 586 solutions converge toward the baseline LES. Unlike the k- ϵ solutions, which are domi-587 nated by subgrid fluxes, the Smagorinsky and the implicit solutions primarily represent 588 turbulent mixing through the resolved flow motions, exhibiting enhanced kinetic energy 589 at the smaller scales. 590

The grav zone simulations are designed to inform modeling strategies on turbulence 591 closures during the intermediate scale between the typical RANS and LES paradigms. 592 Depending on modeling objectives, the consistency of $k - \epsilon$ is valuable for simulations aim-593 ing to resolve processes larger than boundary layer turbulence (e.g., Langmuir cells and 594 submesoscale fronts). Even at gray-zone resolutions with partially resolved turbulent ed-595 dies, $k - \epsilon$ shows good performance in representing bulk mixing in the boundary layer. How-596 ever, the k- ϵ solutions are associated with damped turbulence in the resolved velocity 597 fields. While the absence of resolved turbulence may be acceptable for some submesoscale-598 resolving simulations, it can be problematic for nested simulations that need to traverse 599 the gray zone to achieve realistic LES. For example, in a nested setup, an overly lam-600 inar parent-domain flow can inhibit small-scale turbulence development in the child do-601 main (Zhou et al., 2014). In this case, using $k - \epsilon$ in the gray zone may effectively act as 602 a much lower-resolution simulation than the grid spacing suggests. A potential remedy 603 is to switch from $k \epsilon$ to Smagorinsky or use no closure (implicit method) during inter-604

mediate nesting steps, allowing for more active turbulence at the boundaries. However, 605 this comes with the trade-off of less accurately resolved mean-state evolution, potentially 606 leading to solution drift in larger-scale properties. Given the fast convergence at smaller 607 grid spacing, this trade-off may be acceptable, especially considering that high-resolution 608 nests are typically run for short durations, limiting mean-state error accumulation. Strongly 609 time-dependent problems introduce an additional challenge: turbulence onset exhibits 610 sensitivity to both grid spacing and closure, making it difficult to predict a priori. This 611 sensitivity should be carefully considered when designing nesting strategies. 612

613 Here we have focused on the performance of common RANS and LES closures; however, several open questions remain regarding best practices for modeling in the ocean 614 gray zone. First, further case studies should investigate submesoscale frontal configura-615 tions and the associated instabilities (e.g., symmetric instability), as their length scales 616 likely fall within the gray zone and significantly impact turbulence properties in ways 617 not well captured by RANS closures (Bachman et al., 2017; Dong et al., 2022; Chor et 618 al., 2022). Second, despite our effort to design surface forcing representative of the re-619 alistic ocean, similar sensitivity analyses at the gray zone resolution should be applied 620 to realistic regional simulations formally validated by observational data. Third, the pre-621 scription of open boundary conditions—including issues such as the forcing frequency 622 and nudging strength—likely influences the fidelity of gray zone simulations. While this 623 issue has been recognized in atmospheric simulations (F. Chow et al., 2019), its impact 624 in ocean simulations remains largely unexplored (Scotti, 2010). Finally, while we have 625 tested existing RANS and LES closures in the gray zone outside their intended appli-626 cations, the existence of the gray zone problem ultimately necessitates the development 627 of scale-aware turbulence parameterizations and hybrid RANS/LES methods, such as 628 detached-eddy simulation (Spalart, 2009). High-resolution ocean simulation will bene-629 fit from a uniform turbulence model that can be seamlessly applied across the gray zone 630 without strong trade-offs between mean-state and turbulence representation, while also 631 minimizing the need for extensive sensitivity testing and ad hoc tuning. Efforts in this 632 direction can be guided by existing research on scale-aware turbulence closures for at-633 mospheric modeling (F. Chow et al., 2019; Honnert et al., 2020). 634

⁶³⁵ Open Research Section

The original source code for Coastal and Regional Ocean COmmunity model (CROCO) v1.3.0 used to generate the gray zone simulations is preserved at Zenodo (Auclair et al., 2022). The archiving of model data is underway, and will become publicly available at

⁶³⁹ Zenodo (https://zenodo.org/records/15116237).

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Supporting Information for "Evaluating Turbulence Parameterizations at Gray Zone Resolutions for the Ocean Surface Boundary Layer"

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- 1. Text S1 S2
- 2. Figures S1 S13

Introduction

The supporting information includes supplementary texts and figures from our ocean surface mixed layer simulations, which explore different atmospheric and oceanic forcing

:

scenarios. These additional figures, presented here for completeness, were generated using the same analytical methods and visual formats as the figures in the main text.

Text S1: Convection-shear Case

In the convection-shear case, the vertical velocity in Figure S1 displays patterns of turbulent eddies for all grid spacings and closures tested. Similar to other forcing cases, $k-\epsilon$ damps the small-scale eddy features. For the mean state profiles in Figure S3 and the flux profiles in Figure S4, the gray-zone behavior of $k-\epsilon$ is similar to the convection-only case where the mean states converge to the baseline solution with smaller grid spacings. The flux profiles also indicate a reasonable partition between the resolved and subgrid components. For Smagorinsky and the implicit method, the solutions suggest insufficient mixing, with significant temperature inversion and velocity shear near the surface. The mean-state solutions are inferior to $k-\epsilon$ especially at coarse grid spacing, although they converge to the baseline with grid refinement.

Text S2: Wind-wave Case

The wind-wave case was designed to investigate the gray zone behavior of a classical set up of Langmuir turbulence at $La_t \sim 0.3$ - a process that is often fully parametrized or resolved by ocean modeling studies. Among the temperature and velocity profiles in Figure S6, the most interesting feature is that the sensitivity of k- ϵ solutions to grid spacing, unlike the insensitivity found in the shear-only case. In addition, k- ϵ drives vertical mixing more effectively than Smagorinsky and the implicit method, especially for the coarse grid spacings where the boundary layer eddies could barely be resolved (i.e., $\Delta > 12$ m, relative to a mixed layer depth of 40 m). In Figure S7, the flux profiles by different closures display the similar characteristic as the convection-wind-wave case. The

total fluxes generally agree with the baseline solution, and the subgrid component from the closure reduces slightly in response to larger resolved turbulence at higher resolution.

[드 ⁶⁰⁰ > ₄₀₀

400 200 0

0 250

500 x [m] 1000

Ó

750



0.020

0.015

0.010

0.005

0.000 8

-0.005

-0.010

-0.015

-0.020

k-ε

500 x [m]

750 1000

1. Turbulence Flow Visualization

Figure S1. Snapshots of the vertical velocity (w) at 10 m depth for the convection-shear case.

250

500 x [m] 750

1000 0

250

500 x [m]

750 1000 0 250



Figure S2. Snapshots of the vertical velocity (w) at 10 m depth from the convection-wind-wave case.

2. Mixed Layer Profiles

2.1. Convection-shear



Figure S3. Averaged temperature and u-velocity profiles for the convection-shear case. Note the differences in x-axis scale in the bottom row.



Figure S4. Averaged vertical velocity variance, together with momentum and temperature fluxes from the convection-shear case. In the flux profiles, the solid lines show the total flux and the dotted lines show the subgrid component.



Figure S5. Time series MLD in the convection-shear case by different closures at the gray zone resolutions (Panel a-c). The mean error of the MLD time series relative to the baseline run in the last inertial period are shown in Panel d.



2.2. Wind-wave

Figure S6. Averaged temperature and u-velocity profiles for the wind-wave case of Langmuir turbulence. Note the differences in x-axis scale in the bottom row.





Figure S7. Averaged vertical velocity variance, together with momentum and temperature fluxes from the wind-wave case of Langmuir turbulence. In the flux profiles, the solid lines show the total flux and the dotted lines show the subgrid component.





Figure S8. Time series MLD in the convection-shear case by different closures at the gray zone resolutions (Panel a-c). The mean error of the MLD time series relative to the baseline run in the last inertial period are shown in Panel d.

24

Δ[m]

12

April 29, 2025, 2:53pm

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Depth [m]

Depth [m]

Depth [m]

48

3. Turbulence onset



Figure S9. Time series of vertical velocity variance in the convection-wind-wave case. The variance is calculated as the horizontal average at the middle of the mixed layer depth, which evolves throughout the simulations.

4. Kinetic Energy Spectra



Figure S10. Turbulent kinetic energy spectra from the convection-only case at 20 m depth, derived from the 3D velocity fields and averaged over one inertial period. The red vertical line marks the range of numerical dissipation at 7 Δ , and the characteristic -5/3 slope of the inertial subrange is shown in gray.



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Figure S11. Turbulent kinetic energy spectra from the shear-only case at 10 m depth, derived from the 3D velocity fields and averaged over one inertial period. The red vertical line marks the range of numerical dissipation at 7Δ , and the characteristic -5/3 slope of the inertial subrange is shown in gray.

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Figure S12. Turbulent kinetic energy spectra from the convection-shear case at 20 m depth, derived from the 3D velocity fields and averaged over one inertial period. The red vertical line marks the range of numerical dissipation at 7Δ , and the characteristic -5/3 slope of the inertial subrange is shown in gray.



Figure S13. Turbulent kinetic energy spectra from the wind-wave case at 10 m depth, derived from the 3D velocity fields and averaged over one inertial period. The red vertical line marks the range of numerical dissipation at 7Δ , and the characteristic -5/3 slope of the inertial subrange is shown in gray.