

1 **Spaceborne mineral mapping reduces dust's shortwave**
2 **radiative impact uncertainty**

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31 **Abstract**

32 Mineral dust impacts climate through complex interactions with radiation, which
33 remain poorly quantified due to uncertainties in the amount of light-absorbing iron
34 oxides within dust particles. NASA's EMIT imaging spectrometer, now delivering
35 high-resolution soil mineralogy from the International Space Station, provides the
36 first observational basis to address this gap at a global scale. Using the EMIT data
37 within Earth system model ensembles, we show that surface composition retrievals,
38 especially of iron oxides, reduce uncertainty in the dust shortwave direct radiative
39 effect by over 50% for both present-day and late-21st-century climates. The greatest
40 improvements occur over the Sahara, where the regional dust concentration is high
41 and dust radiative impacts are simulated with improved fidelity. While uncertainties
42 remain, EMIT shifts the primary uncertainty source from mineralogical composition
43 to our imprecise knowledge of the processes controlling the mass concentration of
44 dust particles, especially those related to emission. These findings represent a
45 pivotal step toward mineral-resolved dust aerosol modeling, offering improved
46 insight into how dust alters Earth's energy balance today and in a warming future.

47

48 **Main**

49

50 Mineral dust, Earth's most abundant atmospheric aerosol by mass¹, causes a
51 complex, yet poorly quantified, influence over climate². From perturbing the
52 planet's radiation budget³ to affecting cloud microphysics⁴ and fertilizing marine
53 ecosystems^{5,6}, dust interacts with nearly every Earth system component. Nowhere is
54 this influence more pronounced than in major source regions such as the Sahara,
55 Middle East, and East Asia, where vast quantities of soil particles are lofted into the
56 atmosphere¹. However, a critical obstacle remains in quantifying the dust direct
57 radiative effect (DRE), which strongly depends on the soil mineralogical
58 composition⁷⁻¹⁰ and, by extension, its optical properties.

59

60 Unlike purely scattering aerosols such as sulfates¹¹, dust particles both absorb and
61 scatter shortwave (SW) and longwave (LW) radiation^{3,12}. This complicated
62 interaction renders their net radiative impact especially complex, as the balance
63 between warming and cooling is dictated by multiple factors¹³, including the
64 abundance of iron oxides such as hematite and goethite⁷⁻¹⁰. These minerals strongly
65 absorb SW radiation, and their atmospheric abundance varies dramatically across
66 space and time^{14,15}. Yet until recently, Earth system models (ESMs) have been ill-
67 equipped to capture this mineralogical variability: Existing soil mineral atlases^{14,15}
68 were derived from sparse soil samples, often unrepresentative of soils in dust
69 source regions. This limitation leaves the SW DRE with a large uncertainty range of -
70 0.23 to +0.28 W m⁻², much of which is traced to the treatment of mineralogy⁷.

71

72 EMIT¹⁶ has fundamentally changed this landscape. Using a high-resolution imaging
73 spectrometer operating from its vantage point on the International Space Station
74 (ISS), EMIT has mapped the mineralogical composition of Earth's arid and semiarid

75 surfaces comprehensively at high spatial resolution¹⁷⁻²⁰, providing the first global-
76 scale observational dataset capable of resolving spatial patterns of dust-relevant
77 minerals, including iron oxides^{21,22} (Methods). These improvements are important
78 for understanding the dust DRE, especially in the context of future climate. As
79 warming and land use changes alter wind patterns, vegetation cover, and arid
80 landscapes, the sources and thus the mineralogical composition of dust aerosols are
81 expected to shift^{23,24}, which may cause changes that pre-EMIT soil atlases (C1999¹⁴
82 and J2014¹⁵) cannot capture with high accuracy. EMIT's constraints arrive at a
83 pivotal moment, providing a foundation for developing the next generation of ESMs
84 that explicitly resolve this issue. For example, EMIT can provide the mineralogical
85 composition of projected future dust sources with fidelity. With this information,
86 the scientific community now can redefine how mineral dust is represented in
87 ESMs^{7,25-28}.

88
89 Here, we incorporate EMIT-derived mineralogical atlases²⁹ (see "Data availability")
90 into an ensemble of four ESMs (Methods) to provide the first observationally
91 constrained estimate of the SW DRE under all-sky conditions at the top of the
92 atmosphere (a level definition consistently applied here and elsewhere) for both the
93 present day and the late 21st century. We compare these results with estimates
94 derived from the pre-EMIT soil atlases^{14,15} using the same methodology. A natural
95 question is whether the leap in mineralogical detail provided by EMIT translates to
96 better agreement between models and observations of dust radiative properties. To
97 answer this question, we compare simulations using the EMIT soil atlas against
98 satellite-derived DRE efficiency (DREE; defined as the DRE divided by the dust
99 extinction at visible wavelengths; Methods).

100
101 By grounding simulations in observationally constrained soil mineralogical
102 composition, we demonstrate that EMIT data reduce uncertainty in the dust SW
103 DRE by over 50% for both present-day and late-21st-century climates, which
104 represents a significant advance. This reduction is most pronounced over high-
105 albedo desert regions, where dust SW absorption plays an important role in its
106 radiative impact³⁰. The EMIT data thus both sharpen our understanding of dust's
107 role in present-day climate processes and strengthen our capacity to predict future
108 climate trajectories.

109
110 This work enables a key turning point in Earth system science in the context of dust:
111 the transition from bulk representations of mineralogically homogeneous dust,
112 which fail to capture the spatial variability of dust aerosols' optical properties³¹ and
113 their climatic impacts, to a nuanced, mineral-resolved modeling of its interaction
114 with radiation^{7,27,32}. As ESMs evolve to include ever more biogeochemical realism,
115 including the impact of specific dust minerals upon ocean primary productivity, our
116 findings demonstrate how targeted Earth observations can close longstanding

117 knowledge gaps, and, in doing so, reshape our understanding of one of the most
118 enigmatic agents of atmospheric change.

119

120 **EMIT unlocks the mineral fingerprint of global dust**

121

122 Leveraging measurements from a cutting-edge imaging spectrometer operating
123 from its vantage point on the ISS and the Tetracorder algorithm²⁰, EMIT mosaicked
124 cloud-free observations from August 2022 to October 2024 to generate estimates of
125 key minerals relevant to the dust SW DRE, including hematite and goethite, at the
126 instrument's native 60-meter resolution (Methods). These data were processed with
127 rigorous quality control on input reflectance, particle size attribution, and
128 uncertainty quantification (95% confidence interval; Methods), then aggregated to
129 0.5° spatial resolution for use in ESMs. Validated against in situ measurements^{20,33}
130 and covering nearly all major dust-emitting regions, the EMIT atlas provides a
131 foundation for physically based simulations of dust aerosols and their climate
132 impacts.

133

134 Our analysis suggests that EMIT's first map of soil iron oxides significantly improves
135 our ability to track these minerals in both soils and the atmosphere, which is critical
136 for estimating the dust SW DRE. Where earlier estimates of the soil iron oxides had
137 uncertainties exceeding 75%, EMIT cuts these to under 45% across most regions
138 (Fig. 1b; Supplementary Fig. 1; Methods). This leap in precision ripples through
139 climate simulations: when fed into Community ESM version 2 (CESM2) and ModelE
140 (Methods), EMIT's soil data reduce uncertainty in airborne iron oxide content by
141 over 80% in most dust-prone areas (Fig. 1a).

142

143 **EMIT advances knowledge of dust optical properties**

144

145 The precision and global comprehensiveness of EMIT's hyperspectral retrievals¹⁷
146 are reflected in a significant narrowing of uncertainty. The comparison of the
147 modeled clear-sky SW DREE with satellite-based retrievals at the top of the
148 atmosphere (Fig. 2) also suggests a clear signal: EMIT-derived soil iron oxides allow
149 the models to simulate the dust clear-sky SW DREE with dramatically tighter
150 uncertainty ranges than previously possible (Fig. 2). This improvement is seen over
151 the Sahara Desert, the planet's largest source of dust aerosols¹, which dominates
152 global dust absorption of SW radiation³⁴. In CESM2, EMIT reduces SW DREE error
153 over this desert by 80%, yielding values that fall within observationally constrained
154 ranges. Across the four ESMs with diverse aerosol size treatments and radiative
155 transfer schemes, EMIT fosters convergence in the SW DREE estimates within
156 $\pm 2.0 \text{ W m}^{-2} \tau^{-1}$ (τ , a unitless measure of optical depth at visible wavelengths) over
157 the Sahara Desert, compared to the $\sim 20 \text{ W m}^{-2} \tau^{-1}$ difference among these ESMs
158 using the pre-EMIT data. This accomplishment underscores the dominant role of
159 iron oxides in determining the SW DREE and highlights the more confident
160 estimates of iron oxides made possible by EMIT.

161

162 While we focus on CESM2 for clarity, similar improvements are expected in the

163 other models (e.g., Fig. 2). This tightening of uncertainty ranges is driven by EMIT's
164 improved characterization of iron oxide content (Fig. 1), specifically hematite and
165 goethite, which are powerful absorbers in the SW spectrum. Sensitivity analyses
166 confirm that uncertainty in the iron oxide mass fraction, as represented by pre-
167 EMIT soil atlases, is the primary driver of modeled SW DRE range across
168 experiments (Fig. 3, Supplementary Figs. 2-3).

169
170 On a global scale, EMIT does not drastically alter mean estimates of present-day SW
171 DRE, which remain within previously reported ranges^{7,35} (however, below, we note
172 a reduction in uncertainty due to EMIT). EMIT also does not lead to a significant
173 revision in the projected change in the dust SW DRE from the present day to the late
174 21st century (Fig. 3a). Nonetheless, EMIT introduces notable changes in regional and
175 monthly global values (Supplementary Figs. 4-5), particularly in spring/summer,
176 when Northern Hemisphere dust loading peaks (Supplementary Fig. 5). Over
177 reflective regions of North Africa, EMIT leads to enhanced dust warming by more
178 than 4.0 W m^{-2} ; conversely, it predicts stronger cooling over the Middle East. These
179 changes are tightly linked to EMIT-induced changes in iron oxide content
180 (Supplementary Fig. 6) with corresponding impacts upon dust single-scattering
181 albedo (SSA) (Supplementary Fig. 7).

182

183 **Untangling the sources of dust SW DRE spread across experiments**

184

185 As we delve into understanding the factors that influence the global dust SW DRE,
186 an intriguing picture emerges. Through multiple model analysis, we identify four
187 key drivers that contribute to the dust SW DRE uncertainty, arranged in a clear
188 hierarchy of influence in pre-EMIT simulations (Supplementary Figs. 2, 3, 8): (i) Iron
189 oxide content, (ii) model structural differences, (iii) the contrasting nudged and
190 model winds and their impact upon dust emission and transport, and (iv) emission
191 scenarios derived from shared socioeconomic pathways (SSPs). Among these, iron
192 oxides, represented by the pre-EMIT soil atlases, emerge as the dominant factor
193 (Supplementary Fig. 2). This influence is not a new finding; past studies, both
194 laboratory experiments^{8,36} and modeling efforts^{7,10}, have highlighted the importance
195 of these minerals to the dust SW DRE. What makes this mineralogical impact
196 particularly striking is its present-day relevance. While the composition of dust
197 remains important for future SW DRE, the magnitude of the future SW DRE is
198 increasingly driven by climate-induced changes in dust emission and transport.
199 Unlike the future, these present-day climate conditions are better constrained by
200 observations, providing a more robust basis for model simulations.

201

202 One objective of the EMIT mission was to estimate the dust SW DRE in the future.
203 Despite the common inputs in the ESMs, when examining the future, the models
204 diverge. Both CESM2 and ModelE show similar sensitivity to iron oxide
205 perturbations within observational uncertainties in the present day. However, their
206 projections of the future dust SW DRE differ markedly (Supplementary Fig. 3b),
207 indicating the increasing role of climate-driven factors in shaping dust's radiative
208 impacts on climate.

209

210 One of the most significant climate-driven factors is vegetation. Sensitivity
211 experiments using ModelE suggest that future changes to vegetation, especially the
212 response of plants to increased CO₂ levels, account for an additional 20% (pre-
213 EMIT) or 40% (EMIT) of the SW DRE variability (Supplementary Fig. 2). When
214 vegetation is fixed at present-day levels, the model projects stronger global dust
215 cooling (-0.43 W m^{-2}) compared to future scenarios where vegetation is allowed to
216 change dynamically (-0.24 W m^{-2}). The results indicate that CO₂-induced greening
217 (i.e., increased vegetation) could suppress dust emission in certain regions, aligning
218 with earlier findings^{23,24,37}.

219

220 The differences between models also contribute to divergent projections of the
221 future dust SW DRE. For example, CESM2 and ModelE yield relatively similar
222 present-day values (-0.41 W m^{-2} and -0.57 W m^{-2} , respectively). However, in future
223 climate scenarios, CESM2 projects stronger dust cooling (-0.56 W m^{-2}) than ModelE
224 (-0.24 W m^{-2}).

225

226 **EMIT reduces dust SW DRE uncertainties**

227

228 In the quest to improve the accuracy of the dust SW DRE projections, the EMIT
229 primary mission of detecting surface mineral composition in arid regions plays a
230 pivotal role in reducing uncertainties (Fig. 3). Before the EMIT mineral retrievals
231 were incorporated, uncertainty in the iron oxide content was the dominant factor,
232 accounting for over 90% and 50% of the SW DRE uncertainty estimated using the
233 pre-EMIT soil mineral data in present-day and late-21st-century climates,
234 respectively (Supplementary Fig. 2a). However, with the precise mineral fractions
235 provided by EMIT, this contribution has decreased substantially, dropping to just
236 20% and 3%, respectively. The impact of the EMIT retrievals is thus profound: it
237 significantly reduces the SW DRE uncertainty under both present-day and future
238 climate conditions.

239

240 Much of the uncertainty reduction is due to iron oxide data provided by EMIT. In
241 pre-EMIT estimates, iron oxide uncertainty alone introduces a 0.62 W m^{-2} impact,
242 over four times greater than that of model spread (0.17 W m^{-2}) in the present day or
243 scenario-related change (0.14 W m^{-2}) (Fig. 3). EMIT reduces this contribution
244 threefold to 0.10 W m^{-2} , bringing it in line with, or even below, other sources of
245 uncertainty. Iron oxides are therefore no longer the dominant uncertainty driver in
246 the global SW DRE estimates (Fig. 3b and Supplementary Fig. 2). This reduction in
247 uncertainty is a crucial step forward in refining our understanding of dust's impact
248 on global climate.

249

250 Spatially, EMIT reduces the dust SW DRE uncertainty caused by iron oxide content
251 across most major global dust source regions in the present day, particularly those
252 in the North Africa and Middle East (Fig. 4b). Regional analysis suggests a
253 percentage reduction in uncertainty over 50% for all the nine major sources
254 worldwide in the present day (Fig. 4c). This uncertainty reduction enabled by EMIT

255 remains substantial if all factors are considered in most the major source regions
256 (regional mean reduction still over 50%; Supplementary Fig. 9). The improvement is
257 mainly realized by improved description of iron oxide content compared to that in
258 the pre-EMIT soil mineral atlas (Fig. 1), which consistently indicates that iron oxide
259 content dominates the error sources to the dust SW DRE estimates prior to EMIT.

260

261 Future uncertainty projections of the dust SW DRE largely follow present-day
262 patterns (Supplementary Fig. 4). However, the Sahel and oceanic regions downwind
263 of dust sources in North Africa and the Middle East exhibit significantly greater
264 future uncertainty compared to present-day estimates (Supplementary Fig. 9b). This
265 suggests that although EMIT substantially refines present-day estimates, future
266 uncertainties continue to be influenced by various climate factors, including those
267 that are difficult to predict, such as clouds, droughts, and vegetation changes.
268 Nevertheless, EMIT significantly reduces the regional mean uncertainties in the
269 future projections of dust SW DRE across most of the nine major dust source regions
270 (Supplementary Fig. 10).

271

272 Despite these gains enabled by EMIT, four pivotal limitations persist. First, the
273 relationship between observed and modeled dust SSA continues to show scatter
274 (Supplementary Fig. 11), even with EMIT's comprehensive global data, pointing to
275 the importance of constraining factors beyond the soil mineral abundances, such as
276 particle size distribution. Second, over oceanic regions (Fig. 4b), EMIT offers only a
277 limited constraint, as the SW DRE here is more sensitive to the other factors such as
278 the size distribution and vertical structure of dust that are subject to model biases³⁸.
279 Third, the LW DRE remains poorly constrained due to the lack of direct detection of
280 quartz and feldspar^{9,10}, which lie beyond EMIT's spectral capabilities, and the
281 absence of their radiative impacts in our models. This partially results in a
282 persistent discrepancy between observed and modeled LW DRE (Supplementary
283 Figure 12). Fourth, the influence of vegetation, which accounts for over 20% of the
284 SW DRE spread across experiments (Supplementary Fig. 2b), exhibits strong model
285 dependence and is uncertain. Our ensemble also underrepresents the full range of
286 potential future climates simulating only the Coupled Model Intercomparison
287 Project Phase 6 (CMIP6) SSP2-4.5 scenario³⁹ in the late-21st-century, underscoring
288 the need for broader sensitivity sampling.

289

290 **EMIT shifts dust SW DRE research priorities**

291

292 Our findings recalibrate the research agenda for dust-climate interactions. While
293 iron oxides remain central to the SW DRE estimate, EMIT's introduction shifts the
294 axis of residual uncertainty. In ESMs using pre-EMIT soil atlases, mineralogy is the
295 primary bottlenecks (Fig. 3 and Supplementary Fig. 2). With EMIT, uncertainty is
296 restructured, now with comparable factors other than soil iron oxides, such as
297 particle size distribution, the influence of vegetation on dust emission and its
298 sensitivity to climate, and the LW DRE of unresolved minerals like quartz and
299 feldspar.

300

301 Under warming scenarios, the uncertainty in dust's influence on climate appears to
302 evolve from being dominated by mineral composition to climate-dependent
303 processes. These include uncertainties on processes that shift dust source regions,
304 change vegetation-dust emission sensitivity, and alter dust removal and transport
305 dynamics. These processes are not yet captured by existing ESMs, even though they
306 can represent dust aerosols as component minerals dynamically responsive to
307 climate drivers.

308

309 **Implications and future directions**

310

311 The implications of these refined SW DRE estimates are three-fold. First, the
312 narrowed uncertainty around the SW DRE increases the likelihood that dust exerts a
313 non-negligible SW cooling influence in both the present-day and late-21st-century
314 climates. If future dust loading increases in a manner consistent with estimates from
315 the preindustrial era to the late 20th century², it could partially offset greenhouse
316 warming in climate-sensitive regions such as the Sahel and South Asia. In these
317 areas, dust simultaneously affects air quality⁴⁰, ecosystem function⁴¹, and
318 hydrological regimes^{42,43}. Second, EMIT delivers the first observational benchmark
319 for global-scale dust mineral modeling. This benchmark reveals a pivotal transition:
320 uncertainties once dominated by dust mineralogy now likely hinge more on particle
321 size distribution⁴⁴ and processes governing dust emission². Third, EMIT's limited
322 improvement over oceans, where low surface albedo partially covers iron-oxide
323 signals, highlights the urgency of deploying complementary airborne/shipborne
324 campaigns targeting size-resolved dust radiative properties.

325

326 Looking forward, three strategic priorities emerge. First, future spaceborne
327 missions should expand mineralogical detection into wavelengths that are sensitive
328 to quartz and feldspar, minerals critical for the dust LW DRE⁹, and cloud ice
329 nucleation in mixed-phase clouds⁴⁵, respectively. These components remain
330 effectively invisible to EMIT but are pivotal for resolving dust's full climate footprint.
331 Second, models should move beyond static mineralogical baselines to incorporate
332 spatial and temporal variations in dust composition. Shifting climate conditions via
333 aridification^{46,47}, CO₂-driven vegetation change²⁴, or cryosphere retreat⁴⁸ may not
334 only alter dust emission but also modify dust sources and thus dust mineral
335 composition, particularly the prevalence of iron oxides. These changes can reshape
336 regional absorption patterns and, by extension, local-to-global climate responses.
337 Third, intensive ground-based campaigns are needed in data-sparse regions to
338 ground-truth orbital retrievals and disentangle mineralogical composition from the
339 other drivers of the DRE uncertainty. These efforts are equally vital for refining
340 ESMs and informing local dust mitigation strategies, as dust absorption modulates
341 surface heating, atmospheric stability⁴⁹, and even pollutant dispersion⁴⁰.

342

343 Beyond radiative impacts, EMIT's mineral atlases unlock entirely new frontiers from
344 snow-darkening feedbacks in cryosphere regions⁵⁰ to iron fertilization in marine
345 biogeochemistry^{5,6}, and to cloud microphysics⁴ influenced by ice-nucleating
346 minerals. With dust mineralogical composition now transformed from a poorly

347 constrained boundary condition into a resolvable Earth system component, this
348 work affirms the transformative potential of orbiting imaging spectroscopy. As
349 ESMs increasingly adopt mineral-resolved representations of dust^{7,25-28}, there is a
350 growing opportunity to better integrate mineralogical detail with dynamical
351 processes to improve our understanding of dust's role in past, present, and future
352 climates.

353

354

355 **Methods**

356

357 **The EMIT soil mineral atlases and associated uncertainties**

358

359 NASA's EMIT mission provides the first high-resolution global atlas of soil mineral
360 mass fractions at 0.5° spatial resolution²⁹. These mass fractions are derived from
361 calibrated VSWIR (visible to shortwave infrared) reflectance spectra collected by
362 the EMIT imaging spectrometer¹⁷ from its position aboard the ISS, using a Hapke-
363 like model in combination with the Tetracorder spectral matching algorithm²⁰.

364

365 Key mineral groups, such as chlorite, dolomite, goethite, gypsum, hematite,
366 illite/muscovite, kaolinite, montmorillonite, and vermiculite, are directly mapped,
367 while quartz and feldspar abundances are inferred indirectly due to the absence of
368 diagnostic absorption features in the VSWIR range.

369

370 The Supplementary Sections "The state-of-the-art EMIT imaging spectrometer" and
371 "The first high-resolution soil mineral atlases from EMIT and associated
372 uncertainties" respectively describe the EMIT instrument and the methodology used
373 to generate the soil mineral atlases, including the derivation of 95% confidence
374 intervals for mineral mass fractions. Additional details on the spectrometer design
375 and the full methodology for producing the mineral atlases are provided in prior
376 documents¹⁸⁻²⁰. The publicly available EMIT soil mineral atlases are archived at
377 NASA's Distributed Active Archive Center (DAAC)⁵¹ (see "Data availability"). In this
378 study, these publicly available atlases are incorporated into ESMs to derive
379 observationally constrained estimates of the dust SW DRE for both the present day
380 and the late 21st century.

381

382 A calibration of a free parameter in the EMIT mineral mass fraction derivation (see
383 Supplementary Section "The first high-resolution soil mineral atlases from EMIT
384 and associated uncertainties") helps reduce the global bias between modeled and
385 AEROSOL ROBOTIC NETWORK (AERONET)-based dust SSA compared to pre-EMIT soil
386 mineral atlases (Supplementary Fig. 11). This calibration step ensures consistency
387 between the simulated dust SSA using the retrieved mineralogy and the observed
388 values. However, it limits the use of the AERONET-based dust SSA global bias as an
389 independent validation metric. Consequently, this study emphasizes the reduction
390 in uncertainty enabled by incorporating EMIT-derived soil mineralogy in ESM
391 simulations.

392

393 **Model descriptions**

394

395 Three ESMS and one atmospheric chemistry model, which is also referred to as an
396 ESMS for brevity, are used in this study: Community Atmosphere Model version 6
397 (CAM6) embedded within the CESM2, the NASA Goddard Institute for Space Studies
398 (GISS) Earth System ModelE2.1 (referred to as ModelE), the National Oceanic and
399 Atmospheric Administration (NOAA) Geophysical Fluid Dynamics Laboratory
400 (GFDL) AM4, and the Barcelona Supercomputing Center's (BSC) Multiscale Online
401 Nonhydrostatic Atmosphere Chemistry model (MONARCH). All these ESMS were
402 recently enabled to simulate dust aerosol composition as prognostic mineral
403 mixtures^{7,25-28}. The following provides a brief overview of the calculations of dust
404 optical properties and DRE; Supplementary Section "Descriptions of mineral-
405 resolving Earth system models" describes these calculations and each of the four
406 ESMS in detail.

407

408 *Treatment of optical properties*

409

410 The four ESMS employ distinct approaches to simulating dust mineralogy,
411 particularly for iron oxides (hematite and goethite), which are critical for modeling
412 dust optical properties in the SW spectrum. Differences among ESMS arise from
413 three key aspects: whether hematite and goethite are distinguished, how mineral
414 mixtures are represented, and the choice of iron oxide refractive indices.

415

416 CESM2 and ModelE treat hematite and goethite collectively as a single iron oxide
417 species in their optical property calculations, whereas MONARCH and AM4
418 distinguish between the two minerals, if both are characterized by a soil atlas. For
419 the mixing state, CESM2 represents iron oxides as part of an internal mixture with
420 all other minerals and aerosol species within each dust mode (i.e., all species
421 presented are internally mixed), both during transport and for radiative
422 calculations. In contrast, ModelE, MONARCH, and AM4 simulate iron oxides in two
423 forms during transport: internally mixed with other minerals and as pure
424 (externally mixed) minerals; for radiative calculations, however, they are treated as
425 internally mixed with a host mixture composed of the other (weakly absorbing)
426 minerals. For further details on each model, see Supplementary Section
427 "Descriptions of mineral-resolving Earth system models".

428

429 As to refractive indices of individual minerals, all models use identical data sourced
430 from the same prior study²⁷ for non-iron-oxide minerals. For iron oxides, although
431 the refractive indices in the SW spectrum differ slightly between CESM2 and the
432 other three models, these differences are relatively minor^{31,32}; moreover, the
433 variables related to the dust SW DRE simulated by all the four ESMS are reasonably
434 consistent with retrievals of satellite-based SW DRE at the top of the atmosphere
435 under clear-sky conditions and AERONET-based mid-visible-band (0.44-0.63 μm)
436 SSA (Supplementary Figure 11).

437

438 CESM2 uses the Modal Aerosol Module (MAM4) with three dust modes, while the
439 other three models simulate dust with a binned method. Aerosol optical properties
440 in CESM2 are calculated based on the wet surface mode radius and wet refractive
441 index of each component, with hygroscopic growth estimated using κ -Köhler
442 theory⁵². CESM2 applies the volume mixing rule to derive the refractive indices of
443 the mixture and parameterizes optical properties using Chebyshev polynomials⁵².
444 This approach captures the effects of composition, size, and ambient humidity on
445 the aerosol optical property diagnostics. Because of the use of the parameterized
446 optical properties, CESM2 simulates SSA collectively for all aerosols rather than dust
447 alone; dust-specific values are estimated by selecting pixels where dust contributes
448 more than 50% of the total aerosol optical depth in the mid-visible band (0.44-0.63
449 μm). In contrast, the other models directly simulate SSA for dust aerosols
450 (Supplementary Section “Descriptions of mineral-resolving Earth system models”).
451

452 *Diagnostics of dust DRE and DREE*

453
454 In each of the four ESMs, the dust DRE is diagnosed as the difference in net radiative
455 fluxes with and without dust aerosols, computed for both SW and LW bands. The
456 simulated dust DREE is evaluated against clear-sky observations at the top of the
457 atmosphere. Robust inter-model and model-observation comparisons are
458 complicated by methodological heterogeneity in diagnosing clear-sky DRE across
459 the four ESMs and by sampling limitations specific to ModelE. Accordingly, the
460 analysis relies on all-sky DRE values for ModelE and for models (i.e., MONARCH and
461 AM4), where clear-sky DRE was not explicitly archived. This likely has little impact
462 on the model-observation comparisons over the key dust source region, the Sahara
463 Desert, where low cloud cover means that the difference of clear-sky and all-sky
464 DRE values is not substantial.
465

466 **ESM simulations**

467
468 We conduct 36 ESM simulations using both pre-EMIT and EMIT soil mineral atlases
469 as inputs for the four ESMs (see Supplementary Section “Descriptions of mineral-
470 resolving Earth system models” for model configurations).
471

472 We focus on the dust SW DRE, because EMIT’s contributions to reducing the dust
473 DRE are not uniform across all spectral bands or geographic domains. Over oceanic
474 regions, improvements resulting from EMIT constraints upon dust mineral
475 composition are limited (Figs. 2, 4), due to the greater sensitivity of the SW DRE to
476 factors like dust size and concentrations that result from transport biases³⁸.
477 Similarly, the LW DREE remains largely unaffected (Supplementary Fig. 12), as
478 EMIT does not retrieve the soil fractions of key absorptive minerals at thermal
479 wavelengths, such as quartz^{7,9,10}, which lack prominent spectral signatures in the
480 wavelength range measured by EMIT. Instead, the mass fractions of quartz and
481 feldspar are inferred from the Hapke-like model using grain size retrieved from
482 EMIT reflectance²⁰ (see Supplementary Section “The first high-resolution soil
483 mineral atlases from EMIT and associated uncertainties”). The partitioning of these

484 minerals into clay and silt fractions is subsequently derived from the empirical
485 relationships presented in C1999, applied to the soil texture retrieved from each
486 EMIT scene²⁰.

487

488 The dust SW DRE is strongly influenced by the mineral fraction of iron oxides such
489 as hematite and goethite^{7,10,53}. We incorporate best estimates and 95% confidence
490 intervals for iron oxides from the pre-EMIT and EMIT atlases. These simulations
491 capture key drivers of the simulated DRE, including emission scenarios, vegetation
492 and drought variability, and model-specific sensitivities to the atmospheric
493 circulation.

494

495 Supplementary Tables 1-2 summarize the numerical experiments conducted in this
496 study, which form the basis for quantifying the dust SW DRE and its associated
497 uncertainty under present-day and late-21st-century climate conditions. The
498 following three subsections provide detailed descriptions of these experiments.

499

500 *Present-day dust SW DRE using EMIT*

501

502 The dust aerosol burden or load is intimately related to meteorological processes
503 like winds that raise and transport dust particles. We attempt to match the model
504 winds to those observed by relaxing toward reanalysis values from the Modern-Era
505 Retrospective analysis for Research and Applications dynamics version 2 (MERRA2)
506 or else the National Centers for Environmental Prediction (NCEP) or European
507 Centre for Medium-Range Weather Forecasts (ECMWF) Reanalysis version 5 (ERA5).
508 We use reanalysis values between 2006 and 2011. This period is chosen to allow
509 comparison to previous sensitivity studies⁷ along with observations. Sea surface
510 temperature and sea ice are prescribed from observations during this period.
511 During the first year (2006), the dust cycle “spins up” to statistical equilibrium and
512 is excluded from the data analysis. These experiments, denoted as “(ESM name)-
513 (mineral atlas)_base_(reanalysis product)” in Supplementary Table 1, are our best
514 estimate of the present-day dust DRE.

515

516 Additional experiments simulating the present-day dust cycle are carried out to
517 estimate the uncertainty of our best estimate DRE. These experiments are also listed
518 in Supplementary Table 1. Previous work demonstrates that the soil mass fraction
519 of hematite and goethite makes the largest contribution to the DRE uncertainty⁷ in
520 mineral-resolved ESMs. To bound the SW DRE uncertainty, we carry out two types
521 of experiments (e.g., “EMIT_hi_irox” and “EMIT_low_irox”) using the EMIT soil
522 mineral atlases, perturbing the hematite and goethite soil fractions to their
523 respective retrieved high and low values, based upon the uncertainty of the quartz-
524 feldspar grainsize parameter (see Section “The first high-resolution soil mineral
525 atlases from EMIT and associated uncertainties”).

526

527 In the absence of meteorological reanalysis to drive calculations of future DRE in the
528 late 21st century, we use winds computed prognostically by CESM2 and ModelE. For

529 comparison of present-day and future control experiments, we add present-day
530 experiments “freeWD”, where the dust DRE is calculated with prognostic winds.

531

532 *Uncertainty in present-day dust SW DRE before EMIT*

533

534 To assess the effect of EMIT retrievals of soil composition compared to previous
535 estimates by pre-EMIT soil mineral atlases, we estimate the SW DRE in two
536 additional types of experiments using these prior estimates of soil minerals
537 (Supplementary Table 1: “C1999” and “J2014”).

538

539 Prior to EMIT, spatial variations in soil mineral composition were inferred through a
540 global atlas that compiled observations of regionally varying soil descriptors like
541 soil type¹⁴ or soil unit¹⁵. Each type or unit was assumed to have a characteristic
542 fractional distribution of minerals that was identical at every location where that
543 descriptor was indicated by the atlas. This assumption was necessitated by the
544 extremely limited number of direct measurements of soil mineral composition.
545 Uncertainty of soil mineral fractions thus arose through two effects: the uncertain
546 characteristic mineral fractions for each soil descriptor that were inferred
547 empirically through limited measurements, and uncertainty in the gridded soil
548 descriptor provided by the atlas that was constructed from multiple regional
549 surveys with spatially inhomogeneous sampling density and especially low
550 resolution in the remote arid regions that are prolific dust sources.

551

552 DRE uncertainty prior to the availability of EMIT mineral composition retrievals was
553 estimated by a previous study⁷ using a wide range of sensitivity studies with CESM2,
554 as well as complementary experiments with ModelE, AM4, and MONARCH. These
555 experiments focused on the DRE uncertainty due to poorly constrained knowledge
556 of soil minerals before EMIT. Two different atlases prescribing the mineral content
557 of soils were used (C1999 and J2014). Based on the reported soil mineral ranges¹⁴,
558 the previous study⁷ calculated the 95% confidence interval of iron oxides for each
559 soil type. The DREs were then calculated in sensitivity studies based on the high and
560 low distribution of mineral content in the soils⁷. Their analysis⁷ showed that the SW
561 DRE uncertainty is most dependent on the hematite and goethite mass fractions,
562 with little dependence upon other mineral fractions, especially those not detected
563 by EMIT (i.e., quartz and feldspar). Overall, the range of dust net DRE was calculated
564 to be -0.23 to 0.35 W m⁻², and over 90% of this uncertainty is due to uncertainty in
565 the soil mineralogy⁷. Smaller uncertainty results from different estimates of aerosol
566 mineral composition by different models (despite using the same soil mineral atlas).

567

568 These calculations characterize the uncertainty related to the mineral fractions that
569 were attributed to each soil type or unit, and which were derived based upon
570 limited measurements. This underestimates the true mineral uncertainty by
571 excluding contributions from the mischaracterization of soil type or unit within the
572 atlas at any location due to limited spatial sampling and uncertain interpolation. To
573 account for this “unknown” uncertainty caused by limited measurements based on
574 which the pre-EMIT soil mineral atlases were created and by the challenges in

575 quantifying uncertainty in the soil type atlas used to infer mineral composition, we
576 consider the difference between C1999 and J2014 base simulations as an error
577 source of iron oxides.

578

579 *Future projections of dust, SW DRE, and associated uncertainties*

580

581 Future experiments take atmospheric composition (including CO₂ and
582 concentrations of non-dust aerosol species) and ocean surface conditions (including
583 sea surface temperature and sea ice) from coupled experiments that are part of the
584 standard CMIP6 protocol, carried out separately with CESM2 and ModelE. Here we
585 estimate the dust SW DRE, using EMIT soil atlases and the mineral-resolved version
586 of each model, allowing changes to the present-day source extent due to projected
587 vegetation changes.

588

589 In order to quantify the future dust SW DRE, as a result of changes in dust source
590 area from climate change, CESM2 and ModelE are run with prognostic winds and
591 the best estimate of the EMIT soil mineral atlas to simulate future climate in the late
592 21st century, 2090-2094: experiments denoted by “Future_BASE” in Supplementary
593 Table 2. To diagnose future changes in the SW DRE, based on the EMIT soil atlases,
594 this set of simulation is contrasted to the “Present-day_online” experiment in
595 Supplementary Table 1. As with the present-day experiments, we include
596 simulations spanning the uncertainty range of soil mass fractions of hematite and
597 goethite retrieved by EMIT, denoted by “hi_irox_FU” and “low_irox_FU”, respectively.
598 Separate experiments show the effect of EMIT retrievals compared to using prior
599 soil mineral atlases, denoted by “C1999_base_FU” and “J2014_base_FU”.

600 Uncertainty of future dust DRE is characterized using different scenarios, as
601 projected by CESM2 and ModelE, to at least partially cover the spread among late-
602 21st-century climate projections by CMIP6. A key uncertainty of future DRE
603 projections results from uncertain projections of vegetation. Previous studies
604 suggest that most models predict an increase in leaf area index coverage in the
605 future⁵⁴, indicating a likely decrease in dust source regions, although this is sensitive
606 to the assumptions of CO₂ fertilization³⁷. CESM2 calculates vegetation
607 prognostically, whereas vegetation is prescribed in ModelE. For the standard
608 “base_FU” scenario, ModelE uses the leaf area index from CESM2. To see the effect of
609 leaf area index changes upon dust DRE over the late 21st century, we carry out two
610 additional experiments with 2090-2094 atmospheric composition and present-day
611 or future vegetation (denoted in Supplementary Table 2 by “ModelE_CAMvegPD”
612 and “ModelE_CAMvegFU”, respectively).

613 Finally, we consider the sensitivity of the dust SW DRE to late-21st-century
614 atmospheric composition according to the CMIP6 intermediate-mitigation (SSP2-
615 4.5) and high-emissions (SSP5-8.5) scenarios, using CESM2 whose vegetation
616 responds prognostically to this range of dust DRE (“base_FU245” and “base_FU585”,
617 respectively, in Supplement Table 2). Although fully characterizing the influence of
618 future climate uncertainty is beyond the scope of the EMIT mission, differences

619 among the ESMs with varying climate sensitivities offer some insight into the range
620 of dust-climate interactions in a warming world.

621 Recognizing that some ESMs simulate a relatively flat dust trend from the pre-
622 industrial era to the year 2000, which contrasts with the $\sim 55\%$ increase according
623 to deposition measurements², we conduct an additional experiment to evaluate the
624 implications of this alternative trend on future projections. In this experiment, we
625 increase dust optical depth in the mid-visible band by $\sim 30\%$, extrapolating the rate
626 since the pre-industrial suggested by dust proxies². This amplifies the dust SW DRE
627 to -0.57 W m^{-2} with the EMIT soil mineral atlas and -0.51 W m^{-2} with the pre-EMIT
628 atlases.

629 **Quantifying uncertainty in dust SW DRE**

630

631 We divide the factors affecting the dust SW DRE (ξ) estimate into four
632 subcategories: dust mineral composition, assumed to be dominated by the mass
633 fraction of iron oxides (φ), model spread (σ), future scenarios (ψ), and vegetation
634 dynamics (η). The iron oxide fraction is assumed to dominate the compositional
635 uncertainty ($\Delta\xi$) in the dust SW DRE based upon a prior study⁷. We perturb the iron
636 oxide abundance based on the best estimates described in both EMIT (E) and C1999
637 (C) or J2014 (J) soil mineral atlases (Q). The resultant high- ($\Delta\xi_{Q_h}$) or low-branch
638 ($\Delta\xi_{Q_l}$) uncertainty is obtained by calculating the difference between the SW DRE
639 from the base mineral simulation and that from the paired uncertainty experiments
640 in each model (m) with either increased (φ_{Q_h}) or decreased (φ_{Q_l}) iron oxide
641 fractions.

642

643 We first determine whether the dust SW DRE in a perturbed experiment represents
644 more cooling or warming compared to the base (b) simulation. We then calculate
645 the uncertainty for each branch by finding the difference between the simulated
646 dust SW DRE from the paired simulations in each ESM using the EMIT or C1999 soil
647 mineral atlases. For example, when using the EMIT (E) soil atlas, the high-branch
648 uncertainty in the dust SW DRE, $\Delta\xi(\varphi_{E_h,m})$ with respect to the best estimate of the
649 dust SW DRE, $\xi(\widetilde{\varphi_{E_b,m}})$ in model m (i.e., CESM2 and ModelE) due to higher (h) soil
650 iron oxides (φ_{E_h}) relative to the base value ($\widetilde{\varphi_{E_b,m}}$), is calculated as follows:

651

$$\Delta\xi(\varphi_{E,m})_h = \left| \xi(\varphi_{E_h,m}) - \xi(\widetilde{\varphi_{E_b,m}}) \right|. \quad (1)$$

652

653 There are multiple paired simulations in CESM2 with C1999, which has two
654 simulations representing high-bound iron oxides, one for iron oxides in the clay-
655 sized category (c) and the other for those in the silt-sized category (s). Although
656 each simulation isolates perturbations in a single dust size category, both contribute
657 similarly to the dust SW DRE due to their shared dust source regions, emission
658 processes, and percentage uncertainties. Therefore, we calculate their sum to
659 represent the amplitude for this high branch:

660

$$\Delta\xi(\varphi_{C,m})_h = \left| \xi(\varphi_{C_{hc},m}) - \xi(\widetilde{\varphi_{C_b,m}}) \right| + \left| \xi(\varphi_{C_{hs},m}) - \xi(\widetilde{\varphi_{C_b,m}}) \right|, \quad (2)$$

661

662

where m refers specifically to CESM2.

663

664

The low-branch dust SW DRE uncertainty, $\Delta\xi(\varphi_{Q,m})_l$ with respect to the best estimate of the dust SW DRE in model m due to lower (l) soil iron oxides relative to the base value using the pre-EMIT or EMIT soil atlases, is calculated using the same approach.

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The uncertainty resulting from the pre-EMIT soil atlases is harder to quantify because the soil mineral fractions are inferred indirectly from other observed quantities following multiple assumptions. In the absence of global measurements of soil minerals prior to EMIT, the mineral fractions were derived from a soil descriptor like soil type (C1999) that was available from a global atlas. Each soil type is assumed to have a characteristic mineral composition based upon empirical relations derived from the limited number of soil measurements. Thus, the total uncertainty has a contribution from the imprecise relation between soil type and composition. This is the first term in Equation (3) below. Additional uncertainty results from the use of two alternative soil descriptors to infer the mineral fractions: soil type (C1999) or soil unit (J2014). Thus, our estimate of the pre-EMIT (P) high-branch uncertainty $\Delta\xi(\varphi_{P,m})_h$ additionally includes the magnitude of the difference between C1999 (C) and J2014 (J) as an error source to account for the differing base values of the dust SW DRE resulting from whether soil type or unit is chosen to specify the iron oxide fraction:

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$$\Delta\xi(\varphi_{P,m})_h = \sqrt{\left(\Delta\xi(\varphi_{C,m})_h\right)^2 + \left[\xi(\widetilde{\varphi_{J_b,m}}) - \xi(\widetilde{\varphi_{C_b,m}})\right]^2}. \quad (3)$$

683

684

The two terms in this Equation (3) represent distinct and uncorrelated sources of uncertainty: one captures variability within a single mineral atlas (Gaussian spread), while the other reflects systematic bias between different soil mineral atlases (mean difference).

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There is an additional contribution to pre-EMIT uncertainty that is harder to quantify that results from misidentification in the global atlas of the soil descriptor at a particular location where the descriptor was derived by interpolation rather than direct observation. The global atlas is constructed from interpolation of measurements made with varying spatial resolution that are often concentrated in agricultural regions rather than arid dust sources. This contribution to uncertainty and its spatial variation is difficult to reconstruct; we simply note its existence, while omitting it, and regard Equation (3) as a conservative estimate of the total pre-EMIT uncertainty. Note that this contribution from interpolation, along with the final term

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698 in Equation (3), are absent in the EMIT uncertainty, because EMIT provides direct
 699 retrievals of mineral composition at each location where dust is emitted.

700

701 If the same paired simulations exist in other ESMs like ModelE, we repeat the
 702 calculation for each branch and then average the obtained results across different
 703 models with a soil mineral atlas Q :

704

$$\overline{\Delta\xi(\varphi_Q)_h} = \frac{1}{N_{model}} \sum_{m=1}^{N_{model}} \Delta\xi(\varphi_{Q,m})_h. \quad (4)$$

705

706 This approach assumes that the response to the perturbed parameter in the
 707 different ESMs is correlated: for example, increased iron oxides in the soil compared
 708 to the baseline simulation causes more warming in all the ESMs.

709

710 The half-length uncertainty in the dust SW DRE due to iron oxides is calculated as
 711 follows, ensuring a symmetric uncertainty centered about the base best result:

712

$$\Delta\xi(\overline{\varphi_Q}) = \max_{o=h,l} \overline{\Delta\xi(\varphi_Q)_o}. \quad (5)$$

713

714 The total 95% confidence interval is subsequently derived based on the model
 715 ensemble mean of the best estimates of the dust SW DREs, $\xi(\overline{\varphi_Q})$, across
 716 simulations using the base soil mineral atlases, and $\Delta\xi(\overline{\varphi_Q})$:

717

$$95\%CI_Q = [\xi(\overline{\varphi_Q}) - \Delta\xi(\overline{\varphi_Q}), \xi(\overline{\varphi_Q}) + \Delta\xi(\overline{\varphi_Q})]. \quad (6)$$

718

719 Model spread (σ) in simulating the dust SW DRE is quantified as the standard
 720 deviation (σ_{std}) of the SW DRE values simulated by all the ESMs:

721

$$\sigma = 1.96\sigma_{std}. \quad (7)$$

722

723 We also treat differences caused by the use of prognostic versus reanalyzed winds
 724 within the same model as a type of model spread.

725

726 Similarly, we calculate the dust SW DRE difference, ψ and η , arising from different
 727 emission scenarios in future simulations, if projection uncertainty is considered, and
 728 from vegetation dynamics, respectively.

729

730 To combine these factors, we assume that each factor (iron oxides, emission
 731 scenarios, vegetation, and model spread) is independent of the others and used a
 732 root mean square sum of the uncertainties caused by each factor in the paired
 733 models with a total model number of N_σ , N_ψ , and N_η , respectively:

734

$$\Delta\xi(Q) = \sqrt{\Delta\xi(\varphi_Q)^2 + \sigma^2/N_\sigma + \psi^2/N_\psi + \eta^2/N_\eta}. \quad (8)$$

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Uncertainty in the present-day and late 21st-century dust SW DRE is quantified using the same methodology described above, applied to simulations for each respective period. A combined estimate of uncertainty across both time periods is also derived, incorporating all simulations used to assess the relative importance of contributing factors to dust SW DRE variability.

Quantifying uncertainty in soil and airborne iron oxides

The uncertainty in the model-simulated change in airborne iron oxide mass fractions ($\Delta\varphi_Q$), based on either EMIT (P) or pre-EMIT (E) soil mineral atlases, is calculated using the same procedure outlined in Equations (1-6), substituting the dust SW DRE with the iron oxide mass fraction. The relative reduction in uncertainty (ε) provided by EMIT, compared to pre-EMIT soil mineral atlases, is then computed as follows and shown in Fig. 1a:

$$\varepsilon = -100(\Delta\varphi_P - \Delta\varphi_E)/\Delta\varphi_P. \quad (9)$$

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For soil iron oxides, differences across models are minimal, as all use the same underlying soil mineral atlas. The uncertainties for both EMIT and pre-EMIT atlases are derived directly from the best estimates and associated 95% confidence intervals reported by each product. We then calculate probability distributions across 16 bins, each spanning 5%, from 20% to 100% iron oxide mass fraction, and present the results in Fig. 1b.

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Observational dust SSA

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We derive climatological means of observed dust SSA in the mid-visible band (0.44-0.63 μm) from AERONET Version 3 Level 2.0 Almucantar retrievals for total aerosols^{55,56}. To isolate dust-dominated conditions, we apply stringent filters:

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1. Fine-mode exclusion: Removing retrievals with a fine-mode aerosol volume fraction larger than 10% minimizes contamination from non-dust fine-sized aerosol species (e.g., sulfate, nitrate, and organics, including absorbing carbonaceous aerosols like black/brown carbon).
2. SSA spectral signature: Retaining only retrievals where SSA increases from 0.440 to 0.675 μm distinguishes dust from all other aerosols, including sea salt which exhibits a spectrally flat SSA⁵⁷.
3. Absorption threshold: Excluding data with an imaginary refractive index larger than 0.004286 (averaged over 0.675-1.02 μm)⁵⁸ further reduces carbonaceous aerosol interference.

776 These criteria are applied to hourly data during the 2007-2011 period. Monthly dust
777 SSA is computed as an extinction-weighted mean, considering only months with
778 more than 80 hourly measurements over the 5-year period. While this approach
779 improves the prominence and fidelity of the dust signal, residual non-dust
780 contamination remains difficult to quantify.

781

782 **Observational dust DREE**

783

784 Observational data of the clear-sky SW and LW DREE at the top of the atmosphere,
785 as summarized in Supplementary Tables 3 and 4, respectively, are used to evaluate
786 model performance. The observational DREE were derived in previous studies from
787 space-borne or ground-based observations and, in some cases, including 1-D
788 radiative transfer modeling. While there are observations of surface clear-sky DREE,
789 our study focuses exclusively on those at the top of the atmosphere.

790

791 Most observational estimates are for daytime, while model results represent
792 monthly fully diurnal averages. For a consistent comparison, daytime observations
793 are converted to fully diurnal averages by accounting for the day-night duration
794 during the observation months and assuming zero SW DREE at night. Given the
795 diurnal cycle of the LW DREE, which peaks during the day and remains nonzero at
796 night, a correction is applied to convert daytime estimates to diurnal averages,
797 which are approximately 80% of the original values⁵⁹.

798

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800

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820 generation of the EMIT soil mineral atlases.

821

822

823 **Competing interests**

824

825 The authors declare no competing interests.

826

827 **Contributions**

828

829 L.L. and N.M.M. conceived the study and designed the modeling experiments. L.L.
830 performed the CESM2 simulations (18 in total), carried out the analysis (including
831 results from all the four models), and drafted the initial manuscript. L.L. and N.M.M.
832 co-wrote the experiment design section. R.L.M. contributed to the experimental
833 design, performed the ModelE simulations (13 in total), processed the model
834 results, and provided the description of the ModelE model. C.P.G.-P. and M.G.A. also
835 contributed to the experimental design. V.O. performed the MONARCH simulations
836 (2 in total) and the comparison between models and AERONET dust SSA. M.G.A.
837 processed AERONET data to create the climatological monthly means of dust SSA
838 used in the comparison. V.O. and M.G.A. provided the description of the MONARCH
839 model. P.G. conducted the AM4 simulations (3 in total) and processed the model
840 results. Q.S. and P.G. wrote the description of the AM4 model. D.R.T. wrote the
841 description of the EMIT instrument. P.G.B., M.G.A., R.L.M., C.P.G.-P., and L.L. authored
842 the description of the EMIT soil mineral atlas sections. All remaining authors, along
843 with those previously mentioned, contributed to the generation of the EMIT soil
844 mineral atlases. R.O.G. is the principal investigator of the EMIT project. All authors
845 contributed to editing the manuscript.

846

847 **Data availability**

848

849 The EMIT soil mineralogy atlas can be accessed at the link below.

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851 [https://search.earthdata.nasa.gov/search/granules?p=C2408034484-
852 LPCLOUD&pg\[0\]\[v\]=f&pg\[0\]\[gsk\]=-start_date&g=G3588168095-
853 LPCLOUD&fi=EMIT%2BImaging%2BSpectrometer&fdc=Land%2BProcess%2BDistr
854 ibuted%2BActive%2BArchive%2BCenter%2B%2528LPDAAC%2529&tl=17056987
855 63.639!4!!&lat=-8.210058215496623&long=-
856 47.58799505025262&zoom=4.119445421900885](https://search.earthdata.nasa.gov/search/granules?p=C2408034484-LPCLOUD&pg[0][v]=f&pg[0][gsk]=-start_date&g=G3588168095-LPCLOUD&fi=EMIT%2BImaging%2BSpectrometer&fdc=Land%2BProcess%2BDistributed%2BActive%2BArchive%2BCenter%2B%2528LPDAAC%2529&tl=1705698763.639!4!!&lat=-8.210058215496623&long=-47.58799505025262&zoom=4.119445421900885)

857

858 The ensemble model outputs and pre-EMIT soil mineralogy atlases are available
859 through the link provided below.

860

861 [https://search.earthdata.nasa.gov/search/granules?p=C2408755900-
862 LPCLOUD&pg\[0\]\[v\]=f&pg\[0\]\[gsk\]=-
863 start_date&fi=EMIT%2BImaging%2BSpectrometer&fdc=Land%2BProcess%2BDistr
864 ibuted%2BActive%2BArchive%2BCenter%2B%2528LPDAAC%2529&tl=14595019
865 18.074!5!!](https://search.earthdata.nasa.gov/search/granules?p=C2408755900-LPCLOUD&pg[0][v]=f&pg[0][gsk]=-start_date&fi=EMIT%2BImaging%2BSpectrometer&fdc=Land%2BProcess%2BDistributed%2BActive%2BArchive%2BCenter%2B%2528LPDAAC%2529&tl=1459501918.074!5!!)

866

867 **Code availability**

868

869 The raw model results were processed using NetCDF Operator
870 (NCO; <https://nco.sourceforge.net/>). The codes to carry out the other analyses are
871 available from the first corresponding author.

872

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874

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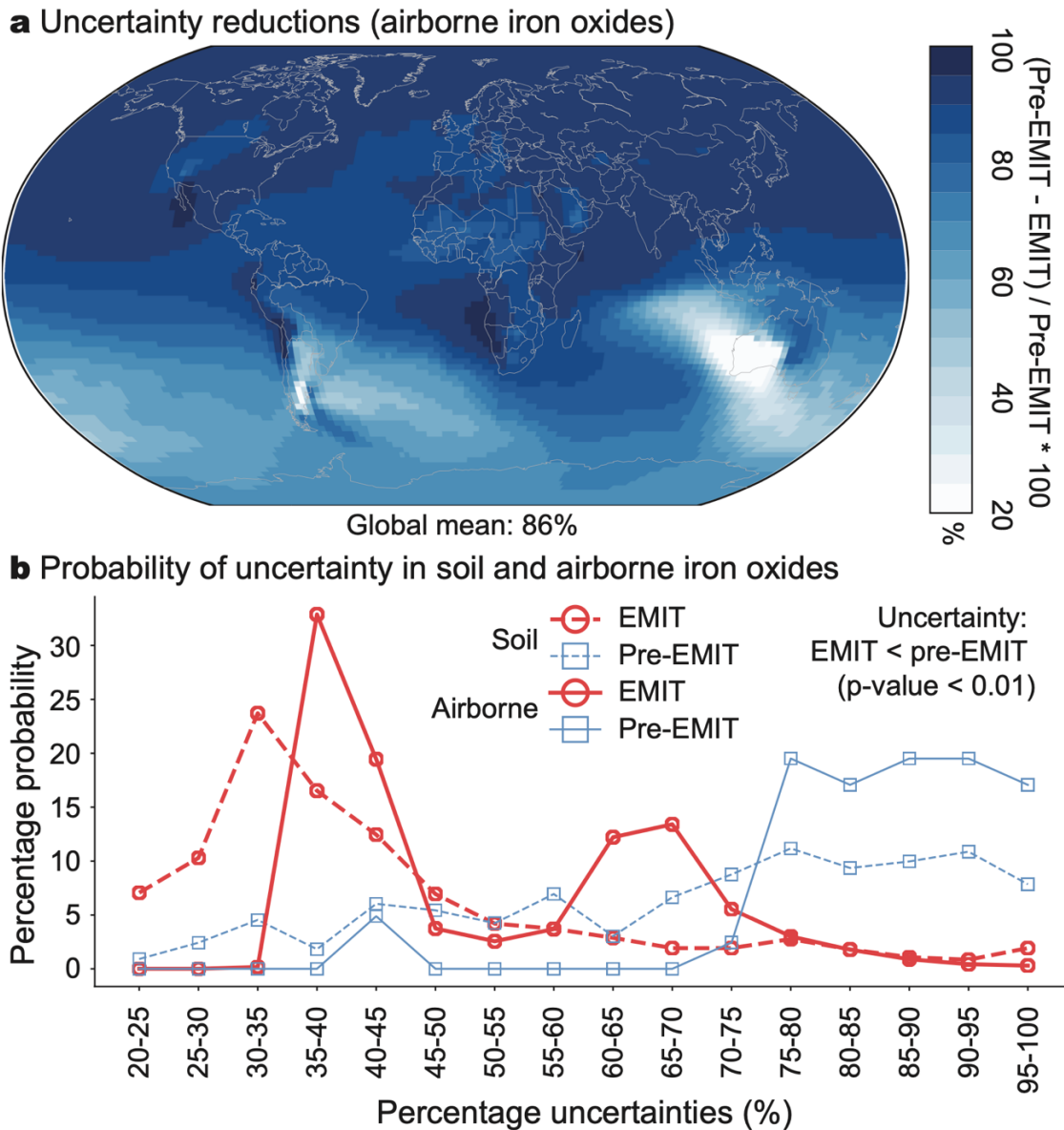
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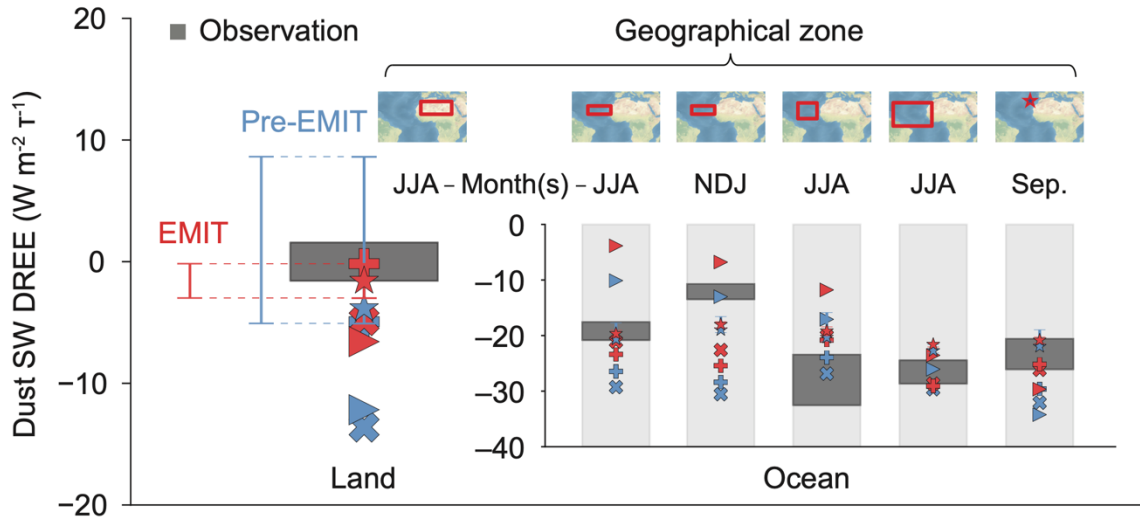
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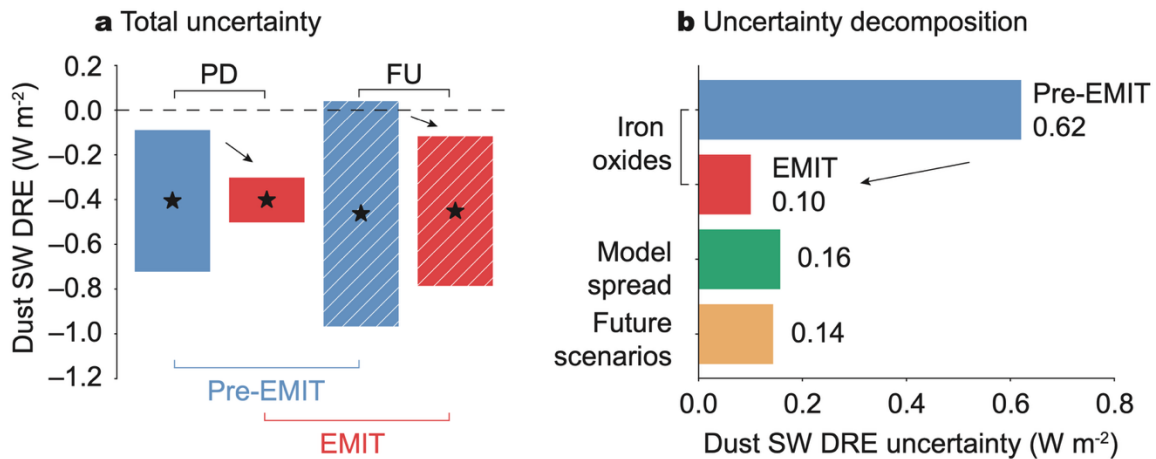
Figure 1. Improved representation of iron oxides with EMIT. **a**, Ensemble-mean 95% confidence intervals for airborne iron oxides in the present-day climate simulated by CESM2 and ModelE, showing the reduction of uncertainty resulting from EMIT compared with pre-EMIT soil mineral atlases from Claquin et al.¹⁴ (C1999) and Journet et al.¹⁵ (J2014). **b**, Probability distributions of uncertainties in soil and simulated airborne iron oxides. The EMIT uncertainty is significantly lower than the pre-EMIT uncertainty (one-sided Mann-Whitney U test, p -value < 0.01).



Model	CESM2	ModelE	MONARCH	AM4
EMIT	★	+	▶	✘
R, RMSE	0.66, 20	0.26, 17	0.60, 25	0.37, 17
Pre-EMIT	★	+	▶	✘
R, RMSE	0.66, 20	0.31, 15	0.60, 21	0.14, 14

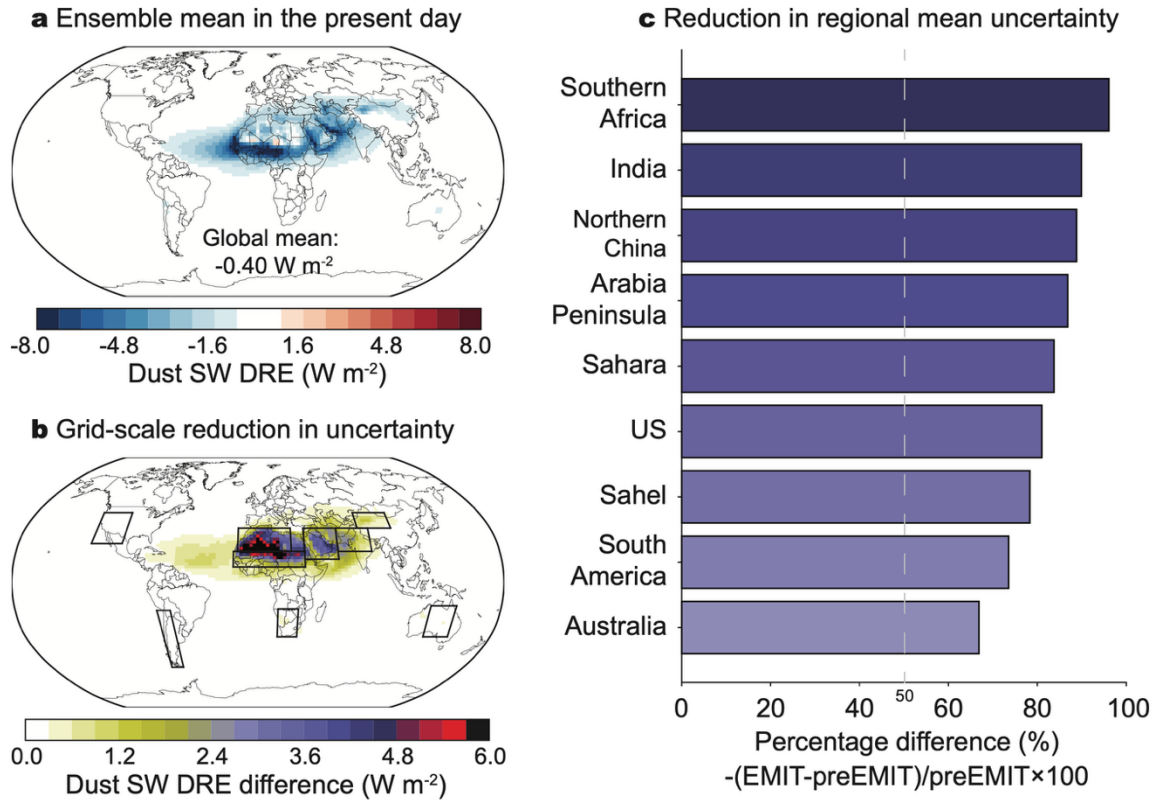
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Figure 2. EMIT improves the modeling of the dust shortwave (SW) direct radiative effect efficiency (DREE). Simulated daily mean dust SW DREE at the top of the atmosphere using EMIT and pre-EMIT soil atlases (clear-sky for CESM2; all-sky for the other ESMS: See Methods), compared with satellite-based retrievals (grey bars). Daytime mean estimates from observations have been adjusted to represent daily means (Methods). Error bars correspond to the CESM2 simulations; they are asymmetric error bars for the simulations using pre-EMIT atlases; this results from more absorbing dust in the simulation with the J2014 soil atlas¹⁵, due to its higher iron oxide content compared to the C1999 atlas¹⁴. Spatial correlation (R) and root mean square error (RMSE; unit: $W m^{-2} \tau^{-1}$) are calculated across the six regions with equal weighting.



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Figure 3. Improved estimates of the dust shortwave (SW) direct radiative effect (DRE). **a**, Ensemble mean (black stars) of global SW DRE at the top of the atmosphere under all-sky conditions in present-day (PD) and future (FU: late 21st century) climates with associated uncertainty (colored bars) for EMIT and pre-EMIT soil mineral atlases. The best estimate is the model ensemble mean; uncertainty reflects dust SW DRE variations from perturbed experiments on key factors such as iron oxide content (irox) and future emission scenarios (Methods). The EMIT uncertainty is significantly lower than the pre-EMIT uncertainty in both the PD and FU climates (one-sided Mann-Whitney U test, p-value < 0.01). **b**, Decomposition of the dust SW DRE uncertainty by individual contributing factors.



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Figure 4. EMIT reduces present-day uncertainty in the shortwave (SW) direct radiative effect (DRE) associated with dust source mineralogy. a, Ensemble mean SW DRE from simulations using EMIT soil mineral atlases. **b**, Spatial distribution of the percentage reduction in dust SW DRE uncertainty enabled by EMIT, relative to pre-EMIT soil mineral atlases. **c**, Percentage reduction in regional mean dust SW DRE uncertainty enabled by EMIT, relative to pre-EMIT soil mineral atlases. The uncertainty in the dust SW DRE is primarily governed by the mass fractions of iron oxides, particularly hematite and goethite, with limited sensitivity to the other mineral components⁷. Accordingly, the reduction in the dust SW DRE uncertainty is largely attributable to EMIT’s improved characterization of soil iron oxide abundance relative to pre-EMIT mineral atlases.